

Now let's show that the radiation fields transport energy. Recall that the energy density in EM fields

$$\text{is } \frac{1}{2}(\epsilon_0 E^2 + \frac{1}{\mu_0} B^2) \equiv \rho \quad (\text{here } \rho = \text{energy density, not charge density})$$

$$\text{Then } \partial_t \left(\int_{\text{Volume}} dV \rho \right) = - \oint_{\text{surf}} (\text{power flux } \vec{S}) \cdot d\vec{A}$$

"energy/area/time"

(Just conservation of energy.)

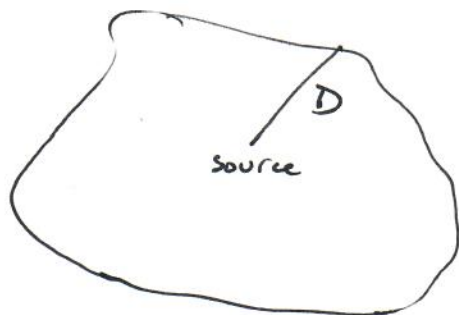
$$\begin{aligned} \text{The LHS is } & \int dV \left(\epsilon_0 \vec{E} \cdot \frac{\partial \vec{E}}{\partial t} + \frac{1}{\mu_0} \vec{B} \cdot \frac{\partial \vec{B}}{\partial t} \right) \\ &= \frac{1}{\mu_0} \int dV \left(\epsilon_0 \vec{E} \cdot \left(\frac{\nabla \times \vec{B}}{\epsilon_0} \right) - \vec{B} \cdot (\nabla \times \vec{E}) \right) \\ &= -\frac{1}{\mu_0} \int dV \nabla \cdot (\vec{E} \times \vec{B}) \\ &= -\frac{1}{\mu_0} \oint_{\text{surf}} d\vec{A} \cdot (\vec{E} \times \vec{B}) \end{aligned}$$

$$\text{So } \vec{S} = \frac{1}{\mu_0} \vec{E} \times \vec{B} \quad \text{is the power flux}$$

"Poynting vector"

Since photons have $E = pc$, the momentum flux rate
is \vec{S}/c

Now look at a surface far away from some source



The total power flowing through the surface is

$$\frac{1}{\mu_0} \int_{\substack{\text{surf} \\ \text{enclosed}}} (\vec{E} \times \vec{B}) \cdot d\vec{A}$$

$$E = E_{\text{rad}} + \mathcal{O}(1/D^2) \sim 1/D$$

$$B = B_{\text{rad}} + \mathcal{O}(1/D^2) \sim 1/D$$

$$\sim \frac{1}{\mu_0} \int (D^2 \sin\theta d\theta d\phi) (E_{\text{rad}} \times B_{\text{rad}} \sim 1/D^2)$$

$$\sim \text{finite, indep of } D \text{ as } D \rightarrow \infty$$

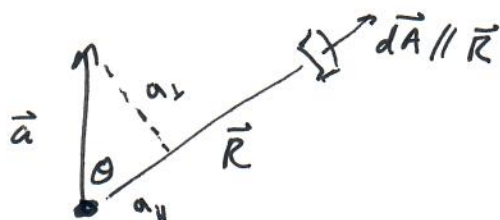
The contribution of any $1/r^2$ fields $\rightarrow 0$ in the limit.

So only the radiation fields can transport energy far from the source.

Let's look at some nonrelativistic ($|\vec{\beta}| = \frac{v}{c} \ll 1$) large distance radiation effects.

Here $\vec{E} \approx \left(\frac{-e\vec{a}_\perp}{4\pi\epsilon_0 c^2 R} \right)_{\text{ret}}$, $\vec{B} = \left(\frac{\hat{n} \times \vec{E}}{c} \right)_{\text{ret}}$

First let's compute the ~~large~~ power radiated through a solid angle. Say that at t' the acceleration is along \hat{z} :



Here $|\vec{a}_\perp| = |\vec{a}| \sin \theta$

and since $\vec{E} \perp \vec{B}$, $\vec{B} \perp \vec{R}$, $\vec{E} \perp \vec{R}$,
 $(\vec{E} \times \vec{B}) \cdot d\vec{A} = |\vec{E}||\vec{B}|R^2 d\Omega$

So $\frac{dP}{d\Omega} = \frac{1}{\mu_0} |\vec{E}||\vec{B}|R^2 = \frac{e^2 a^2 \sin^2 \theta}{4\pi\epsilon_0 \cdot 4\pi c^3}$

maximized
 @ $\theta = \pi/2$, $R \perp a$,
 with $\vec{E} \parallel \vec{a}$.

This is the power/d Ω emitted at $t' = t - R/c$

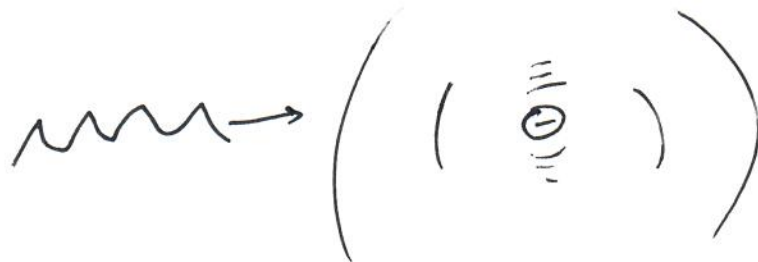
$$\text{Using } \int_0^\pi \sin^2 \theta (\sin \theta d\theta) = \int_{-1}^1 (1 - \cos^2 \theta) d \cos \theta \\ = 2 - \frac{2}{3} = \frac{4}{3},$$

We obtain the Larmor formula,

$$\boxed{P = \frac{2e^2 a^2}{4\pi\epsilon_0 \cdot 3 c^3}} \quad \left(\frac{e^2 a^2}{6\pi\epsilon_0 c^3} \right)$$

This is the power radiated by the charge at some t' when its acceleration was $|\ddot{a}|$. It passes through the sphere at $t = t' + R/c$.

Thomson scattering



an electromagnetic wave strikes a free electron, causing it to accelerate and radiate.

the acceleration is $\vec{a} = e\vec{E}/m_e$
(Newton's law + Coulomb's law)

So the total power radiated is $\frac{2e^2 a^2}{4\pi\epsilon_0 \cdot 3 c^3} = \frac{2e^4 |\vec{E}|^2}{3m_e^2 c^3 \cdot 4\pi\epsilon_0}$