

18.369 Problem Set 3 Solutions

Problem 0: (6+6+5+5+5 points)

(a) Solutions:

(i) From $\frac{\partial \psi}{\partial t} = \hat{C}\psi - \frac{\partial \phi}{\partial t} - \xi$, $\frac{\partial \psi_{\text{inc}}}{\partial t} = \hat{C}\psi_{\text{inc}} - \xi$, and $\psi = \psi_{\text{inc}} + \psi_{\text{scat}}$, it follows that

$$\frac{\partial \psi_{\text{scat}}}{\partial t} = \hat{C}\psi_{\text{scat}} - \frac{\partial \phi}{\partial t}$$

where $\phi = \chi * \psi = \chi * (\psi_{\text{inc}} + \psi_{\text{scat}})$.

Because of the dependence on ϕ , we can't solve the ψ_{scat} equation without first solving the ψ_{inc} equation. Dividing the solutions into $\psi = \psi_{\text{inc}} + \psi_{\text{scat}}$ doesn't make Maxwell's equations easier to solve, but it does give us a new perspective on the solutions (and leads to some numerical approaches such as integral-equation formulations).

(ii) The Fourier transform merely changes $\frac{\partial}{\partial t} \rightarrow -i\omega$, giving

$$\begin{aligned} \frac{\partial \hat{\psi}_{\text{inc}}}{\partial t} &= \hat{C}\hat{\psi}_{\text{inc}} - \hat{\xi} \\ \frac{\partial \hat{\psi}_{\text{scat}}}{\partial t} &= \hat{C}\hat{\psi}_{\text{scat}} - \hat{\phi}, \end{aligned}$$

where the convolution $\phi = \chi * \psi$ becomes a multiplication $\hat{\phi} = \hat{\chi}\hat{\psi}$.

(b) Note: when we apply Poynting's theorem (from section 3 of the notes) to a time-harmonic field $\psi(\mathbf{x}, t) = \hat{\psi}(\mathbf{x})e^{-i\omega t}$, then the $\frac{\partial}{\partial t} [\frac{1}{2} \langle \psi, \psi \rangle]$ term is zero. In general, the $e^{-i\omega t}$ terms cancel from all of the products thanks to the complex conjugations.

If we apply the remaining of Poynting's theorem *for the volume* Ω to the $\hat{\psi}_{\text{inc}}$ equation, there is no ϕ term in the equation and $\hat{\xi} = 0$ inside Ω so the $\langle \hat{\psi}_{\text{inc}}, \hat{\xi} \rangle$ integral vanishes as well! So the only remaining term is $-\iint_{\partial\Omega} \text{Re}[\hat{\mathbf{E}}_{\text{inc}}^* \times \hat{\mathbf{H}}_{\text{inc}}] \cdot d\mathbf{A}$ on the right-hand side, with the left-hand-side equal to zero, so $\iint_{\partial\Omega} \text{Re}[\hat{\mathbf{E}}_{\text{inc}}^* \times \hat{\mathbf{H}}_{\text{inc}}] \cdot d\mathbf{A} = 0$ as desired.

Applying Poynting's theorem to the $\hat{\psi}$ equation in Ω , the $\langle \hat{\psi}, \hat{\xi} \rangle$ term is still $= 0$ for an integral over Ω because $\hat{\xi} = 0$ there, but we *do* have a ϕ term. As shown in section 3.2 of the notes, the ϕ term in the frequency domain simplifies to $\langle \hat{\psi}, \text{Im}[\omega \hat{\chi}] \hat{\psi} \rangle$, and as we explained in the notes one always has $\text{Im}[\omega \hat{\chi}] \geq 0$ (positive semidefinite) for a passive material. Hence $\langle \hat{\psi}, \text{Im}[\omega \hat{\chi}] \hat{\psi} \rangle \geq 0$, and it follows from Poynting's theorem that

$$\langle \hat{\psi}, \text{Im}[\omega \hat{\chi}] \hat{\psi} \rangle = -\iint_{\partial\Omega} \text{Re}[\hat{\mathbf{E}}^* \times \hat{\mathbf{H}}] \cdot d\mathbf{A} = P_{\text{abs}} \geq 0$$

as desired.

