

(Problems with) MMC Data Fits

Reminder

- Spectral lines described by Voigtian

$$V(x; \sigma, \gamma) = (G * L)(x) = \int G(\tau) L(x - \tau) d\tau$$

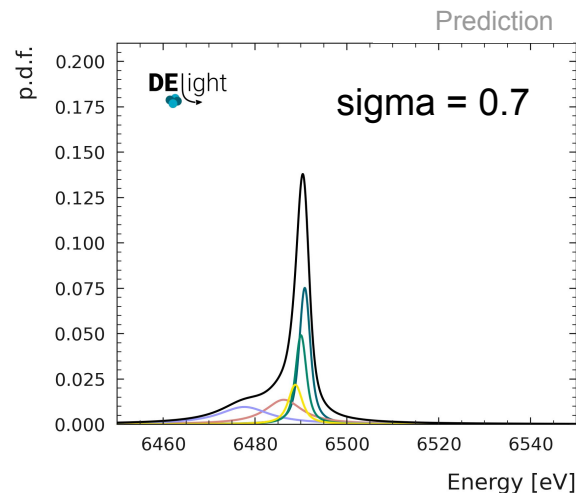
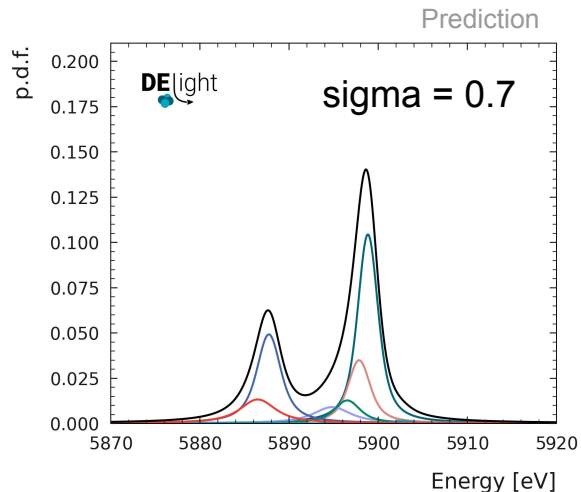
$$G(x; \sigma) = \frac{e^{-x^2/(2\sigma^2)}}{\sigma\sqrt{2\pi}} \quad \leftarrow \text{Detector}$$

$$L(x; \gamma) = \frac{\gamma}{\pi(x^2 + \gamma^2)} \quad \leftarrow \text{Source}$$

