

# PHYS 427 - Thermal and Statistical Physics - Discussion 04 - Solutions

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1. **Geometric series:** Repeatedly in statistical mechanics one encounters the sum:

$$S = \sum_{n=0}^N x^n. \quad (1)$$

[For instance, you'll encounter this sum on the upcoming homework!]

(a) Show that for  $x \neq 1$ :

$$\sum_{n=0}^N x^n = \frac{1 - x^{N+1}}{1 - x} \quad (2)$$

[Hint: explicitly write out the following sum:  $(1 - x)S$ . You should notice that many terms cancel.]

(b) By taking a derivative of this result, show that for  $x \neq 1$ :

$$\sum_{n=0}^N nx^n = \frac{x}{(1 - x)^2} [1 - (N + 1)x^N + Nx^{N+1}] \quad (3)$$

As an aside, it is very common to encounter this sum with  $N \rightarrow \infty$  and  $|x| < 1$ . In this case  $\lim_{N \rightarrow \infty} x^N = 0$  and the results above simplify considerably:

$$\sum_{n=0}^{\infty} x^n = \frac{1}{1 - x} \quad , \quad \text{if } |x| < 1 \quad (4)$$

$$\sum_{n=0}^{\infty} nx^n = \frac{x}{(1 - x)^2} \quad , \quad \text{if } |x| < 1 \quad (5)$$

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(a) Following the hint:

$$(1 - x)S = S - xS \quad (6)$$

$$= (1 + \cancel{x + x^2 + \cdots + x^N}) - (\cancel{x + x^2 + \cdots + x^N} + x^{N+1}) \quad (7)$$

$$= 1 - x^{N+1} \quad (8)$$

Therefore

$$S = \frac{1 - x^{N+1}}{1 - x} \quad (9)$$

(b) Differentiating this result with respect to  $x$  gives:

$$\frac{\partial}{\partial x} \left( \sum_{n=0}^N x^n \right) = \frac{\partial}{\partial x} \left( \frac{1 - x^{N+1}}{1 - x} \right) \quad (10)$$

$$\sum_{n=0}^N \frac{\partial}{\partial x} x^n = \frac{-(N+1)x^N}{1-x} + \frac{1 - x^{N+1}}{(1-x)^2} \quad (11)$$

$$\sum_{n=0}^N nx^{n-1} = \frac{-(N+1)x^N(1-x) + 1 - x^{N+1}}{(1-x)^2} \quad (12)$$

$$= \frac{1 - (N+1)x^N + Nx^{N+1}}{(1-x)^2} \quad (13)$$

where we applied the product rule on the right hand side and then collected terms by using  $(1-x)^2$  as a common denominator. Multiplying both sides by  $x$  gives the final result:

$$\sum_{n=0}^N nx^n = \frac{x}{(1-x)^2} [1 - (N+1)x^N + Nx^{N+1}] \quad (14)$$

2. **2D non-interacting ultra-relativistic gas:** Consider a gas of  $N$  highly-relativistic particles confined to a square of area  $A = L^2$ . The particles have energy  $\varepsilon_{\vec{p}} = |\vec{p}|c$  where the momentum is  $\vec{p} = \hbar\vec{k}$  and the wavevectors are quantized as  $\vec{k} = \pi\vec{n}/L$  with  $n_x, n_y = 1, 2, \dots, \infty$ .

(a) Show that as  $L \rightarrow \infty$  we can approximate a momentum sum by an integral:

$$\sum_{\vec{p}} F(\varepsilon_{\vec{p}}) \sim \frac{L^2}{(\pi\hbar)^2} \int_0^\infty dp_x \int_0^\infty dp_y F(\varepsilon_{\vec{p}}) \quad (15)$$

where  $F(\varepsilon_{\vec{p}})$  is an arbitrary function of the dispersion  $\varepsilon_{\vec{p}}$ .

(b) Write the integral in polar coordinates  $(p, \theta)$ ; then make a change of variables to show that:

$$\sum_{\vec{p}} F(\varepsilon_{\vec{p}}) \sim \frac{L^2}{2\pi(\hbar c)^2} \int_0^\infty d\varepsilon \varepsilon F(\varepsilon) \quad (16)$$

hence, we can approximate momentum sums by energy integrals!

- (c) Compute the partition function of a single particle,  $Z_1$ .  
 (d) Compute the partition function of  $N$  indistinguishable particles,  $Z_N$ .  
 (e) Assuming the  $N$  particle system is in thermal equilibrium with a reservoir, compute its energy. You should find that

$$U = 2Nk_B T \quad (17)$$

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(a) Here a sum over momentum states  $\vec{p} = \pi\hbar\vec{n}/L$  is a sum over the positive integers  $n_x = 1, 2, \dots, \infty$  and  $n_y = 1, 2, \dots, \infty$ .

