





Joint ICTP-IAEA Workshop on Monte Carlo Radiation Transport and Associated Data Needs for Medical Applications

28 October – 8 November 2024 ICTP, Trieste, Italy

Lecture 8

Electron physics and parameters

Gouvernement du Canada

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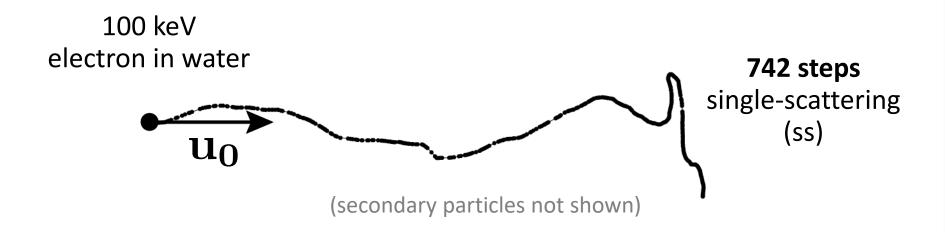
Electrons (positrons) interact via 4 processes

- 1. Emission of a bremsstrahlung photon in the nuclear field or in the field of atomic electrons (electron-electron bremsstrahlung)
- 2. Inelastic collision with atomic electrons (knock-on electron)
- 3. Elastic collision with nuclei and atomic electrons (deviation)
- **4. Positron annihilation** (emission of a photon pair)

The main challenge in modelling the transport of electrons through matter is that the **cross-sections are very large**, because charged particles interact essentially "all the time" with the atoms they encounter.

Challenge: electrons interact "all the time"

The number of interactions as an electron slows down in a material is very large, for example:

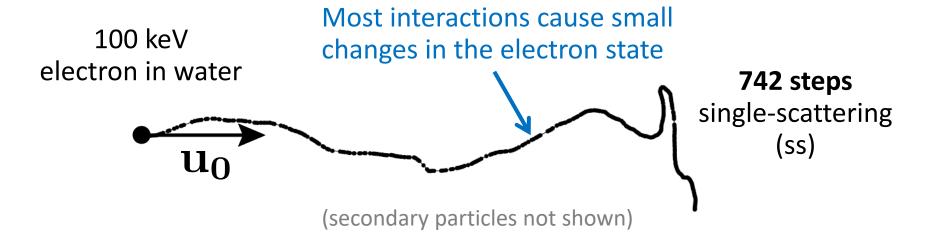


In the radiotherapy (MeV) range, a single incident electron induces $\sim 10^6$ interactions before all energy is absorbed: ~ 1 s of CPU time.

Simulations typically require $\sim 10^9$ histories, which implies more than 10000 days of computation time! Analog simulation is not practical.

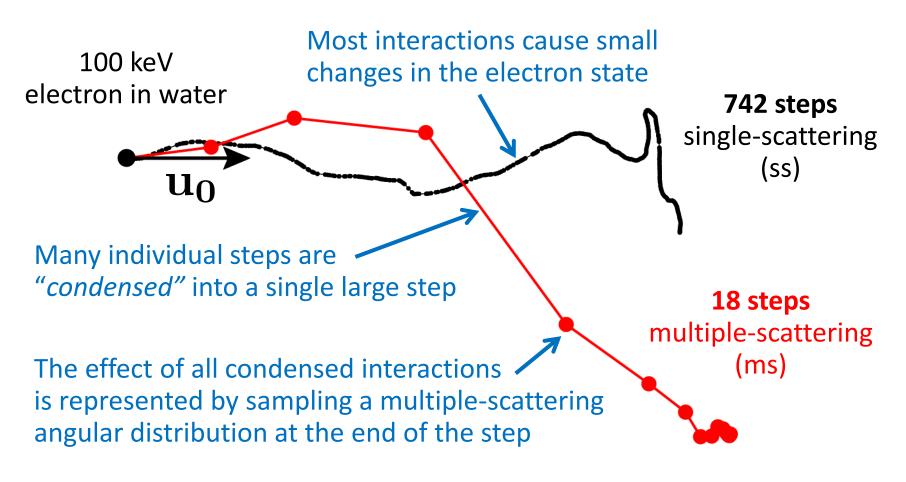
Solution: condensed history technique (CH)

To curb long simulation times (in 1963!) for electron transport, **Martin Berger** introduced the **condensed history technique**.



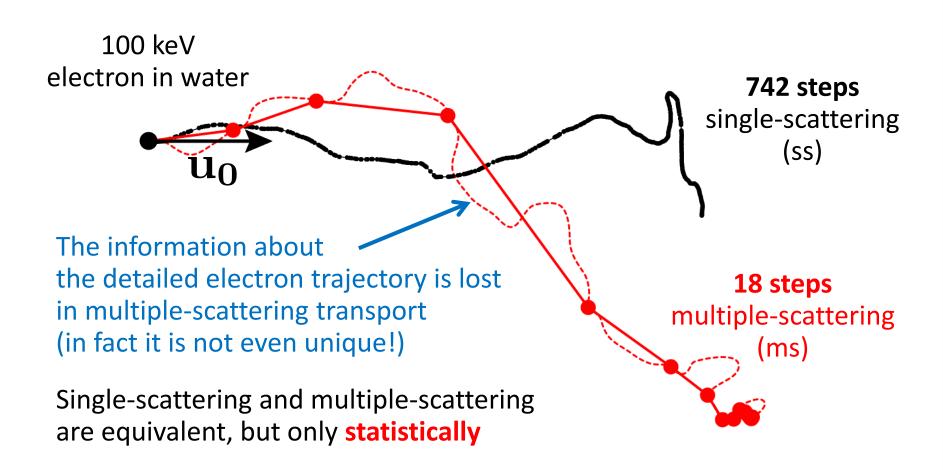
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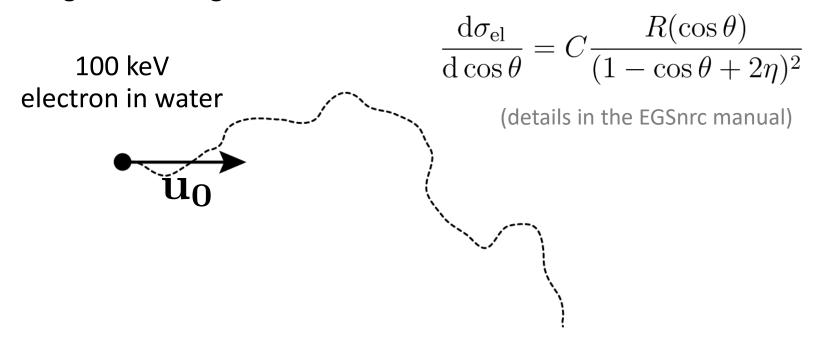
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To curb long simulation times (in 1963!) for electron transport, **Martin Berger** introduced the **condensed history technique**.



Condensing elastic scattering interactions

EGSnrc uses the screened Rutherford cross section to model elastic single-scattering interactions:



Condensing elastic scattering interactions

EGSnrc uses the screened Rutherford cross section to model elastic single-scattering interactions:

100 keV electron in water

$$\frac{d\sigma_{el}}{d\cos\theta} = C \frac{R(\cos\theta)}{(1-\cos\theta+2\eta)^2}$$

(details in the EGSnrc manual)



Answer: Goudsmit-Saunderson distribution (with $u = \cos\theta$):

$$F_{\text{MS}}(s, u) = \sum_{l=0}^{\infty} \left(l + \frac{1}{2} \right) e^{-G_l(s, u)} P_l(u)$$

 G_1 (roughly twice the average angle squared) is not allowed to exceed 0.5 in EGSnrc (ximax parameter)

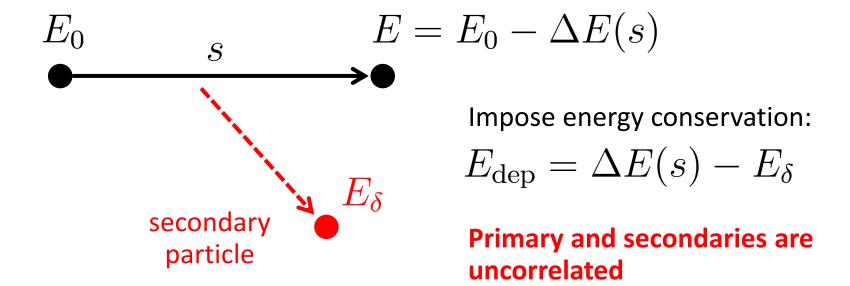
Question: what is the probability to travel along polar angle θ , after a path length s?

Condensing inelastic scattering interactions

Martin Berger defined two condensed history classes:

Class I:

The multiple-scattering step size is sampled from pre-determined distributions of energy loss fractions, and secondary particle emission is sampled independently (ETRAN, ITS, MCNP).

