# Deep ou Bayesiano? Estimando parâmetros de lentes gravitacionais

Análise de Big Data & Astroinformática Prof.: Clécio R. de Bom João Paulo

#### Conteúdo

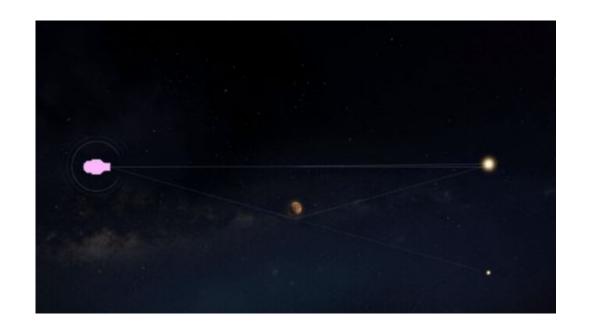
- Por que estimar parâmetros de lentes gravitacionais?
- Duas vertentes interessantes para análise de sistemas SL
- Modelagem
  - Métodos Bayesianos
  - Deep Learning
  - Resultados/comparação
- Conclusão

Por que estimar parâmetros de Lentes gravitacionais?

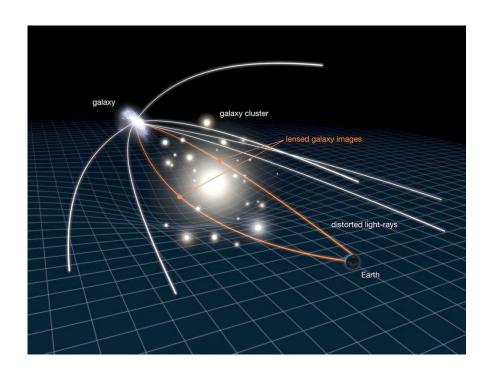
# Detecção direta de matéria escura



Um telescópio gravitacional...



Vínculo em parâmetros cosmológicos



Duas vertentes interessantes para análise de sistemas SL

> Detecção de sistemas

A&A 611, A2 (2018)

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## Deep convolutional neural networks as strong gravitational lens detectors

C. Schaefer<sup>1</sup>, M. Geiger<sup>1</sup>, T. Kuntzer<sup>1</sup>, and J.-P. Kneib<sup>1,2</sup>

> Modelagem em massa

Published: 31 August 2017

# Fast automated analysis of strong gravitational lenses with convolutional neural networks

Yashar D. Hezaveh ™, Laurence Perreault Levasseur ™ & Philip J. Marshall

Nature 548, 555–557 (2017) Cite this article

8432 Accesses | 97 Citations | 346 Altmetric | Metrics

> Modelagem em massa

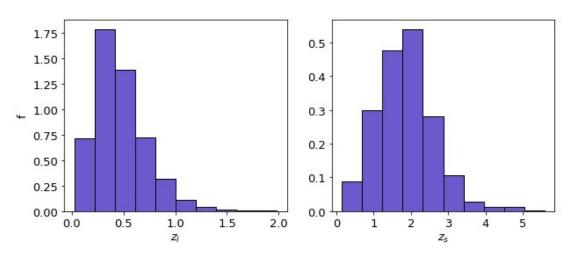
Deep Learning in Wide-field Surveys: Fast Analysis of Strong Lenses in Ground-based Cosmic Experiments

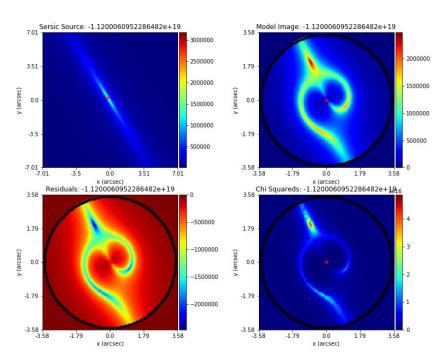
C. R. Boma, J. Pohd, B. Nordc, M. Blanco-Valentinb, L. O. Diasb

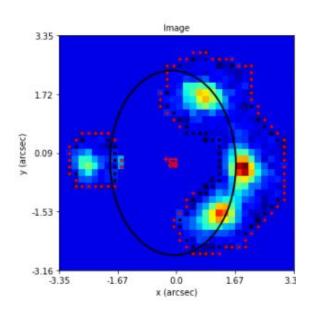
Modelagem via métodos bayesianos e deep learning