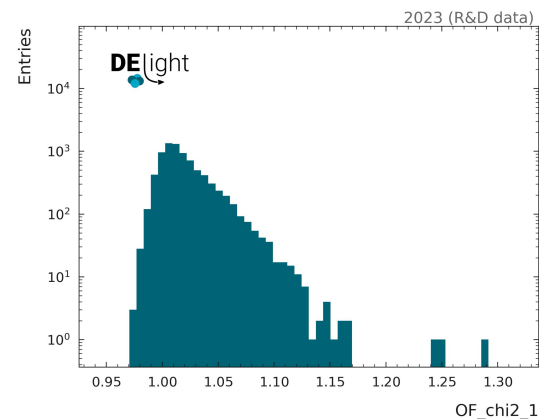
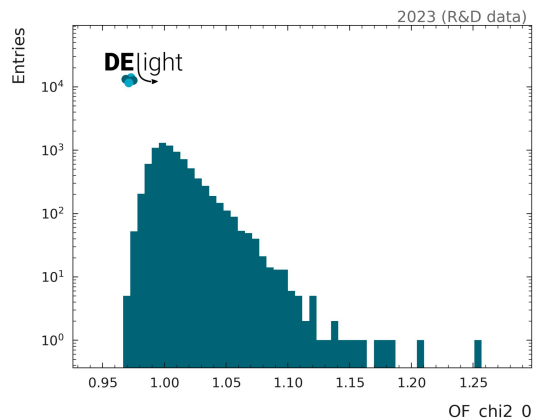
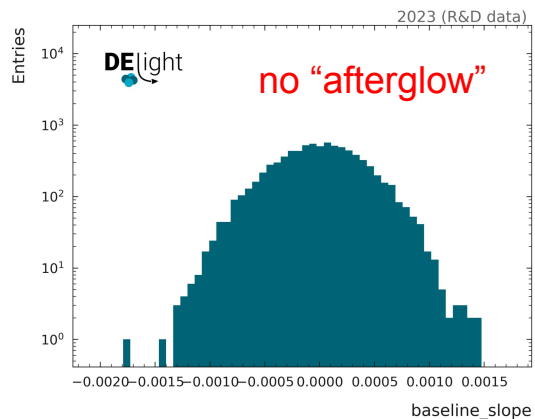
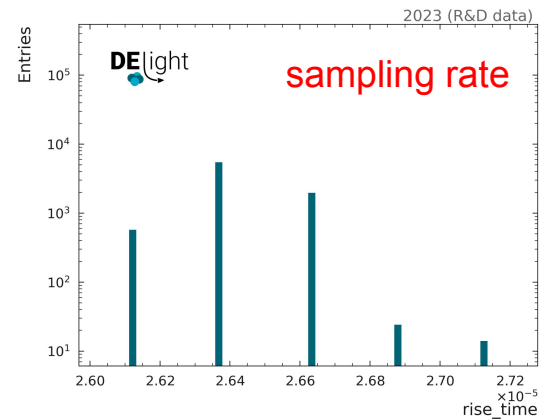
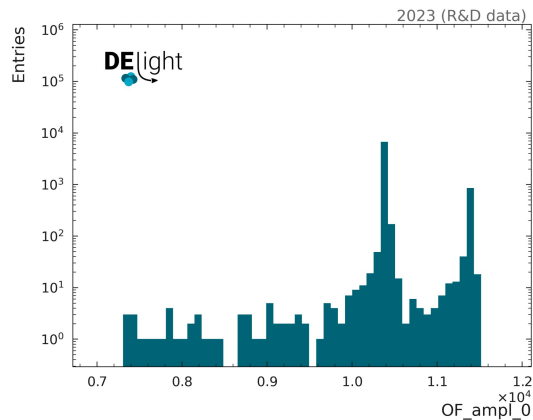
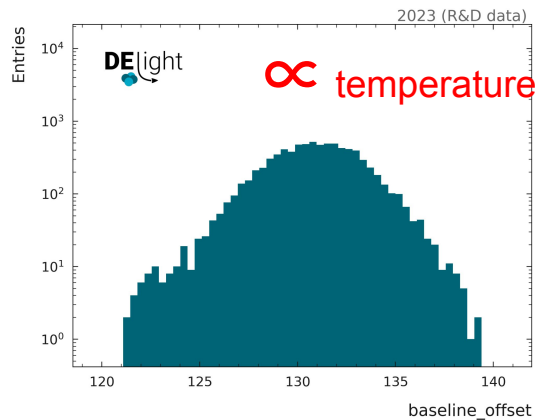
No analytic solution, but good approximations exist

$$\text{PDF} = k_{\alpha} * \text{Sum}(V_{i_{K\alpha}}) + (1 - k_{\alpha}) * \text{Sum}(V_{i_{K\beta}}) \quad , \quad k_{\alpha} = \text{rel. intensity}$$

- Here I present the current status and things I've tried to estimate and improve the resolution