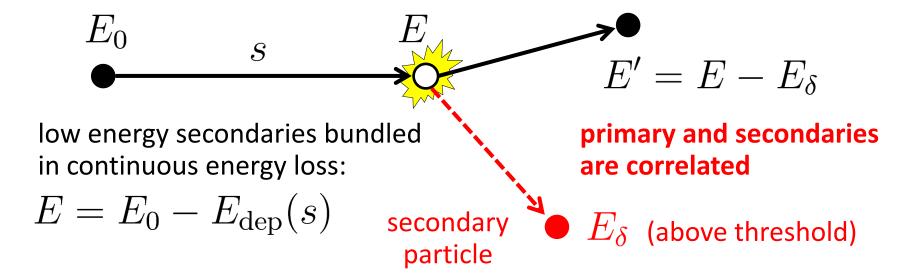


EGSnrc implements a "class II" CH scheme

Martin Berger defined two condensed history classes:

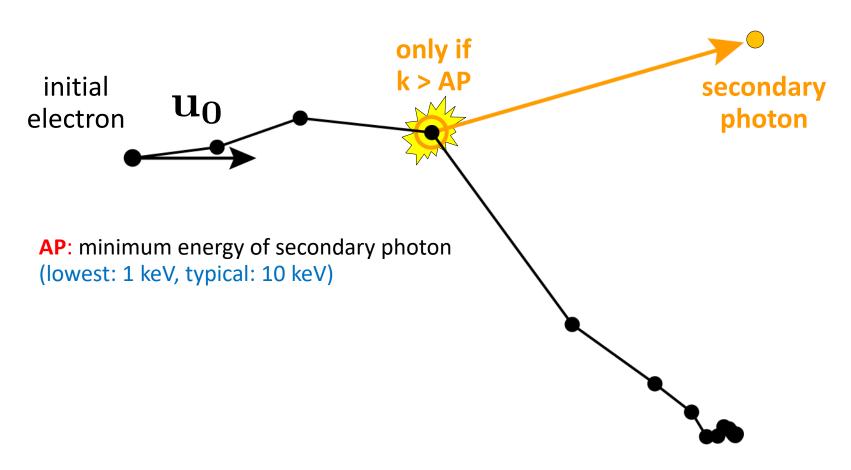
Class II:

Collisions with energy losses below a certain **threshold** are grouped in the multiple-scattering step, and occasional "catastrophic" collisions with energy losses above the **threshold** are sampled explicitly (EGS4, EGSnrc, Penelope, Geant4, vmc++)



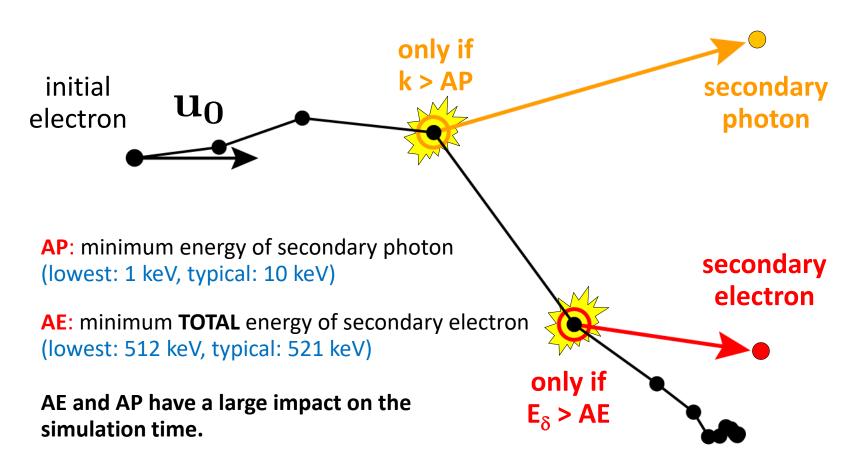
The catastrophic thresholds are AE and AP

"Hard" events are sampled explicitly if the secondary particle has energy above AE (electrons) or AP (photons).



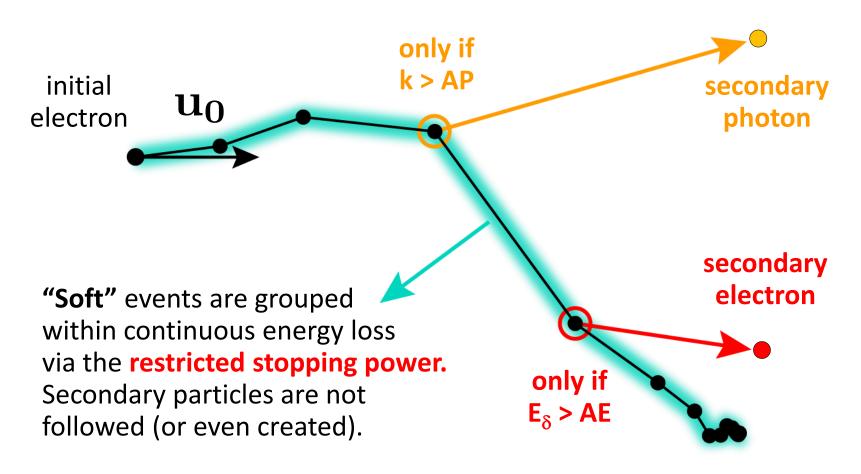
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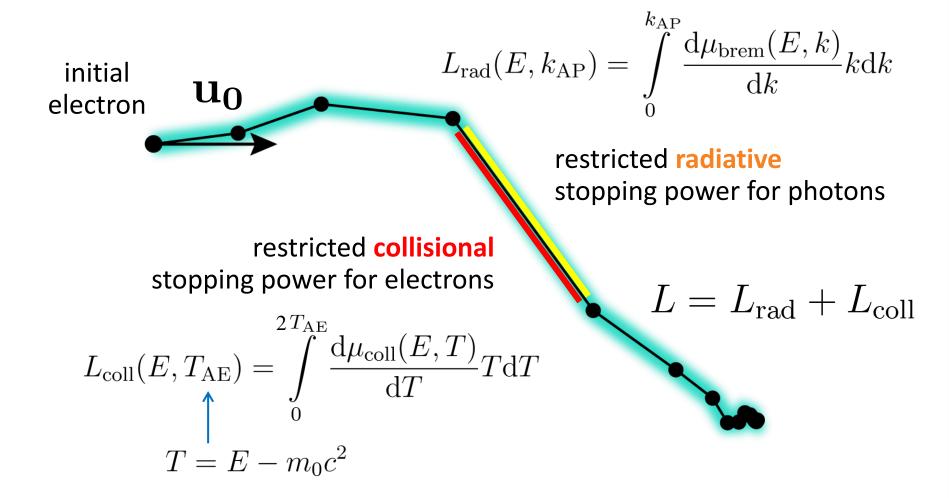
The catastrophic thresholds are AE and AP

"Hard" events are sampled explicitly if the secondary particle has energy above AE (electrons) or AP (photons).



Restricted stopping power below AE and AP

The restricted stopping powers (up to AE and AP) give the average energy lost to sub-threshold events, per unit length.



Can we now transport electrons?

Recall the generic procedure, adjusting for condensed history:

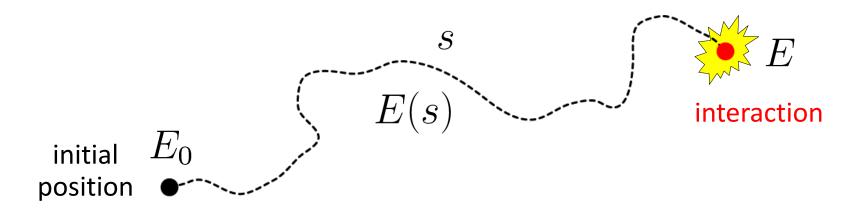
- 1. sample distance to the next **catastrophic** interaction (using the total cross sections for secondaries above AE and AP).
- transport electron to interaction site (or to the next boundary), deposit energy from sub-threshold events along the way, and sample the multiple-scattering angle.
- 3. sample which interaction occurs, according to the relative cross sections.
- 4. sample energies and angles of scattered particles, using the differential cross section.

(rinse and repeat)

Step 1 issue: continuous energy loss

1. sample distance to the next catastrophic interaction:

electrons are loosing energy continuously to sub-threshold events: cross sections and stopping powers are not constant!

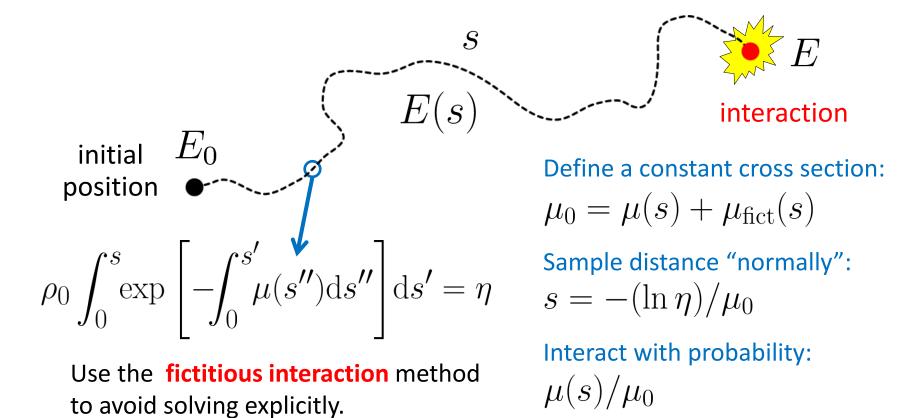


Recall: $\mu s = -\ln \eta$

Step 1 issue: continuous energy loss

1. sample distance to the next catastrophic interaction:

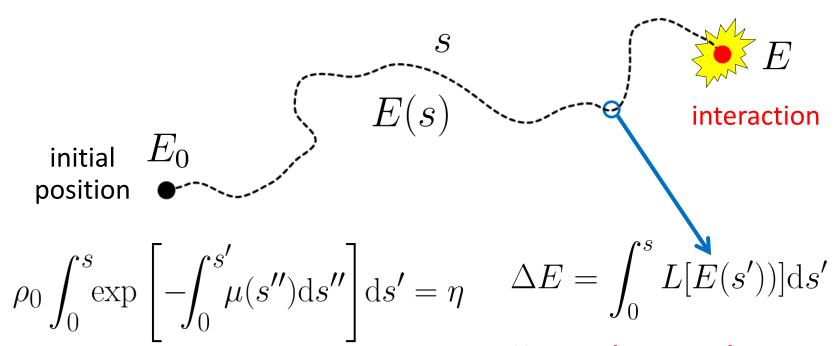
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Step 1 issue: continuous energy loss

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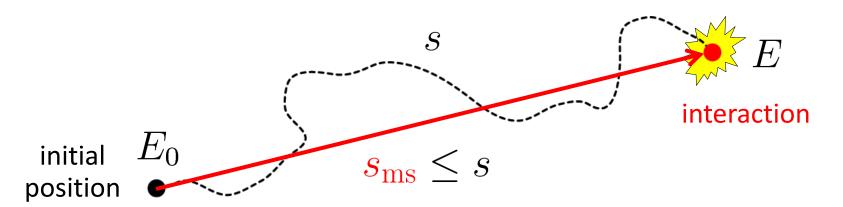
Use the **fictitious interaction** method to avoid solving explicitly.

