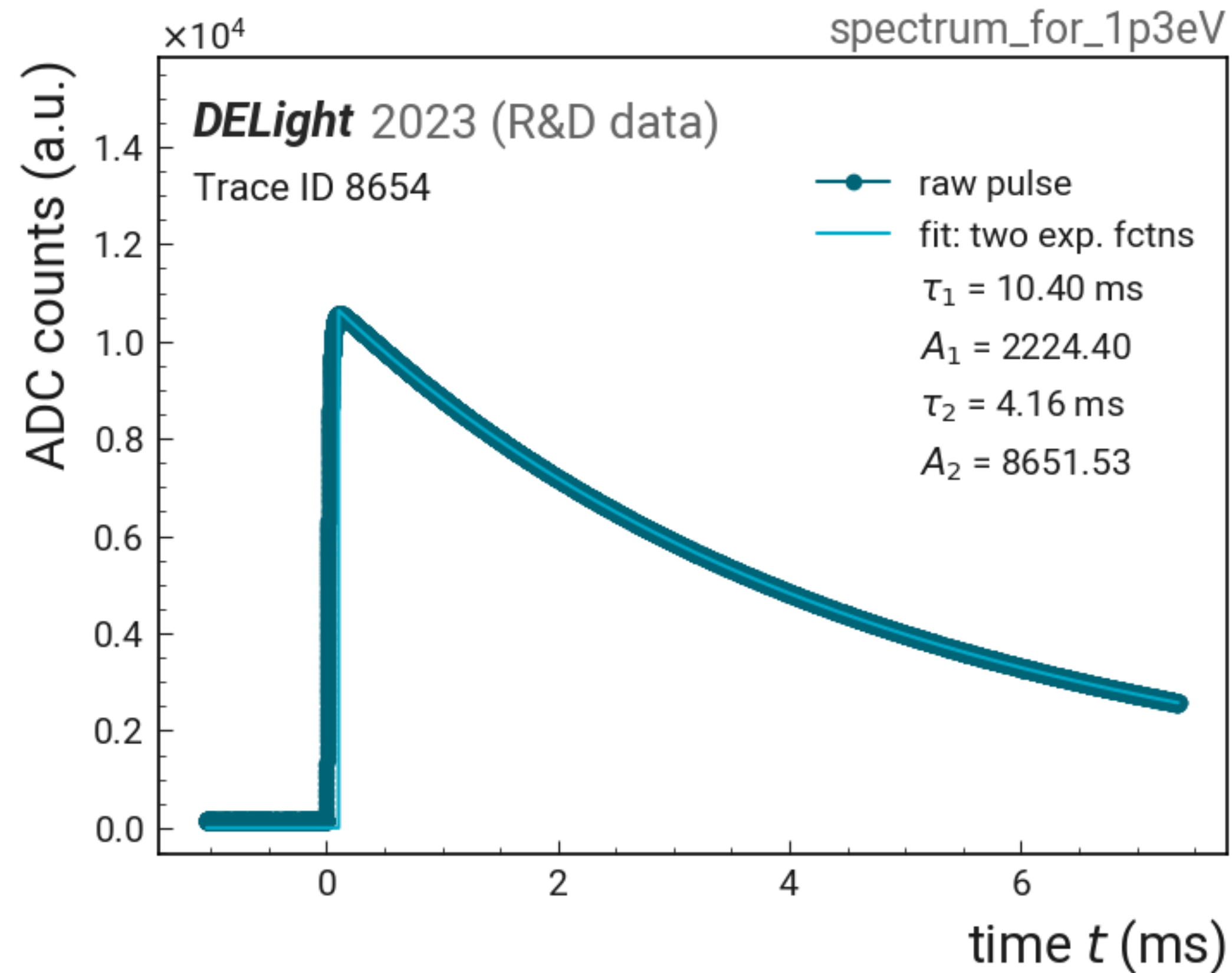
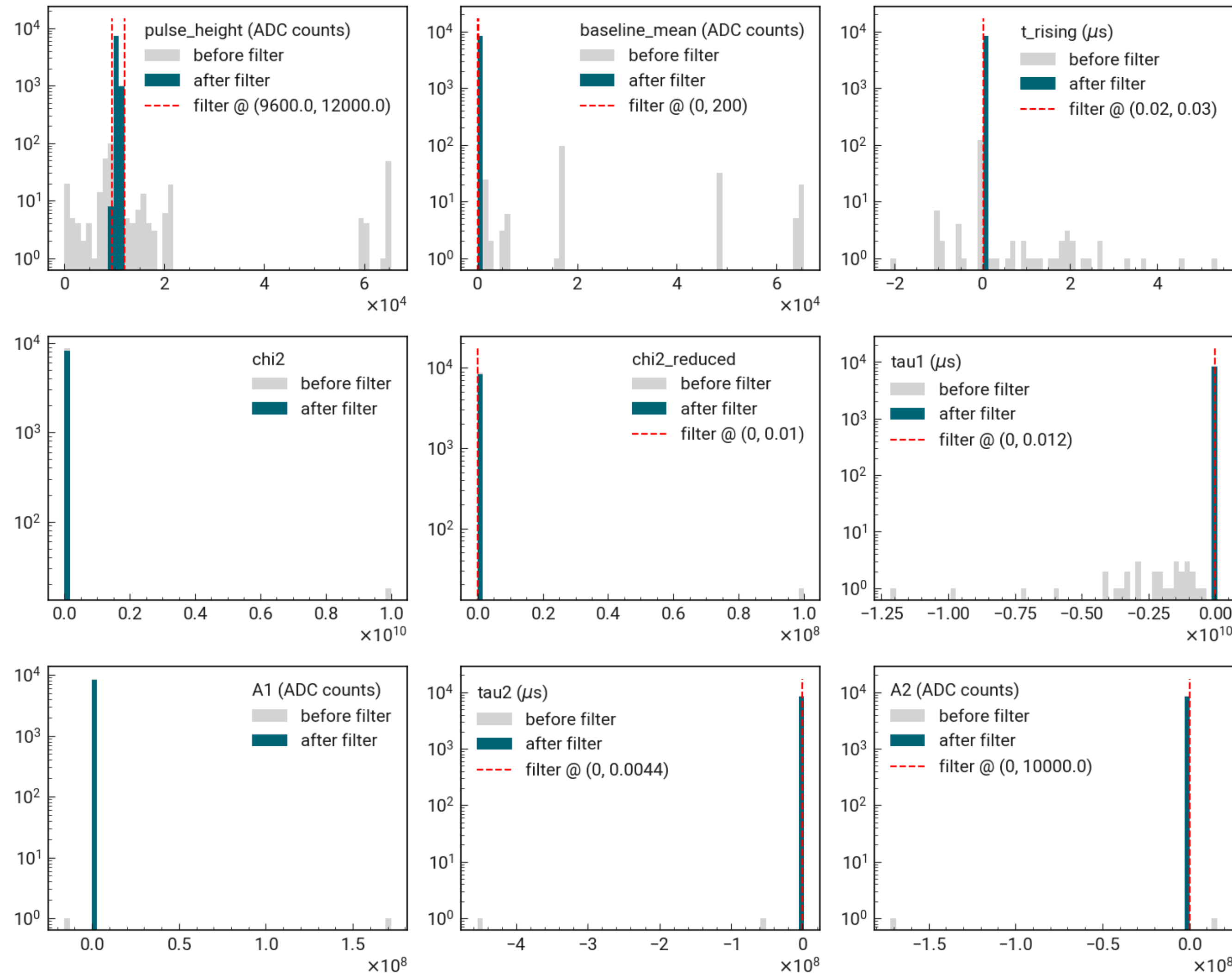