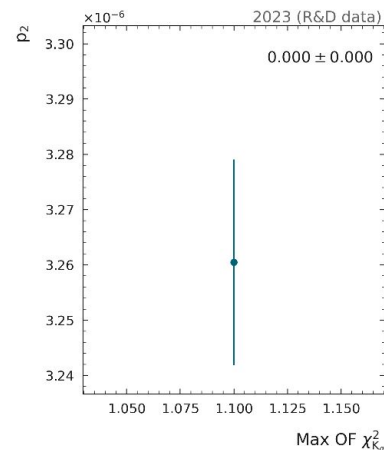
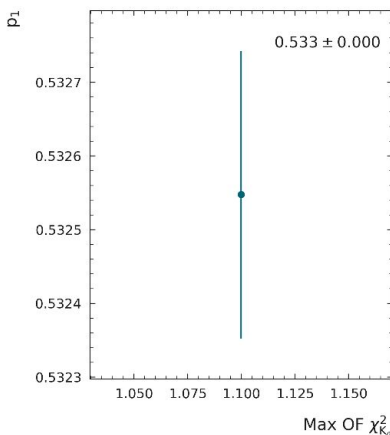
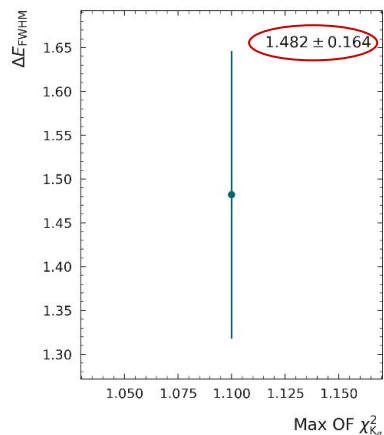
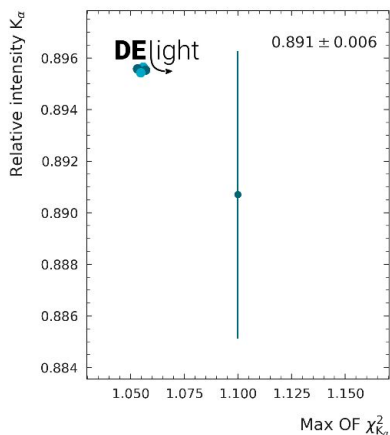
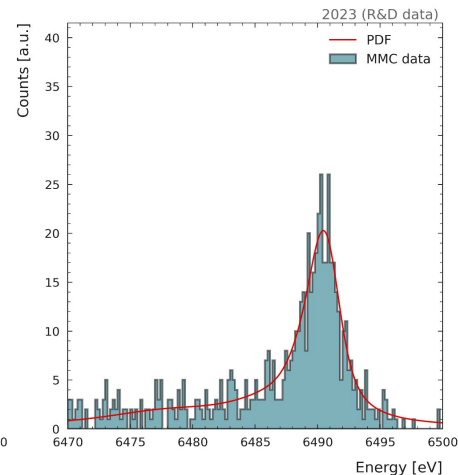
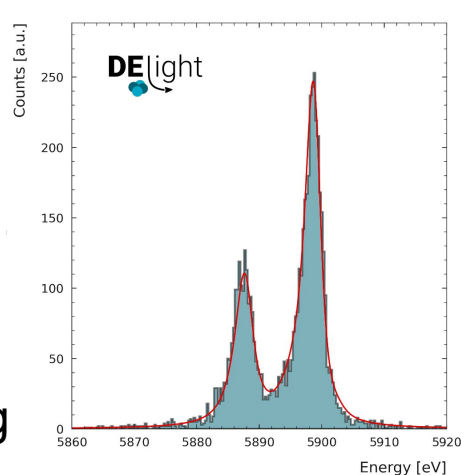


Data after Cleaning + Optimal Filtering



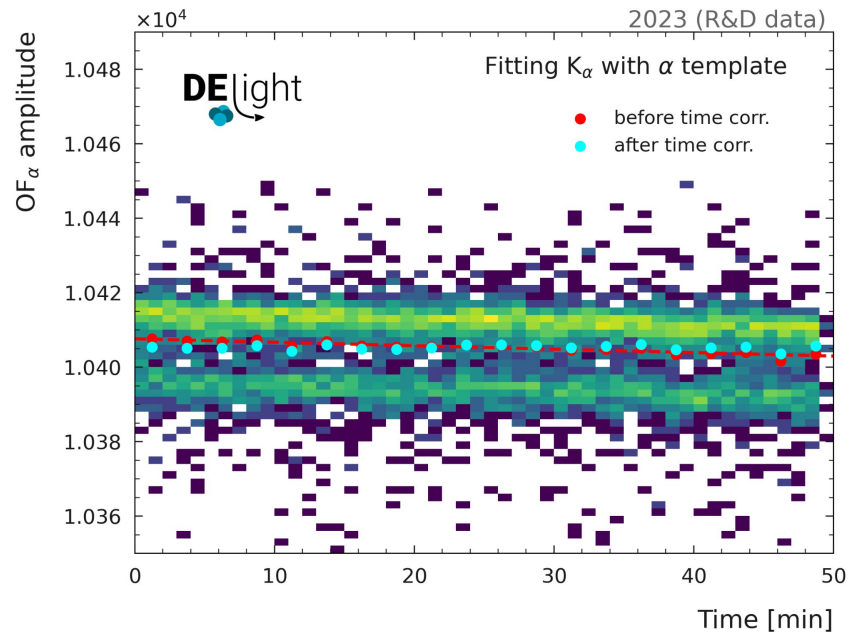
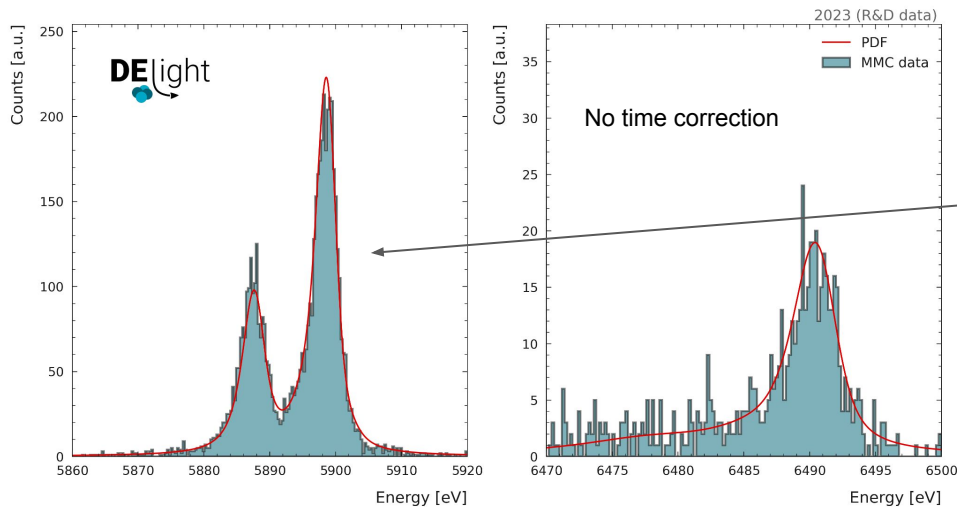
Fitting the full range