Use a **series expansion** to express *E* as a function of *s*.

Step 2 issues: curvilinear electron path

2. transport electron to the interaction site:

a) the distance s is along the real (curved) electron path: what is the corresponding straight line step?

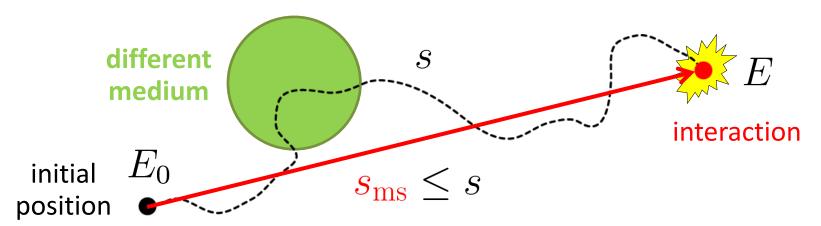


We need an electron step algorithm

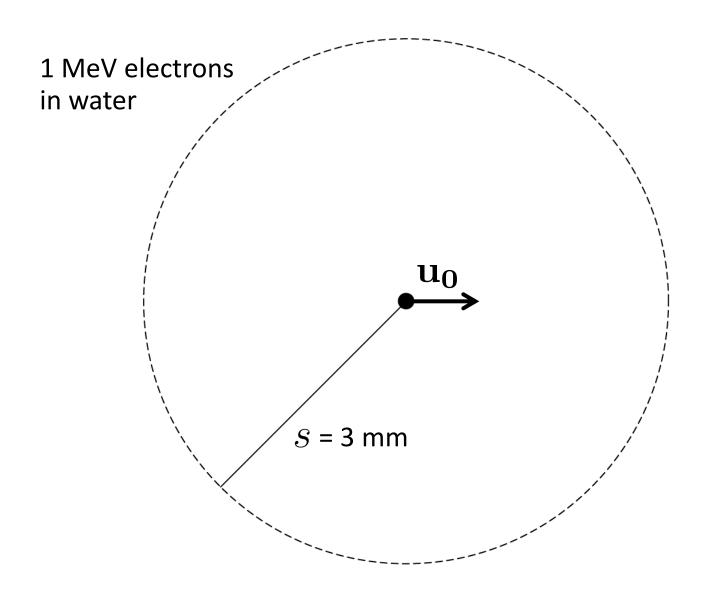
Step 2 issues: curvilinear electron path

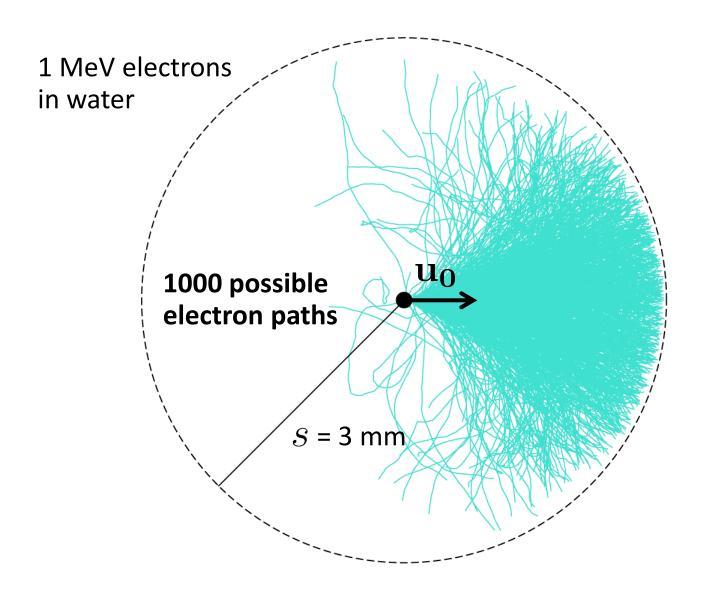
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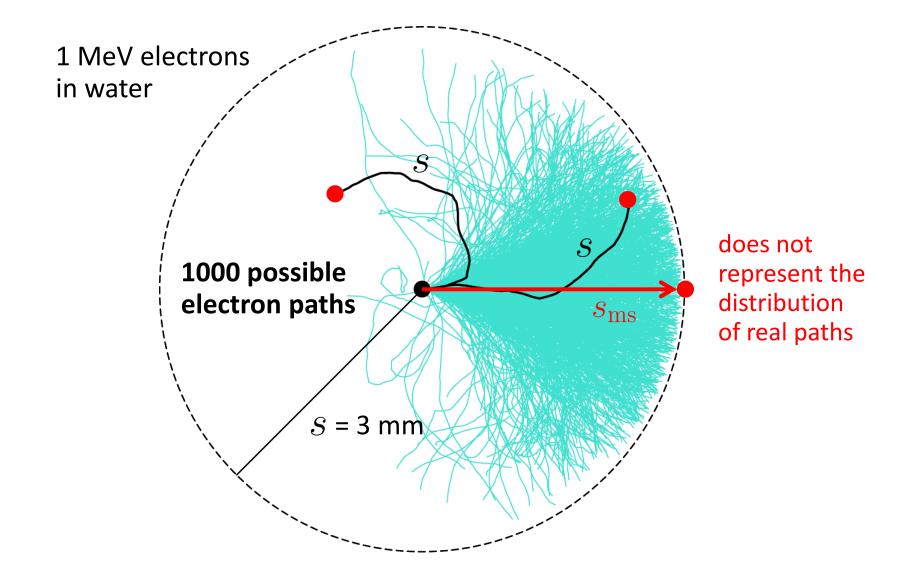
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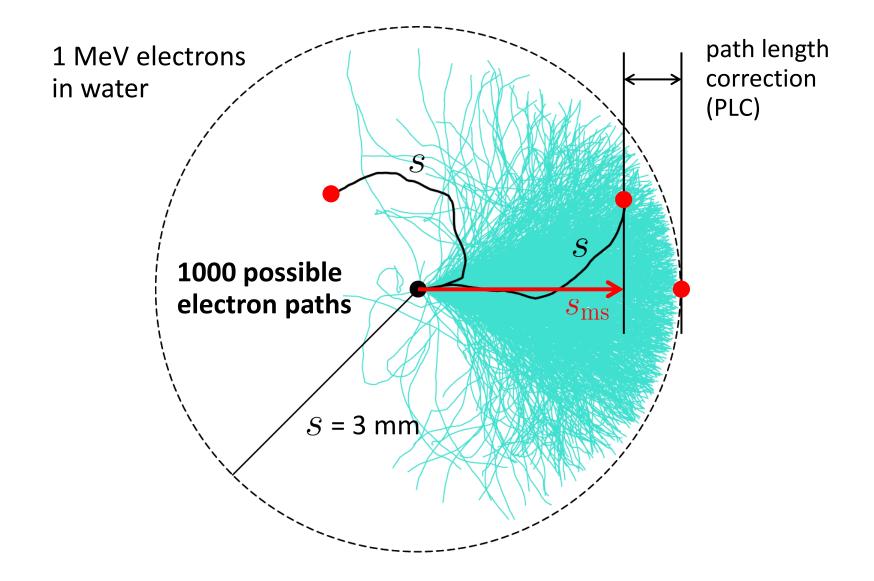


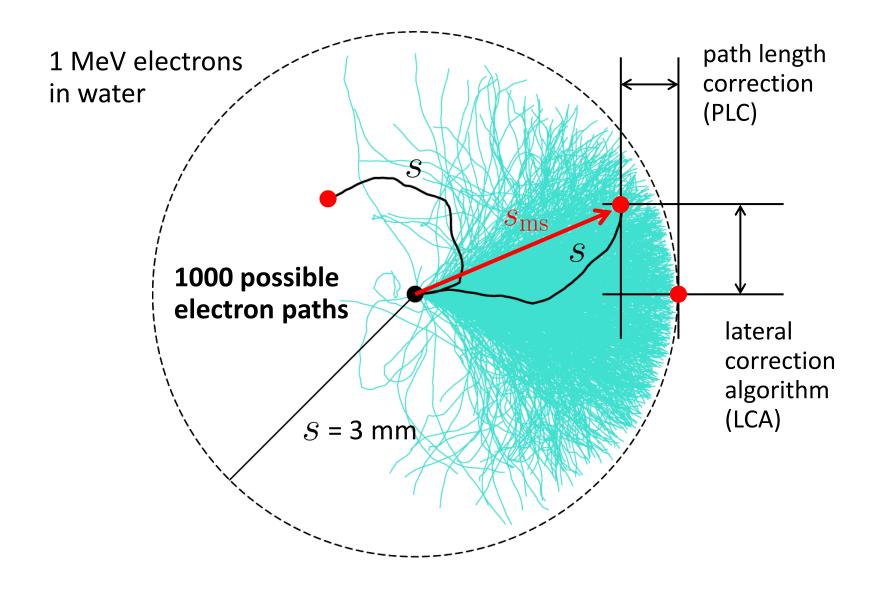
- b) ensure the real (curved) electron path could not have crossed a boundary: consider distance to boundaries in all directions.
- c) enforce limits on fractional energy loss (ESTEPE), multiplescattering angle (ximax), and magnetic deflection (EM ESTEPE).



