PyNeb_manual_4

April 18, 2017

0.1Recombination lines: the RecAtom class

This object is similar to **Atom**, but some methods differ due to the particularities of the recombination spectrum. It is instantiated with the following command:

```
In [1]: %matplotlib inline
        %config InlineBackend.figure_format = 'svg'
        import numpy as np
        import matplotlib.pyplot as plt
        import pyneb as pn
In [2]: H1 = pn.RecAtom('H', 1)
```

The data are read from a fits file and interpolated in temperature and density. The name of the file is retrievable through:

```
In [3]: pn.atomicData.getDataFile('H1', 'rec')
Out[3]: 'h_i_rec_SH95.hdf5'
```

In case the instantiation takes too much time (on some systems, the reading the fits fils is slow), you can use the data store in the hdf5 file using:

```
In [4]: pn.atomicData.setDataFile('h_i_rec_SH95.hdf5')
In [5]: H1 = pn.RecAtom('H', 1)
```

The ions for which recombination lines are available are listed with the command:

```
In [6]: pn.atomicData.getAllAtoms(coll=False, rec=True)
Out[6]: ['H1',
          'He1',
          'He2',
          'C1',
          'C2',
          'C3',
          'C4',
          'N1',
          'N2',
          'N3',
          'N4',
          '01',
          02',
          ,03,
          '04',
          '05',
          'N5',
```

'06']

You can list all the available data files for a given ion by:

2.85691188359

The argument used to identify the various lines in the data file depend on whether the atom is a hydrogenoid or not. In the first case, the transition can be specified either as a pair of levels **lev_i**, **lev_j**, or as a label (labels are descriptive of level pairs):

As in the case of collisional lines, setting **product=False** modifies the input array of temperature and density values:

```
In [11]: print(H1.getEmissivity([1e4, 1.8e4], [1e4, 1e2], label='4_2', product=False))
[ 1.24000000e-25     7.38740000e-26]
```

In the case of non-hydrogenoids, the transition can be specified either as a wavelength or as a label (labels are descriptive of wavelengths):

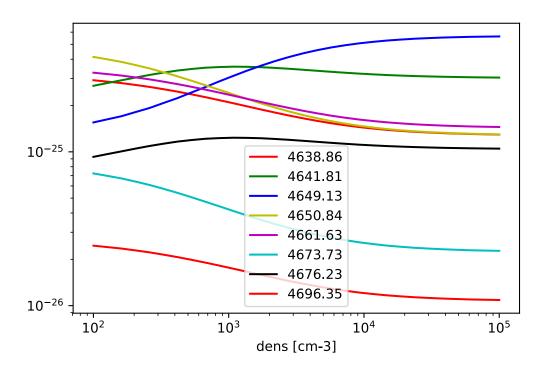
Use Atom.labels to display the valid labels for both kind of recombiation atoms, i.e.:

```
In [13]: H1.labels[0:10]
Out[13]: ('2_1', '3_1', '3_2', '4_1', '4_2', '4_3', '5_1', '5_2', '5_3', '5_4')
In [14]: He1.labels[0:10]
```

```
Out[14]: ('2945.0',
          '3188.0',
          '3614.0',
          3889.0',
          '3965.0',
          '4026.0',
          '4121.0',
          '4388.0',
          '4438.0',
          '4471.0')
   As an example application, you can easily generate a 2D table of H alpha/ H Beta as a function of Te
and Ne:
In [15]: tem = np.linspace(5000, 20000, 100)
         den = np.logspace(2, 6, 50)
         im_Hab = (H1.getEmissivity(tem, den, label='3_2') / H1.getEmissivity(tem, den, label='4_2'))
         print(im_Hab.shape)
         print(np.min(im_Hab))
         print(np.max(im_Hab))
(100, 50)
2.72452373753
3.04092769441
   By default, H1 emissivities are computed under case B; to compute Case A emissivities, the atomic data
file must be changed:
In [16]: pn.atomicData.setDataFile('h_i_rec_SH95-caseA.hdf5')
In [17]: H1_A = pn.RecAtom('H', 1)
  Note that the previously defined H1 atom is still available and use the default values.
In [18]: im_caseAoverB = (H1_A.getEmissivity(tem, den, label='4_2')/H1.getEmissivity(tem, den, label='4
   plt.imshow(im_caseAoverB) plt.colorbar();
0.1.1 Emissivities of OII recombination lines
In [20]: 02 = pn.RecAtom('0', 2)
In [29]: V1_mult = ('4638.86', '4641.81', '4649.13', '4650.84', '4661.63', '4673.73', '4676.23', '4696...
In [31]: cols = 'rgbymckrgbymckr'
         temps = 1e4
         dens = np.logspace(2, 5, 30)
         f, ax = plt.subplots()
         for j, label in enumerate(V1_mult):
             ax.loglog(dens, 02.getEmissivity(temps, dens, label=label), label=label, c=cols[j])
         ax.set_xlabel('dens [cm-3]')
```

ax.legend(loc='best')

Out[31]: <matplotlib.legend.Legend at 0x7fbd49d23a20>



0.1.2 Ionic abundances from collisional and recombination lines, ADF