(c) Consider a ball B of radius R (the interior of the sphere ∂B of radius R) centered on some point in Ω , for any R large enough so that B encloses Ω (i.e. $\Omega \subseteq B$), and apply Poynting's theorem to the volume $B \setminus \Omega$ that lies *between* $\partial\Omega$ and ∂B . Poynting's theorem immediately implies that

$$\begin{aligned} 0 &= -\iint_{\partial(B \setminus \Omega)} \text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{scat}}] \cdot d\mathbf{A} = \iint_{\partial\Omega} \text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{scat}}] \cdot d\mathbf{A} - \iint_{\partial B} \text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{scat}}] \cdot d\mathbf{A} \\ &\implies P_{\text{scat}} = \iint_{\partial\Omega} \text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{scat}}] \cdot d\mathbf{A} = \iint_{\partial B} \text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{scat}}] \cdot d\mathbf{A}, \end{aligned}$$

where $d\mathbf{A}$ pointing *out* of Ω and B for $\partial\Omega$ and ∂B , respectively. More generally, by a similar argument one can show that P_{scat} is equal to the Poynting flux through *any* surface enclosing Ω .

Since outgoing boundary conditions require $R^2 \text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{scat}}]$ to point radially outward for large R , it follows that the *integrand* $\text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{scat}}] \cdot \mathbf{n}$ of the $\oint_{\partial B}$ integral must be ≥ 0 for a sufficiently large R . If the integrand is nonnegative, then the integral $\oint_{\partial B} \text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{scat}}] \cdot d\mathbf{A}$ must also be ≥ 0 . Since $P_{\text{scat}} = \oint_{\partial B} \text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{scat}}] \cdot d\mathbf{A}$ for *all* radii R that enclose Ω , it follows that $P_{\text{scat}} \geq 0$.

(d) We just plug in the definitions in terms of Poynting-flux integrals:

$$P_{\text{ext}} = - \underbrace{\oint_{\partial\Omega} \text{Re}[\hat{\mathbf{E}}^* \times \hat{\mathbf{H}}] \cdot d\mathbf{A}}_{P_{\text{abs}}} + \underbrace{\oint_{\partial\Omega} \text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{scat}}] \cdot d\mathbf{A}}_{P_{\text{scat}}}.$$

Then, we substitute the identities $\mathbf{E} = \mathbf{E}_{\text{inc}} + \mathbf{E}_{\text{scat}}$ and $\mathbf{H} = \mathbf{H}_{\text{inc}} + \mathbf{H}_{\text{scat}}$, and note that the minus sign in the P_{abs} term leads to several cancellations. Another term vanishes due to the $\oint_{\partial\Omega} \text{Re}[\hat{\mathbf{E}}_{\text{inc}}^* \times \hat{\mathbf{H}}_{\text{inc}}] \cdot d\mathbf{A} = 0$ property that we showed above. The only remaining terms are the scat-inc cross terms:

$$P_{\text{ext}} = - \oint_{\partial\Omega} \text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{inc}} + \hat{\mathbf{E}}_{\text{inc}}^* \times \hat{\mathbf{H}}_{\text{scat}}] \cdot d\mathbf{A}.$$

(e) From the first part above (or section 3.2 of the notes), we have $P_{\text{abs}} = \text{Re}\langle \hat{\psi}, -i\omega\hat{\phi} \rangle = \text{Im}\langle \hat{\psi}, \omega\hat{\phi} \rangle$. (The absorbed power is the work done **on** the polarization currents $-i\omega\hat{\phi}$ **by** the total field $\hat{\psi}$.) In the same way, applying Poynting's theorem to $\hat{\psi}_{\text{scat}}$ yields

$$\text{Re}\langle \hat{\psi}_{\text{scat}}, -i\omega\hat{\phi} \rangle = - \oint_{\partial\Omega} \text{Re}[\hat{\mathbf{E}}_{\text{scat}}^* \times \hat{\mathbf{H}}_{\text{scat}}] \cdot d\mathbf{A} = -P_{\text{scat}}.$$

(The scattered power is the work done **by** the polarization currents $-i\omega\hat{\phi}$ **on** the scattered field $\hat{\psi}_{\text{scat}}$.) Combining these two, we obtain

$$P_{\text{ext}} = \underbrace{\text{Re}\langle \hat{\psi}, -i\omega\hat{\phi} \rangle}_{P_{\text{abs}}} - \underbrace{\text{Re}\langle \hat{\psi}_{\text{scat}}, -i\omega\hat{\phi} \rangle}_{P_{\text{abs}}} = \boxed{\text{Re}\langle \hat{\psi}_{\text{inc}}, -i\omega\hat{\phi} \rangle}.$$