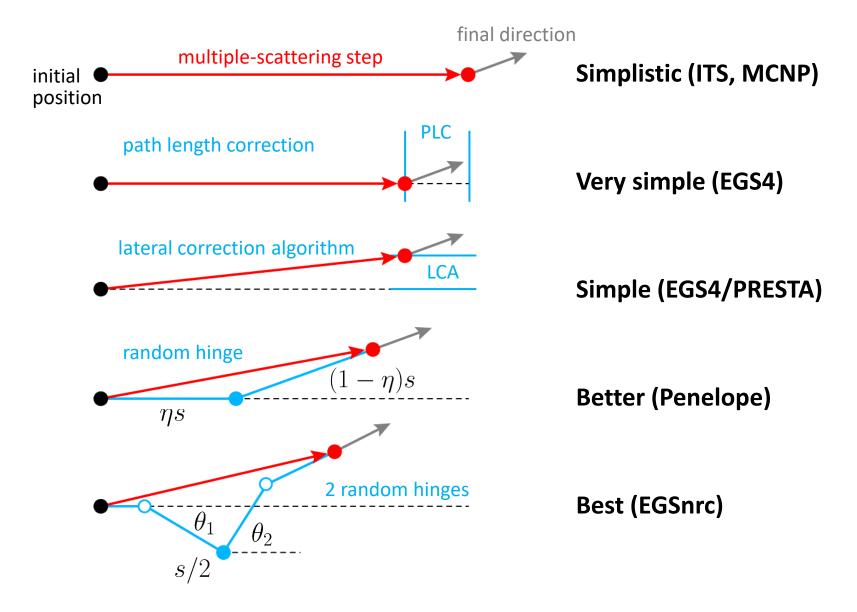




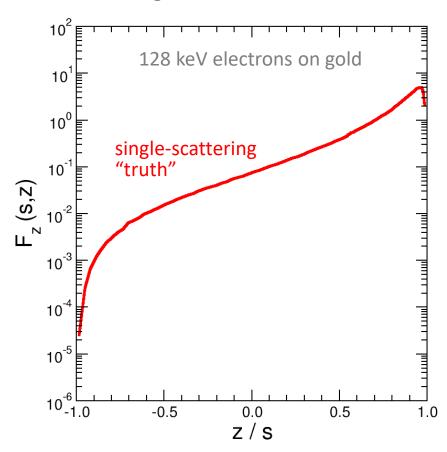




Electron step algorithms have evolved

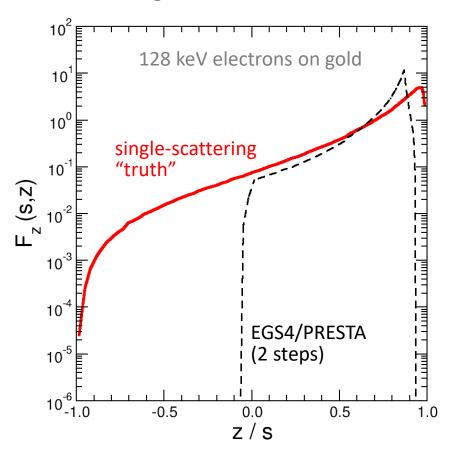


longitudinal distribution



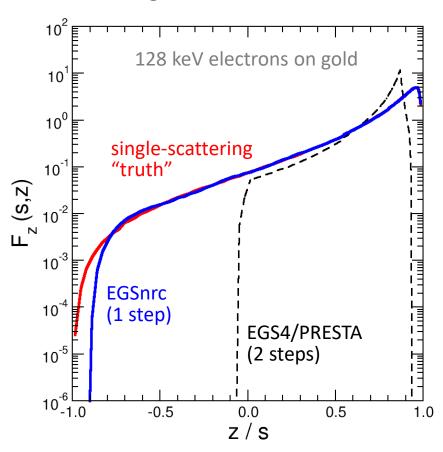
Probability to find the electron at position z after a path length s.

longitudinal distribution



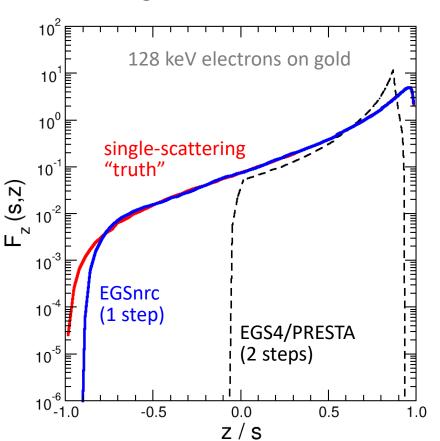
Probability to find the electron at position z after a path length s.

longitudinal distribution



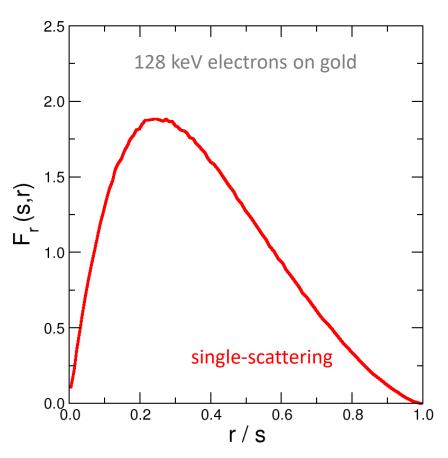
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longitudinal distribution



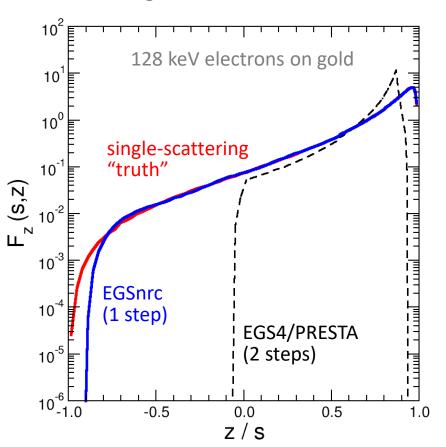
Probability to find the electron at position z after a path length s.

transverse distribution



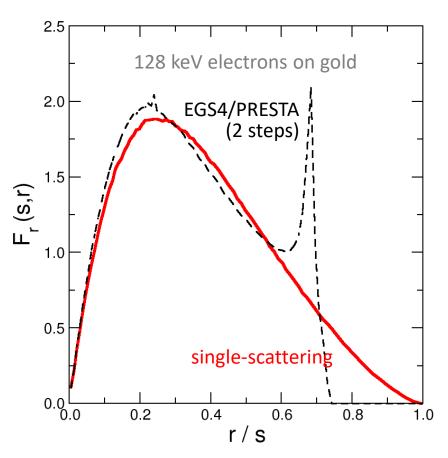
Probability to find the electron at radial position r after a path length s.

longitudinal distribution



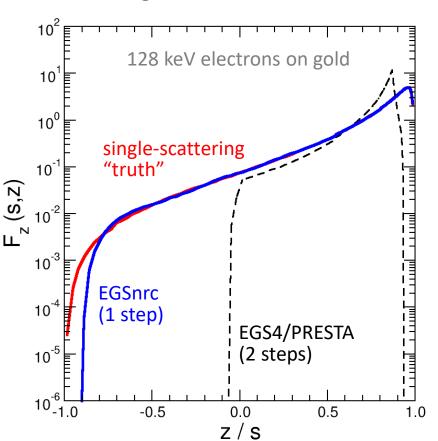
Probability to find the electron at position z after a path length s.

transverse distribution



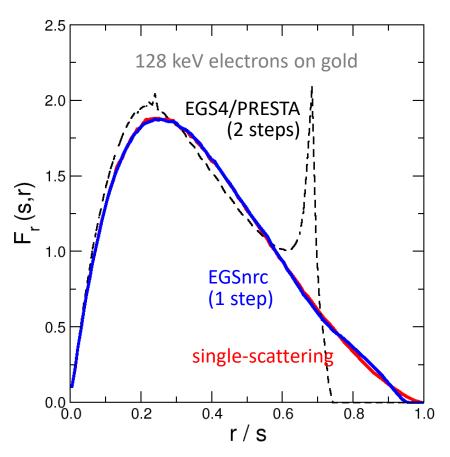
Probability to find the electron at radial position r after a path length s.

longitudinal distribution



Probability to find the electron at position z after a path length s.

transverse distribution

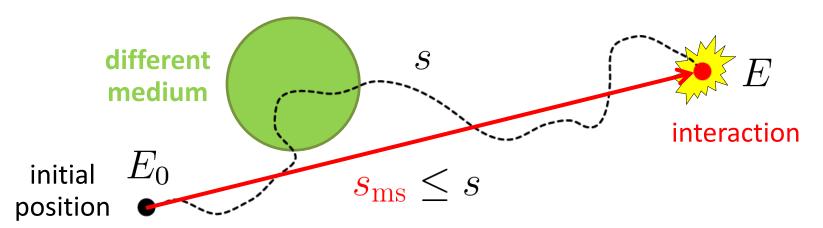


Probability to find the electron at radial position r after a path length s.

Step 2 issues: curvilinear electron path

2. transport electron to the interaction site:

a) the distance s is along the real (curved) electron path: what is the corresponding straight line step?



b) ensure the real (curved) electron path could not have crossed a boundary: consider distance to boundaries in all directions.

