

HypPy Scripting Manual¹

command-line interface (CLI)

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Abstract

This document is meant for users of the command-line applications in the Hyperspectral Python (HypPy)¹ package.

Prerequisites

The programs described in this document were developed using the Python programming language version 2.6. Check the User Manual for references to additional modules that are needed.

The core of these programs forms the `envi2` module (see the HypPy Programmers' Guide), which relies heavily on the array processing functionality of the Numerical Python (NumPy) package.

The continuum removal program uses a modified version of the Quick Hull algorithm for finding the convex hull.

Some programs use the `Image` and `ImageTk`² modules from the Python Imaging Library for reading and writing images, and `Matplotlib` (a.k.a. `PyLab`) for making 2D graphs of spectra.

The thermal correction uses the least squares estimator from the Scientific Python (SciPy) package.

¹In Finland a “hyppy” is a crest in the road that sends a car flying through the air.

²On Linux, `Image` and `ImageTk` are two separate modules!

Chapter 1

Scripting with HypPy

1.1 Example script on Linux

Example on Linux for processing one OMEGA scene:

```
#!/bin/sh

PATH=$PATH:/home/bakker/Python/HypPy

echo mask_noisy_bands
mask_noisy_bands.py -i ORB0422_4_jdat -t 25.0 -p -v

echo geo_correction
geo_correction.py -f -p -i ORB0422_4_jdat -o ORB0422_4_jdat_Gcor
-g ORB0422_4_geocube -m omega

echo solar_correction
solar_correction.py -f -i ORB0422_4_jdat_Gcor
-o ORB0422_4_jdat_Gcor_Scor -s ORB0422_4_specmars.txt

echo thermal_correction
thermal_correction.py -f -i ORB0422_4_jdat_Gcor_Scor
-o ORB0422_4_jdat_Gcor_Scor_Tcor -t ORB0422_4_jdat_Gcor_Scor_T

echo atmo_correction
atmo_correction.py -f -i ORB0422_4_jdat_Gcor_Scor
-o ORB0422_4_jdat_Gcor_Scor_Acor
-t /home/bakker/Python/HypPy/Atmosphere/transmission.dat
-a ORB0422_4_jdat_Gcor_Scor_alpha

echo logresiduals
logresiduals.py -s -b -f -i ORB0422_4_jdat_Gcor_Scor_Acor
-o ORB0422_4_jdat_Gcor_Scor_Acor_lr
-a ORB0422_4_jdat_Gcor_Scor_Acor_albedo
-r ORB0422_4_jdat_Gcor_Scor_Acor_rlub.txt -p
```

```

echo median
median.py -f -i ORB0422_4_jdat_Gcor_Scor_Acor_lr
        -o ORB0422_4_jdat_Gcor_Scor_Acor_lr_med

echo hull
hull.py -f -i ORB0422_4_jdat_Gcor_Scor_Acor_lr_med
        -o ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_cr -c 3.5

echo summary_products
summary_products.py -f -i ORB0422_4_jdat_Gcor_Scor_Acor_lr_med
        -o ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_sp
        -c ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_cr

echo tokml
tokml.py -i ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_sp
        -R 12 -G 13 -B 14

# extra stuff...
echo minwavelength
minwavelength.py -f -i ORB0422_4_jdat_Gcor_Scor_Acor_lr_med
        -o ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_wav -w 2.0 -W 2.4

echo wavemap
wavemap.py -f -i ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_wav
        -o ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_wav_map -w 2.0 -W 2.4
        -d 0.0 -D 0.15

echo Google Earth
googleearth /data/tmp/ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_sp.kml

