

# Muon Physics Prelab

Sunday, May 12, 2019

12:21 PM

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muon: basic particle, unstable heavy charge, denote as  $\mu$

Source of muon:

Cosmic rays collide with high energy particles in atmosphere.

$\pi^- \rightarrow \mu^- \bar{\nu}_\mu$ ,  $\mu^-$  muon, travel through atmosphere and lose kinetic energy

weak and electromagnetic interaction

produced:  $15 \text{ km/c}$   $4 \text{ GeV}$   $1/\text{cm}^2 \cdot \text{minute}$  sea level  $\left\{ \begin{array}{l} \text{spontaneous decay} \\ \text{time dilation.} \end{array} \right.$

muon decay math: exponential.

$$\int \frac{-dN}{N_0} = \lambda \exp(-\lambda t) dt$$

$$\Rightarrow D(t) = \lambda \exp(-\lambda t), \tau = 1/\lambda$$

detector PMT photomultiplier.

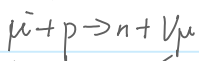


scintillator

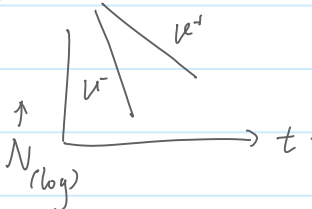
$\downarrow$  (P) may particles that can excite the scintillator, not only muon  
ionization, atomic excitation  $\rightarrow$  emit blue/UV light

PMT: photo multiplier tube, decay: muon  $\rightarrow$  electron neutrino, anti-neutrino

proton  $\leftarrow$  from any nuclei



$\mu^-$  life less than  $\mu^+$ ,  $\propto Z^4$



$$\tau_{obs} = \frac{\tau^- N^- + \tau^+ N^+}{N^+ + N^-}$$

$$\rho = \frac{\tau^+}{\tau^-} \left( \frac{\tau^- - \tau_{obs}}{\tau^+ - \tau_{obs}} \right)$$

stop  $\rightarrow$  decay  
 $\downarrow$  signal  $\downarrow$  signal  
 $\Delta t$  time interval

Fermi Couple Constant  $G_F$

$\mu$  decay via weak effect

$$\text{lifetime } \tau = \frac{192 \pi^3 \hbar}{G_F^2 m^5 c^{10}}$$

$G_F$  (from PDG)

# Muon Physics Prelab

Sunday, May 12, 2019

3:42 PM

a) typical muon energy for the detector size: 160 MeV

b) energy decay rate

$$S = \rho x \quad \text{thickness } g/cm^2.$$

Bethe-Bloch equation.

$$\frac{dE}{dx} = - \frac{4\pi N_e}{mc^2 \beta^2} Z^2 \left( \ln \frac{mc^2 \beta^2 \gamma^2}{I} - \beta^2 \right) \approx 2 \text{ MeV}/(g/cm^2) \cdot \rho_{air}$$

energy loss with height

$N_e$ : electron density,  $Z$ : atomic number,  $\beta$ : velocity of projectile,  $\gamma = 1/\sqrt{1-\beta^2}$ ,  $I$ : excitation energy of medium atom

$$\tau' = \int dt \quad \text{proper time}$$

$$= \int_H^0 \frac{dh}{c\beta(h)\gamma(h)} \quad | \quad dt = \frac{dh}{c\beta(h)}$$

$$\frac{dE}{dh} = \rho C_0 \Rightarrow dE = \rho C_0 dh.$$

$$E = mc^2 \gamma \quad dE = mc^2 d\gamma.$$

$$\Rightarrow dh = \frac{dE}{\rho C_0} = \frac{mc^2}{\rho C_0} d\gamma.$$

$\left( \frac{1}{\gamma} \right)$  average thickness of detector  $S \approx 10 g/cm^2$ .