Limit path length to nearest boundary distance

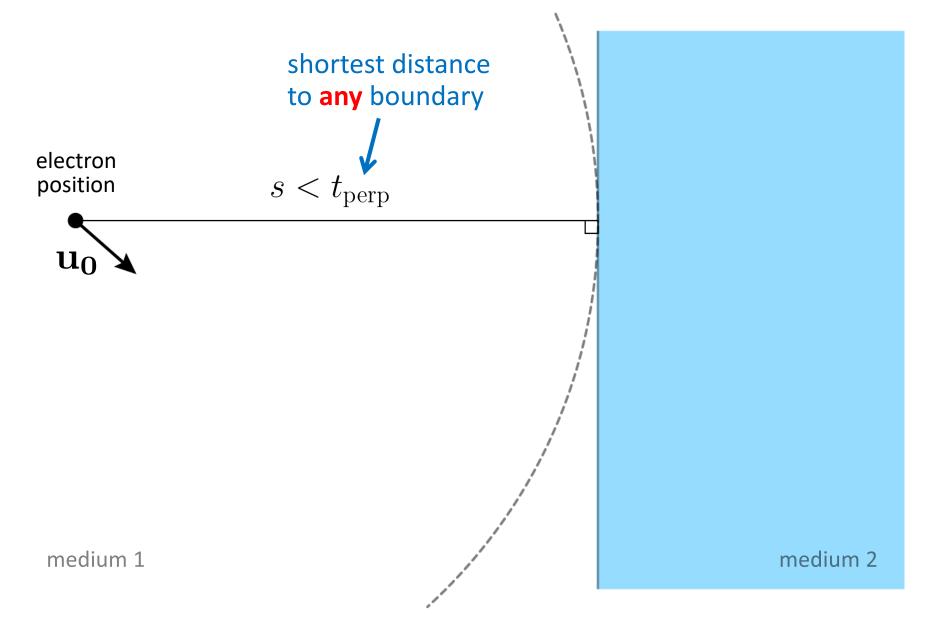
electron position



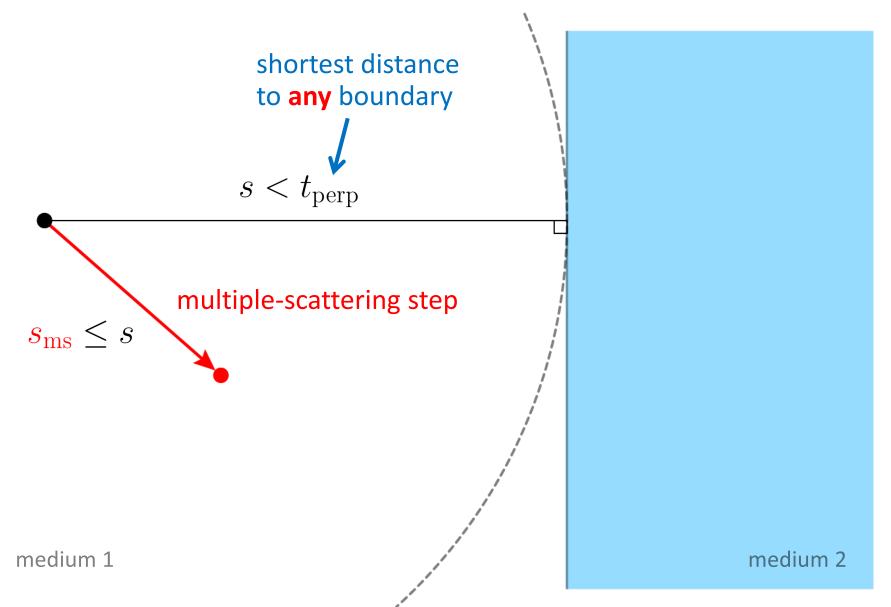
medium 1

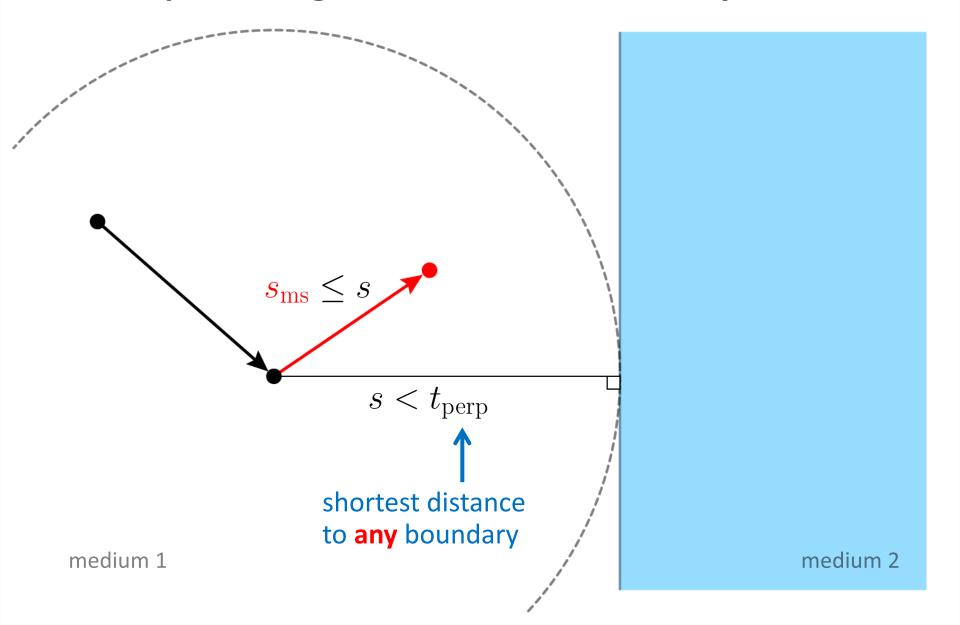
medium 2

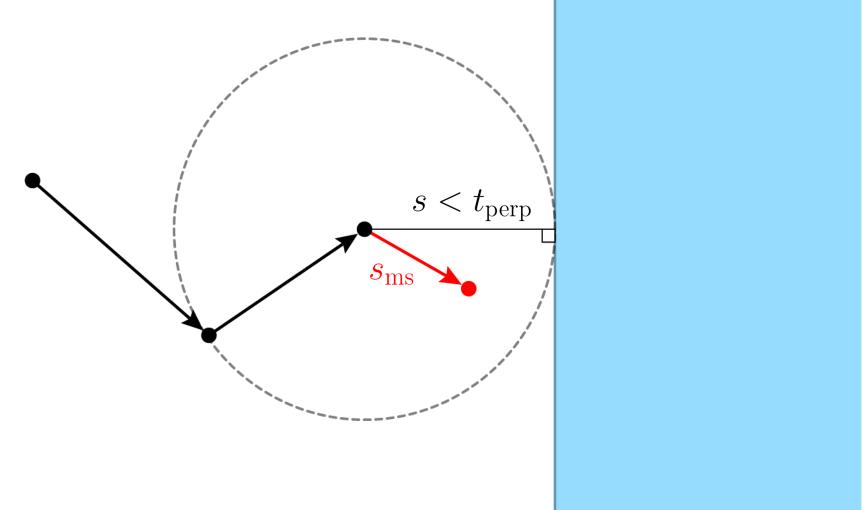
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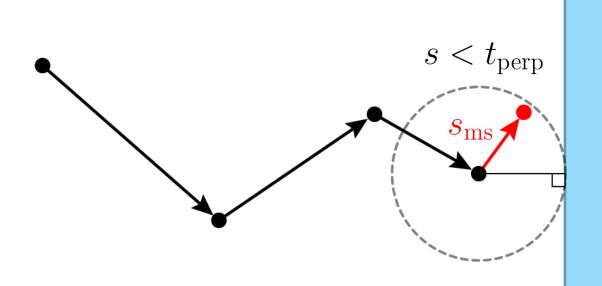