- $E = p_0 + p_1 x + p_2 x^2 + p_3 x^3$, $x = \text{ampl}$
- Resolution: 1.48 GeV
- Marginal improvement over last time (1.49 eV) coming from outlier cleaning
- Referred to as “default” hereafter



Time correction

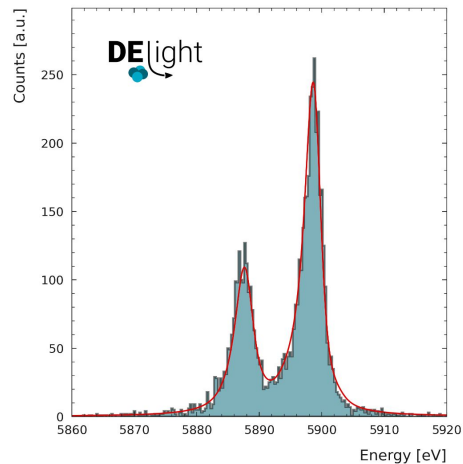
- The default has a time-dependent correction of the fitted amplitudes based on the mean of the trend
- Correction derived on K_{α} , but applied also to K_{β}



No time correction clearly
adverse effect on fit (2.18 eV)

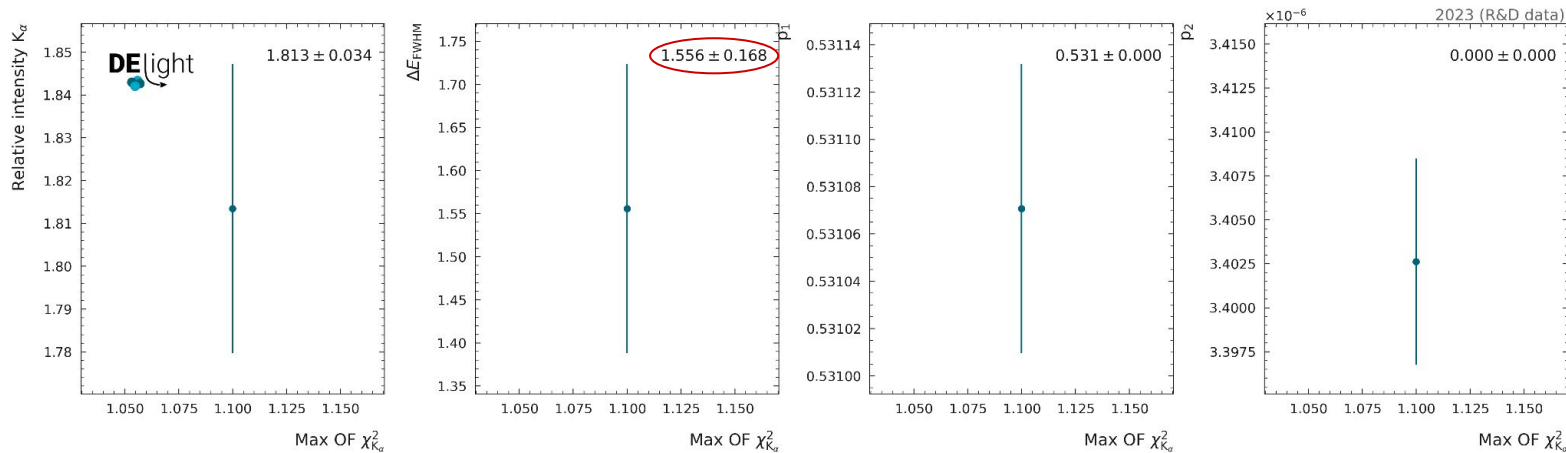
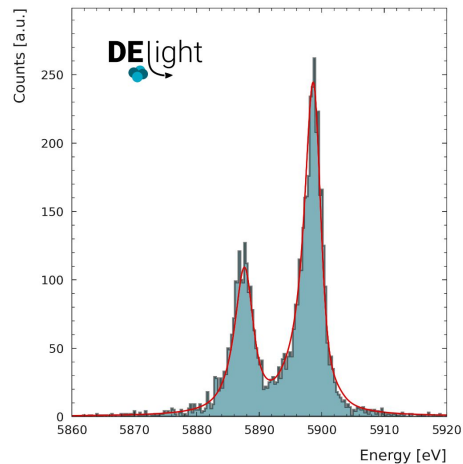
Masking the K_β line