Comment: The optical theorem is often phrased in terms of a “forward-scattering amplitude.” Recall that $\hat{\psi}_{\text{scat}}$ is produced by the polarization currents $\frac{\partial\hat{\phi}}{\partial t} = -i\omega\hat{\phi}$. If $\hat{\psi}_{\text{inc}}$ is a planewave $\sim e^{i\mathbf{k}\cdot\mathbf{x}}$, then $\langle \hat{\psi}_{\text{inc}}, -i\omega\hat{\phi} \rangle$ is a Fourier transform: the $e^{i\mathbf{k}\cdot\mathbf{x}}$ *Fourier component* of the polarization current. From class, a current $\sim e^{i\mathbf{k}\cdot\mathbf{x}}$ in vacuum (like the $\hat{\psi}_{\text{scat}}$ equations given $\hat{\phi}$) generates fields $\sim e^{i\mathbf{k}\cdot\mathbf{x}}$, which are the “forward-scattered” fields. So $\langle \hat{\psi}_{\text{inc}}, -i\omega\hat{\phi} \rangle$ is proportional to the amplitude of the forward-scattered fields.

Comment: It is also worth commenting on the physical intuition behind the optical theorem. The basic idea is that any scattered or absorbed light leaves a shadow, so one can determine P_{ext} by looking at the darkness of the shadow. The shadow, a region on the “forward” side where $\hat{\psi}$ is small, corresponds to a forward-scattered $\hat{\psi}_{\text{scat}}$ that partially *cancels* $\hat{\psi}_{\text{inc}}$. Hence, P_{ext} is determined by the $\hat{\psi}_{\text{scat}}$ that is generated in the forward direction (the same direction as $\hat{\psi}_{\text{inc}}$), i.e. the forward-scattering amplitude.

Problem 1: (10+15 points)

In both parts of this problem, similar to class, we need to prove that the Rayleigh quotient satisfies $\langle H, \hat{\Theta}_{\mathbf{k}} H \rangle / \langle H, H \rangle < k^2$ for some trial function H , or equivalently that

$$\int_0^a \int_{-\infty}^{\infty} (1 - \Delta) |(\nabla + i\mathbf{k}) \times \mathbf{H}_{\mathbf{k}}|^2 dx dy - k^2 \int_0^a \int_{-\infty}^{\infty} |\mathbf{H}_{\mathbf{k}}|^2 dx dy < 0$$

for the trial Bloch envelope $\mathbf{H}_{\mathbf{k}} = \mathbf{H}e^{-i\mathbf{k}\cdot\mathbf{x}}$, $\mathbf{k} = k\hat{\mathbf{x}}$, and $\varepsilon^{-1} = 1 - \Delta$.

- (a) We will choose $u(x, y) = e^{-|y|/L}$ for some $L > 0$, exactly as in class—that is, it is the simplest conceivable periodic function of x , a constant. Thus, $\int |u|^2 = 2a \int_0^\infty e^{-2y/L} dy = aL$ over the unit cell. In this case, the variational criterion above becomes, exactly as in class except for the factor of a :

$$\begin{aligned} \int_0^a \int_{-\infty}^\infty (1 - \Delta) (k^2 + L^{-2}) e^{-2|y|/L} dx dy - k^2 aL &< 0 \\ &= \frac{a}{L} - \int_0^a \int_{-\infty}^\infty \Delta \cdot (k^2 + L^{-2}) e^{-2|y|/L} dx dy, \end{aligned}$$

which becomes negative in the limit $L \rightarrow \infty$ thanks to our assumption that $\int_0^a \int_{-\infty}^\infty \Delta(x, y) dx dy > 0$. Note that the fact that $\int |\Delta| < \infty$ ensures that we can interchange the limits and integration, via the dominated convergence theorem discussed in class.