# Pulse decay fit

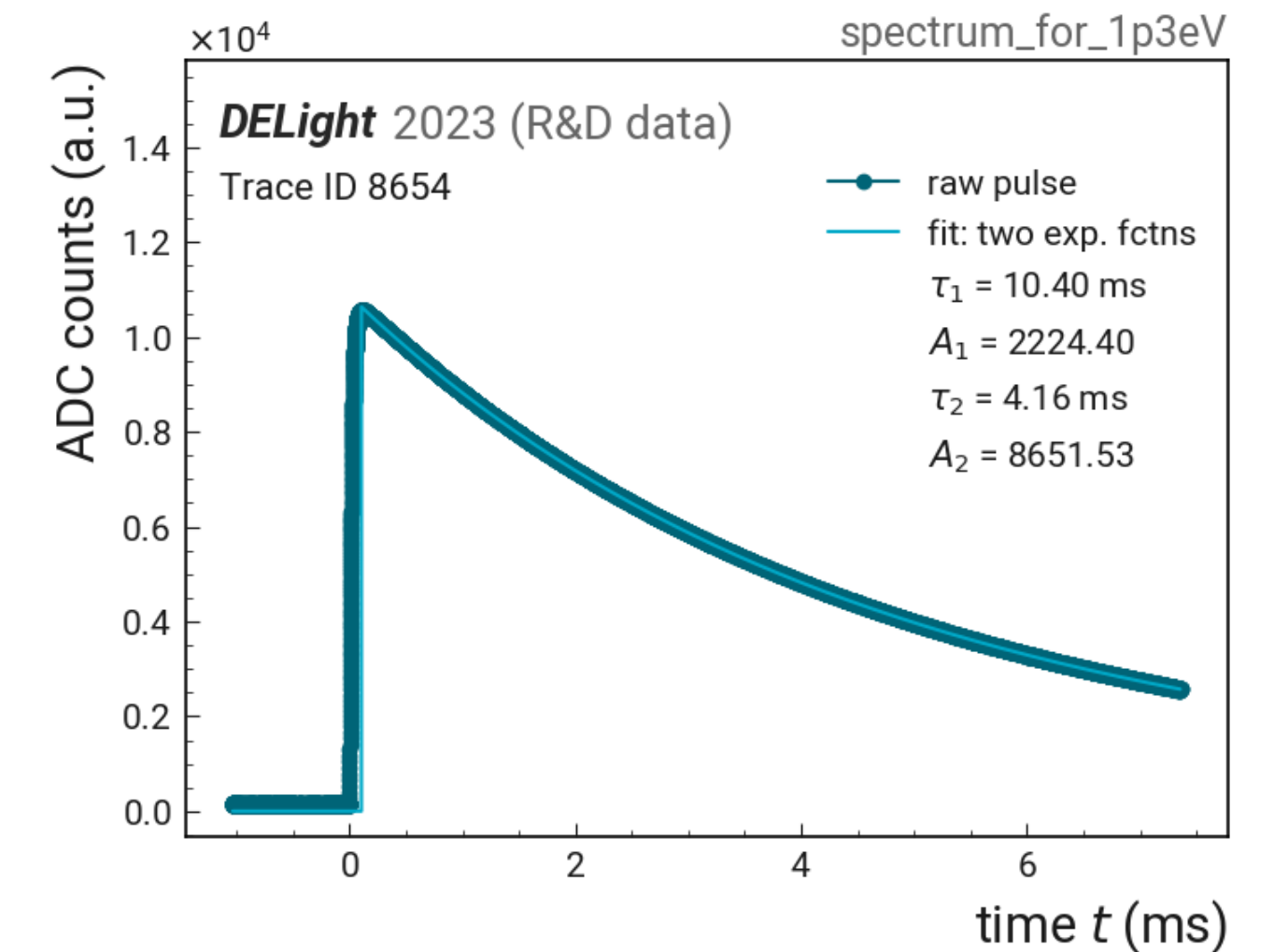


- pulse decay fit using
$$f(t) = A_1 \cdot e^{-t/\tau_1} + A_2 \cdot e^{-t/\tau_2}$$
- determination of chi-square  $\chi^2$  and chi-square per degree of freedom ( $\nu$ )  $\chi^2_\nu$   
with  $\nu = n - m$  equals the number of observations  $n$  minus the number of fitted parameters  $m$