- Only considering the K_α lines
- Idea: chemical composition of K_β source may be different and the spectra off
- Fit not converging when removing K_β term from PDF
- Used full PDF and just masked events in K_β region



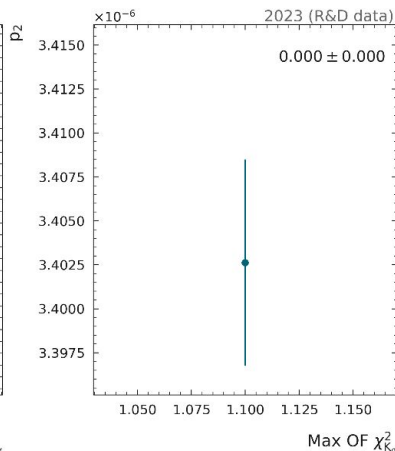
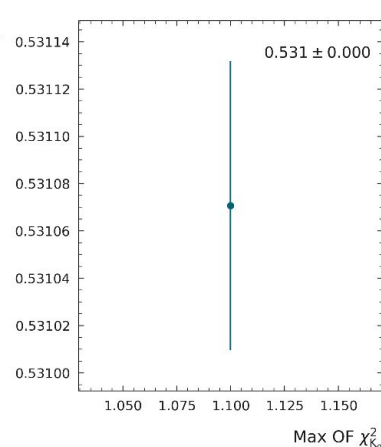
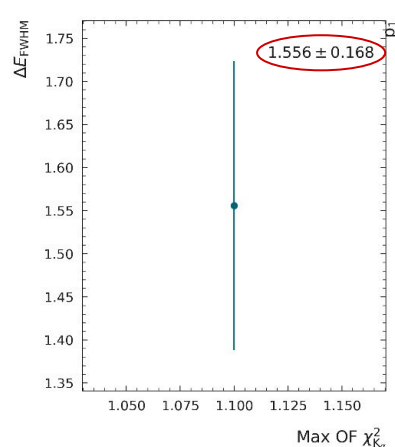
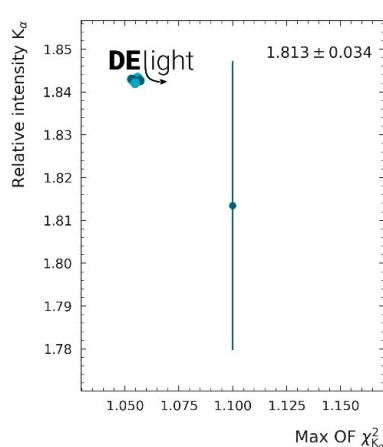
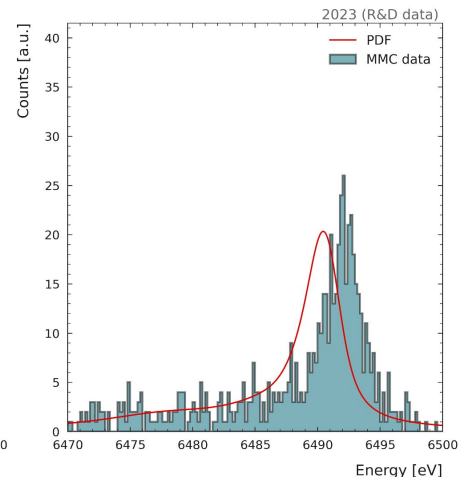
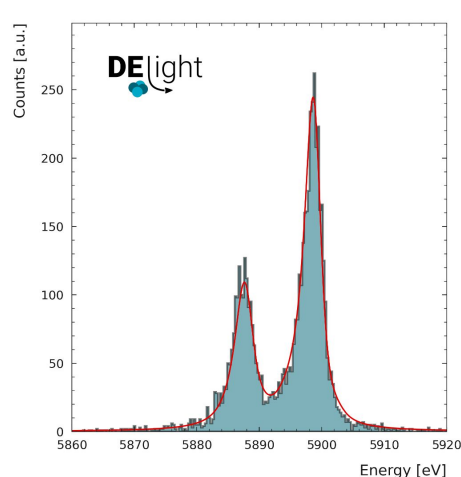
Masking the K_β line