- (b) Let us guess that we can choose $u(y)$ and $v(y)$ to be functions of y only (i.e., again the trivial constant-function periodicity in x). The fact that $\nabla \cdot \mathbf{H} = 0$ implies that $(\nabla + i\mathbf{k}) \cdot [u(y)\hat{\mathbf{x}} + v(y)\hat{\mathbf{y}}] = 0 = iku + v'$, and therefore $u = iv'/k$. Therefore, it is convenient to choose $v(y)$ to be a smooth function so that u is differentiable. Let us choose

$$v(y) = e^{-y^2/2L^2}$$

in which case $u(y) = -\frac{iy}{kL^2} e^{-y^2/2L^2}$. Recall the Gaussian integrals $\int_{-\infty}^\infty e^{-y^2/L^2} dy = L\sqrt{\pi}$ and $\int_{-\infty}^\infty y^2 e^{-y^2/L^2} dy = L^3\sqrt{\pi}/2$. So, $\int |\mathbf{H}|^2 = a \int |u|^2 + |v|^2 = aL\sqrt{\pi}[1 + \frac{1}{k^2L^2}]$. Also, $(\nabla + i\mathbf{k}) \times [u(y)\hat{\mathbf{x}} + v(y)\hat{\mathbf{y}}] = (ikv - u')\hat{\mathbf{z}}$. So,

$$|\nabla \times \mathbf{H}|^2 = |(\nabla + i\mathbf{k}) \times \mathbf{H}_{\mathbf{k}}|^2 = |u'|^2 + k^2|v|^2 = k^2 \left[1 + \frac{1}{k^4L^4} \left(1 - \frac{y^2}{L^2} \right) \right] e^{-y^2/L^2}.$$

Then, if we look at our variational criterion, we have two terms: $\int |\nabla \times \mathbf{H}|^2$ and $-\int \Delta \cdot |\nabla \times \mathbf{H}|^2$. Again, we can swap limits with integration in the latter by the dominated convergence theorem. Combining the former with the $-k^2 \int |\mathbf{H}|^2$ term in the variational criterion, we get:

$$\begin{aligned} \int |\nabla \times \mathbf{H}|^2 - k^2 \int |\mathbf{H}|^2 &= a \int_{-\infty}^\infty k^2 \left[1 + \frac{1}{k^4L^4} \left(1 - \frac{y^2}{L^2} \right) \right] e^{-y^2/L^2} dy - k^2 aL\sqrt{\pi} \left[1 + \frac{1}{k^2L^2} \right] \\ &= a \int_{-\infty}^\infty \frac{k^2}{k^4L^4} \left(1 - \frac{y^2}{L^2} \right) e^{-y^2/L^2} dy - \frac{k^2 aL\sqrt{\pi}}{k^2L^2} \\ &= \frac{a}{k^2L^4} L\sqrt{\pi} \left(1 - \frac{L^2}{2L^2} \right) - \frac{a\sqrt{\pi}}{L}, \end{aligned}$$

which goes to zero as $L \rightarrow \infty$. Thus:

$$\int (1 - \Delta) |(\nabla + i\mathbf{k}) \times \mathbf{H}_{\mathbf{k}}|^2 - k^2 \int |\mathbf{H}_{\mathbf{k}}|^2 \rightarrow -k^2 \int_0^a \int_{-\infty}^\infty \Delta(x, y) dx dy < 0.$$

as $L \rightarrow \infty$. Q.E.D.

Problem 2: (5+15 points)

- (a) Maxwell's equations are (in terms of \mathbf{H}) given by the eigen-equation $\nabla \times \frac{1}{\varepsilon} \nabla \times \mathbf{H} = \frac{\omega^2}{c^2} \mathbf{H}$. Suppose that we replace ε by $\alpha\varepsilon$ where α is some constant. By inspection, one obtains the *same* eigensolution \mathbf{H} with ω replaced by $\omega/\sqrt{\alpha}$ (we just divided both sides by α). Thus, scaling epsilon everywhere by a constant just trivially scales the eigenvalues. [We could have alternatively rescaled the geometry and fields: $\varepsilon(\mathbf{x}) \rightarrow \varepsilon(\mathbf{x}\sqrt{\alpha})$ and $\mathbf{H}(\mathbf{x}) \rightarrow \mathbf{H}(\mathbf{x}\sqrt{\alpha})$.] Therefore, we can set $\varepsilon_{lo} = 1$ (that is, $\alpha = 1/\varepsilon_{lo}$), without loss of generality.

