Before we continue with the fields of wrowing point charges, let's take stock & analyze the structure of Jefimento's eq.

Story so far!

Maxwell eq

Wave eq

$$II\vec{E} = -P_A/E_0 - \partial_{\xi} M_0 \vec{j}$$
 $II\vec{B} = M_0 \vec{J}^{**}$ 

In Lorenzz garge

Jefinenko eq

 $\vec{E} = \vec{I}^{-1} (-V_A/E_0 - \partial_{\xi} M_0 \vec{j})$ 
 $\vec{B} = \vec{I}^{-1} (M_0 \vec{j})$ 

or

 $\vec{A} = \vec{I}^{-1} (M_0 \vec{j})$ 

"Retarded potential gravity in the superposition principle.

Here 
$$\prod_{\hat{x}\hat{x}}^{-1} f(x^n) = \int d^3x' dt' \, G(x^n, x^{n'}) \, f(x^{n'})$$

$$G = \frac{\sigma(t - t' - t'c)}{4\pi r} \quad r = |\vec{x} - \vec{x}'|$$

Here we focussed on the sourced wave eq. The vacuum wave eq is  $\prod Y = \left(\frac{1}{c^2} \partial_t^2 - \nabla^2\right) + 20$ 

Which has place wave solutions  $f = f = f = \frac{\pm i(\vec{k} \cdot \vec{x} - \omega t)}{(-\omega^2/c^2 + |\vec{k}|^2)^4}$ 

- o propagation speed = C= / Tuoe.
- · propagation direction k

The finite propagation speed is reflected in Jefimenko's eg by the retarded time! tret = t-1/c eg  $E(x,t) = \frac{1}{4\pi\epsilon_0} \int d^2x' \left[ \frac{\hat{r}}{r^2} p(\bar{x}', t'_{nt}) + ... \right]$ E, B depend on all the sources p, J throughout all of space at a very specific set of thus determined by cowsality. observation points
where we want to
where we want to
have E(x,t), B(x,t) eg a bunch of electrons: Spacetine diagram diagram diagram

The fields at x,t depend on what the sources were doing, when they were on the past lightcome - only these points propagate information to x,t at the speed of light.

For static sources, Jefinanko reduces to Coulomb and Bist-Savart!

$$\int (x',t') \rightarrow \rho(x')$$

$$\int (x',t') \rightarrow \int (x')$$

$$\vec{E}(\vec{r},t) = \frac{1}{4\pi} \int \vec{J}_{x'} \frac{\hat{r}_{x}}{r^{2}} \rho(x')$$

$$\vec{B}(\vec{x},t) = -\frac{\hbar}{4\pi} \int \vec{J}_{x'} \frac{\hat{r}_{x}}{r^{2}} \frac{\hat{r}_{x}}{r^{2}}$$

Jebinesho goversizes this to arbitrary time-dependent sources. The really interesting thing about the new terms is they tall off like 1/r instead of 1/r. In brief, wave energy density or (Amplitude).

At long distances, an amplitude that goes like 1/r can transmit finite energy (1/1) (1/1) or finite energy

We will discuss in more defail.

Back to Jefinenko. Let's evaluate on point charge sources!  $p = e^{5/3}(\vec{x}' - \vec{s}(t'))$   $\vec{j} = e^{\sqrt{3}}(\vec{x}' - \vec{s}(t'))$ 

Now plug into 
$$\vec{E} \nmid \vec{B}$$
 (Jetanto):

 $\vec{E}(\vec{x},t) = \frac{1}{4\pi\epsilon_0} \int_0^3 x' \left\{ \frac{\hat{r}}{r^2} \ln t + \frac{\hat{r}}{\epsilon r} \frac{\partial \epsilon}{\partial \epsilon} \ln t - \frac{1}{\epsilon r} \frac{\partial \epsilon}{\partial \epsilon} \ln t \right\}$ 

Let's start with the first term.

 $\frac{1}{4\pi\epsilon_0} \int_0^3 x' \frac{\hat{r}}{r^2} e^{-\delta G}(\vec{x}' - \vec{s}(t-\frac{|\vec{x}-\vec{x}'|}{\epsilon}))$ 

The  $\vec{x}'$  dependence in  $\vec{s}$  means  $\int_0^3 x' \frac{\partial G}{\partial \epsilon} (...)$  is not so shaple.

Recall moth:  $\int_0^3 x h(\vec{x}) \frac{\partial G}{\partial \epsilon} (\vec{f}(\vec{x})) = \frac{2}{\epsilon}$ 

change vars  $\vec{k} \rightarrow \vec{f}$ :

 $= \int_0^3 f \left[ \frac{\partial f}{\partial x_0} \right]^{-1} h(\vec{k}) \frac{\partial G}{\partial \epsilon} (\vec{f})$ 

Jacobian deformant

 $= \sum_{\substack{b \in \mathcal{N} \\ \text{of}}} \frac{h(\vec{x})}{\hat{r}} \int_0^3 h(\vec{k}) \frac{\partial G}{\partial \epsilon} (\vec{f})$ 

easy to mes the Jacobian of the root care to  $|\vec{f}|$ .