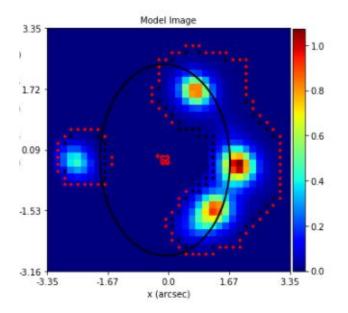
#### > Data set

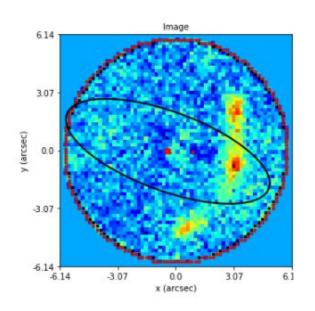
#### ~ 19000 sistemas

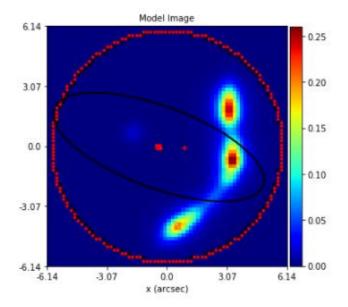


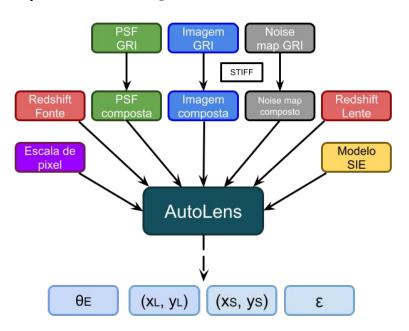






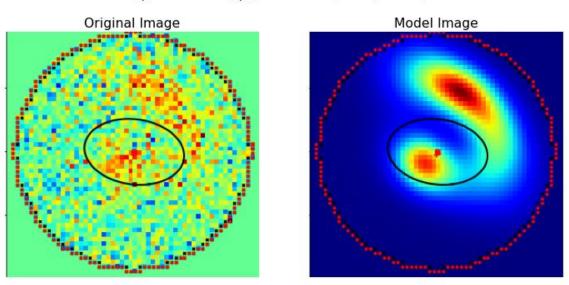






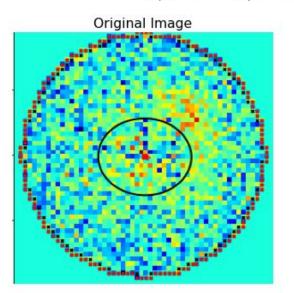
## > AutoLens (Distribuição de brilho da fonte)

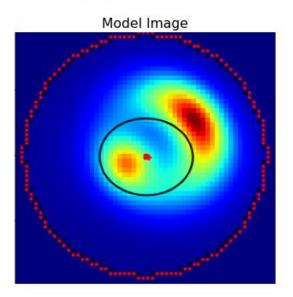
 $\theta_{E,t}$ : 2.47  $\theta_{E,m}$ : 2.42 ± (0.17, 0.15)



### > AutoLens (Distribuição de brilho da fonte)

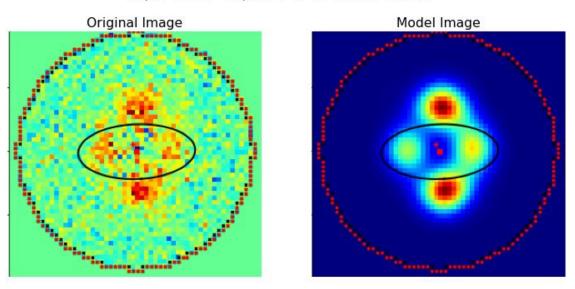
 $\theta_{E,t}$ : 2.3  $\theta_{E,m}$ : 2.51 ± (0.23, 0.35)





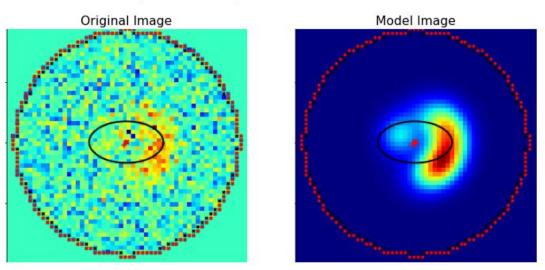
#### > AutoLens (Distribuição de brilho da fonte)