```

1.2 Example script on Windows

Example batch file on Windows for processing one OMEGA scene:

```

rem Change the following to your Python executable
set PYTHON="O:\GroupData\ESATools\PortablePython_1.1_py2.6.1\AppData\python.exe"

rem Change the following to your HypPy directory
set HYPPY="O:\GroupData\ESATools\HypPy"

rem Change the following to your Google Earth executable
set GOOGLEEARTH="C:\Program Files (x86)\Google\Google Earth\client\googleearth.exe"

echo mask_noisy_bands
%PYTHON% %HYPPY%\mask_noisy_bands.py -i ORB0422_4_jdat -t 25.0 -p -v

echo geo_correction
%PYTHON% %HYPPY%\geo_correction.py -f -p -i ORB0422_4_jdat
        -o ORB0422_4_jdat_Gcor -g ORB0422_4_geocube -m omega

```

```

echo solar_correction
%PYTHON% %HYPPY%\solar_correction.py -f -i ORB0422_4_jdat_Gcor
    -o ORB0422_4_jdat_Gcor_Scor -s ORB0422_4_specmars.txt

echo thermal_correction
%PYTHON% %HYPPY%\thermal_correction.py -f -i ORB0422_4_jdat_Gcor_Scor
    -o ORB0422_4_jdat_Gcor_Scor_Tcor
    -t ORB0422_4_jdat_Gcor_Scor_T

echo atmo_correction
%PYTHON% %HYPPY%\atmo_correction.py -f
    -i ORB0422_4_jdat_Gcor_Scor -o ORB0422_4_jdat_Gcor_Scor_Acor
    -t transmission.dat -a ORB0422_4_jdat_Gcor_Scor_alpha

echo logresiduals
%PYTHON% %HYPPY%\logresiduals.py -s -b -f
    -i ORB0422_4_jdat_Gcor_Scor_Acor
    -o ORB0422_4_jdat_Gcor_Scor_Acor_lr
    -a ORB0422_4_jdat_Gcor_Scor_Acor_albedo
    -r ORB0422_4_jdat_Gcor_Scor_Acor_rlub.txt -p

echo median
%PYTHON% %HYPPY%\median.py -f -i ORB0422_4_jdat_Gcor_Scor_Acor_lr
    -o ORB0422_4_jdat_Gcor_Scor_Acor_lr_med

echo hull
%PYTHON% %HYPPY%\hull.py -f
    -i ORB0422_4_jdat_Gcor_Scor_Acor_lr_med
    -o ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_cr -c 3.5

echo summary_products
%PYTHON% %HYPPY%\summary_products.py -f
    -i ORB0422_4_jdat_Gcor_Scor_Acor_lr_med
    -o ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_sp
    -c ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_cr

echo tokml
%PYTHON% %HYPPY%\tokml.py
    -i ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_sp -R 12 -G 13 -B 14

rem Extra stuff...
echo minwavelength
%PYTHON% %HYPPY%\minwavelength.py -f
    -i ORB0422_4_jdat_Gcor_Scor_Acor_lr_med
    -o ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_wav -w 2.0 -W 2.4

echo wavemap
%PYTHON% %HYPPY%\wavemap.py -f
    -i ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_wav
    -o ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_wav_map -w 2.0 -W 2.4

```

```
-d 0.0 -D 0.15
```

```
echo Google Earth (untested!)
```

```
%GOOGLEEARTH% ORB0422_4_jdat_Gcor_Scor_Acor_lr_med_sp.kml
```


Chapter 2

Command Line Usage

2.1 Atmospheric Correction

Usage: `atmo_correction.py -b -f -i input -t transmission -o output`
 `-a alpha`

Atmospheric correction.

Options:

<code>-h, --help</code>	show this help message and exit
<code>-b</code>	use bad band list from the header
<code>-f</code>	force overwrite on existing output files
<code>-i INPUT</code>	input file name
<code>-t TRANS</code>	input transmission file name
<code>-o OUTPUT</code>	output file name
<code>-a ALPHA</code>	output optical depth (alpha) file name

2.2 Band Depths

Usage: `banddepths.py -s -b -f -i input -o output -d delta`

Calculate Band Depths.

