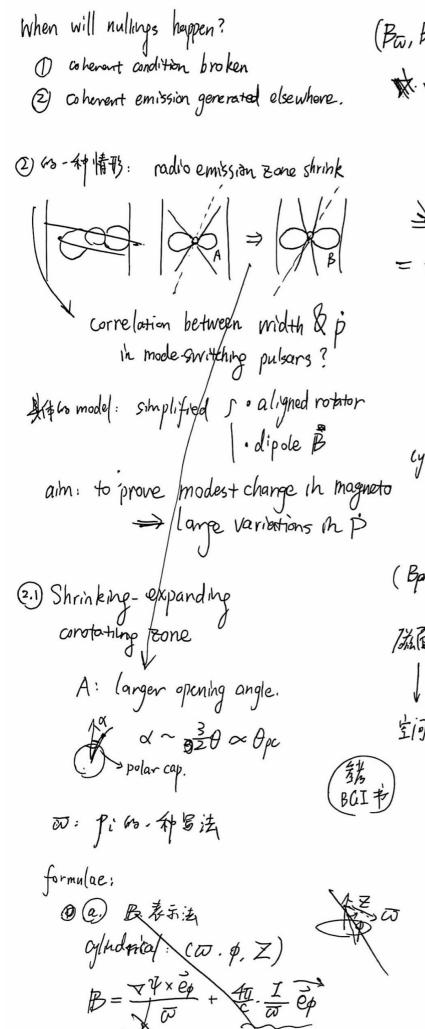
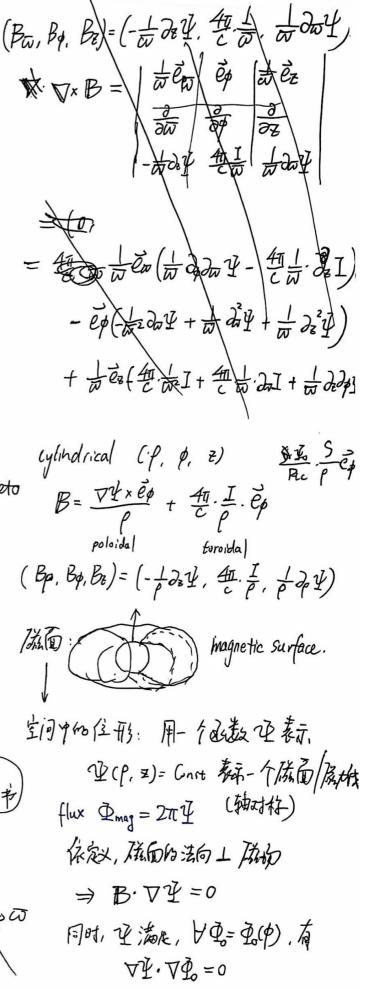
A model for nulling & mode changing in pulsars. ⇒Ω的就不能仅因的有无radio emission导致 A.N. Timokhin. (less when redio out visible Phenomena? mode changing: 不同时数, integrated prof 变 nulling: 有知时起 pulse 清末到 > (days of nulling) many other psrs: ~ hours or less Same ? Single class of reasons P支化 混剂 (+iming) 117 ti. a few pulse periods properties -> dihours/days... alobular changes: in magnetosphere nulling 子轻微 严重 · Recap: psr magnetosphere accelerating E reasons for nulling i microphysics in radiation mechanism only in geometrically very small regions Whole magnetosphere's change most magnetosphere: force-free PSR B1931+24 & J1832+0029 如何影响户: 改变 E 大學 (Kramer et al. 2006, Lyne 2009) → Pス同, 在nulling Q非 nulling 态 版為能形成及current sto 至少有一部分 nulling 伴随着磁层整体变化。 modeswitch & nulling: This paper: qualitative (semi-quantitative) Switches between some quasi-stable model to explain milling & mode change, States. J O TIP For closed field like some and the -p change. @ 7.13 bs open field like & Current distribution · Recap: "sph-down 能量分質乙 Esphdown = $\frac{1}{2}I\Omega$, $\dot{E} = I\Omega\Omega$ $\int_{1}^{2} \langle 0^{-3} \rangle radio emission$ => HE emission/cold





Euler potential g = g(P, z), $J = \frac{C}{c} I B_{pol}$ B= 77 x 72. + g. VI. general form

