

PHY 410 - Reference Sheet

Stirling's approximation - for very large N:

$$\log N! \approx N \log N - N$$

$$N! \approx \sqrt{2\pi N} N^N e^{-N}$$

Fractional uncertainty of \mathbb{X} is uncertainty of expected value per particle.

$$\frac{\Delta \mathbb{X}}{N} = \frac{\sqrt{\langle \mathbb{X}^2 \rangle - \langle \mathbb{X} \rangle^2}}{N}$$

Boltzmann's constant

 $k_B = 1.380649 \times 10^{-23} \mathrm{m}^2 \, \mathrm{s}^{-2} \, \mathrm{K}^{-1}$

Entropy $S = k_B \sigma$, $\sigma_{TOT} = \sigma_1 + \sigma_2$ Temperature $T = \tau/k_B$

Microcanonical Ensemble

Multiplicity function

$$g = \#$$
 of microstates, $\mathcal{P}(n) = \frac{1}{q}$

Expected value of $\mathbb X$ is the average across all microstates.

$$\langle \mathbb{X} \rangle = \sum_{n} \mathbb{X}(n) \mathcal{P}(n) = \frac{1}{g} \sum_{n} \mathbb{X}(n)$$

Entropy can be written in terms of the multiplicity function.

$$\sigma(N,T,U,V,P) \equiv \log[g(N,T,U,V,P)]$$

Binary System

A binary system is a system of N particles where each particles has two possible states. Let N_{\uparrow} is the number of particle in the up state and N_{\perp} be the number of particles in the down state.

$$g(N, N_{\uparrow}) = \frac{N!}{N_{\uparrow}!(N - N_{\uparrow})!}, \quad \sum_{N_{\uparrow}=0}^{N} g(N, N_{\uparrow}) = 2^{N}$$

The binary system can be rewritten in terms of the difference between up states and down states this is the **spin excess**.

$$2S = N_{\uparrow} - N_{\downarrow}$$

$$g(N, S) = \frac{N!}{(\frac{N}{2} + S)!(\frac{N}{2} - S)!}$$

$$\sum_{S = -\frac{N}{2}}^{N} g(N, N_{\uparrow}) = 2^{N}$$

Applying Stirling's approximation to the binary model, for large N the multiplicity function and fractional uncertainty are

$$g(N,S) \approx g(N,0)e^{-2s^2/N}$$

 $g(N,S) \approx \sqrt{\frac{2}{\pi N}} 2^N e^{-2s^2/N}$
 $\frac{\Delta S}{N} \approx \frac{1}{\sqrt{N}}$

An example of a binary system is N spin 1/2 particles in an external **magnetic field** B. The total energy U and magnetization M of the system are

$$U = \sum_{i=1}^{N} -\vec{m_i} \cdot \vec{B} = -(N_{\uparrow} - N_{\downarrow})mB = -2SmB$$

$$\begin{split} M &= 2Sm = -U/B \\ g(N,U) &= \frac{N!}{\left(\frac{N}{2} - \frac{U}{2mB}\right)!\left(\frac{N}{2} + \frac{U}{2mB}\right)!} \\ \sigma(N,S) &\approx -\left(\frac{N}{2} + S\right)\log\left(\frac{1}{2} + \frac{S}{N}\right) - \\ \left(\frac{N}{2} - S\right)\log\left(\frac{1}{2} - \frac{S}{N}\right) \\ M &= Nm\tanh(mB/\tau) \end{split}$$

Einstein Solid

An **einstein solid** is a system of N atoms where each atom is modeled as a harmonic oscillator the energy of the system is determined by the number of atoms n oscillating at frequency ω .

$$U = n\hbar\omega$$

$$g(N, n) = \frac{(n + N - 1)!}{n!(N - 1)!}$$

$$g(N, n) \approx \frac{\left(\frac{n+N}{n}\right)^n \left(\frac{n+N}{n}\right)^N}{\sqrt{2\pi n(n+N)/N}}$$