- (b) Since we are looking for “TM” solutions $E_z(x, y) = e^{ikx}E_k(y)$, i.e. with \mathbf{E} in the z direction, then we already saw from the last problem set that the eigen-equation simplifies to $-\nabla^2 E_z = \frac{\omega^2}{c^2} \epsilon E_z$, and when we plug in the e^{ikx} form we get:

$$-\frac{d^2}{dy^2} E_y = (\omega^2 \epsilon - k^2) E_y$$

(where I have chosen $c = 1$ units for simplicity).

- (i) In any region where ϵ is constant, the above equation is solved simply by sines and cosines if $\omega^2 \epsilon - k^2 > 0$ and by exponentials otherwise. Since we have a $y = 0$ mirror plane, the solutions can be chosen either even or odd, and therefore in the $|y| < h/2$ region we have solutions $E_k = A \cos(k_\perp y)$ or $A \sin(k_\perp y)$, where

$$k_\perp = \sqrt{\omega^2 \epsilon_{hi} - k^2}.$$

If k_\perp is imaginary, these become cosh and sinh solutions, but we will see below that this won't happen. In the $|y| > h/2$ region, since we are looking for solutions below the light line ($\omega^2 \epsilon_{lo} < k^2$), we must have exponentials...and requiring the solutions to be finite at infinity we must have $E_k = Be^{-\kappa y}$ for $y > h/2$ and $\pm Be^{\kappa y}$ for $y < -h/2$ (with \pm depending on whether the state is even or odd, where:

$$\kappa = \sqrt{k^2 - \omega^2 \epsilon_{lo}} = \sqrt{k^2(1-f) - k_\perp^2 f},$$

where we define $f = \epsilon_{lo}/\epsilon_{hi} < 1$ (the dielectric contrast), and we have used the definition of k_\perp from above.

- (ii) Let's consider first the *even* solutions (cosine). Continuity of E_k implies that $A \cos(k_\perp h/2) = Be^{-\kappa h/2}$, and continuity of $E'_k \sim H_x$ implies that $-k_\perp A \sin(k_\perp h/2) = -\kappa B e^{-\kappa h/2}$. Dividing these two equations, we find:

$$\tan(k_\perp h/2) = \frac{\kappa}{k_\perp} = \frac{\sqrt{k^2(1-f) - k_\perp^2 f}}{k_\perp}.$$

Similarly, for the *odd* solutions (sine), we obtain:

$$\cot(k_\perp h/2) = -\frac{\sqrt{k^2(1-f) - k_\perp^2 f}}{k_\perp}.$$

These are transcendental equations for k_\perp . We plot the left and right hand sides of these two equations in figure 1, where the intersections of the curves give the guided-mode solutions.

What about imaginary k_\perp solutions? In this case, the left hand side (tan or cot) would be purely imaginary, while the right hand side would also be purely imaginary, so it seems like there might be some such solutions. Consider the even mode (tan) equation. The tangent of an imaginary k_\perp is always imaginary with the *same* sign as the imaginary part of k_\perp , whereas the right hand side will be imaginary with the *opposite* sign ($1/i = -i$)—because of that, the two curves will *never* intersect for imaginary k_\perp and there will be no solution. Conversely for the odd-mode case. So, there are no imaginary k_\perp solutions, as promised—this means that the guided modes must always be *above* the light line for ϵ_{hi} , which makes physical sense (they must correspond to *propagating* modes in the ϵ_{hi} region and *evanescent* modes in the ϵ_{lo} regions).

- (iii) We can see immediately that the right-hand side of the transcendental equations is a real number only when $k_\perp \leq |k| \sqrt{\frac{1}{f} - 1} = k_\perp^{max}$. Furthermore, we will clearly have an intersection for *every* branch of the tangent/cotangent curve that passes through zero *before* k_\perp^{max} . The tangent curves pass through zero whenever $k_\perp h/2$ is an integer multiple of π , and the cotangent curves pass

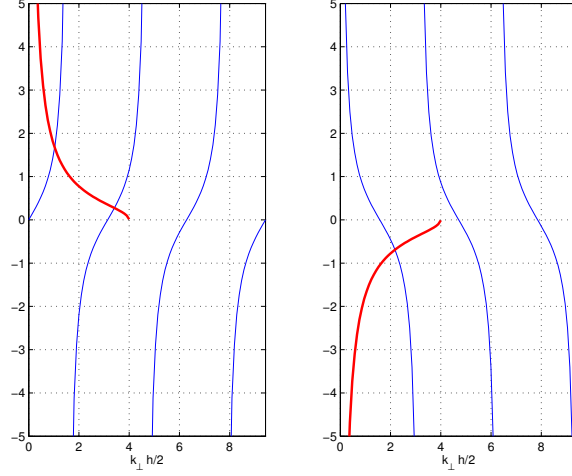


Figure 1: Plot of the two transcendental equations for even modes (left plot) and odd modes (right plot) as a function of $k_{\perp}h/2$. The thick lines show the right hand sides, while the thin lines show the left hand sides (tan or cot) of the equations, and the intersections correspond to guided-mode solutions. This plot is for the particular case of $f = 0.1$ and $kh/2 = 2$.

