PS3a: N-Body Simulation (static)

Due: Thursday, June 2, 11:59pm

In 1687, Sir Isaac Newton formulated the principles governing the motion of two particles under the influence of their mutual gravitational attraction in his famous $Principia\ Mathematica$. However, Newton was unable to solve the problem for three particles. Indeed, in general, solutions to systems of three or more particles must be approximated via numerical simulations. Your challenge is to write a program to simulate the motion of n particles in the plane, mutually affected by gravitational forces, and animate the results. Such methods are widely used in cosmology, semiconductors, and fluid dynamics to study complex physical systems. Scientists also apply the same techniques to other pairwise interactions including Coulombic, Biot-Savart, and van der Waals.

Average time to complete assignment: ~ 5 hours.

1 Details

For Part A of this assignment, we will create a program that loads and displays a static universe. In Part B, we will add the physics simulation and animate the display.

- Make sure to download the universe specification files and images files from nbody.zip
- You should build a command-line app which reads the universe file (e.g. planets.txt) from stdin. Name your executable NBody, so you would run it with e.g.

```
./NBody < planets.txt
```

The < planets.txt construct is known as an input redirect.

1.1 Reading in the universe

The input format is a text file that contains the information for a particular universe (in SI units). The first value is an integer n which represents the number of particles. The second value is a real number r which represents the radius of the universe, used to determine the scale of the drawing window. Finally, there are n rows and each row contains 6 values. The first two values are the x and y coordinates of the initial position; the next pair of values are the x and y components of the initial velocity; the fifth value is the mass; the last value is a string that is the name of an image used to display the particle. As an example, planets.txt contains data for our own solar system (up to Mars):

```
...mass...
...xpos...
             ...ypos...
                         ...xvel...
                                      ...yvel...
                                                               filename
1.4960e+11
            0.0000e+00
                         0.0000e+00
                                      2.9800e+04
                                                  5.9740e+24
                                                               earth.gif
2.2790e+11
            0.0000e+00
                         0.0000e+00
                                      2.4100e+04
                                                  6.4190e+23
                                                               mars.gif
                                      4.7900e+04
5.7900e+10
            0.0000e+00
                         0.0000e+00
                                                  3.3020e+23
                                                               mercury.gif
0.0000e+00
            0.0000e+00
                         0.0000e+00
                                      0.0000e+00
                                                  1.9890e+30
                                                               sun.gif
1.0820e+11
            0.0000e+00
                         0.0000e+00
                                      3.5000e+04
                                                  4.8690e+24
                                                               venus.gif
```

You should read in exactly as many rows of body information as are indicated by n, the first value in the file.

The planets.txt universe file contains the Sun and the first four planets, with the Sun at the center of the universe (x = 0, y = 0) and the four planets in order toward the right (as below). When this is working, you should be rewarded with:



For this project, you'll implement a Universe class which will contain all celestial bodies, and a CelestialBody class representing the celestial bodies.

- The class Universe should create CelestialBodys.
- The CelestialBody should have the following features:
 - It must be sf::Drawable, with a private virtual void method named draw (as required of sf::Drawable subclasses).
 - Each instance of the class should contain all properties needed for the simulation (e.g. x and y position, x and y velocity, mass, and image data).
 - It may contain a sf::Sprite object (as well as the sf::Texture object needed to hold the Sprite's image).
- You must override the input stream operator >> for both Universe and CelestialBody and use it to load parameter data into an object.
- You must override the output stream operator << for both Universe and CelestialBody and use it to write the state back out in the same format as the input.

2 Extra Credit

You can earn extra credit by drawing a background image. This image should appear behind all of the planets. If you do any of the extra credit work, make sure to describe exactly what you did in Readme-ps3a.md.

3 Development Process

There are a lot of parts to this assignment. We'd suggest the following incremental development process:

- 1. Create a bare-bones version of your Universe class.
 - (a) Override the << operator to read just the number of planets and universe radius.
 - (b) Override the >> operator to print the information back.
 - (c) Verify that both work as intended.
- 2. Create a bare-bones version of your CelestialBody class.
 - (a) Override the << operator to read the planet data.
 - (b) Override the >> operator to print the information back.
 - (c) Read one planet from the file and verify that it works as intended.

- 3. Setup the main draw loop in your main file.
 - Hint: Your class will need to know both the universe and viewport window dimensions.
 - Hint: The universe's center is (0, 0), but for SFML, (0, 0) is in the upper-left corner.
- 4. Implement the draw method in the CelestialBody class. Draw this in your loop.
- 5. Move control of the CelestialBody into the Universe. Implement the draw method in your Universe class.
- 6. Read the rest of the planets into your Universe.

4 What to turn in

Your makefile should build a program named NBody.

Submit a tarball to Blackboard containing:

- Your main file main.cpp.
- Your Universe (Universe.cpp, Universe.hpp) and CelestialBody (CelestialBody.cpp, CelestialBody.hpp) classes.
- The makefile for your project. The makefile should have targets all, NBody, lint and clean. Make sure that all prerequisites are correct.
- Your Readme-ps3a.md that includes
 - 1. Your name
 - 2. Statement of functionality of your program (e.g. fully works, partial functionality, extra credit)
 - 3. Key features or algorithms used
 - 4. Any other notes
- Any other source files that you created.
- Any images or other resources not provided (such as for the extra credit)
- A screenshot of program output

Make sure that all of your files are in a directory named ps3a before archiving it and that there are no .o or other compiled files in it.

5 Grading rubric

Feature	Points	Comment
Core Implementation	10	Full & Correct Implementation
	2	Contains all code files and builds
	1	Contains Universe and CelestialBody classes
	1	Both classes are Drawable with private draw() methods
	1	Loads universe from stdin
	2	Draws the planets to the screen correctly
	1	Overrides << and >> operators for both classes
	1	Supports arbitrary number of CelestialBody objects
	1	Scales to different universe sizes
Makefile	2	
	1	Has targets NBody, lint, all, and clean
	1	all builds NBody
Screenshot	1	
Readme	2	Complete
	1	Describes the algorithms or data structures used.
	1	Describes how the features were implemented.
Extra Credit	1	
	+1	Shows background image
Penalties		
	-2	Linting problems
	-1	Non-private fields
	-1	Submission includes . o files
	-1	Submission not in a directory named ps3a
	-10%	Each day late
Total	15	