Options:

<code>-h, --help</code>	show this help message and exit
<code>-s</code>	sort bands on wavelength
<code>-b</code>	use bad band list from the header
<code>-f</code>	force overwrite on existing output file
<code>-i INPUT</code>	input file name
<code>-o OUTPUT</code>	output file name
<code>-d DELTA</code>	delta, shift between bands (default=1)

2.3 Class statistics

usage: classstats.py [-h] -c CLASSNAME [-s] [-b] [-i INPUT] [-o OUTPUT] [-l SPECLIB]
[-r REPORT] [-t TEXREPORT]

Get class statistics.

optional arguments:

- h, --help show this help message and exit
- c CLASSNAME input ENVI classification file
- s sort bands on wavelength
- b use bad band list from the header
- i INPUT input ENVI image
- o OUTPUT output plot file (.png or .pdf)
- l SPECLIB output directory for ascii spectral library
- r REPORT output report (text file)
- t TEXREPORT output report (LaTeX file)

2.4 Colorize

usage: colorize.py [-h] [-s] [-b] [-f] -i INPUT1 -j INPUT2 -o OUTPUT
[-I BAND1] [-J BAND2] -w WSTART -W WEND -d DSTART -D DEND
[-c {rainbow,steps}] [-C COLORFILE] [-l]

Make color/intensity map using color table and HSV transform

optional arguments:

- h, --help show this help message and exit
- s sort bands on wavelength
- b use bad band list from the header
- f force overwrite on existing output file
- i INPUT1 input file name for color
- j INPUT2 input file name for intensity
- o OUTPUT output file name
- I BAND1 input band for color
- J BAND2 input band for intensity
- w WSTART lower stretch value for color
- W WEND upper stretch value for color
- d DSTART lower stretch value for intensity
- D DEND upper stretch value for intensity
- c {rainbow,steps} color table: rainbow (default) or steps
- C COLORFILE input color table file
- l save legend in a .png file

2.5 Convert EOS HDF

usage: converthdf.py [-h] [-f] -i INPUT -o OUTPUT

Convert OES HDF image to ENVI format.

optional arguments:

```
-h, --help  show this help message and exit
-f          force overwrite on existing output file
-i INPUT    input file name
-o OUTPUT    output file name
```

2.6 Convert to ENVI format

usage: `convert.py [-h] [-f] -i INPUT -o OUTPUT`

Convert image to ENVI format.

optional arguments:

```
-h, --help  show this help message and exit
-f          force overwrite on existing output file
-i INPUT    input file name
-o OUTPUT    output file name
```

2.7 Convert ENVI decision tree to ASCII decision tree

usage: `converttree.py [-h] [-f] [-d] -i INPUT -o OUTPUT`

Convert ENVI or ASCII Decision Tree to ASCII Decision Tree or dot graph

optional arguments:

```
-h, --help  show this help message and exit
-f          force overwrite of output file
-d          convert tree to dot graph
-i INPUT    input tree
-o OUTPUT    output tree
```

2.8 Convert / Resample Spectral Library

usage: `resamplespeclib.py [-h] [-s] [-b] [-r] [-m WMULTIPLIER] -i INPUT [-t TOSPEC] (-o OUTPUT | -e ENVIOUTPUT)`

Convert / Resample Spectral Library.

optional arguments:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-r          recursively scan input directory
-m WMULTIPLIER input wavelength multiplier
-i INPUT    input spectral library
```

```
-t TOSPEC      resample to
-o OUTPUT      output ASCII directory
-e ENVIOUTPUT  output ENVI Speclib
```

2.9 Convex Hull Removal

Usage: hull.py -s -b -f -i input -o output -m {div|sub} -c cutoff

Convex hull removal.

Options:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite on existing output file
-i INPUT    input file name
-o OUTPUT    output file name
-m MODE     mode: div (divide, default), sub (subtract)
-c CUTOFF   cutoff wavelength (or band)
```