medium 1

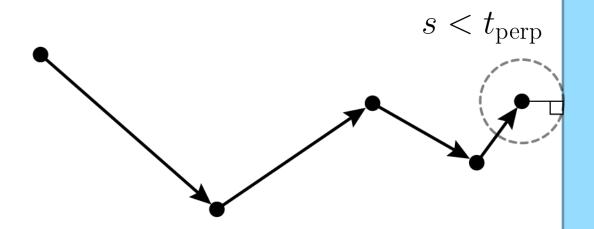
medium 2



medium 1

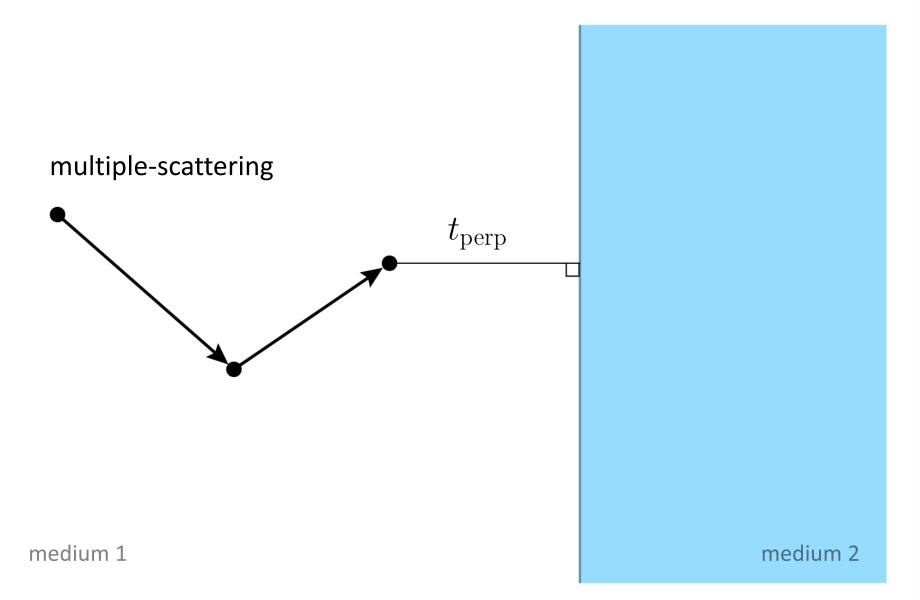
medium 2

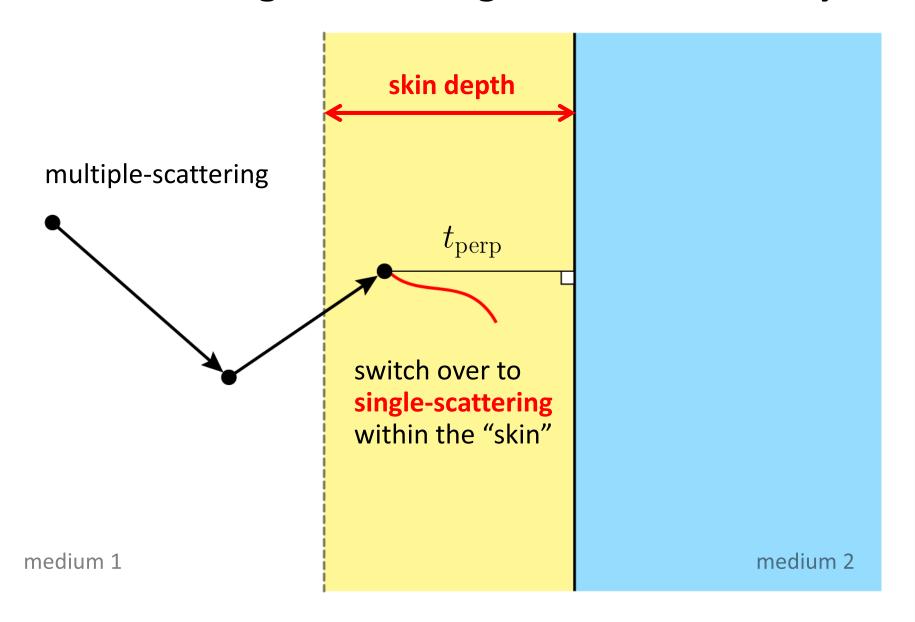
"Zeno's algorithm": particle never reaches the boundary!

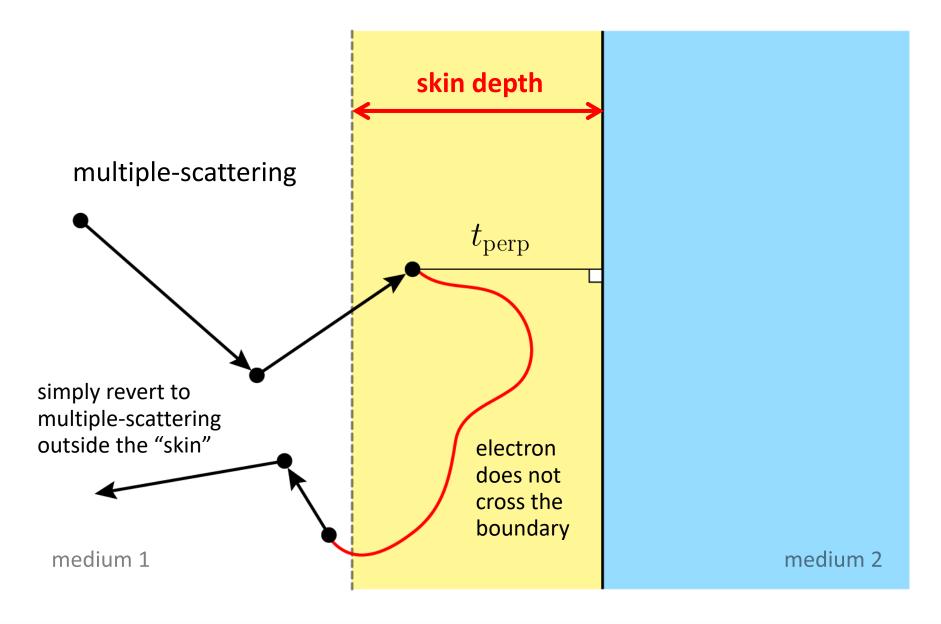


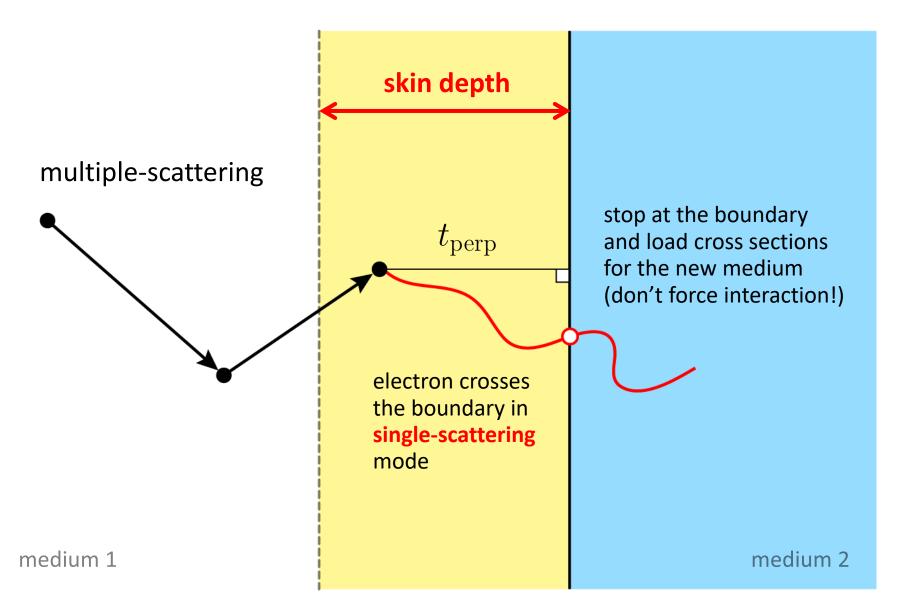
To move across the boundary, we need a boundary crossing algorithm (BCA)

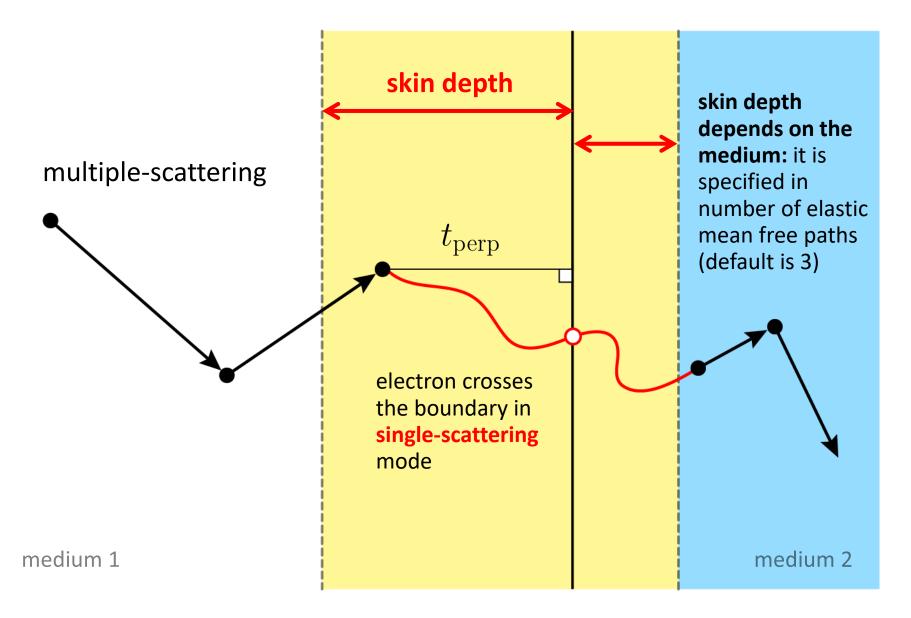
medium 1





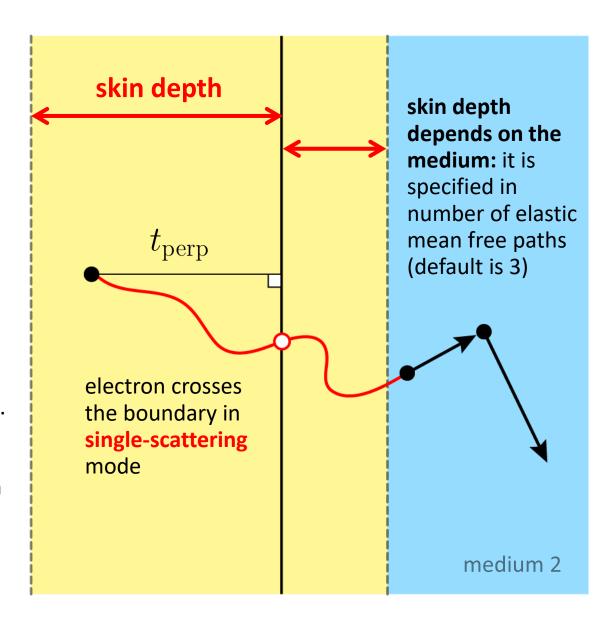






This is called the "Exact BCA" in EGSnrc

- In EGSnrc this is called the Exact BCA. There is little reason to use PRESTA-I.
- The default skin depth is
 3 elastic mean free paths
 because a single step in multiple-scattering "costs" about as much as 3 steps in single-scattering.
- To run an entire simulation in single-scattering mode, specify a "huge" skin-depth.
- EGSnrc converges to its single-scattering mode with default parameters (large multiple-scattering steps).