- Degraded resolution extracted compared to fit on full range



Masking the K_β line

- Slightly degraded resolution extracted compared to fit on full range
- Clearly, estimate off for K_β as no constraints at high energies

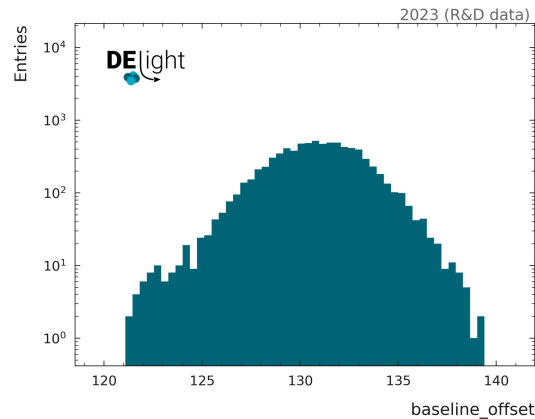


Some fits that did not converge (yet)

- Anything with RooFit
- Only considering the K_α traces + just the K_α PDF term
- Parametrizing (as polynomial) the detector resolution
 - $\text{Sigma}_0 \rightarrow \text{sigma} = \text{sigma}_0 + \text{sigma}_1 * x$
- 3rd order polynomial for $x \rightarrow E$
 - Currently a 2nd order polynomial is used

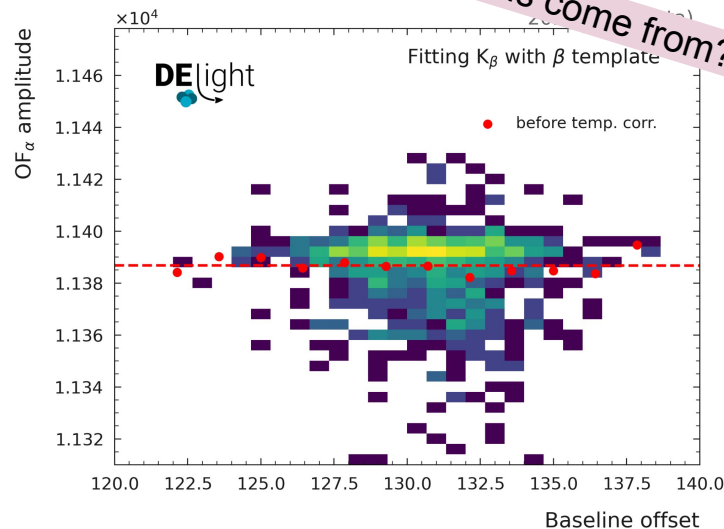
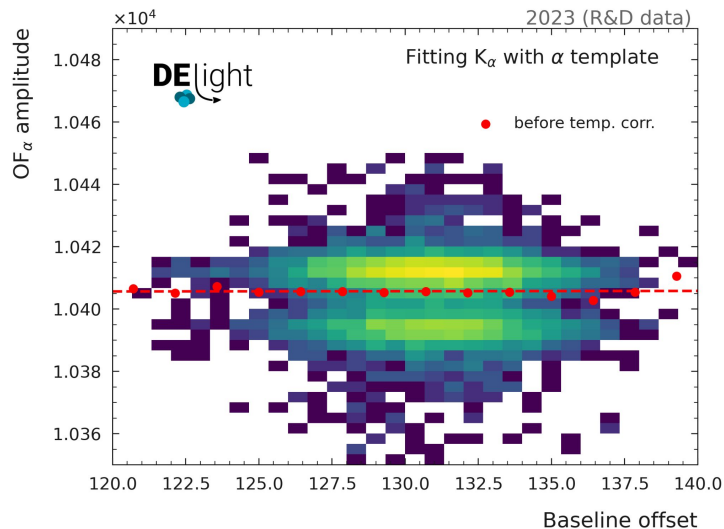
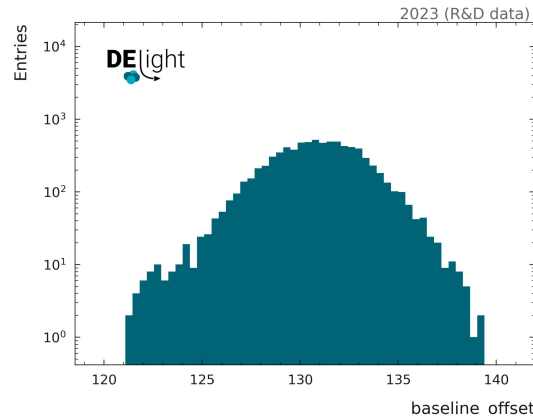
Temperature dependence