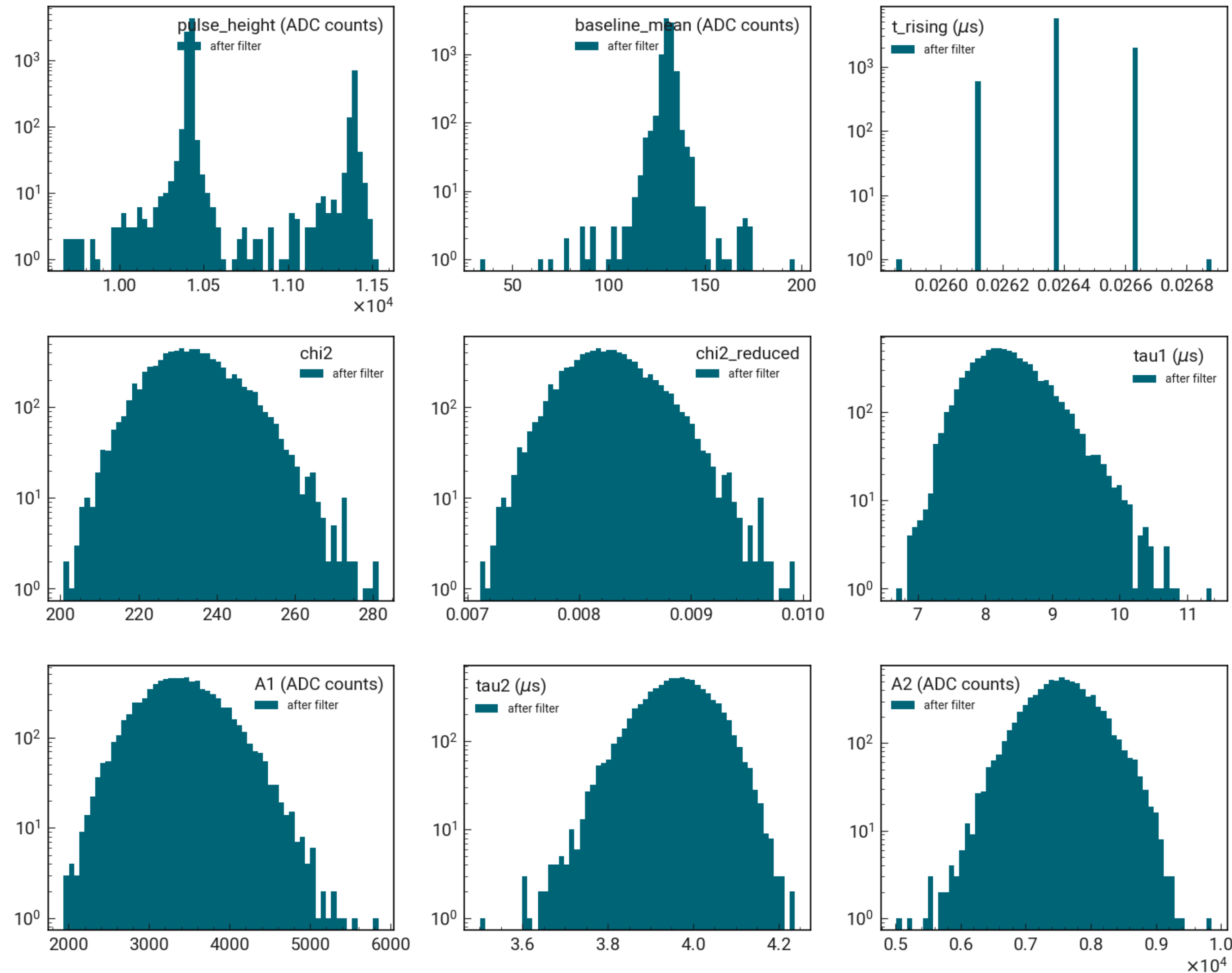
# Data filtering: outlier cuts



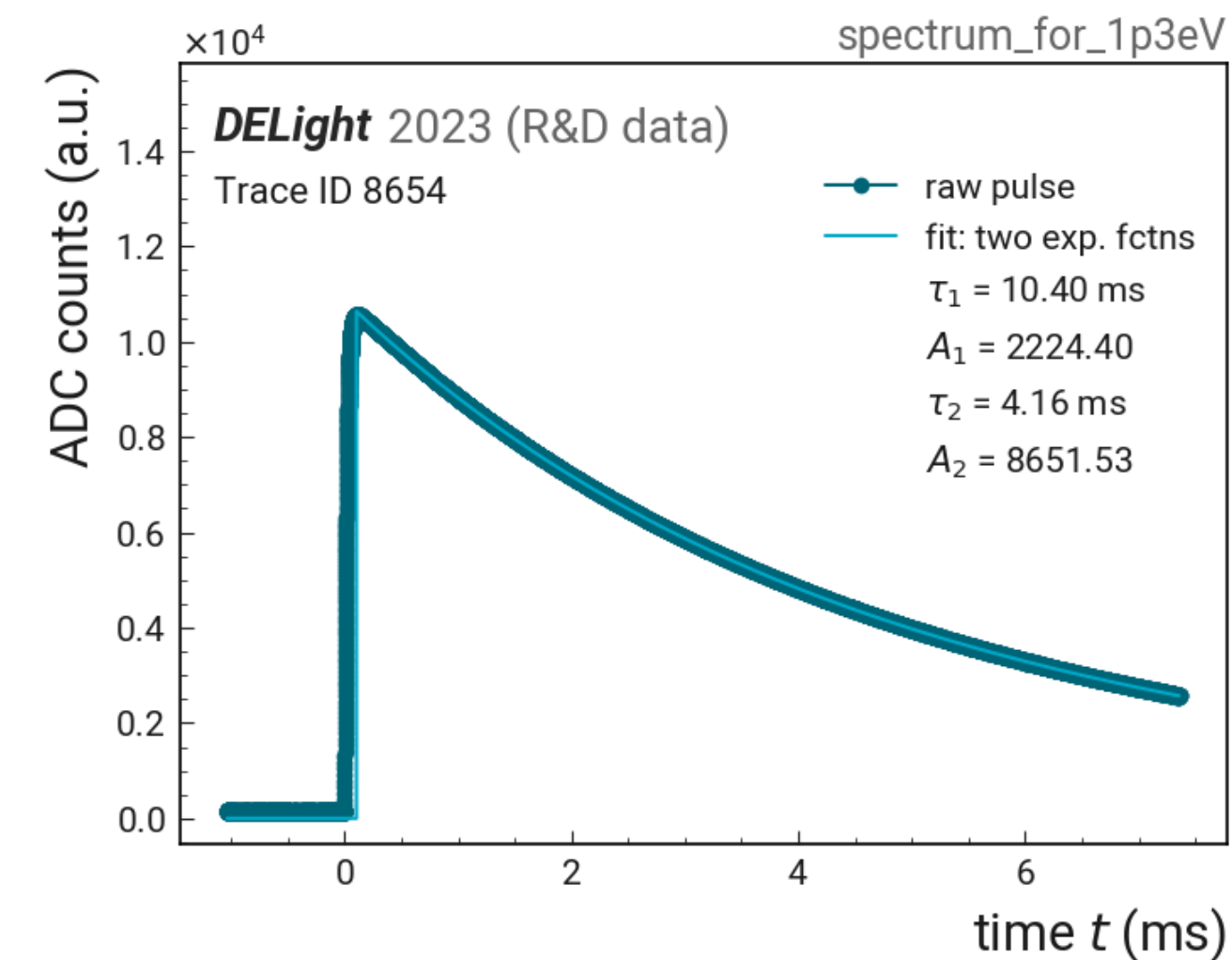
- remove all events with features outside of expected distribution
- pulse height:** maximal value - baseline
- baseline:** mean baseline value
- pulse rising time:** time between 20-80% of pulse height
- $\chi^2, \chi^2_{\nu}, \tau_1, A_1, \tau_2, A_2$  from fit