2.10 Convolve

Usage: convolve.py -f -i input -o output -k kernel
 --bias bias --offset offset

Linear filter

Options:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite on existing output file
-i INPUT    input file name
-o OUTPUT    output file name
-k KERNEL    kernel: smooth, laplace or 9 weights
--bias=BIAS  Constant for bias (default 1.0)
--offset=OFFSET Constant for offset (default 0.0)
```

2.11 Dark & White Reference Correction

```
usage: darkwhiteref.py [-h] [-f] [-m MANIFEST] [-i INPUT] [-d DARKREF]
                        [-w WHITEREF] [-o OUTPUT]
                        [-t {uint8, int16, int32, float32, float64, complex64,
                        complex128, uint16, uint32, int64, uint64}]
                        [-r]
```

Dark & white reference correction

optional arguments:

```
-h, --help      show this help message and exit
-f             force overwrite on existing output file
-m MANIFEST     manifest xml file name
-i INPUT       input file name
-d DARKREF      dark reference file name
-w WHITEREF     white reference file name
-o OUTPUT       output file name
-t {uint8,int16,int32,float32,float64,complex64,complex128,uint16,uint32,int64,uint64}
               output data type (default float32)
-r             robust white reference
```

2.12 Decision Tree

```
usage: decisiontree.py [-h] -t TREE -o OUTPUT [-b1 image band]
                        [-b2 image band] [-b3 image band] [-b4 image band]
                        [-b5 image band] [-b6 image band] [-b7 image band]
                        [-b8 image band] [-b9 image band]
```

Execute ENVI Decision Tree

optional arguments:

```
-h, --help      show this help message and exit
-t TREE         input ENVI decision tree
-o OUTPUT       output image file name
-b1 image band  variable b1
-b2 image band  variable b2
-b3 image band  variable b3
-b4 image band  variable b4
-b5 image band  variable b5
-b6 image band  variable b6
-b7 image band  variable b7
-b8 image band  variable b8
-b9 image band  variable b9
```

2.13 Destriping and Illumination Correction

```
Usage: destripe.py -s -b -f -i input -o output -d {hor, ver}
           -m {sub, div} -p {0, 1, 2, ...} -F
```

Destriping (p=0) and illumination correction (p>0)

Options:

```
-h, --help      show this help message and exit
-s             sort bands on wavelength
-b            use bad band list from the header
-f            force overwrite on existing output file
-i INPUT       input file name
```

```
-o OUTPUT      output file name
-d DIRECTION   direction: hor (horizontal, default) or
               ver (vertical)
-m MODE        correction method: sub (subtract, default) or
               div (divide)
-p ORDER       polynomial order of correction: destriping
               (p=0, default) or illumination correction (p>0)
-F            fast version, keeps entire image in memory!
```

2.14 Hyperspectral Edge Filtering

Usage: `edgy.py -s -b -f -i input -o output -m {SAM|ED|ID|BC|SID}`

Hyperspectral edge filter

Options:

```
-h, --help    show this help message and exit
-s           sort bands on wavelength
-b           use bad band list from the header
-f           force overwrite on existing output file
-i INPUT     input file name
-o OUTPUT    output file name
-m MODE      mode: SAM (spectral angle, default),
              ED (Euclidean distance), ID (intensity difference),
              BC (Bray-Curtis), SID (spectral
              information divergence)
```

2.15 Fix Missing Data

Usage: `fixnodata.py -s -b -f -i input -o output -n string`

Set nodata values to NaN

Options:

```
-h, --help    show this help message and exit
-s           sort bands on wavelength
-b           use bad band list from the header
-f           force overwrite on existing output file
-i INPUT     input file name
-o OUTPUT    output file name
-n NODATA    list of nodata values (=string!)
```

2.16 Fix 8th pixel striping in SWIR camera

usage: `fixswir.py [-h] [-s] [-b] [-f] -i INPUT -o OUTPUT`

Fix unruly samples of SWIR camera.

optional arguments:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite on existing output file
-i INPUT    input image file name
-o OUTPUT    output image file name
```