Thermal Equilibrium Temperature

$$\frac{1}{\tau} = \left(\frac{\partial \sigma}{\partial U}\right)_{N,V}$$

Thermal Equilibrium

$$\left(\frac{\partial \sigma_1}{\partial U_1}\right)_{N_1, V_1} = \left(\frac{\partial \sigma_2}{\partial U_2}\right)_{N_2, V_2}$$
$$\frac{1}{\tau_1} = \frac{1}{\tau_2}$$

2nd law of thermo - Change in entropy ≥ 0 . **Sharpness of Equilibrium** For a two binary systems, the number of states in a configuration of deviation δ from equilibrium is

$$g_1g_2 = (g_1g_2)_{max}e^{\left(-\frac{2\delta^2}{N_1} - \frac{2\delta^2}{N_2}\right)}$$

Canonical Ensemble

Partition Function - partition by energy levels for a fixed temperature

$$z = \sum_{n} e^{-\varepsilon_n/\tau}, \quad \mathcal{P}(n) = \frac{1}{z} e^{-\varepsilon_n/\tau}$$

$$z = \sum_{\alpha} g(\varepsilon_{\alpha}) e^{-\varepsilon_{\alpha}/\tau}$$
, for degeneracy $g(\varepsilon_{\alpha})$

Expected Value of X is the average across all energies (Thermal Average).

$$\langle \mathbb{X} \rangle = \sum_n \mathbb{X}(n) \mathcal{P}(n) = \frac{1}{z} \sum_n \mathbb{X}(n) e^{-\varepsilon_n/\tau}$$

Expected Energy in the canonical ensemble is

$$U = \langle \varepsilon \rangle = \frac{1}{z} \sum_{n} \varepsilon_n e^{-\varepsilon_n/\tau}$$

$$U = \langle \varepsilon \rangle = \tau^2 \frac{1}{z} \frac{\partial z}{\partial \tau} = \tau^2 \frac{\partial}{\partial \tau} \log z$$

The total partition function and expected value for N non-interacting particles is simply

$$z_N = z_1^N$$

$$\langle \mathbb{X} \rangle_N = N \langle \mathbb{X} \rangle_1 \quad \Rightarrow \quad U_N = NU_1$$

Helmholtz Free Energy

$$F = U - \tau \sigma = U - ST = -\tau \log z$$

$$dF = -\sigma d\tau - PdV$$

Entropy
$$\sigma = -\left(\frac{\partial F}{\partial \tau}\right)_V$$
, $S = k_B \sigma$

Temperature
$$\tau = \left(\frac{\partial U}{\partial \sigma}\right)_V$$

Pressure

$$P = -\left(\frac{\partial U}{\partial V}\right)_{\sigma} = \tau \left(\frac{\partial \sigma}{\partial V}\right)_{U} = -\left(\frac{\partial F}{\partial V}\right)_{\tau}$$

Energy
$$U = -\tau^2 \frac{\partial}{\partial \tau} \left(\frac{F}{\tau} \right)$$

Ideal Gas

DeBroglie Thermal Wavelength is the wavelength of the wave functions of matter at a given temperature.

$$\lambda_T = \sqrt{\frac{2\pi\hbar^2}{m\tau}}$$

Concentration of a system is the inverse of the volume

$$n = \frac{1}{V}$$

Quantum Concentration is the density of quantum state per particle. It is used to define when a system will behave classically $(n \ll n_Q)$ and when a system will be dominated by quantum effects $(n \gg n_Q)$.

$$n_Q = \frac{1}{\lambda_T^3}$$

Single Particle Ideal Gas is a system in the canonical ensemble consisting of a signle particle in a box of side lengths L. The energy levels , partition function and average energy are

$$\begin{split} E_{n_x,n_y,n_z} &= \frac{\hbar^2}{2m} \left(\frac{\pi}{L}\right)^2 (n_x^2 + n_y^2 + n_z^2) \\ z_1 &= \frac{V}{\lambda_T^3} \\ U_1 &= \frac{3}{2} \tau \\ \sigma_1 &= \log \left(\frac{V}{\lambda_T^3}\right) + \frac{3}{2} \end{split}$$