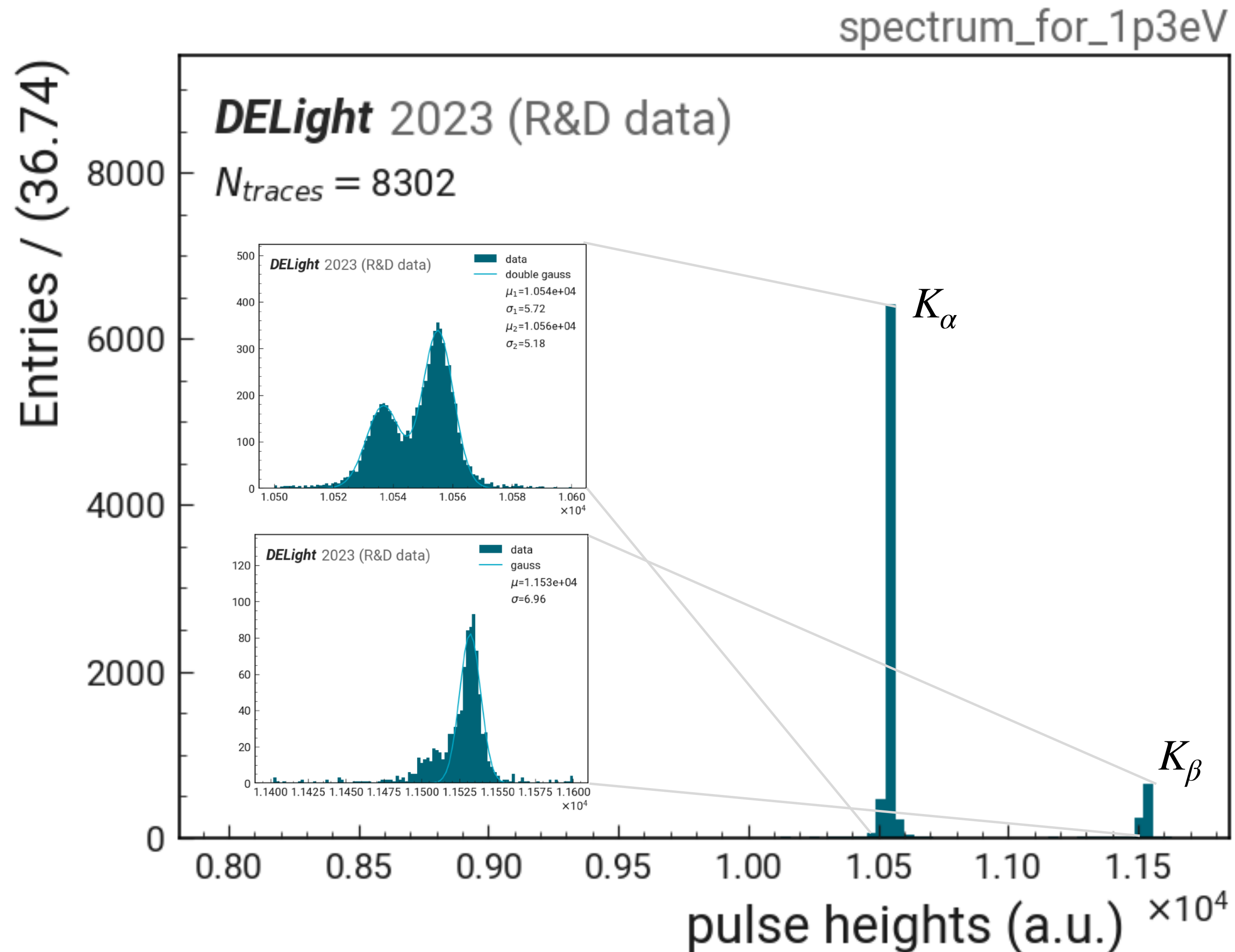
# Data filtering: outlier cuts



- remove all events with features outside of expected distribution
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- $\chi^2, \chi^2_{\nu}, \tau_1, A_1, \tau_2, A_2$  from fit



