

PHY 410 - Reference Sheet

Stirling's approximation - for very large N the factorial can be very accurately approximated with the following

$$\log N! \approx N \log N - N$$
$$N! \approx \sqrt{2\pi N} N^N e^{-N}$$

Fractional uncertainty of X is uncertainty of expected value per particle.

$$\frac{\Delta \mathbb{X}}{N} = \frac{\sqrt{\langle \mathbb{X}^2 \rangle - \langle \mathbb{X} \rangle^2}}{N}$$

Boltzmann's constant $k_B=1.380649\times 10^{-23}\mathrm{m}^2\,\mathrm{s}^{-2}\,\mathrm{K}^{-1}$ Entropy $S=k_B\sigma,~\sigma_{TOT}=\sigma_1+\sigma_2$ Temperature $T=\tau/k_B$

Microcanonical Ensemble Multiplicity function

$$g = \#$$
 of microstates, $\mathcal{P}(n) = \frac{1}{g}$

Expected value of $\mathbb X$ is the average across all microstates.

$$\langle \mathbb{X} \rangle = \sum_{n} \mathbb{X}(n) \mathcal{P}(n) = \frac{1}{g} \sum_{n} \mathbb{X}(n)$$

Entropy can be written in terms of the multiplicity function.

$$\sigma(N, T, U, V, P) \equiv \log[g(N, T, U, V, P)]$$

Binary System

A binary system is a system of N particles where each particles has two possible states. Let N_{\uparrow} is the number of particle in the up state and N_{\downarrow} be the number of particles in the down state.

$$g(N, N_{\uparrow}) = \frac{N!}{N_{\uparrow}!(N - N_{\uparrow})!}, \quad \sum_{N_{\star}=0}^{N} g(N, N_{\uparrow}) = 2^{N}$$

The binary system can be rewritten in terms of the difference between up states and down states this is the **spin excess**.

$$2S = N_{\uparrow} - N_{\downarrow}$$

$$g(N, S) = \frac{N!}{(\frac{N}{2} + S)!(\frac{N}{2} - S)!}$$

$$\sum_{S = -\frac{N}{2}} g(N, N_{\uparrow}) = 2^{N}$$

Applying Stirling's approximation to the binary model, for large N the multiplicity function and fractional uncertainty are

$$g(N,S) \approx g(N,0)e^{-2s^2/N}$$

 $g(N,S) \approx \sqrt{\frac{2}{\pi N}} 2^N e^{-2s^2/N}$
 $\frac{\Delta S}{N} \approx \frac{1}{\sqrt{N}}$

An example of a binary system is N spin 1/2 particles in an external **magnetic field** B. The total energy U and magnetization M of the system are

$$U = \sum_{i=1}^{N} -\vec{m_i} \cdot \vec{B} = -(N_{\uparrow} - N_{\downarrow})mB = -2SmB$$

$$M = 2Sm = -U/B$$

$$g(N,U) = \frac{N!}{(\frac{N}{2} - \frac{U}{2mB})!(\frac{N}{2} + \frac{U}{2mB})!}$$

Einstein Solid

An **einstein solid** is a system of N atoms where each atom is modeled as a harmonic oscillator the total energy of the system is determined by the number of atoms n oscillating at frequency

$$U = n\hbar\omega$$

$$g(N, n) = \frac{(n+N-1)!}{n!(N-1)!}$$

Thermal Equilibrium Temperature

$$\frac{1}{\tau} = \left(\frac{\partial \sigma}{\partial U}\right)_{N,V}$$

Thermal Equilibrium

$$\left(\frac{\partial \sigma_1}{\partial U_1} \right)_{N_1, V_1} = \left(\frac{\partial \sigma_2}{\partial U_2} \right)_{N_2, V_2}$$

$$\frac{1}{\tau_1} = \frac{1}{\tau_2}$$

2nd law of thermo - Change in entropy ≥ 0 . Sharpness of Equilibrium For a two binary systems, the number of states in a configuration of deviation δ from equilibrium is

$$g_1 g_2 = (g_1 g_2)_{max} e^{\left(-\frac{2\delta^2}{N_1} - \frac{2\delta^2}{N_2}\right)}$$