2.17 Flip Image in X, Y or Z

usage: flip.py [-h] [-s] [-b] [-f] -i INPUT -o OUTPUT [-m {x,y,z}]

Flip image

optional arguments:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite of existing output file
-i INPUT    input image file name
-o OUTPUT    output image file name
-m {x,y,z}  flip mode: x (flip left and right (default)), y (flip up and
              down), z (flip spectrum)
```

2.18 Geo-correction

Usage: geo_correction.py -s -b -f -p -i input -o output
 -g latlon|geocube -m {latlon|omega}

Geocorrection using latlon or geocube file. OMEGA files should have 352 bands in the original order.

Options:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite on existing output file
-i INPUT    input file name
-o OUTPUT    output file name
-p          plot geographic extent
-g GEOCUBE  input latlon or geocube file name
-m MODE     mode: latlon (use latitude/longitude file, default),
              omega (use geocube)
```

2.19 GLT correction, forward and backward

usage: glt.py [-h] [-s] [-b] [-f] -i INPUT -g GLT -o OUTPUT
 [-m {forward,backward}]

Forward or backward GLT transform.

optional arguments:

```
-h, --help      show this help message and exit
-s             sort bands on wavelength
-b             use bad band list from the header
-f             force overwrite on existing output file
-i INPUT       input image file name
-g GLT         input GLT file name
-o OUTPUT      output image file name
-m {forward,backward}
                mode forward (default) or backward
```

2.20 Get Transmittance

Usage: `transmittance.py -s -b -f -i input -m mola -t transmittance`

Calculate Atmospheric Transmittance from OMEGA scene.

Options:

```
-h, --help      show this help message and exit
-s             sort bands on wavelength
-b             use bad band list from the header
-f             force overwrite on existing output file
-i INPUT       input file name
-m MOLA        MOLA input file name
-t TRANSMITTANCE transmittance output file name
```

2.21 Hyperspectral Gradient Filtering

Usage: `gradient.py -s -b -f -i input -o output`
`-m {SAM|ED|ID|BC|SID} -a`
`X Y XY U D UD XYUD E4 E8 SOBX SOBY SOBEL`

Hyperspectral edge detection and gradient filters.

Options:

```
-h, --help      show this help message and exit
-s             sort bands on wavelength
-b             use bad band list from the header
-f             force overwrite on existing output file
-i INPUT       input file name
-o OUTPUT      output file name
-a             select all gradient filters
-m MODE        mode: SAM (spectral angle, default),
                ED (Euclidean distance), ID (intensity difference),
                BC (Bray-Curtis), SID (spectral information)
```


divergence)

2.22 Linear Filter

See Convolve.

2.23 Log Residuals

Usage: logresiduals.py -s -b -f -k -i input -o output -a albedo
-r rlub -p -n stddevs

Normalize image data using Log Residuals or Kwik Residuals

Options:

- h, --help show this help message and exit
- s sort bands on wavelength
- b use bad band list from the header
- f force overwrite on existing output file
- i INPUT input file name
- o OUTPUT output file name
- k use kwik residuals instead of log residuals
- a ALBEDO albedo output file name
- r RLUB rlub output file name (text file)
- p plot RLUB
- n N number of standard deviations for maximum
(default=3.0)

2.24 Make Legend

usage: makelegend.py [-h] -i INPUT [-u]

Create a legend for ENVI Classification image

optional arguments:

- h, --help show this help message and exit
- i INPUT input file
- u Suppress multiple unclassified legend entries

2.25 Mask Noisy Bands

Usage: mask_noisy_bands.py -v -p -i input -t threshold

Mask noisy bands in bad band list (BBL) of the header.