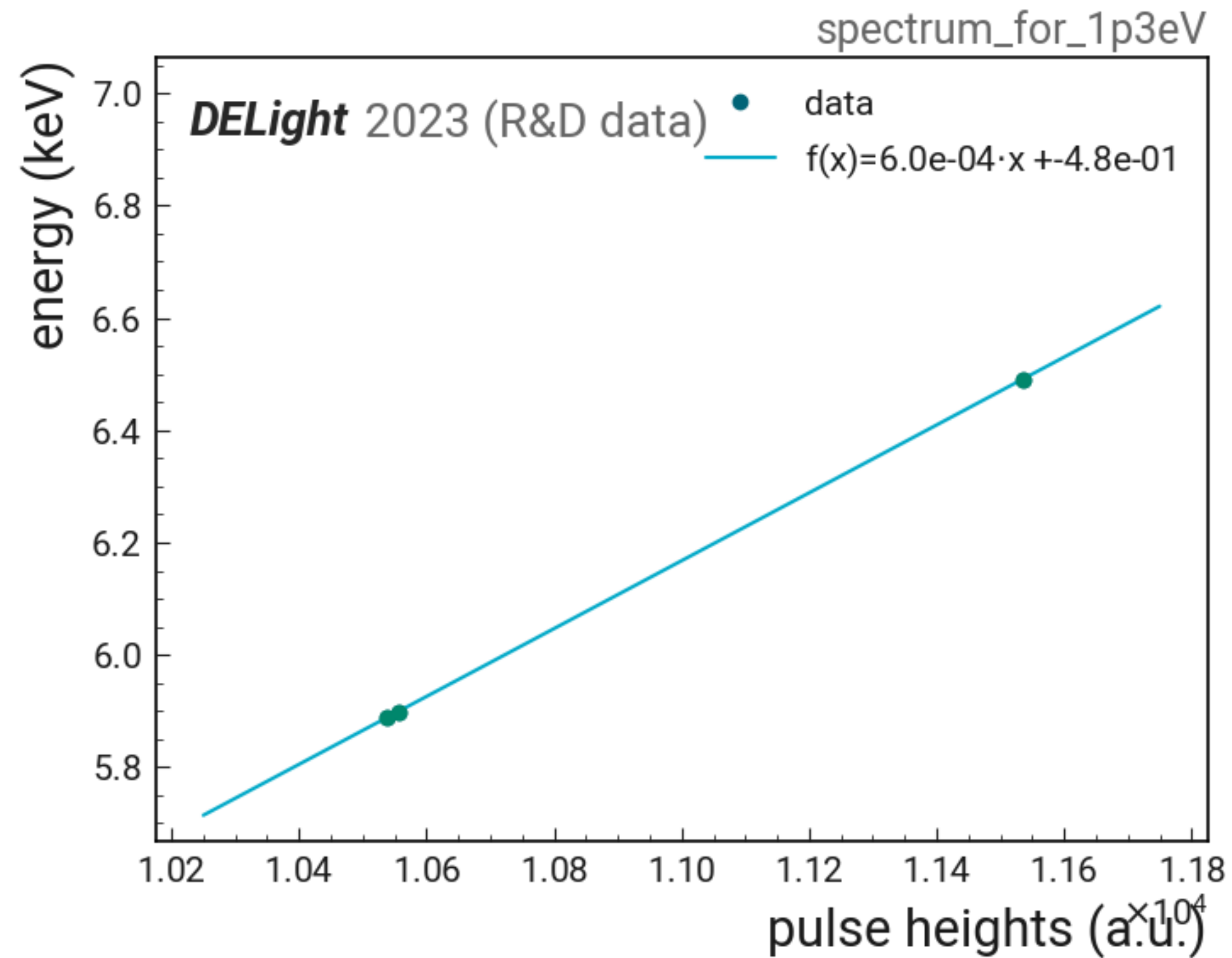
# Energy calibration by pulse heights



- distribution of pulse heights corresponds to natural line shape of the  $^{55}\text{Mn}$   $K_\alpha$  and  $K_\beta$  lines
- line shape can be used to calibrate the energy

