

reading7

November 16, 2025

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[1]: import numpy as np
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0.0.1 a) Calculate the Fried parameter r_0 at 500 nm given seeing of 0.5 arcseconds

0.0.2 The relationship between seeing (FWHM) and Fried parameter is:

0.0.3 seeing (arcsec) $0.98 * \lambda / r_0$

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[2]: lam = 500e-9          # wavelength 500 nm
FWHM_arcsec = 0.5
FWHM_rad = np.deg2rad(FWHM_arcsec/3600.0) # convert arcsec to radians
r0 = 0.98 * lam / FWHM_rad
print("Fried parameter r0 at 500 nm: {:.3f} m ({:.1f} cm)".format(r0, r0*100))
```

Fried parameter r_0 at 500 nm: 0.202 m (20.2 cm)

0.0.4 b) Calculate r_0 at 1.65 microns using Kolmogorov turbulence scaling

0.0.5 For Kolmogorov turbulence: $r_0 \propto \lambda^{6/5}$

0.0.6 Therefore: $r_0(\lambda_2) = r_0(\lambda_1) * (\lambda_2 / \lambda_1)^{6/5}$

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[3]: lam_ir = 1.65e-6 # wavelength 1.65 microns
r0_ir = r0 * (lam_ir/lam)**(6/5)
print("Fried parameter r0 at 1.65 m: {:.3f} m ({:.1f} cm)".format(r0_ir,
    ↪r0_ir*100))
```

Fried parameter r_0 at 1.65 m: 0.847 m (84.7 cm)

0.0.7 c) Calculate number of sub-apertures for a 10 m telescope

0.0.8 Number of sub-apertures $(D / r_0)^2$

0.0.9 where D is the telescope diameter

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[4]: D = 10.0 # telescope diameter in meters
subapertures = (D / r0_ir)**2
print("Number of sub-apertures (approx): {:.0f}".format(subapertures))
```

Number of sub-apertures (approx): 139