Options:

- h, --help show this help message and exit
- i INPUT input file name

```
-v          verbose: print useful info
-p          plot signal to noise ratio and threshold
-t THRESHOLD threshold for bad bands
```

2.26 Hyperspectral Median Filter

Usage: median.py -f -i input -o output -m {med7, med27}

3D-median filter, 7-neighborhood and 27-neighborhood

Options:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite on existing output file
-i INPUT    input file name
-o OUTPUT    output file name
-m MODE     mode: med7 (7-neighborhood, default) or
            med27 (27-neighborhood)
```

2.27 Wavelength of Minimum

Usage: minwavelength.py -b -f -i input -o output -m {div|sub|none}
 -w startwav -W endwav

Determine wavelength of minimum plus other parameters

Options:

```
-h, --help  show this help message and exit
-b
-f
-i INPUT
-o OUTPUT
-w START
-W END
-m MODE
```

2.28 Normalize Bands

Usage: normalize.py -s -b -f -i input -o output -a stddevs

Normalize bands.

Options:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite on existing output file
```

```
-i INPUT      input file name
-o OUTPUT     output file name
-a ADD_STDEV  add standard deviations (default=0.0)
```

2.29 Band Ratio

Usage: ratio.py -s -b -f -i input -o output -w numerator -W denominator

Calculate Band Ratio.

Options:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite on existing output file
-i INPUT    input file name
-o OUTPUT    output file name
-w WAV1     wavelength (or band) of numerator
-W WAV2     wavelength (or band) of denominator
```

2.30 Band Ratios

Usage: ratios.py -s -b -f -i input -o output -d delta

Calculate Band Ratios.

Options:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite on existing output file
-i INPUT    input file name
-o OUTPUT    output file name
-d DELTA    delta, shift between bands (default=1)
```

2.31 Replace Values

usage: replace_values.py [-h] [-s] [-b] [-f] -i INPUT -o OUTPUT -n NODATA
[-v NEWDATA]

Replace nodata values to new value

optional arguments:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite on existing output file
-i INPUT    input file name
```

```
-o OUTPUT    output file name
-n NODATA    list of nodata values (=string!)
-v NEWDATA   new output value(s), can be NaN or another image file
```

2.32 Resample Image

usage: resample.py [-h] [-s] [-b] [-f] -i INPUT -o OUTPUT [-p STEPSIZE]

Resampling using step size

optional arguments:

```
-h, --help    show this help message and exit
-s            sort bands on wavelength
-b            use bad band list from the header
-f            force overwrite on existing output file
-i INPUT      input file name
-o OUTPUT     output file name
-p STEPSIZE   stepsize (default 2)
```

2.33 Smile Measurement

usage: smile.py [-h] [-s] [-b] -i INPUT -o OUTPUT

Determine smile

optional arguments:

```
-h, --help    show this help message and exit
-s            sort bands on wavelength
-b            use bad band list from the header
-i INPUT      input image file name
-o OUTPUT     output smile file name
```

2.34 Solar Correction

Usage: solar_correction.py -b -f -i input -o output -s solarspec

Solar correction using Solar spectrum.

Options:

```
-h, --help    show this help message and exit
-b            use bad band list from the header
-f            force overwrite on existing output file
-i INPUT      input file name
-o OUTPUT     output file name
-s SOLARSPEC  input Solar spectrum file name
```

2.35 Sort Bands by Wavelength

Usage: sortchannels.py -b -f -i input -o output

Sort bands by wavelength.

Options:

- h, --help show this help message and exit
- b use bad band list from the header
- f force overwrite on existing output file
- i INPUT input file name
- o OUTPUT output file name

2.36 Spatial Binning

usage: spatialbinning.py [-h] [-s] [-b] [-f] -i INPUT -o OUTPUT [-x BINSIZE]

Spatial binning using bin size

options:

- h, --help show this help message and exit
- s sort bands on wavelength
- b use bad band list from the header
- f force overwrite on existing output file
- i INPUT input file name
- o OUTPUT output file name
- x BINSIZE bin size (default 3)

2.37 Spatial Spectral Binning

usage: spatialspectralbinning.py [-h] [-s] [-b] [-f] -i INPUT -o OUTPUT
[-x XYBINSIZE] [-z ZBINSIZE]

Spatial spectral binning using xy and z bin sizes

options:

- h, --help show this help message and exit
- s sort bands on wavelength
- b use bad band list from the header
- f force overwrite on existing output file
- i INPUT input file name
- o OUTPUT output file name
- x XYBINSIZE xy bin size (default 3)
- z ZBINSIZE z bin size (default 3)

2.38 Spectral Binning

usage: spectralbinning.py [-h] [-s] [-b] [-f] -i INPUT -o OUTPUT [-z BINSIZE]

Spectral binning using bin size

options:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite on existing output file
-i INPUT    input file name
-o OUTPUT    output file name
-z BINSIZE  bin size (default 3)
```