Gibbs Resolution states that for systems in the classical regime the partition function for an ideal gas with N particles is

$$z_N = \frac{1}{N!} (z_1)^N$$

$$U_N = \frac{3}{2} N \tau$$

$$\sigma_N = N \left[\log \left(\frac{V}{N \lambda_T^3} + \frac{5}{2} \right) \right]$$

N-Particle Ideal Gas - by applying Gibbs resolution and properties of expected values we can find the classical ideal gas results

$$\begin{aligned} PV &= N\tau \\ U &= \frac{3}{2}N\tau \\ \sigma &= N \left[\log \left(\frac{V}{N\lambda_T^3} \right) + \frac{5}{2} \right] \end{aligned}$$

Thermal Radiation

Single Frequency Photon Gas is a system in the canonical ensemble that considers photons of a specific frequency ω .

$$\varepsilon_s = s\hbar\omega, \quad s = 0, 1, 2, 3, \dots$$

$$z = \sum_{s=0}^{\infty} e^{-s\hbar\omega/\tau} = \frac{1}{1 - e^{-\hbar\omega/\tau}}$$

$$\mathcal{P}(s) = \frac{e^{-s\hbar\omega/\tau}}{z}$$

$$\langle s \rangle = \frac{1}{z} \sum_{s=0}^{\infty} s e^{-s\hbar\omega/\tau} = \frac{1}{e^{\hbar\omega/\tau} - 1}$$

Photon Gas is an expansion of the single frequency photon gas that considers all the possible cavity modes. The modes are 2 fold degenerate for the 2 independent polarizations.

$$\omega_n = \frac{c\pi}{L} \sqrt{n_x^2 + n_y^2 + n_z^2} = \frac{c\pi n}{L}$$

$$U = \langle \epsilon \rangle = 2 \sum_{n} \frac{\hbar \omega_n}{e^{\hbar \omega_n / \tau} - 1} = \frac{\pi^2 V}{15(\hbar c)^3} \tau^4$$

Stefan-Boltzmann Law

$$\frac{U}{V} = \frac{\hbar}{\pi^2 c^3} \int_0^\infty \frac{\omega^3}{e^{\hbar \omega / \tau} - 1} d\omega = \frac{\pi^2}{15(\hbar c)^3} \tau^4$$

Spectral Density Function

$$\frac{\partial}{\partial \omega} \frac{U}{V} = \frac{\hbar}{\pi^2 c^3} \frac{\omega^3}{e^{\hbar \omega / \tau} - 1}$$

Flux Density(σ_B =Stefan-Boltzmann constant)

$$J_{\mu} = \frac{1}{4} \frac{cU}{V} = \sigma_B \tau^4 = \frac{\pi^2}{60(\hbar c)^3} \tau^4$$

Phonons in a Solid (Debye Model)

Phonons in a solid is a system in the canonical ensemble that is very similar to thermal radiation except there is 3 fold degeneracy from 3 polarizations of phonons and an upper cutoff frequency ω_D due to the separation distance between atoms.

$$\omega_n = \frac{\pi c_S}{L} \sqrt{n_x^2 + n_y^2 + n_z^2} = \frac{\pi c_s}{L} n$$

Debye cutoff frequency

$$\omega_D = c_S \left(\frac{6\pi^2 N}{V}\right)^{1/3}, \quad \omega_D = \frac{\pi c_S}{L} n_D$$

Grand Canonical Ensemble Chemical Potential

$$\mu = \left(\frac{\partial F}{\partial N}\right)_{\tau,V}$$

$$\mu = \tau \log \left(\frac{N\lambda_T^3}{V}\right) = \tau \log \left(\frac{n}{n_Q}\right)$$

$$\mu = \left(\frac{\partial U}{\partial N}\right)_{\sigma,V} = -\tau \left(\frac{\partial \sigma}{\partial N}\right)_{U,V}$$