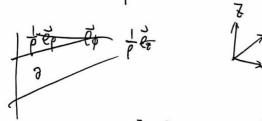
V·B= V·(VI×VI)+ FIE V(g·VI)

$$\nabla \vec{l} = \frac{\partial \vec{l}}{\partial \rho} \vec{e}_{\rho} + \frac{\partial \vec{l}}{\partial \rho} \vec{e}_{\phi} + \frac{\partial \vec{l}}{\partial z} \vec{e}_{z}$$

$$= \frac{\partial \vec{l}}{\partial \rho} \vec{e}_{\rho} + 0 + \frac{\partial \vec{l}}{\partial z} \vec{e}_{z}$$

$$\nabla \vec{l} = \frac{\partial \vec{l}}{\partial \rho} \vec{e}_{\rho} + 0 + \frac{\partial \vec{l}}{\partial z} \vec{e}_{z}$$

$$\nabla \vec{l} = \frac{\partial \vec{l}}{\partial \rho} \vec{e}_{\rho} + 0 + \frac{\partial \vec{l}}{\partial z} \vec{e}_{z}$$



$$\Delta_{i}(-...) = \frac{1}{i} \stackrel{\text{def}}{\Rightarrow} \left(\frac{24}{34} \cdot \frac{34}{34} \right) + \frac{35}{3} \left(\frac{134}{134} \cdot \frac{34}{34} \right)$$

$$= -\frac{1}{\sqrt{94}} \cdot \frac{3695}{3\sqrt{5}} + \frac{1}{\sqrt{5}} \cdot \frac{3695}{3\sqrt{5}} \cdot \frac{36}{3\sqrt{5}} = 0$$

$$\nabla \cdot (\vec{a} \cdot \Delta \vec{b}) = \Delta \vec{a} \cdot \Delta \vec{b} \cdot \vec{b} \cdot \vec{b} = \vec{b} \cdot \vec{b} \cdot \vec{b} \cdot \vec{b} = \vec{b} \cdot \vec{b} \cdot \vec{b} \cdot \vec{b} = \vec{b} \cdot \vec{b} \cdot \vec{b} \cdot \vec{b} \cdot \vec{b} = \vec{b} \cdot \vec$$

$$E = -\frac{(2 \times 1)}{c} \times B = -\frac{1}{c} \left(2 e_{z} \times (p e_{p} + z e_{z}) \right)$$

$$= -\frac{1}{c} \Omega p \cdot e_{p} \times \left(\frac{1}{p} \nabla I \times e_{p} + \frac{1}{p} e_{p} \right)$$

$$= -\frac{1}{c} \Omega \nabla I$$

$$\vec{P}_{pl} = \cdot \frac{\Omega}{c} I \cdot \frac{1}{r} \left(\partial_z \vec{\Psi} \vec{e}_p - \partial_z \vec{\Psi} \cdot \vec{e}_z \right)$$

$$\vec{P}_r = -\frac{I\Omega p}{c} \frac{1}{r} \cdot \left(\vec{Z} \cdot \partial_r \vec{\Psi} - \vec{P} \cdot \partial_z \vec{\Psi} \right)$$

Energy loss:
$$dW = -r^2 Pr. d\omega \Rightarrow solid angle$$

$$\Rightarrow \frac{dW}{d\omega} = \frac{I\Omega}{c} \cdot \frac{r}{r} \cdot (z \cdot \partial_r \psi - \rho \partial_z \psi)$$

$$W = \int_{ATI} \frac{dW}{dw} \cdot dw$$

$$= \int_{ATI} \frac{dW}{dw} \cdot \frac{\sinh \theta}{\sinh \theta} d\theta d\theta$$

$$= \int_{E^2 + \rho^2} \frac{dW}{dw} \cdot \frac{\rho}{\int_{E^2 + \rho^2}}$$

$$\beta = \gamma \sin \theta$$
, $\gamma = \sqrt{\beta^2 + \xi^2}$
 $\phi = \phi$ $\theta = a \arctan \frac{\beta}{z}$
 $z = \gamma \cos \theta$ $\phi = \phi$

$$d\theta = \frac{1}{1 + \left(\frac{\rho}{z}\right)^{2}} \left(\frac{d\rho}{z} + \frac{d\rho}{z^{2}}\right)$$

$$= \frac{1}{z^{2} + \rho^{2}} \left(zd\rho - \rho d\rho\right)$$

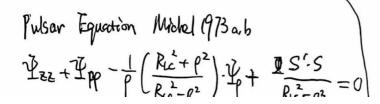
$$t = \frac{1}{2z\rho}$$

$$\begin{aligned} \mathcal{Y}_{last} &= \mathcal{Y}|_{z=0} \\ &= \mu \cdot \frac{1}{\rho} \end{aligned}$$