2.39 Spectral Gradient

usage: spectralgradient.py [-h] [-b] [-s] [-f] -i INPUT -o OUTPUT

Calculate spectral gradient

optional arguments:

```
-h, --help  show this help message and exit
-b          use bad band list
-s          sort wavelengths
-f          force overwrite of output file
-i INPUT    input file
-o OUTPUT    output file
```

2.40 Spectral Math

usage: specmath.py [-h] [-s] [-b] -o OUTPUT [-t DATA_TYPE] -e EXPRESSION
image [image ...]

Spectral math

positional arguments:

```
image          input filenames
```

optional arguments:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-o OUTPUT    output image file name
-t DATA_TYPE output data type (double, float32, int32, uint16, ...)
-e EXPRESSION Python expression
```

2.41 Split into Separate Bands

Usage: split.py -s -b -i input -m {ENVI|JPEG|TIFF|PNG}

Split into separate bands.

Options:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-i INPUT    input file name
-m MODE     output format: ENVI, JPEG (default), TIFF, PNG
```

2.42 Statistics

```
usage: statistics.py [-h] [-b] [-s] [-hist] [-stats] -i INPUT [-band BAND]
                  [-w WAVELENGTH] [-bandname BANDNAME] [-n NUMBINS]
```

Calculate statistics

optional arguments:

```
-h, --help          show this help message and exit
-b                 use bad band list
-s                 sort wavelengths
-hist              calculate histogram
-stats             calculate statistics
-i INPUT           input file
-band BAND         select band
-w WAVELENGTH      select band by wavelength
-bandname BANDNAME select band by bandname
-n NUMBINS         number of bins to use for histogram (default 256)
```

2.43 Subset Image

```
Usage: subset.py -s -b -f -i input -o output -x x0 -X x1
              -y y0 -Y y1 -B bandlist -m {bip|bil|bsq}
```

Make subset

Options:

```
-h, --help  show this help message and exit
-s          sort bands on wavelength
-b          use bad band list from the header
-f          force overwrite on existing output file
-i INPUT    input file name
-o OUTPUT   output file name
-x X0       first sample
-X X1       last sample
-y Y0       first line
-Y Y1       last line
-B BANDLIST band list (string!),
            e.g. "<2 4 6-8 >10" for 0 1 2 4 6 7 8 10 11 12...
```

-m MODE format of output file: bip, bil or bsq (default)

2.44 Summary Products, Vivian Beck and other indices

Pelkey Summary Products

Usage: summary_products.py -s -b -f -i input -c continuumremoved
 -o output -u {mic|nan} -l

Summary Products, mineral and other indices.