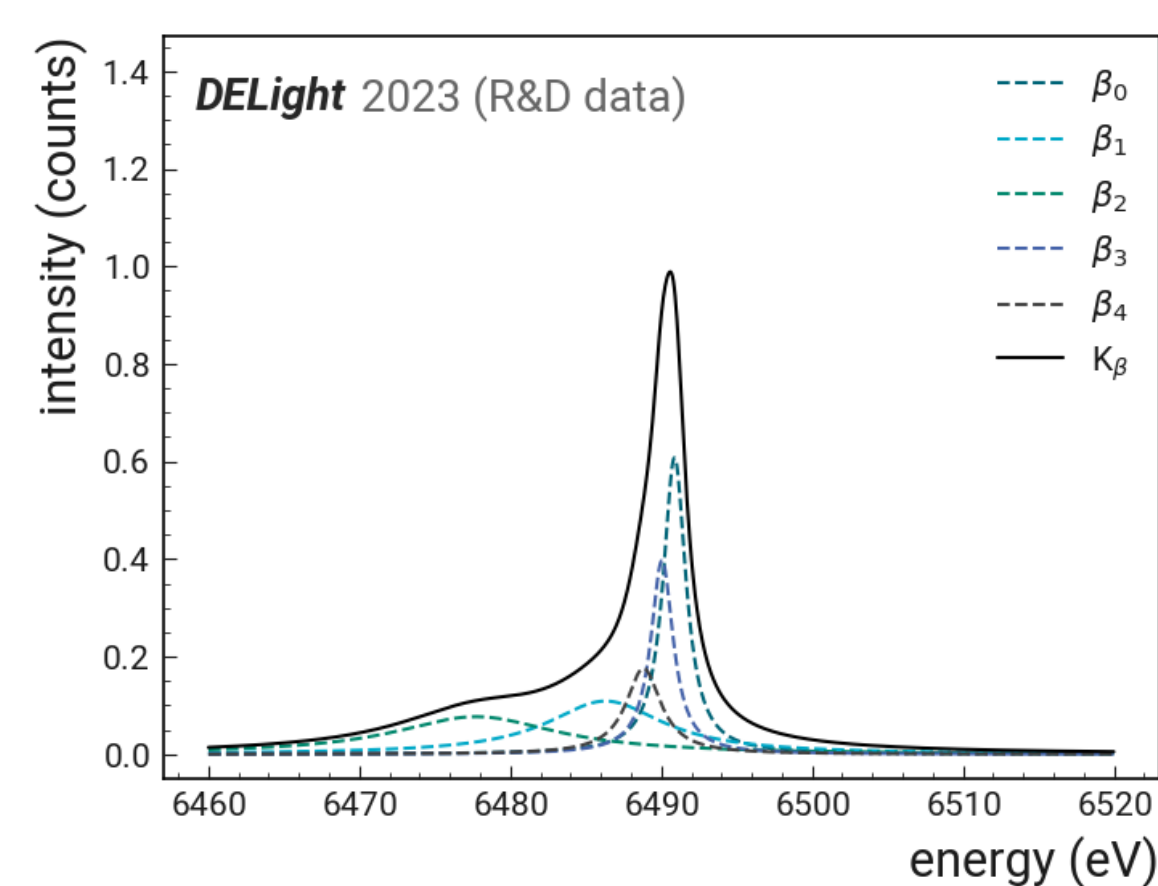
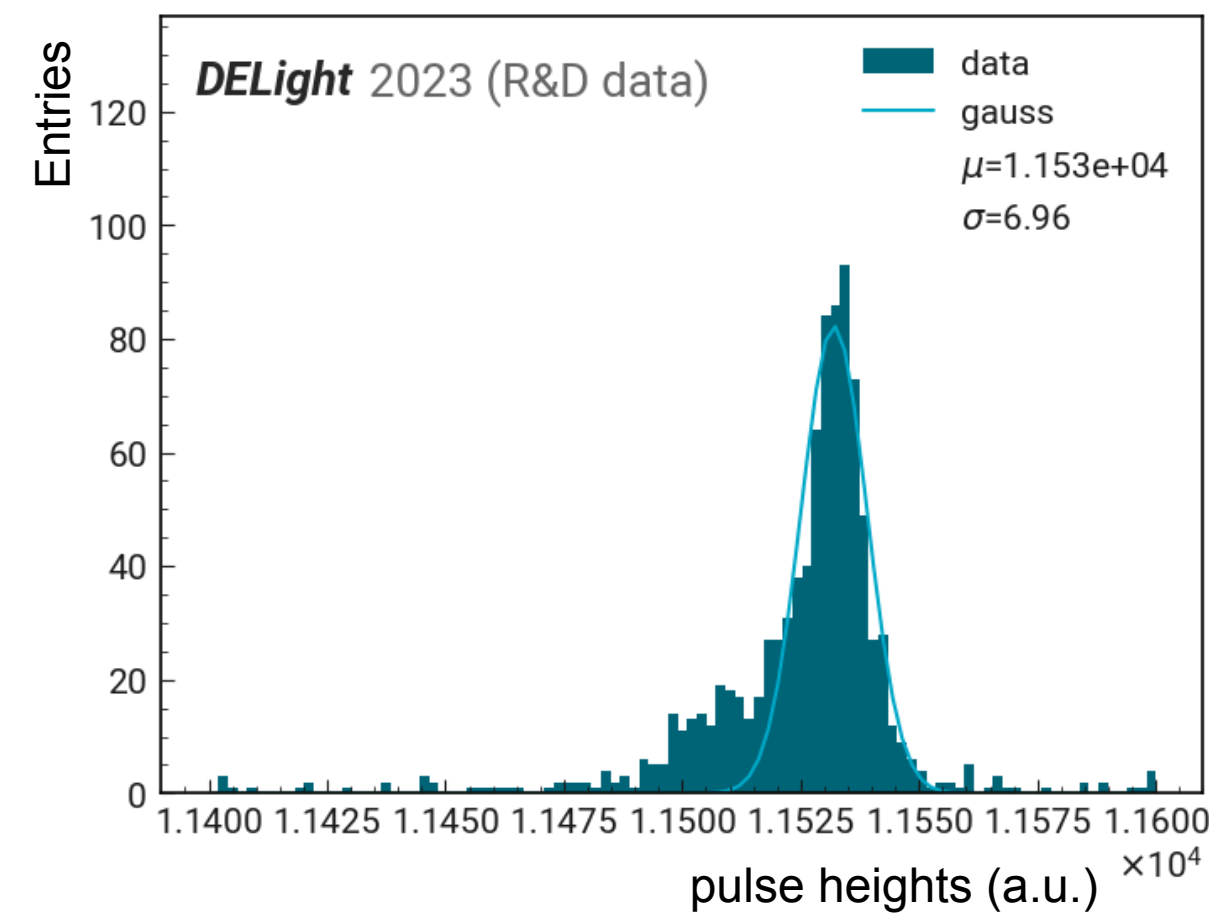