W= San (Z) - P21) - (Zd) - Paz) dp=roso.do small dip Z= forthy VOSO $\partial_{1} \mathcal{L} = (\partial_{1} \mathcal{L}) \frac{\partial r}{\partial \rho} + (\partial_{0} \mathcal{L}) \frac{\partial \theta}{\partial \rho} = (\partial_{1} \mathcal{L}) \frac{1}{\sin \theta} + (\partial_{0} \mathcal{L}) \frac{1}{\cos \theta}$ P= rsihe 224= (244) - dr + (204) - de = (2r4) - (204) - (204) - (201) W= SATT C SIND (rosd. 24 - 13MB 24) sind do do = Jan In (r. cost orf + dof or r. sold orf + dof) do do = 2 \ \frac{127}{27} I\Omega_F d\frac{1}{2} \\
I = I(\frac{1}{2}) pen B TA3 W=[Wind] [" bust Sdy S=-42-42-4) I ~ - 1/2- 1/2) Michel 19/36 - Solast XC2- Xinst) JX - 9- 1/ 1 + 1 . Vian → ₩~ -@ Ø Wast last S=-(Years - 7/2) Ope = PNS (W= Ope & d+ Man - 1 (P=I·U) 2.2 Plast ret, je > W~ sie & d JJ.DB.rdk/c S= # (1) ~ J. 2 8 - 1162 $\propto J \cdot \Omega$



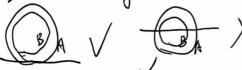


Jac D (Blandford & Znajek 1977)

a Diseussion;

1) If miling/mode change pure geometrical?

> Not really ...



For this kind of psr, change of magnetosphoric paras may cease radio emission...

(2) mode switching pulsars Pchange? measurements...

with high timing noise.

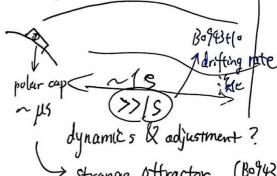
(3) Why states are quasi-stable?

total energy $E \Rightarrow E(R_Y, J, ...)$

Ry local minima? different states.

With per aging, V polarcap V potential small Wariation leads to larger changes

(4) psr magnetosphere: highly non-linear system.



> strange attractor (B0943?)

Gaasi-peniodically

(b) RRAT: pure goometrical? → 統計/約-T. RRAT 63/19?

https://rratalog.github.io/rratalog 超高温人世…

J distribution change (Violarca) change



$$\frac{1}{\sqrt{1 + z^{2}}} = \frac{1}{\sqrt{1 + z^{2}}} = \frac{1}{\sqrt$$

$$\frac{dS^{2}}{d\Psi} = \frac{\partial \Psi}{\partial \Psi} \cdot (I - \frac{\Psi^{2}}{\Psi^{2}})$$

$$S = A \Psi^{2} + B \Psi + C$$

$$S' = 2A \Psi + B$$

$$S' \cdot S' = 2A^{2} \Psi^{2} + AB \Psi^{2} + 2AB \Psi^{2} + B\Psi$$

$$+ 2AC\Psi + BC$$

$$= 2A^{2} \Psi^{3} + 3AB \Psi^{2} + (B^{2} + 2AC)\Psi$$

$$+ BC$$

$$= 2A^{2} \Psi^{3} + 3AB \Psi^{2} + (B^{2} + 2AC)\Psi$$

$$+ BC$$

$$\Rightarrow S = \frac{1}{4} \Psi^{2} - \Psi_{0}$$

$$\Rightarrow S = \frac{1}{4} \Psi^{2} - \frac{1}{4} \Psi^{2} + \frac{1}{4} \Psi^{2} \Psi^{2} +$$