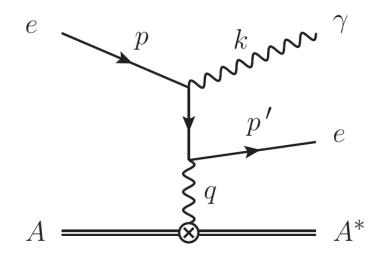
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- 2. Inelastic collision with atomic electrons (knock-on electron)
- 3. Elastic collision with nuclei and atomic electrons (deviation)
- 4. Positron annihilation (emission of a photon pair)

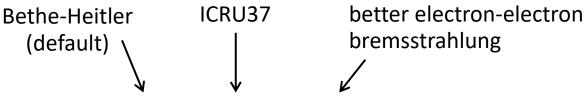
The main challenge in modelling the transport of electrons through matter is that the **cross-sections are very large**, because charged particles interact essentially "all the time" with the atoms they encounter.

Bremsstrahlung emission

- Simulated only for T > AP (k > AP)
- Restricted stopping powers for sub-threshold, "soft" bremsstrahlung
- Dominates at high energies
- Scales as Z²



Options:

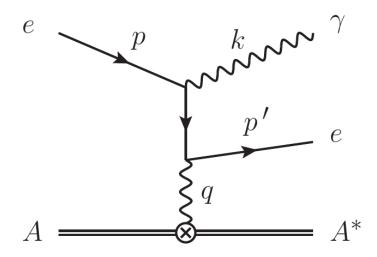


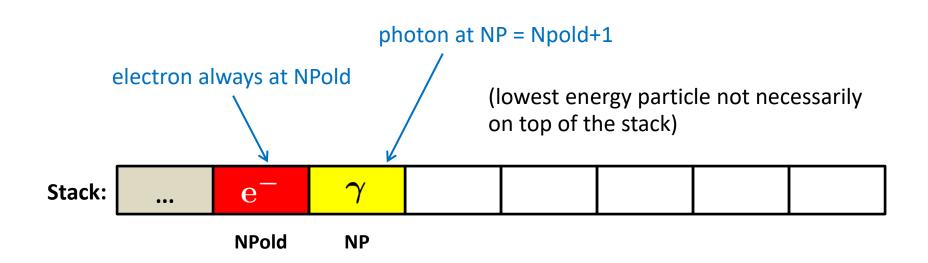
Brems cross sections = BH | NIST | NRC

Brems angular sampling = Simple | KM ← modified Koch-Motz (default)

Bremsstrahlung emission

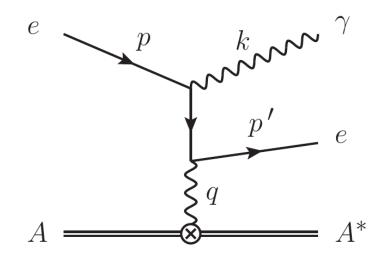
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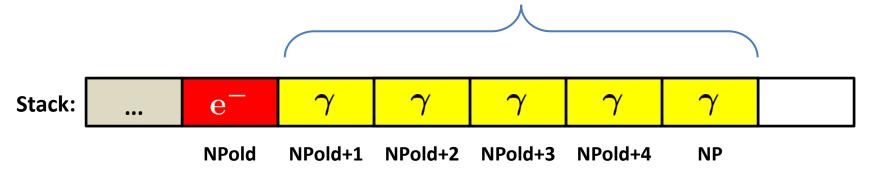


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when **splitting** is used, photons are between NPold+1 and NP (not ordered by energy)

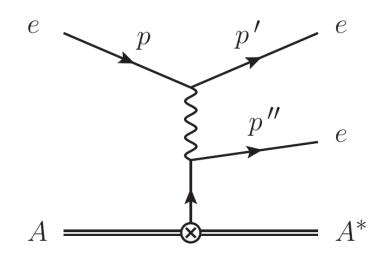


Electron inelastic collision (Møller scattering)

- Simulated only for $T > 2T_{AE}$ because electrons are indistinguishable.
- Restricted stopping powers for sub-threshold, "soft" collisions
- Generally, dominates at low and intermediate energies

NPold

NP



- Scales as Z
- Ignores binding

 secondary, lower energy electron
 is at NP = NPold+1

 Stack:

 ...

 e

 e

 e

Electron impact ionization (Eii)

Incident electron knocks out a **bound electron**, and **atom relaxes** emitting fluorescent photons, Auger and Coster-Kronig electrons. Essential for accurate kilovoltage X-ray spectra.

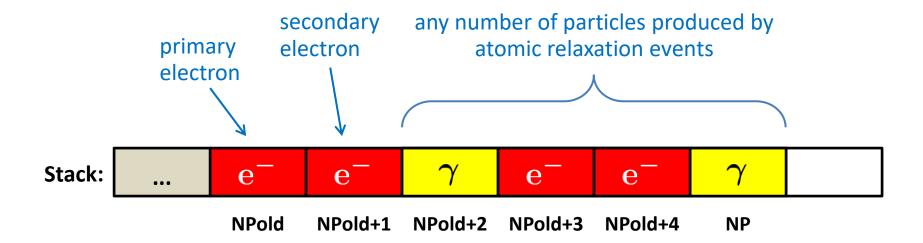
- Treated within the Møller routine, as a type of inelastic collision.
- Simulated for K and L shells with binding energies $E_b > T_{AE}$.
- When Eii is turned on, Eii interactions are based on the bound electron cross section. All other interactions are treated as Møller.

Options:

Electron impact ionization (Eii)

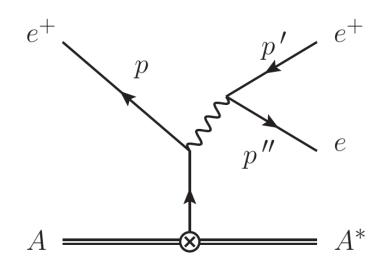
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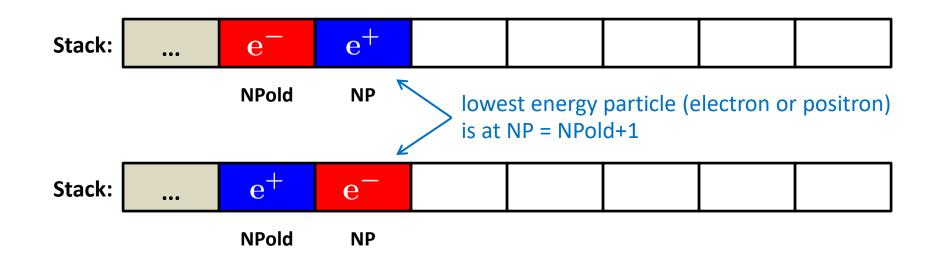
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Positron inelastic collision (Bhabha scattering)

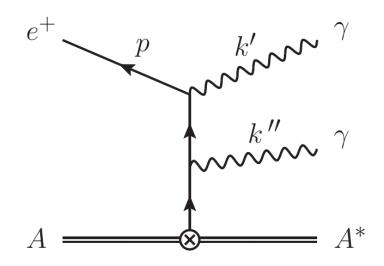
- Simulated only for $T > T_{AE}$ because particles are distinguishable.
- Generally, dominates at low and intermediate energies
- Scales as Z
- Ignores binding





Positron annihilation

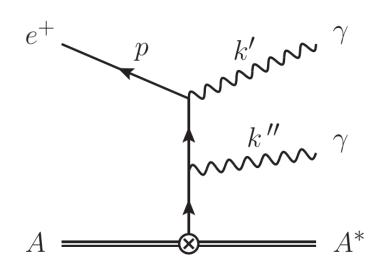
- Only two-photon annihilation is simulated, in first Born approximation.
- Two-photon annihilation is by far the dominant process.
- Important only at low energies: positrons always annihilate at rest, if not before!



Scales as Z
resulting photons at NPold and NPold+1
Stack: \(\gamma \gam

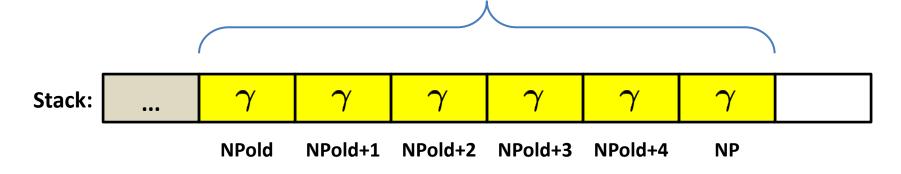
Positron annihilation

- Only two-photon annihilation is simulated, in first Born approximation.
- Two-photon annihilation is by far the dominant process.
- Important only at low energies; positrons always annihilate at rest, if not before!