$$\sum_{\vec{p}} = \sum_{n_x=1}^\infty \sum_{n_y=1}^\infty \quad (18)$$

As the length approaches infinity, the spacing between two consecutive momentum levels becomes infinitesimal:  $\Delta p_x = \pi\hbar/L \rightarrow 0$  and  $\Delta p_y = \pi\hbar/L \rightarrow 0$  as  $L$  blows up! Hence, the momentum approaches a continuous variable. In this case, the error in approximating a sum over positive integers by the following integral is minimal:

$$\sum_{\vec{p}} = \sum_{n_x=1}^\infty \sum_{n_y=1}^\infty \approx \int_1^\infty dn_x \int_1^\infty dn_y, \text{ as } L \rightarrow \infty \quad (19)$$

Since  $\vec{p} = \pi\hbar\vec{n}/L$  (or rearranging  $\vec{n} = \vec{p}L/\pi\hbar$ ) the integral becomes

$$\sum_{\vec{p}} \approx \frac{L^2}{(\pi\hbar)^2} \int_0^\infty dp_x \int_0^\infty dp_y, \text{ as } L \rightarrow \infty \quad (20)$$

An alternative way to see this is to realize that  $1 = \Delta p_x L/\pi\hbar$  and  $1 = \Delta p_y L/\pi\hbar$ . Hence

$$\sum_{p_x} \sum_{p_y} = \frac{L^2}{(\pi\hbar)^2} \sum_{p_x} \Delta p_x \sum_{p_y} \Delta p_y \sim \frac{L^2}{(\pi\hbar)^2} \int_0^\infty dp_x \int_0^\infty dp_y, \text{ as } L \rightarrow \infty \quad (21)$$

(b) We're interested in applying this result to functions of the dispersion  $\varepsilon_{\vec{p}} = |\vec{p}|c$ .

$$\sum_{\vec{p}} F(\varepsilon_{\vec{p}}) \sim \frac{L^2}{(\pi\hbar)^2} \int_0^\infty dp_x \int_0^\infty dp_y F(\varepsilon_{\vec{p}}) \quad (22)$$

Notice this dispersion is **isotropic**, i.e. it only depends on the magnitude of the momentum and not on its orientation. Writing the integral in polar coordinates,  $p_x = p \cos \theta$  and  $p_y = p \sin \theta$ , gives

$$\frac{L^2}{(\pi\hbar)^2} \int_0^\infty dp_x \int_0^\infty dp_y F(\varepsilon_{\vec{p}}) = \frac{L^2}{(\pi\hbar)^2} \int_0^\infty dp \int_0^{\pi/2} d\theta p F(pc) \quad (23)$$

$$= \frac{L^2}{(\pi\hbar)^2} \frac{\pi}{2} \int_0^\infty dp p F(pc) \quad (24)$$

We now make the change of variables  $\varepsilon = pc$  which gives the desired result

$$\sum_{\vec{p}} F(\varepsilon_{\vec{p}}) \sim \frac{L^2}{(\pi\hbar)^2} \int_0^\infty dp_x \int_0^\infty dp_y F(\varepsilon_{\vec{p}}) = \frac{L^2}{2\pi(\hbar c)^2} \int_0^\infty d\varepsilon \varepsilon F(\varepsilon) \quad (25)$$

(c) The single-particle partition function is

$$Z_1 = \sum_{\vec{p}} e^{-\beta\varepsilon_{\vec{p}}} \quad (26)$$

$$\sim \frac{L^2}{2\pi(\hbar c)^2} \int_0^\infty d\varepsilon \varepsilon e^{-\beta\varepsilon} \quad , \text{ let } x = \beta\varepsilon \quad (27)$$

$$= \frac{L^2}{2\pi(\hbar c)^2} (k_B T)^2 \int_0^\infty dx x e^{-x} \quad (28)$$

Notice that the remaining integral is just an **unimportant** dimensionless constant! We don't actually need to evaluate it — all the physics is contained in the fact that  $Z_1$  scales with the length squared and temperature squared.

Nevertheless, this is an example of a standard integral known as the [gamma function](#):

$$\Gamma(\nu) = \int_0^\infty dx x^{\nu-1} e^{-x} \quad , \quad \text{Re}(\nu) > 0 \quad (29)$$

Thus, our integral is  $\Gamma(2) = 1! = 1$  (you could quickly evaluate this integral by using integration by parts also). Therefore

$$Z_1 = \frac{L^2}{2\pi(\hbar c)^2} (k_B T)^2 \quad (30)$$

(d) The partition function of  $N$  indistinguishable particles with this dispersion is given by

$$Z_N = \frac{1}{N!} (Z_1)^N = \frac{1}{N!} \left( \frac{L^2}{2\pi(\hbar c)^2} (k_B T)^2 \right)^N \quad (31)$$

(e) The energy of the gas is given by

$$U = -\frac{\partial \ln Z}{\partial \beta} = -\frac{1}{Z} \frac{\partial Z}{\partial \beta} \quad (32)$$

$$= -\frac{1}{\frac{1}{N!} \left( \frac{L^2}{2\pi(\hbar c)^2} \right)^N (k_B T)^{2N}} \frac{1}{N!} \left( \frac{L^2}{2\pi(\hbar c)^2} \right)^N \frac{\partial \beta^{-2N}}{\partial \beta} \quad (33)$$

$$= -\frac{1}{(k_B T)^{2N}} (-2N) (k_B T)^{2N+1} \quad (34)$$

$$= 2N k_B T \quad (35)$$