## Vacuum Maxwell Eg $\vec{\nabla} \cdot \vec{E} = P/\epsilon_0$ Gauss $\vec{\nabla} \cdot \vec{E} = -\partial_t \vec{B}$ Faraday ▽京=0 No mondpoles (yet!) VxB-MoEo DE = Moj Ampère-Maxwell madration electromagnet dp = q [E + VxB] Lorente force

E. : vacuum permittivity. How well a charge ereates E.

Mo : vacuum permeability: How well a current creates IS.

We'll come back to the displacement felds D& H later.

If the sources do not extend to infinity, we may est. te  $\vec{E} = -\vec{\nabla} \phi + \vec{\nabla} \times \vec{F}$ Just a convention w.l.e.g. Helmholtz thun-simplest proof uses Fourter transform - see wikipedia of is a scalar under 3D notations & F is a vector. sm.larly  $\vec{B} = -\vec{\nabla}S + \vec{\nabla} \times \vec{A}$ S scalar, A vector Thur:  $\vec{\nabla} \cdot \vec{B} = 0 \implies \vec{\nabla}^2 S = 0$   $\implies S = 0 \quad \text{if fields} \quad \xrightarrow{\text{if }} 0$ And: Faraday => D(D.F)-VF = DS-VXA So  $\vec{\nabla}_{x}\vec{F} = -\vec{A} + \vec{\nabla}\hat{\lambda}$  for scalar  $\lambda$ . We can absorb i into of and write  $\vec{E} = -\nabla \phi - \vec{A}$ w.log. B = TxA

of \$ \$ \$ are the scalar & vector potentials.

The presence of I means the decomposition is not unique. The transformations b > d + \ "gauge symnety" A > A + D1 leave E & B unique. Only E&B are observable locally - as well as nonlocal objects like of Andx" -So this garge symmetry is not a real symmetry in the sense of taking one solution to the EOM into another, physically disthet solution. Rather it means there are many equivalent descriptions of the same electromagnetic theory and we

are free to choose & in different problems

to simplify calculations.

Inserting the potentials into the Gauss & Ampère laws, - \nabla^2 \psi - 2 \nabla \overline{A} = \beta/60  $\vec{\nabla}(\vec{\nabla}\cdot\vec{A})-\vec{\nabla}^2\vec{A}-\mu_0\epsilon_0(-\vec{\nabla}\vec{\phi}-\vec{A})=\mu_0\hat{j}$ A convenient gauge is to choose I so that Mo€o \$ - \$\overline{\nabla} \overline{A} = 0 "Lorenz gauge" Then the eg above smplity to: M. E. & - 72 d = p/e. Mo Eo Ā - Paā = Moj (Decoupling A & p) It is convenient to define 4-vectors  $A^{m}=\left(\frac{1}{c}\phi,A\right)$ with 2= 1 no 6.  $J^{n} = (c_{f}, \vec{j})$ and I = ( = 2 2 - \( \nabla\_2 \) then DA = M. Ju

2 A =0 5 Lorenz gauge EOM for Am

This is the (inhomogeneous, or "sourced") wave equation.
-) Radiation,
Plug in to definitions of \( \vec{E}\) and \( \vec{B}\):
□ = - ▽(ロφ) - 2(ロス)
and $\Box \vec{B} = \nabla \times (\Box \vec{A})$ elkethre source
and $\Box \vec{B} = \nabla \times (\Box \vec{A})$
= Tx (noj)
Again sourced wave equations, with sources given
by derivatives of s and J. Garge invariant
Well, let's solve!
We have an equation of the form
L Y = f
Lisa 2nd ender partial differential operato - (II)
I is a Luciton of spacetime we want to know
I is a given "source" finedien of spacetime
Obviously 4= L'f!

Although Maxwell's equations were a triumph of matching theory to experiment, they were philosophically troubling in the late 19th century. The reason goes back to Galileo and Newton.