Scales as Z

when **splitting** is used, there are nbr_split photon pairs on the stack



There are 3 sets of energy thresholds

1. Lower thresholds AE and AP: lowest energies of secondary electrons and photons. Below these energies, interactions are grouped as part of the continuous slowing down approximation.

secondary electron energy

		csda	secondary electrons created
0	0.51	1 4	E (512 keV minimum)

secondary photon energy

csda	secondary photons created
------	---------------------------

O **AP** (1 keV minimum)

There are 3 sets of energy thresholds

2. Upper thresholds UE and UP: upper limit on the energy of electrons and photons (simply set those above the maximum energy of source particles).

secondary electron energy

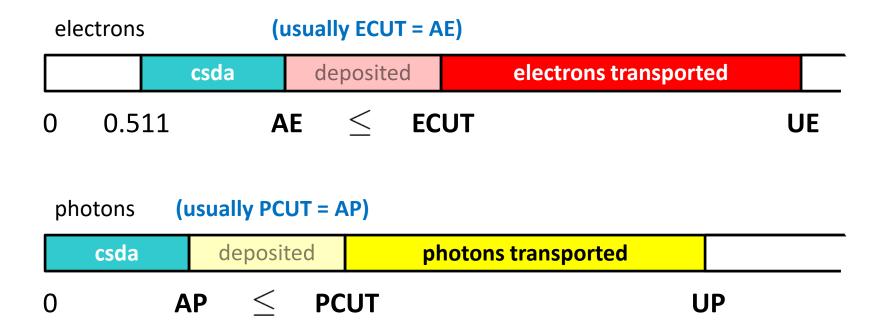
		csda		secondary electrons created		
0	0.52	11	AE		U	Ε

secondary photon energy



There are 3 sets of energy thresholds

3. Transport cutoffs ECUT and PCUT: energy threshold below which electron and photon transport is stopped, and remaining energy is deposited locally (approximation).



```
start media definition: # or in PEGS file!

AE = 0.521  # (MeV), minimum is 0.512
AP = 0.010  # (MeV), minimum is 0.001

UE = 50.511  # (MeV)
UP = 50  # (MeV)

# ...

stop media definition:
```

:start MC transport parameter:

```
Global ECUT = 0.521 # AE by default (MeV)
Global PCUT = 0.010 # AP by default (MeV)
```

:start MC transport parameter:

```
Global ECUT = 0.521 # AE by default (MeV)
Global PCUT = 0.010 # AP by default (MeV)

Spin effects = On # [On], Off
```

Spin effects incur a 30–100% penalty in simulation time. But it always has an impact on results (when electrons transport matters). You must have some very strong argument to turn this off!

:start MC transport parameter:

```
Global ECUT = 0.521  # AE by default (MeV)
Global PCUT = 0.010  # AP by default (MeV)

Spin effects = On  # [On], Off
Brems angular sampling = KM  # [KM], Simple
Brems cross sections = NRC  # [BH], NIST, NRC
```

Not the default option

BH only accurate for energies above ~ 50 MeV

:start MC transport parameter:

```
Global ECUT = 0.521 # AE by default (MeV)
Global PCUT = 0.010 # AP by default (MeV)

Spin effects = On # [On], Off

Brems angular sampling = KM # [KM], Simple

Brems cross sections = NRC # [BH], NIST, NRC

Electron Impact Ionization = On # [Off], On ...
```

Not the default option

Required for accurate characteristic X-rays

:start MC transport parameter:

```
Global ECUT
                              = 0.521 # AE by default (MeV)
Global PCUT
                              = 0.010 # AP by default (MeV)
                                        # [On], Off
Spin effects
                              = 0n
Brems angular sampling
                                        # [KM], Simple
                              = KM
Brems cross sections
                                        # [BH], NIST, NRC
                              = NRC
Electron Impact Ionization
                                        # [Off], On ...
                              = On
Electron-step algorithm
                              = EGSnrc # [EGSnrc], PRESTA-I
Boundary crossing algorithm
                              = Exact # [Exact], PRESTA-I
Skin depth for BCA
                              = 3 # [3]
FSTFPF
                              = 0.25 # [0.25]
                              = 0.5  # [0.5]
ximax
```

:stop MC transport parameter:

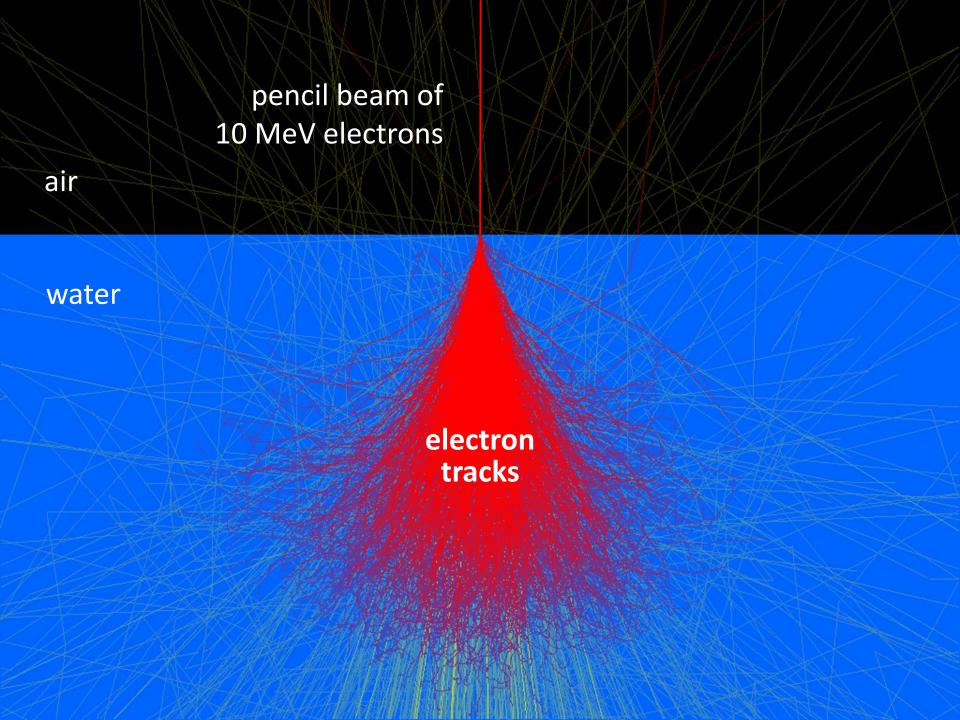
EGSnrc can model electromagnetic fields

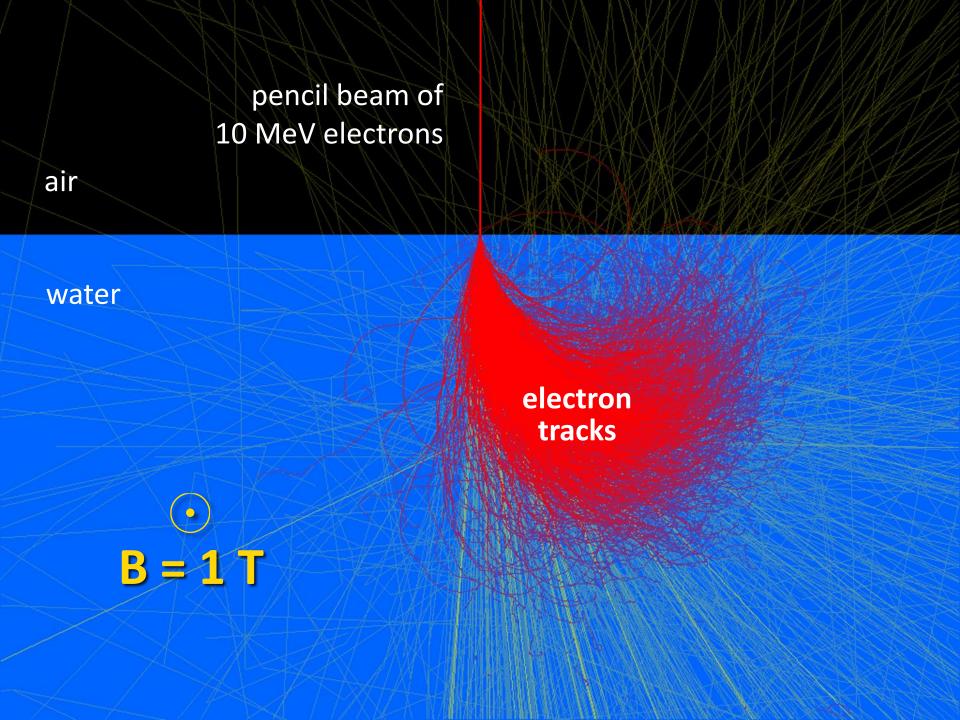
- Static magnetic and electric field components can be set by input
- Electric fields are not fully consistent (electron is discarded if its energy drops below AE upon changing direction!)
- A refined implementation has been developed by Malkov and Rogers, and will be added to EGSnrc: Med. Phys. 43(7), 4447.

:start MC transport parameter:

```
Magnetic field = 0, 0, 0.5 \# (x, y, z) components in Tesla Electric field = 0, 0.1, 0 \# (x, y, z) components in V/cm \# EM ESTEPE = 0.02 \# [0.02]
```

```
:stop MC transport parameter:
```





Suggested parameter selection strategy

- 1. Start with the **most accurate** electron transport parameters available (defaults, except NRC brems cross section, and Eii turned on if relevant).
- 2. Consider **faster options** if simulation is too long; but don't turn off spin effects!
- 3. In most situations, proper selection of AE, AP, ECUT and PCUT proves the most critical parameters affecting simulation time.
- 4. Rule of thumb: set ECUT so that the electron range at ECUT is less than 1/3 the dimension of interest (e.g., voxel size).
- 5. Remember that higher thresholds imply an approximation, so convince yourself that your results are valid by running a representative simulation with lower thresholds. (Better yet: use range rejection and variance reduction!)

Conclusion

- Analog simulation (every single interaction) of charged particle interaction is costly because cross-sections are very large.
- The condensed history technique solves the efficiency problem, by grouping particle steps into large, statistically equivalent steps.
- EGSnrc implements a class II scheme: "soft" events are grouped within the restricted stopping power, while "catastrophic" events above energy thresholds AE and AP are handled explicitly.
- Multiple-scattering introduces new complexities due to the unknown "real" particle path: continuous energy loss, electron step algorithm, and boundary crossing algorithm.
- The thresholds AE and AP, and transport cutoffs ECUT and PCUT, have the most impact on simulation efficiency (and accuracy!)