Grand Partition Function - partition by energy levels for a fixed temperature and all possible values of N

$$\mathbf{z} = \sum_{N} \sum_{n(N)} e^{-(\varepsilon_n^N - \mu N)/\tau}$$

$$\mathcal{P}(N, \varepsilon_n) = \frac{1}{\mathbf{z}} e^{-(\varepsilon_n^N - \mu N)/\tau}$$

Fugacity

$$\lambda = e^{\mu/\tau}$$

$$\mathbf{z} = \sum_{N} \lambda^{N} \sum_{s(N)} e^{-\varepsilon_{s}^{N}/\tau} = \sum_{N} \lambda^{N} z_{N}$$

Expected Value of $\mathbb X$ is the average across all energies (Diffusive Average).

$$\langle \mathbb{X} \rangle = \frac{1}{\mathsf{z}_{\mathsf{L}}} \sum_{N} \sum_{s} \mathbb{X}(N, s) e^{-(\varepsilon_{s}^{N} - \mu N) / \tau}$$

Expected Number of Particles in the grand canonical ensemble is

$$N = \langle N \rangle = \tau \frac{\partial}{\partial \mu} \log z = \lambda \frac{\partial}{\partial \lambda} \log z$$

Expected Energy in the grand canonical

$$U = \langle \varepsilon \rangle = \frac{1}{\mathbf{Z}} \sum_{N} \sum_{n(N)} \varepsilon_{n}^{N} e^{-(\varepsilon_{n}^{N} - \mu N)/\tau}$$

$$U = \langle \varepsilon \rangle = \tau^2 \left(\frac{\partial}{\partial \tau} \log \mathbf{z} \right)$$

Concentration and DeBroglie Wavelength

$$n=\frac{N}{V}, \quad n_Q=\frac{1}{\lambda_T^3}, \quad \lambda_T=\sqrt{\frac{2\pi\hbar^2}{m\tau}}$$

Grand Potential

$$\Omega = U - \sigma \tau - \mu N$$

$$\Omega = -\tau \log z$$

$$\sigma = \left(\frac{-\partial \Omega}{\partial \tau}\right)_{V,\mu} P = \left(\frac{-\partial \Omega}{\partial V}\right)_{\tau,\mu} N = \left(\frac{-\partial \Omega}{\partial \mu}\right)_{\tau,V} Heat Capacity measures the change in heat System of Non-interacting Particles energy per unit temperature$$

System of Non-interacting Particles

The grand partition function for a system with M energy states where n_{α} is the number of particles occupying a state is

$$\mathbf{z} = \prod_{\alpha=1}^{M} \mathbf{z}_{\alpha}, \quad \mathbf{z}_{\alpha} = \sum_{n_{\alpha}} e^{-n_{\alpha} (\varepsilon_{\alpha} - \mu) / \tau}$$

$$U = \sum_{\alpha=1}^{M} \varepsilon_{\alpha} f(\varepsilon_{\alpha}), \quad N = \sum_{\alpha=1}^{M} f(\varepsilon_{\alpha})$$

Fermions

$$\begin{aligned} n_{\alpha} &= 0, 1 \\ \mathbf{z}_{\alpha} &= 1 + e^{-(\varepsilon_{\alpha} - \mu)/\tau} = 1 + \lambda e^{-\varepsilon_{\alpha}/\tau} \end{aligned}$$

Fermi-Dirac Distribution is the expected number of a particles in a particular energy ε_{α} .

$$\langle n_{\alpha} \rangle = f(\varepsilon_{\alpha}) = \frac{1}{e^{(\varepsilon_{\alpha} - \mu)/\tau} + 1} = \frac{1}{\lambda^{-1} e^{\varepsilon_{\alpha}/\tau} + 1}$$