 $\theta_{E,t}$ : 2.32  $\theta_{E,m}$ : 2.48 ± (0.21, 0.08)



## > AutoLens (Distribuição de brilho da fonte)

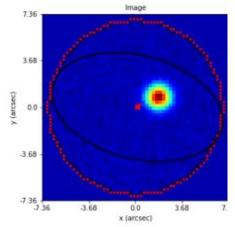
 $\theta_{E,t}$ : 1.71  $\theta_{E,m}$ : 1.83 ± (0.21, 0.2)

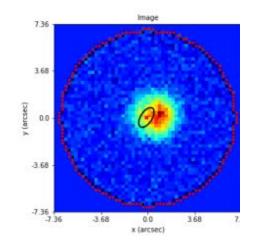


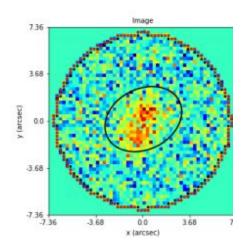
> AutoLens (Distribuição de brilho da fonte)

30 sistemas modelizados...

#### ...mas 15 outliers

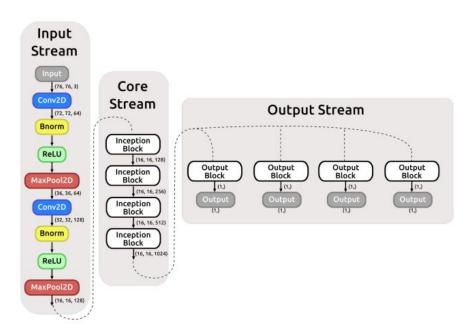






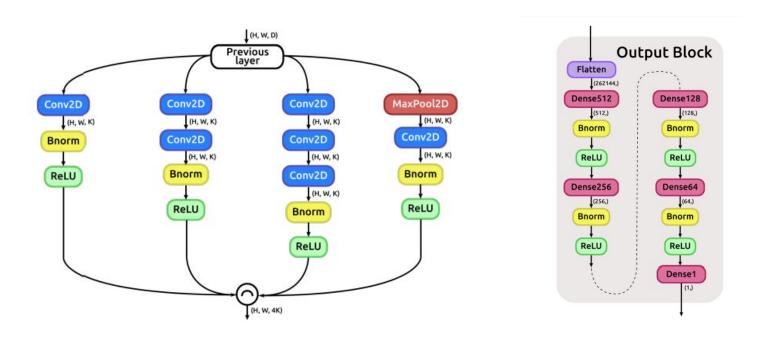
### Modelagem via Deep Learning - Inception

> Análise baseada em 1911.06341 (Bom, et al)



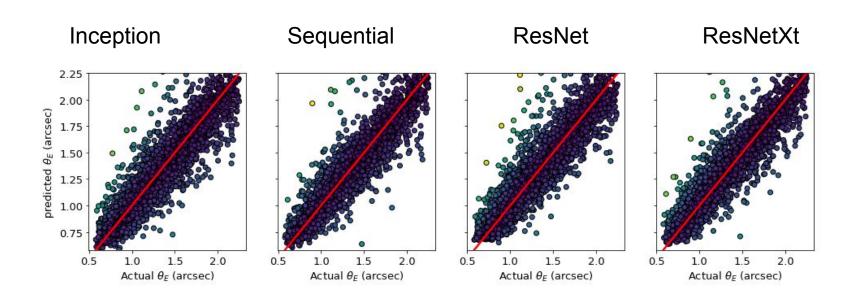
#### Modelagem via Deep Learning - Inception

> Análise baseada em 1911.06341 (Bom, *et al*)



#### Modelagem via Deep Learning - Inception

> Análise baseada em 1911.06341 (Bom, *et al*)



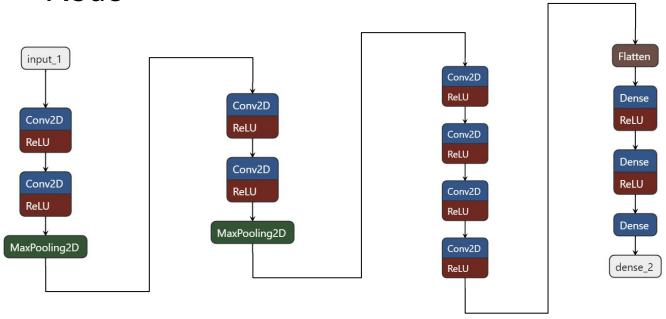
#### 

Estimating parameters of strong gravitational lenses with convolutional neural networks.