Options:

-h, --help show this help message and exit
 -s sort bands on wavelength
 -b use bad band list from the header
 -f force overwrite on existing output files
 -i INPUT input file name
 -c CR input continuum removed file name
 -o OUTPUT output file name
 -u UNITS input wavelength units: mic (default, micrometers)
 or nan (nanometers)
 -l create a logfile

Viviano-Beck Summary Products

usage: viviano_beck.py [-h] [-s] [-b] [-f] -i INPUT -o OUTPUT [-u {mic,nan}]
 [-l]

Summary Products, mineral and other indices.

optional arguments:

-h, --help show this help message and exit
 -s sort bands on wavelength
 -b use bad band list from the header
 -f force overwrite on existing output files
 -i INPUT input file name
 -o OUTPUT output file name
 -u {mic,nan} input wavelength units: mic (default, micrometers) or nan
 (nanometers)
 -l create a logfile

Other Indices

usage: otherindices.py [-h] [-s] [-b] [-f] -i INPUT -o OUTPUT [-u {mic,nan}]
 [-l]

Summary Products, mineral and other indices.

optional arguments:

```
-h, --help    show this help message and exit
-s           sort bands on wavelength
-b           use bad band list from the header
-f           force overwrite on existing output files
-i INPUT     input file name
-o OUTPUT     output file name
-u {mic,nan} input wavelength units: mic (default, micrometers) or nan
              (nanometers)
-l           create a logfile
```

2.45 Thermal Correction

Usage: `thermal_correction.py -b -f -i input -o output -t thermal`

Thermal correction.

Options:

```
-h, --help    show this help message and exit
-b           use bad band list from the header
-f           force overwrite on existing output files
-i INPUT     input file name
-o OUTPUT     output file name
-t THERMAL   output thermal file name
```

2.46 Convert to PNG, PGW and KML

Usage: `tokml.py -s -b -e -z -i input -R red -G green -B blue
-m {NO|MM|1P|SD}`

Convert data to .png, .pgw and .kml in one go

Options:

```
-h, --help    show this help message and exit
-s           sort bands on wavelength
-b           use bad band list from the header
-i INPUT     input file name
-e           strip edges
-z           strip zeros
-R RED       red band, wavelength or band number
-G GREEN     green band
-B BLUE     blue band
-m MODE      stretch mode: NO (none), MM (min-max),
              1P (1 percent, default), SD (2 standard deviation)
```

2.47 Get Transmittance

See Get Transmittance.

2.48 Wavelength Mapping

Usage: wavemap.py -f -i input -o output -w startwav -W endwav
 -d startdepth -D enddepth
 -c {rainbow|steps} -l

Make wavelength/depth map using color table and HSV transform

Options:

-h, --help	show this help message and exit
-f	force overwrite on existing output file
-i INPUT	input file name
-o OUTPUT	output file name
-w WSTART	lower wavelength
-W WEND	upper wavelength
-d DSTART	lower depth value
-D DEND	upper depth value
-c COLORTABLE	color table: rainbow (default) or steps
-l	save legend in a .png file

2.49 WMS Get

Usage: wmsget.py -u url -l layer -s layerstyle -p reference
 -x centerx -y centery -r resolution
 -i imageformat -a width -b height -o output -e

Get image via Web Map Server (WMS).

Options:

-h, --help	show this help message and exit
-u URL	URL of the WMS
-l LAYER	selected layer
-s STYLE	selected layer style
-p SRS	selected spatial reference, for instance "EPSG:28992"
-x CENTERX	selected center x coordinate
-y CENTERY	selected center y coordinate
-r RESOLUTION	required resolution
-i IMFORMAT	output image format, typically "image/jpeg" etc.
-a WIDTH	output width in pixels
-b HEIGHT	output height in pixels
-o OUTPUT	input file name
-e	flag for also generating ENVI format image

2.50 Zonal Statistics

usage: zonalstatistics.py [-h] -z ZONES -i INPUT [-b] [-s]
 [-m {max,mean,median,min,std,sum,m-2s,m+2s}]

Calculate zonal statistics (as tables)

optional arguments:

-h, --help	show this help message and exit
-z ZONES	zones file
-i INPUT	input file
-b	use bad band list (on input)
-s	sort wavelengths (on input)
-m {max,mean,median,min,std,sum,m-2s,m+2s}	method (default 'mean')