For
$$\tau \to 0$$
: $f(\varepsilon_{\alpha}) = \theta(\varepsilon_{\alpha} - \mu)$

Bosons (Bonsons)

$$\mathbf{z}_{\alpha} = \frac{1}{1 - e^{-(\varepsilon_{\alpha} - \mu)/\tau}} = \frac{1}{1 - \lambda e^{-\varepsilon_{\alpha}/\tau}}$$

Boson Distribution is the expected number of a particles in a particular energy ε_{α} .

$$\langle n_{\alpha} \rangle = f(\varepsilon_{\alpha}) = \frac{1}{e^{(\varepsilon_{\alpha} - \mu)/\tau} - 1} = \frac{1}{\lambda^{-1} e^{\varepsilon_{\alpha}/\tau} - 1}$$

Ideal Gas

Both fermions and bosons behave identically at the classical limit $\varepsilon_{\alpha} - \mu >> \tau$.

$$\varepsilon_n = \frac{\hbar^2 \pi^2}{2mL^2} n^2 = \frac{\hbar^2 \pi^2}{2mL^2} (n_x^2 + n_y^2 + n_z^2)$$

$$\frac{\pi \hbar c}{2mL^2} \frac{\pi \hbar c}{2mL^2} \left(\frac{\pi \hbar c}{2mL^2} + \frac{\pi \hbar c}{2mL^2} \right)$$

$$\varepsilon_n = pc = \frac{\pi \hbar c}{L} n = \frac{\pi \hbar c}{L} \sqrt{n_x^2 + n_y^2 + n_z^2}$$
$$\langle n_\alpha \rangle = f(\varepsilon_\alpha) = e^{-(\varepsilon_\alpha - \mu)/\tau}$$

$$\mathbf{z} = \sum_{N} \lambda^{N} z_{N} = \sum_{N} \lambda^{N} \frac{1}{N!} z_{1}^{N} = e^{\lambda z_{1}}$$

$$\lambda = \frac{n}{n_Q}, \quad PV = N\tau, \quad U = \frac{3}{2}N\tau, \quad \mu = \tau\log\frac{n}{n_Q}$$

$$\sigma = N \left[\log \frac{n_Q}{n} + \frac{5}{2} \right], \quad F = N \tau \left[\log \frac{n}{n_Q} - 1 \right]$$

energy per unit temperature

$$C_P > C_V, \quad C_V = \left(\frac{\partial U}{\partial T}\right)_V = \tau \left(\frac{\partial \sigma}{\partial T}\right)_V$$

$$C_P = \left(\frac{\partial U}{\partial T}\right)_P + P\left(\frac{\partial V}{\partial T}\right)_P = \tau \left(\frac{\partial \sigma}{\partial T}\right)_P$$

Monoatmc gas $C_V = \frac{3}{2}Nk_B$, $C_P = \frac{5}{2}Nk_B$ Isothermal Expansion $\sigma_f - \sigma_i = N \log \frac{V_f}{V_c}$ $Q = N\tau \log \frac{V_f}{V_f}$

Isoentropic Expansion $\frac{\tau_f}{\tau_i} = \left(\frac{V_i}{V_s}\right)^{2/3}$

Internal Excitations

Expansion of the ideal gas to take into account the additional energy states from internal

$$z_{int} = \sum_{\alpha} e^{-\varepsilon_{\alpha}/\tau}, z = 1 + \lambda z_{int} e^{-\varepsilon_{n}/\tau}$$

Internal Excitation Corrections

$$\lambda = \frac{n}{n_Q z_{int}}, \mu = \tau \left(\log \frac{n}{n_Q} - \log z_{int} \right)$$

$$F = N\tau \left[\log \frac{n}{n_Q} - 1 \right] - N\tau \log z_{int}$$

$$\sigma = N \left[\log \frac{n}{n_Q} + \frac{5}{2} \right] - \left(\frac{\partial F_{int}}{\partial \tau} \right)_{V}$$

Fermi Gas

DOG (bork)







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