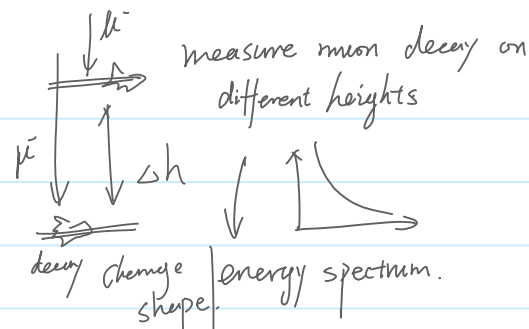
$$\tau = \frac{mc}{\rho C_0} \int_{\gamma_1}^{\gamma_2} \frac{d\gamma}{\beta \gamma} \quad \begin{cases} \gamma_2: \text{before entering the detector. } \gamma_2 \approx 1.5 \\ \gamma_1: \text{at height } H \text{ in atmosphere} \end{cases} \quad E_1 = E_2 + \Delta E, \quad E_2 = 160 \text{ MeV}$$

Hence the stopping rate  $R$  of muon is different at different height.

$$R = \exp(-t'/\tau) \quad \tau = \text{lifetime}$$

Measurement 1: muon life time.

Measurement 2: flux of muon at sea level.



$\Rightarrow$  lose energy when travelling vertically

energy loss: 2 MeV/g/cm<sup>2</sup>.

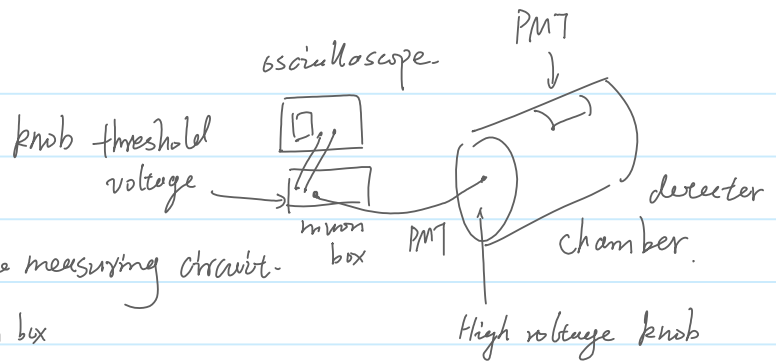
$$\text{here: } \rho_{air} = - \frac{dP}{dh}, \quad P = P_0 \exp(h/h_0) \quad \begin{cases} P_0 = 1030 g/cm^2 \\ h_0 = 8.4 \text{ km} \end{cases}$$

# Lab Note Setup

Monday, May 13, 2019

1:15 PM

Partner: Moshirfatemi Nastazia



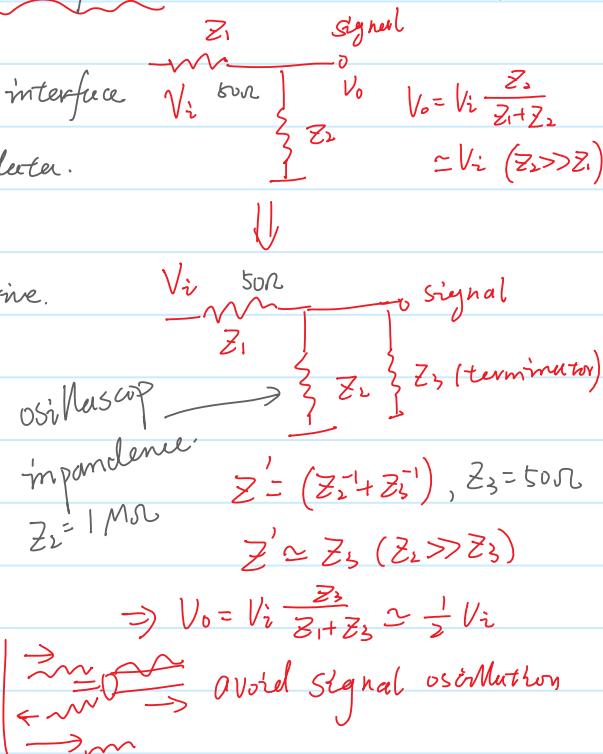
- 1) check all apparatus.
- 2) connect wires and tables to build the measuring circuit.
  - connect PMT from detector to muon box
  - connect the serial output to computer to collect data
- 3) set voltage, HV at 1kV-12kV, threshold at 180-220 mV

- 4) install terminator on the BNC cable, set a 50Ω impedance terminator is important
- 5) open muon-physics software and learn about the interface
- 6) set the configuration and start recording data.
  - config: com3 port, 6uSec, 20 bin
- 7) save data file and upload to google drive.