- Temperature hiding in the baseline offset
- We expect to see an amplitude dependence on the temperature, but how strong is it?



Temperature dependence

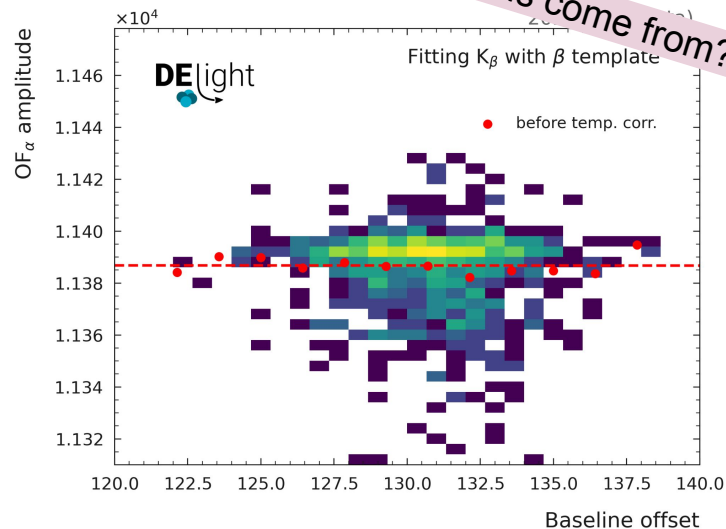
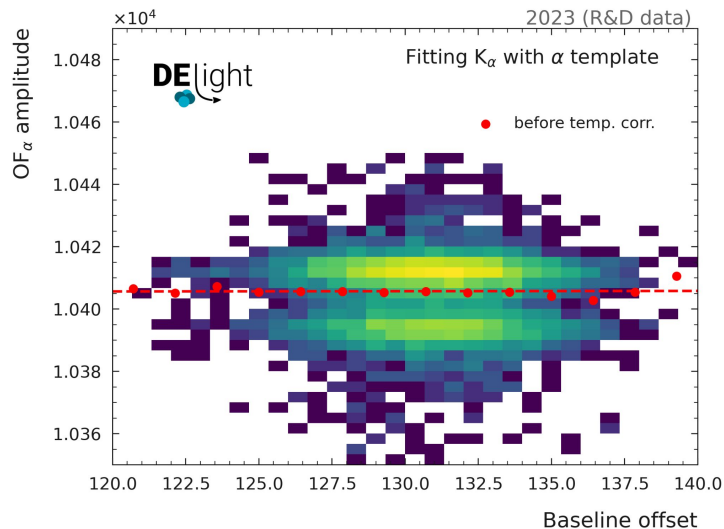
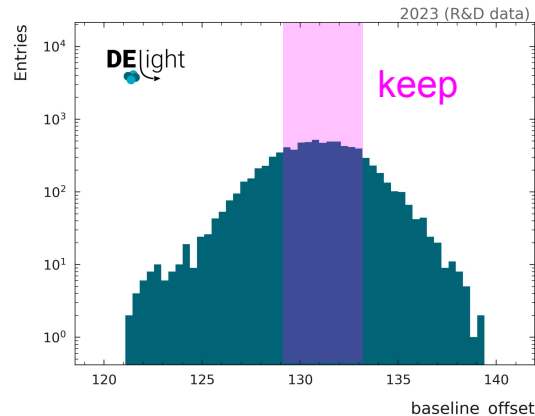
- Temperature hiding in the baseline offset
- We expect to see an amplitude dependence on the temperature, but how strong is it?
- \sim sinusoidal dependence?



Where does this come from? Ideas?