The Jacobian is 
$$J = \det\left(\frac{\partial}{\partial x^{ij}}(x'^{i} - s'(t - \frac{1x - x'}{x^{i}}))\right)$$

The first term is easy!  $\frac{\partial x'^{i}}{\partial x'^{i}} = 5$ :

Second term!  $\frac{\partial}{\partial x^{i}} s'(t - \frac{1x - x'}{x^{i}}) = -\frac{1}{C} \frac{\partial s'(t')}{\partial t'} \frac{\partial [x - x']}{\partial x^{i}}$ 

but  $\frac{\partial [x - x']}{\partial x^{i}} = \frac{\partial [x - x']}{\partial x^{i}} = \frac{1}{[x - x']}(x^{i} - x^{i}) = (r^{i})^{i}$ 

So  $J = \det\left(\frac{\partial}{\partial x^{i}} - \frac{1}{C} \frac{\partial s'(t')}{\partial t'} |_{t = t - \gamma_{c}}\right)$ 

When  $\det\left(\frac{\partial}{\partial x^{i}} - \frac{\partial}{\partial x^{i}} - \frac{\partial}{\partial x^{i}} \frac{\partial}{\partial x^{i}}\right) = [-A \cdot B]$ 

So  $J = [-B \cdot R]$  where  $B = [-A \cdot B]$ 

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The last fame velocity of the particle at the retarded time.

So the 1st term in  $E$  is  $\frac{1}{\sqrt{11}} = \frac{1}{\sqrt{11}} = \frac{1$ 

Note that this definition of the setwaded time is quite implient. We are supposed to evalvate  $\vec{R}(t') = \vec{x} - \vec{s}(t')$  and  $\vec{B}(t') = \frac{\partial \vec{s}(t')}{\partial t'}$  at an ealier time t' = t - R(t')/c, i.e. when the distance from the charge to the observation point  $\vec{x}$  was c(t-t').

R(t')  $\overline{S}(t')$  c(t-t') = R(t')  $\overline{R}(t')$ 

the signal emitted @ t' is arriving now at t.

Continuing with the other terms in E. We still have to evaluate

$$p_{y} = p(x',t') = e^{s(s)}(\vec{x}'-\vec{s}(t'))$$
  
 $\vec{j}(x',t') = e^{\vec{p}(t')}c^{s(s)}(\vec{x}'-\vec{s}(t'))$ 

So these terms are

$$\frac{e}{4\pi\epsilon_0} \partial_{\epsilon} \left[ \int d^3x' \left( \frac{\hat{r}}{cr} - \frac{\vec{B}(t - 7/c)}{cr} \right) \delta^{(3)}(\vec{x}' - \vec{S}(t - 7/c)) \right]$$

The integrals can be done exactly as before, giving

$$\vec{R}(t') = \vec{\chi} - \vec{S}(t')$$

$$K(t') = 1 - \vec{\beta} \cdot \hat{R}$$

$$\vec{\beta}(t') = \frac{1}{c} \underbrace{\partial \vec{S}(t')}_{\partial t'}$$

$$t'_{ret} = t - R(t'_{ret})/c$$

In total, for a point charge,

$$\vec{E}_{(k,6)}^{-} = \frac{e}{4\pi\epsilon_0} \left\{ \left[ \frac{\hat{R}}{KR^2} \right]_{\text{ret}} + \frac{1}{c} \frac{\partial}{\partial t} \left[ \frac{\hat{R} - \vec{p}}{KR} \right]_{\text{ret}} \right\}$$

5 milarly,

It is possible to evalute the derivatives, giving a useful form in which E&B are explicitly just bunchlous of P, P, and B. The algebra is a little lengthy, so we won't work it all out. Here is the result:

$$\vec{E} = \frac{c}{4\pi\epsilon_0} \left\{ \begin{bmatrix} \hat{\mathbf{n}} - \hat{\mathbf{p}} \\ Y^2 (1 - \hat{\mathbf{p}} \cdot \hat{\mathbf{n}})^3 R^2 \end{bmatrix}_{\text{ret}} + c \begin{bmatrix} \hat{\mathbf{n}} \times [(\hat{\mathbf{n}} - \hat{\mathbf{p}}) \times \hat{\mathbf{p}}] \\ (1 - \hat{\mathbf{p}} \cdot \hat{\mathbf{n}})^3 R \end{bmatrix}_{\text{ret}} \right\}$$

$$\vec{B} = \frac{1}{c} \begin{bmatrix} \hat{\mathbf{n}} \times \hat{\mathbf{E}} \end{bmatrix}_{\text{ret}} \qquad (Y = \frac{1}{1 - \mathbf{p}^2})$$
These are the fields, at  $\hat{\mathbf{x}} \neq \hat{\mathbf{t}}$ , generated by a point charge that had velocity  $c\vec{\mathbf{p}}(t')$  and accef.  $c\vec{\mathbf{p}}(t')$  at  $t'_{\text{ret}} = t - \frac{1}{X} - \hat{\mathbf{s}}(t'_{\text{ret}})|$ 

$$c$$

$$(Y = \frac{1}{1 - \mathbf{p}^2})$$
Thus  $\hat{\mathbf{p}} = 0$  (no accel), fields  $\hat{\mathbf{n}} \times (R^2)$ 
If  $\hat{\mathbf{p}} \neq 0$ , there is a  $\frac{1}{2}$  component.
This is called the "radiation field"

Let 
$$\vec{a} = \vec{\beta}c$$
. Then  $\vec{a}_r = (\vec{a} \cdot \vec{h})\vec{n}$ 

$$\vec{a}_{\perp} = (\vec{a} - \vec{a}_{r})\vec{\beta} = \frac{1}{2}(\vec{a} - \vec{a}_{\perp} + \vec{a}_{\perp})$$
(algebra)
$$(algebra) = \left(-\frac{e(\vec{a}_{\perp} - \vec{h} \times (\vec{a} \times \vec{B}))}{4\pi\epsilon_{0} C R (1 - \vec{B} \cdot \vec{h})^{3}}\right) ret$$

$$\vec{B}_{rad} = \left(\frac{\hat{u} \times \vec{E}_{rad}}{c}\right)_{ret}$$