The bushess of physics is to predict the future. Given the state of affairs now, what will we have in ten milliseconds? ten billion years? Wenton made the solution quantitative! e.g.  $\frac{M_N d^2 \vec{x}_N}{dt} = \vec{F}_N \qquad \text{for a system of purities}$ with  $\vec{F}_N = G \sum_{m} \frac{m_N m_M (\vec{x}_M - \vec{x}_N)}{|\vec{x}_M - \vec{x}_N|^3}$  for a system of purities.

These equations respect the Principle of Celifean Relativity. They are exactly the same for any observer using coordinates related by the Galikan Group:

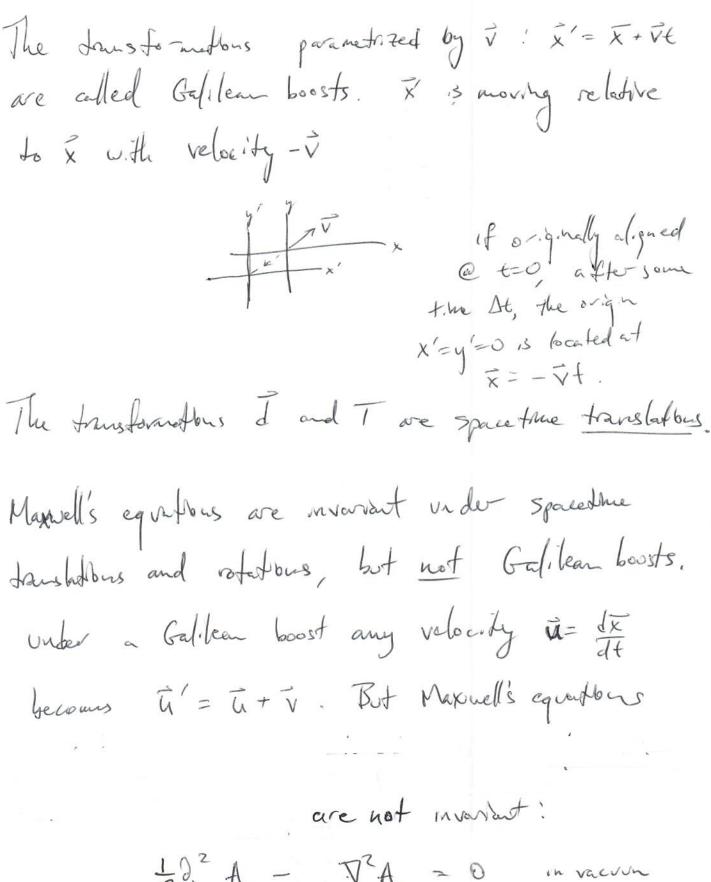
X'= RX+V+J

X= RX+V+1 L= ++ T

- o V, d'are real constant 3- vecto-s (6 param)
- R is a real orthogonal 3x3 medition (3 puram)
- · Tis a real constant (1 param)

A 10 - parameter group

The R's notate coordinates. We can change to a rotated set of coord. hetes  $\chi'' = \sum_{j} R'_{j} \chi^{j}$ This is just a linear xform, or (Rij)(xi) "Enstern notation": we drop E, with understanding that repeated indices are sunned. When we rotate the coordinates, the force vector also rotates, F'i = R'; Fi. Easy to check this manifestly to - the law of gravitation above. (The denominator | x-x, |3 = [(xm-xn)(xm-xn)\dij)= depends only or distances, which should be unaffected by rotations - try to show of mathematically!) So we find  $m\frac{d^2\vec{x}}{1+2} = \vec{F}' - the$ Same form in the rotated system.



 $\frac{1}{2} \frac{\partial^2}{\partial t} A_n - \nabla^2 A_n = 0 \quad \text{in vacuum}$   $\frac{\partial}{\partial t} \left( \frac{\partial}{\partial t'} - \vec{\nabla} \cdot \frac{\partial}{\partial \vec{x}'} \right) \left( \frac{\partial}{\partial t'} - \vec{\nabla} \cdot \frac{\partial}{\partial \vec{x}'} \right) A_n - c^2 \nabla^2 A_n = 0$