through zero when $k_{\perp}h/2 + \pi/2$ is an integer multiple of π . Therefore, the number of even modes is simply the number of zero crossings before k_{\perp}^{max} , namely:

$$\# \text{ even modes} = \left\lfloor \frac{|k|h\sqrt{\frac{1}{f}-1}}{2\pi} \right\rfloor + 1,$$

where the $+1$ is for the first branch of the tangent (which has a zero crossing at $k_{\perp} = 0$ and therefore *always* intersects the right-hand-side at least once). Here, by $\lfloor x \rfloor$ we mean the greatest¹ integer $\leq x$. Similarly, the number of odd modes is also given by the number of zero crossings:

$$\# \text{ odd modes} = \left\lfloor \frac{|k|h\sqrt{\frac{1}{f}-1} + \pi}{2\pi} \right\rfloor,$$

where in this case we see that we will not have *any* odd guided modes for $|k|h\sqrt{\frac{1}{f}-1} < \pi$. Therefore, as $k \rightarrow 0$ we get exactly one (even) guided mode.

Just for fun, let's look at the TE polarization (\mathbf{H} in the $\hat{\mathbf{z}}$ direction). For the $H_z = H_k e^{ikx}$ polarization, we have very similar equations except that the boundary conditions are that H_k is continuous and H'_k/ϵ is continuous (since $H'_k \sim D_x = D_{\parallel}$). Thus, for example for the $\cos(k_{\perp}y)$ mode (the *odd* mode, since \mathbf{H} is a pseudovector), we have $-k_{\perp}A \sin(k_{\perp}h/2)/\epsilon_{hi} = -\kappa B e^{-\kappa h/2}/\epsilon_{lo}$. Therefore, both the tan and cot in the transcendental equations get multiplied by $f = \epsilon_{lo}/\epsilon_{hi}$. What effect does this have on the solutions? Multiplying by $f < 1$ *decreases* the tangent curves, but does *not* change the locations of their zeros. Therefore, the *number* of modes at a given k is *unaffected*. However, the intersection point is clearly pulled towards *larger* values of k_{\perp} when the tan/cot is shrunk, which corresponds to *smaller* values of κ , the decay rate. Therefore, the modes are *less* strongly confined for the H_z (TE) polarization. (Later in the class, we will see how this generally follows from the boundary conditions and the variational theorem.)

¹There is some ambiguity about whether to define the mode as guided when the argument of $\lfloor x \rfloor$ here is exactly an integer, because that corresponds to the case where the mode is exactly on the light line and hence has $\kappa = 0$. If we don't call that a guided mode, then we have to modify our formula by one in that case, but since this situation has measure zero in the parameter space, the question has no practical significance.

Problem 3: (10+10+10 points)

- (a) Trivially from the given identity,

$$\mathbf{J} = \frac{a}{2\pi} \int_0^{2\pi/a} \left[\sum_{n=-\infty}^{\infty} \delta(x-na) \delta(y) e^{ikna-i\omega t} \hat{z} \right] dk,$$

where the term $[\dots]$ is Bloch-periodic. Because irrep is conserved, from class and homework, the resulting steady-state/time-harmonic \mathbf{E} field from each Bloch-periodic term is also Bloch-periodic (from periodicity) and TM-polarized (from $z \leftrightarrow -z$ mirror symmetry), i.e. the field of the current $[\dots]$ is:

$$\mathbf{E}_k(x, y, t) = E_k(x, y) e^{ikx-i\omega t} \hat{z},$$

where $E_k(x+a, y) = E_k(x, y)$. By linearity, we can simply add up the solutions \mathbf{E}_k from the integrand of \mathbf{J} to get the total field by superposition:

$$\mathbf{E} = \frac{a}{2\pi} \int_0^{2\pi/a} \mathbf{E}_k dk = \hat{z} \frac{a}{2\pi} \int_0^{2\pi/a} E_k(x, y) e^{ikx-i\omega t} dk.$$

The field \mathbf{E}_k satisfies $(\nabla \times \nabla \times - \omega^2 \epsilon) \mathbf{E}_k = i\omega \mathbf{J}$, and since both \mathbf{E}_k and \mathbf{J} are Bloch-periodic we can trivially reduce the domain to the unit cell $(x, y) \in [0, a] \times (-\infty, \infty)$ with Bloch-periodic boundary conditions.