Canonical Ensemble

 $\begin{tabular}{ll} \textbf{Partition Function} & - \text{ partition by energy levels} \\ \text{for a fixed temperature} \\ \end{tabular}$

$$z = \sum_{n} e^{-\varepsilon_n/\tau}, \quad \mathcal{P}(n) = \frac{1}{z} e^{-\varepsilon_n/\tau}$$

$$z = \sum_{\alpha} g(\varepsilon_{\alpha}) e^{-\varepsilon_{\alpha}/\tau}$$
, for degeneracy $g(\varepsilon_{\alpha})$

Expected Value of X is the average across all energies (Thermal Average).

$$\langle \mathbb{X} \rangle = \sum_n \mathbb{X}(n) \mathcal{P}(n) = \frac{1}{z} \sum_n \mathbb{X}(n) e^{-\varepsilon_n/\tau}$$

Expected Energy in the canonical ensemble is

$$U = \langle \varepsilon \rangle = \frac{1}{z} \sum_{n} \varepsilon_n e^{-\varepsilon_n/\tau}$$

$$U = \langle \varepsilon \rangle = \tau^2 \frac{1}{z} \frac{\partial z}{\partial \tau} = \tau^2 \frac{\partial}{\partial \tau} \log z$$

The total partition function and expected energy for N non-interacting particles is simply

$$z_N = z_1^N$$

$$U_N = \langle \varepsilon \rangle_N = NU_1 = N \langle \varepsilon \rangle_1$$

(this also applies for expected value of any X)

Theromodynamic Relations 1st Law of Thermo

$$dU = dQ + dW = \tau d\sigma - P dV$$

$$d\sigma = \frac{1}{\tau} dU + \frac{P}{\tau} dV$$

Temperature $\tau = \left(\frac{\partial U}{\partial \sigma}\right)_{V}$

Quasi-static Compression Equilibrium the equilibrium condition for quasi-static compression is

$$\left(\frac{\partial U_1}{\partial V_1}\right)_{\sigma_1} = \left(\frac{\partial U_2}{\partial V_2}\right)_{\sigma_2}$$

Helmholtz Free Energy

$$F = U - \tau \sigma = U - ST = -\tau \log z$$
$$dF = -\sigma d\tau - PdV$$

Entropy $\sigma = -\left(\frac{\partial F}{\partial \tau}\right)_V$, $S = k_B \sigma$

$$P = -\left(\frac{\partial U}{\partial V}\right)_{\sigma} = \tau \left(\frac{\partial \sigma}{\partial V}\right)_{U} = -\left(\frac{\partial F}{\partial V}\right)_{\tau}$$

Energy

$$U = -\tau^2 \frac{\partial}{\partial \tau} \left(\frac{F}{\tau} \right)$$

Ideal Gas

DeBroglie Thermal Wavelength is the wavelength of the wave functions of matter at a given temperature.

$$\lambda_T = \sqrt{\frac{2\pi\hbar^2}{m\tau}}$$

 $\begin{tabular}{ll} \textbf{Concentration} & \text{of a system is the inverse of the volume} \\ \end{tabular}$

$$n = \frac{1}{V}$$

Quantum Concentration is the density of quantum state per particle. It is used to define when a system will behave classically $(n \ll n_Q)$ and when a system will be dominated by quantum effects $(n \gg n_Q)$.

$$n_Q = \frac{1}{\lambda_T^3}$$

Single Particle Ideal Gas is a system in the canonical ensemble consisting of a signle particle in a box of side lengths L. The energy levels, partition function and average energy are

$$E_{n_x,n_y,n_z} = \frac{\hbar^2}{2m} \left(\frac{\pi}{L}\right)^2 (n_x^2 + n_y^2 + n_z^2)$$

$$z_1 = \frac{V}{\lambda_T^3}$$

$$U_1 = \frac{3}{2}\tau$$

$$\sigma_1 = \log\left(\frac{V}{\lambda_T^3}\right) + \frac{3}{2}$$

Gibbs Resolution states that for systems in the classical regime the partition function for an ideal gas with N particles is

$$z_N = \frac{1}{N!} (z_1)^N$$

$$U_N = \frac{3}{2} N \tau$$

$$\sigma_N = N \left[\log \left(\frac{V}{N \lambda_T^3} + \frac{5}{2} \right) \right]$$

N-Particle Ideal Gas - by applying Gibbs resolution and properties of expected values we can find the classical ideal gas results

$$\begin{aligned} PV &= N\tau \\ U &= \frac{3}{2}N\tau \\ \sigma &= N \left[\log \left(\frac{V}{N\lambda_T^3} \right) + \frac{5}{2} \right] \end{aligned}$$

DOG (bork)