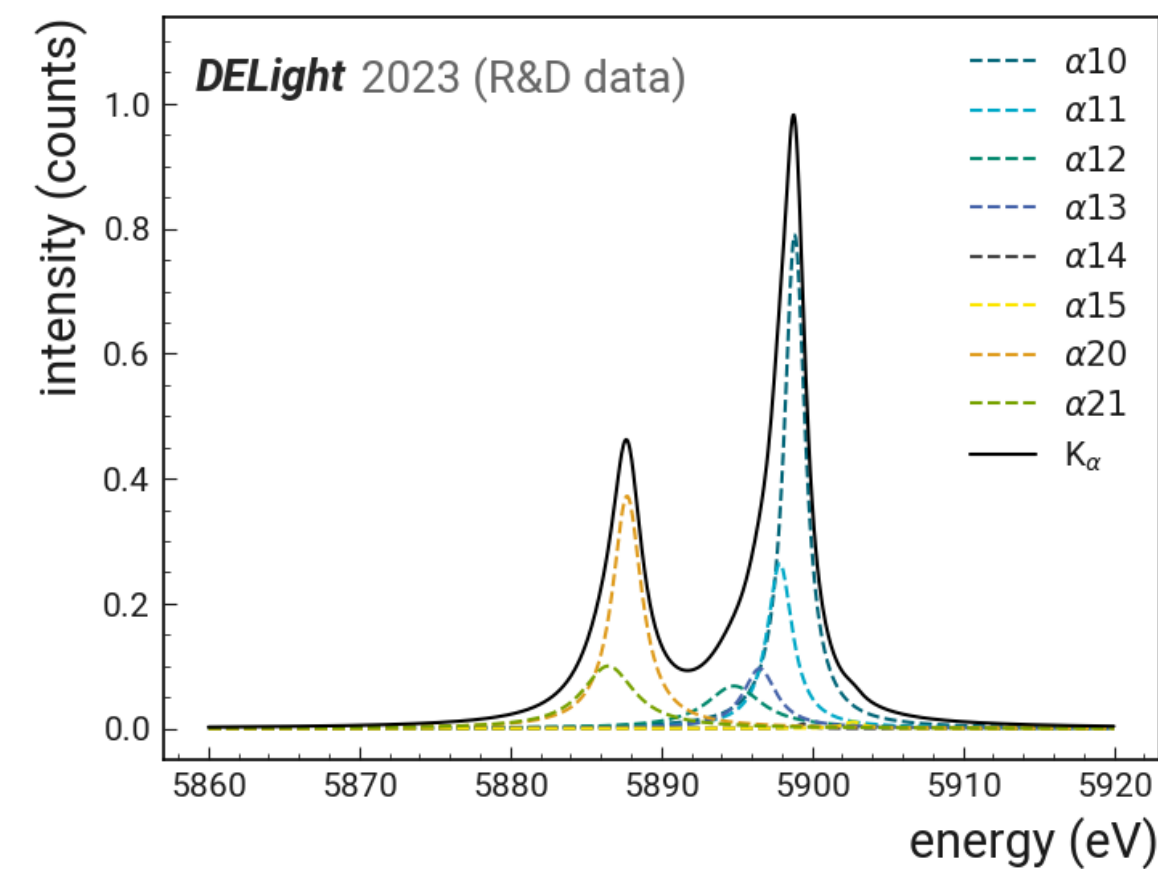
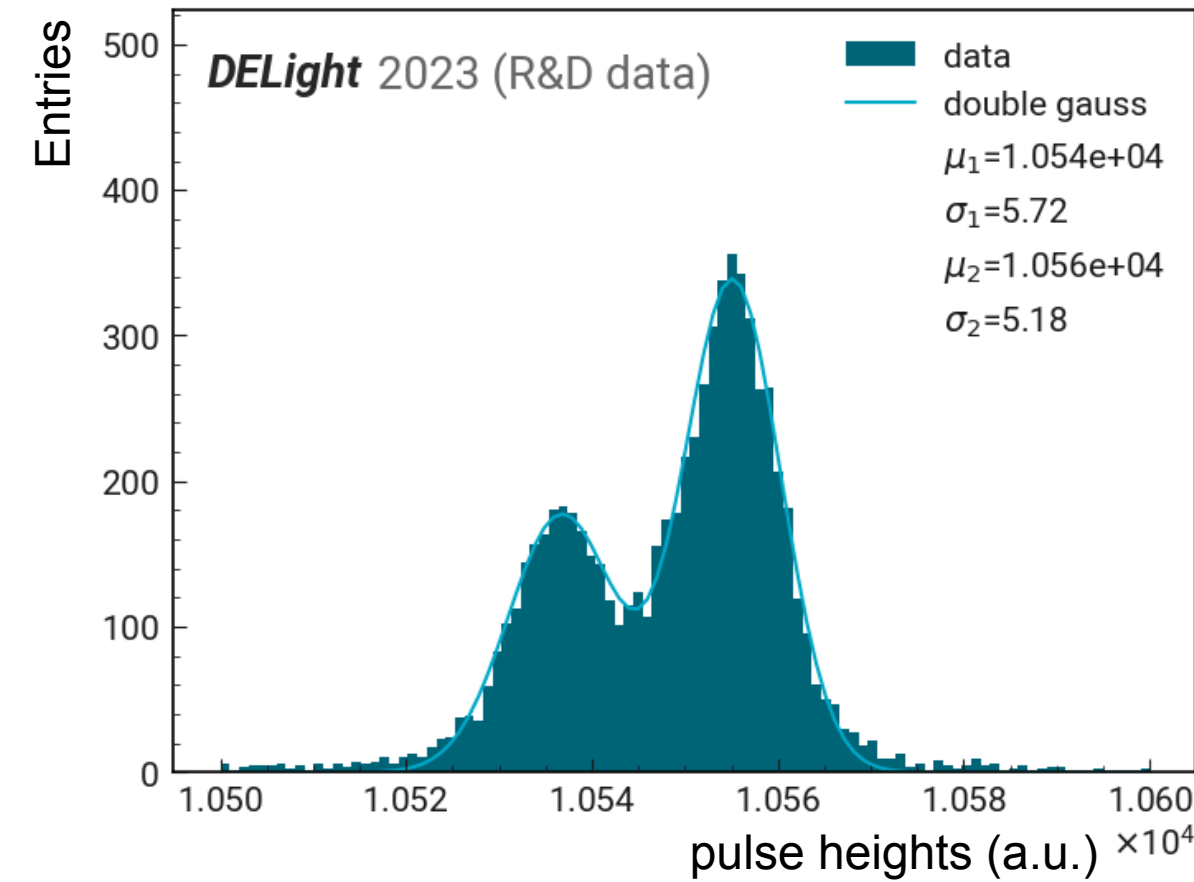