- (b) For such a supercell we get a subset of the Bloch solutions, only $k_m = \frac{2\pi}{Na} m$ for integers $m = 0, \dots, P-1$, since these are the only k that satisfy periodic boundary conditions (kNa is an integer multiple of 2π). Equivalently, one can easily show:

$$\mathbf{J} = \frac{1}{M} \sum_{m=0}^{M-1} \left[\sum_{n=0}^{M-1} \delta(x-na) \delta(y) e^{ik_m na-i\omega t} \hat{z} \right],$$

$$\mathbf{E} = \hat{z} \frac{1}{M} \sum_{m=0}^{M-1} E_{k_m}(x, y) e^{ik_m x-i\omega t}$$

for the *same* E_k solutions as above.

- (c) To get the total power, we also need

$$\mathbf{H} = \frac{1}{i\omega} \nabla \times \mathbf{E} = \frac{a}{2\pi} \int_0^{2\pi/a} \mathbf{H}_k(x, y) e^{ikx-i\omega t} dk,$$

where $\mathbf{H}_k = \nabla_k \times (E_k \hat{z})$ is a periodic function. Hence, \mathbf{H} , like \mathbf{E} , is a superposition of Bloch-periodic functions. Because partners of different irreps are orthogonal, the $\int dx$ of $\hat{y} \cdot (E_k \hat{z} e^{ikx})^* \times (\mathbf{H}_{k'} e^{ik'x})$ must necessarily be zero unless $k = k'$, hence the total power P will be a superposition of terms P_k that are integrals of $\hat{y} \cdot (E_k \hat{z} e^{ikx})^* \times (\mathbf{H}_k e^{ikx})$: the Poynting flux of one k at a time.

Performing the integrals in the infinite domain is a bit tricky, so it is easier to consider the Poynting flux for our periodic supercell of N periods from the previous part. The Poynting flux is then

$$\begin{aligned} P &= \frac{1}{2} \int_0^{Na} \hat{y} \cdot \text{Re} [\mathbf{E}^*(x, y_0) \times \mathbf{H}(x, y_0)] dx \\ &= \frac{1}{2} \hat{y} \cdot \text{Re} \int_0^{Na} \left[\hat{z} \frac{1}{N} \sum_{m=0}^{N-1} E_{k_m}(x, y_0) e^{ik_m x-i\omega t} \right]^* \times \left[\frac{1}{N} \sum_{m'=0}^{N-1} \mathbf{H}_{k_{m'}}(x, y_0) e^{ik_{m'} x-i\omega t} \right] dx. \end{aligned}$$

As above, the $m \neq m'$ cross terms must integrate to zero (since they are partner functions of different irreps of the symmetry group: translations by na for $n = 0, \dots, N-1$). What remains is

$$\begin{aligned}
P &= \frac{1}{2} \hat{y} \cdot \text{Re} \frac{1}{N^2} \sum_{m=0}^{N-1} \int_0^{Na} [\hat{z} E_{k_m}(x, y_0)]^* \times [\mathbf{H}_{k_m}(x, y_0)] dx \\
&= \frac{1}{2} \hat{y} \cdot \text{Re} \frac{1}{N} \sum_{m=0}^{N-1} \int_0^a [\hat{z} E_{k_m}(x, y_0)]^* \times [\mathbf{H}_{k_m}(x, y_0)] dx \\
&= \frac{1}{N} \sum_{m=0}^{N-1} P_{k_m},
\end{aligned}$$

where we have used the periodicity of E_k and \mathbf{H}_k and defined P_k as above. Finally, by multiplying and dividing by $\Delta k = \frac{2\pi}{Na}$, we can take the $N \rightarrow \infty$ limit to recover the integral

$$P = \frac{a}{2\pi} \int_0^{2\pi/a} P_k dk.$$