Temperature dependence

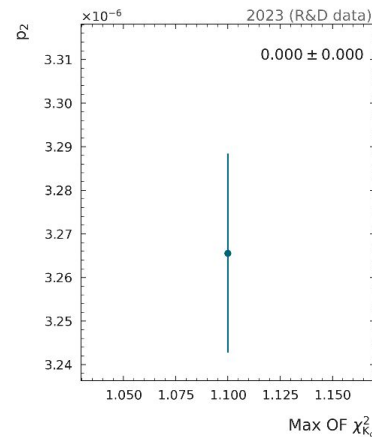
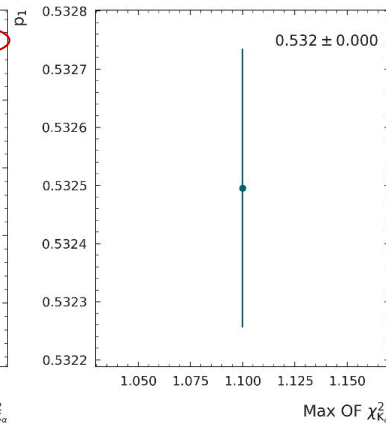
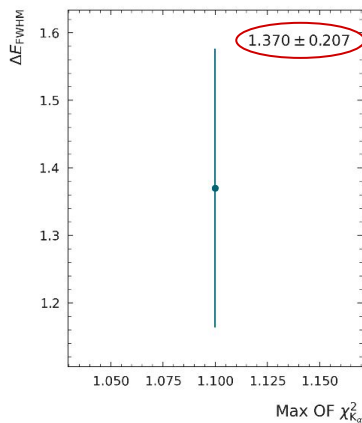
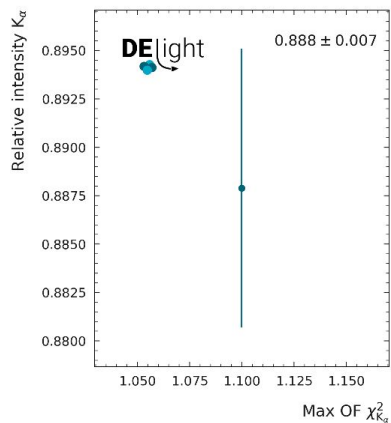
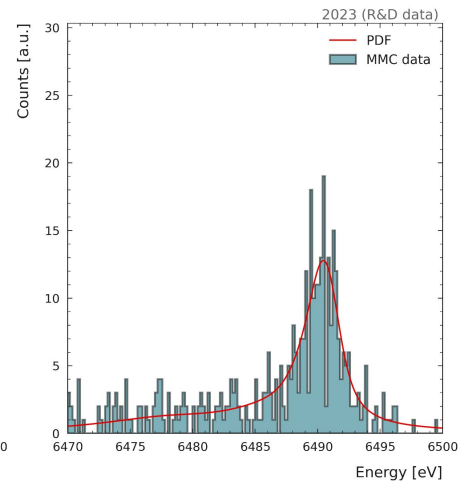
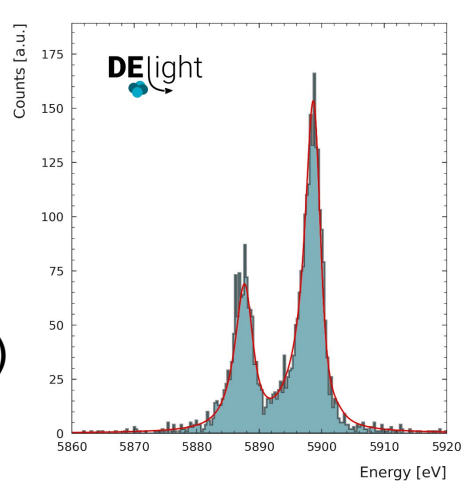
- Temperature hiding in the baseline offset
- We expect to see an amplitude dependence on the temperature, but how strong is it?
- \sim sinusoidal dependence?



Where does this come from? Ideas?

Last slide

- Removing all events with $t =]129,133[$
- Much improved resolution (within unc)
- Asks for closer inspection of temperature effects & correction



Overview

- Run 0: masking beta region
- Run 1: default
- Run 2: no time correction
- Run 3: time correction + correction offset shift
- Run 4: Beta 1 time correction for Kbeta region
- Run 5: Run 3 + Run 4
- Run 6: Run 4 + temperature correction
- Run 7: temperature correction
- Run 8: temperature selection via skim.config
- Run 9: flip sigma and aw_ij