This code uses convolutional neural networks (with tensorflow) to estimate the parameters of strong gravitational lenses. Unfortunately we're not very good at coding, so you'll find that the code is messy, not well documented, and crazily written. However, that shouldn't discourage you from trying it out. Because it's a pretty pretty cool thing: The code can recover the parameters of gravitational lenses in a fraction of a second. Something that used to take hundreds of hours!

https://github.com/yasharhezaveh/Ensai

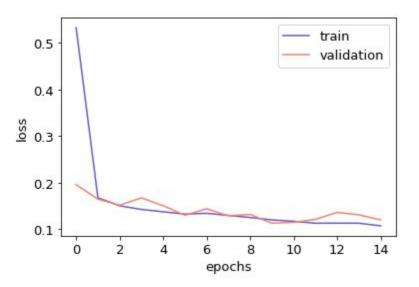
#### > Rede

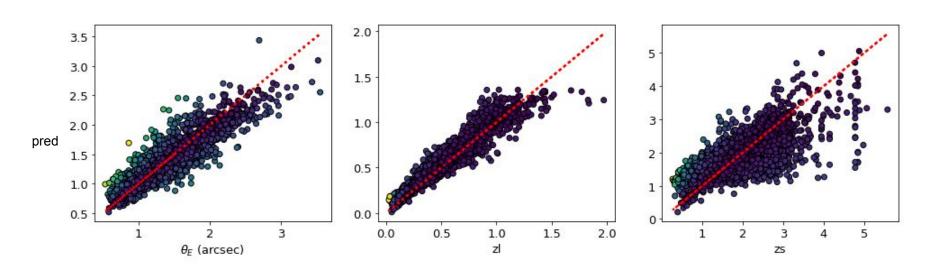


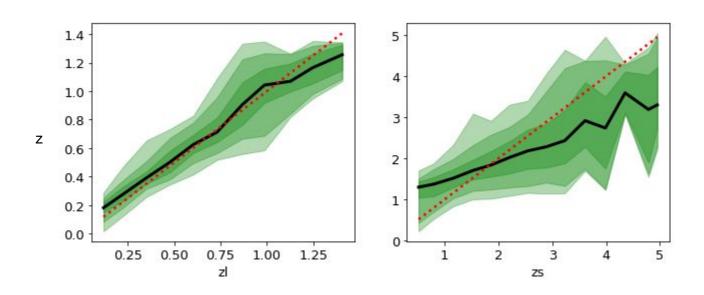
#### > Rede

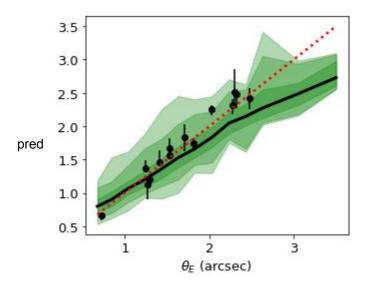
- > Treinamento (~0.7), validação (~0.2) e teste (~0.1)
- > 15 Épocas
- > Batch Size 32
- > Raio de Einstein, z∟ e zs como output

#### > Loss









	Artigo				Este trabalho
	Inception	Sequential	ResNet	ResNetXt	Ensai
RMSE	0.026	0.021	0.023	0.020	0.032

$$RMSE = \sqrt{rac{1}{n}\sum_i (x_i^p - x_i^a)}$$

> Uma curiosidade...

Bayes (30 sistemas):

 $10 \min \times 30 + 10 \min \approx 5h$ 

DL (3720 sistemas):

2.19 s

#### Conclusões

- Análise Bayesiana
  - Mais intuitiva
  - Análise cruzada com visual
  - Análise de outliers (depende do método)
  - Processo menos prático e lento
- Análise via DL
  - Mais prático
  - Mais rápido
  - Tão complexo quanto você queira
  - Sensível a overfit e underfit



## Obrigado!