# Energy calibration



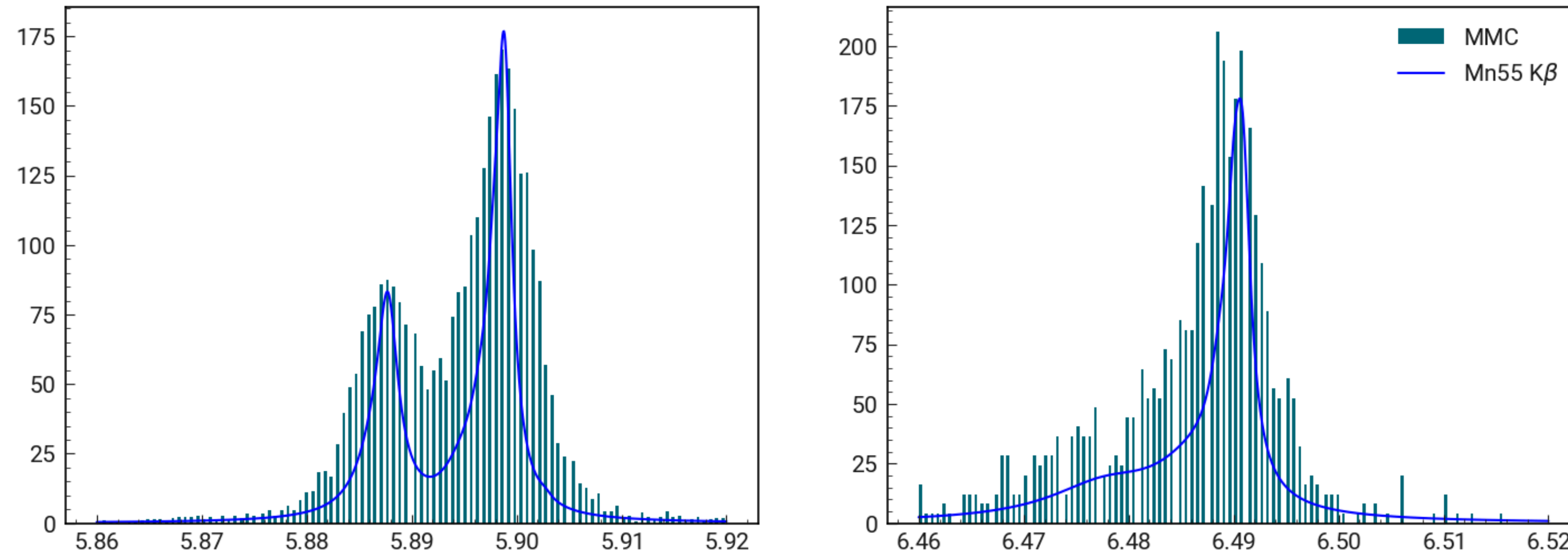
■ plot pulse height gauss peak positions over line shape energies and fit

# Natural line shape of $^{55}\text{Mn}$



spectra from [Ka1,2 and Kb1,3 x-ray emission lines of the 3d transition metals](#) (G. Hölzer et.al.)

# Comparison energy calibration vs. natural line shape



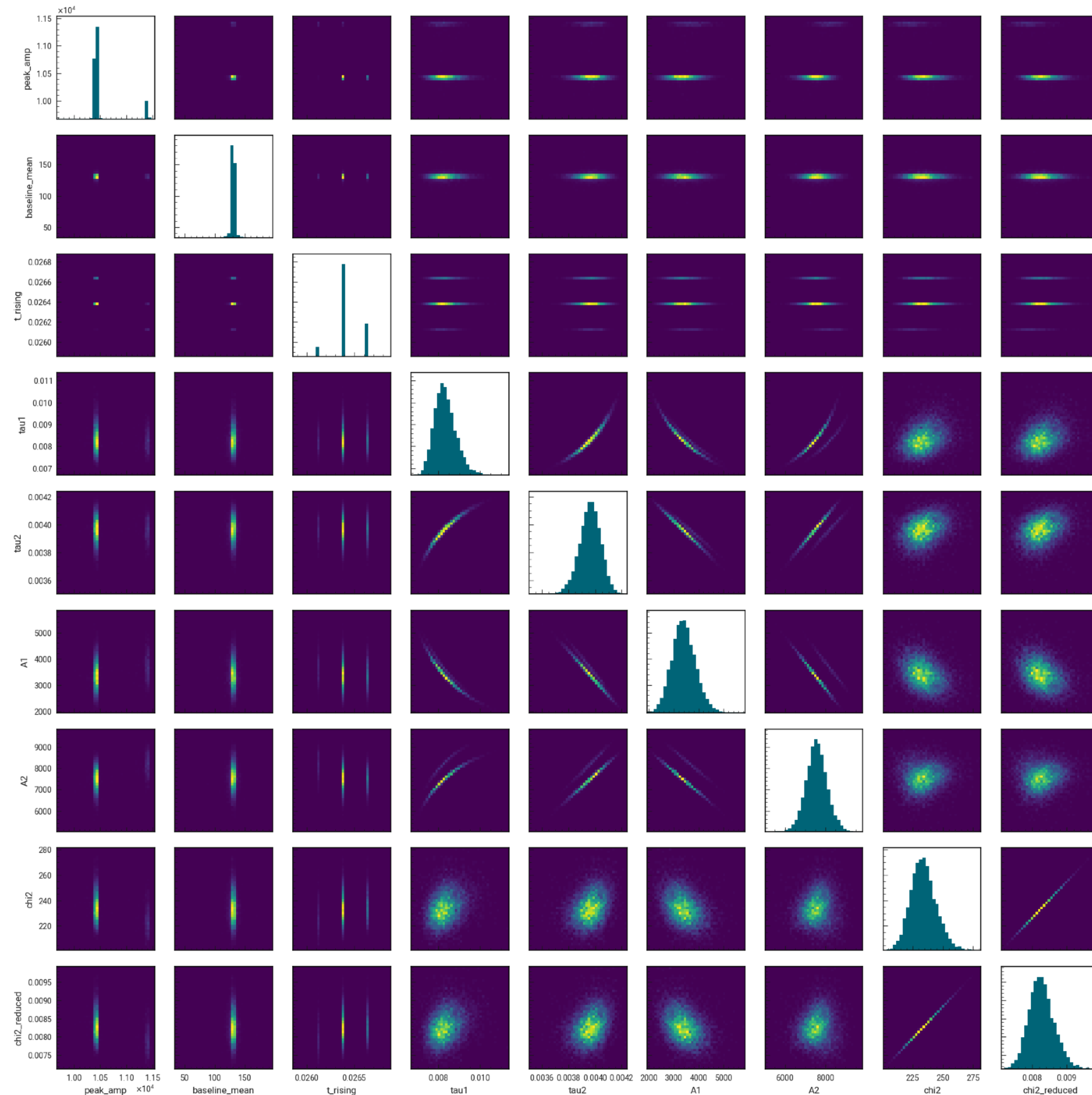
■ positions match very well

■ width needs to be broadened by convolution with gauss (instrument resolution)

# Backup



# Feature correlations



before filtering: correlations hard to identify

after filtering:

features nearly gauss distributed

strong correlations between fit parameters ( $A_1, A_2, \tau_1, \tau_2$ ) and between  $\chi^2$  and  $\chi^2_{\nu}$

baseline mean, pulse amplitude and rising edge time independent from other features