Some problem:

The muon detected frequency should be 6Hz at proper setting. However, the detected muon number are 0-3 every second. much less than expected

possible reason: noise from cosmic rays and fluctuation of electronics



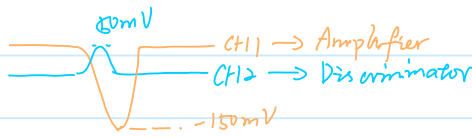
observation of pulse signal:

- 1) connect PMT cable to muon box
- 2) connect AMPLIFIER and DISCRIMINATE output to channel 1 and channel 2 of oscilloscope
- 3) make sure that CH1 and CH2 input cable are installed with terminator to achieve a 50Ω impedance.
- 4) set vertical scale at about 50 mV, trigger level -17 mV, trigger type "pulse".

# Fitting Method and Error

Wednesday, May 15, 2019

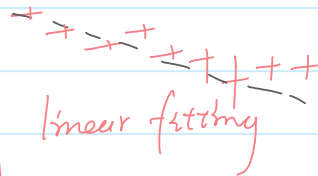
4:23 PM



⇒ important

the actual peak value is 150 mV < 180 ~ 220 mV, smaller  
the recommended threshold from Lab manual  
most muon events may be filtered out by the overestimated threshold value

counts. (log scale)



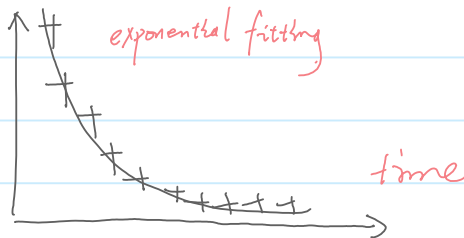
error of histogram of decay time.

1. for each bin in histogram, if the count number in that bin is  $N$ , the uncertainty should be  $\sqrt{N}$

2. horizontal error should be the standard deviation of the data allocated to the bin, plus the least resolution of measurement (20 nSec)  $(\sigma_{bin}^2 + \sigma_{min-e})^{\frac{1}{2}}$

Counts

exponential fitting



$$f(t) = A \exp(-\lambda t) + B$$

$\lambda$ : decay rate  $B$ : offset value of noise (should be 0 in ideal condition)

3. two ways of fitting

① non-linear fitting with exponential decay model

② linear fitting on log scale data

space overage  
muon life time

$$T = 1/\lambda$$

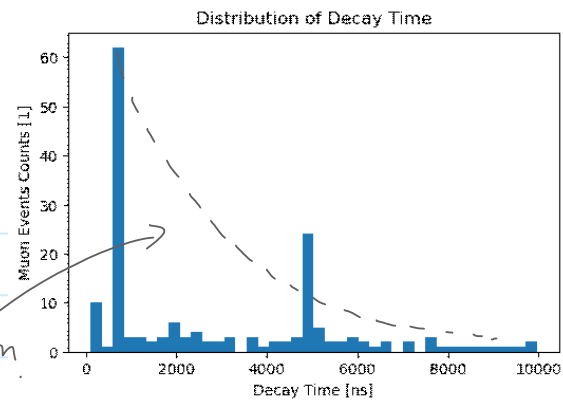
The error bar in log scale is not symmetric, it actually looks shorter at larger value

# First Data Analysis

Wednesday, May 15, 2019 4:23 PM

Problems of the first-day data

1) completely not a exponential distribution



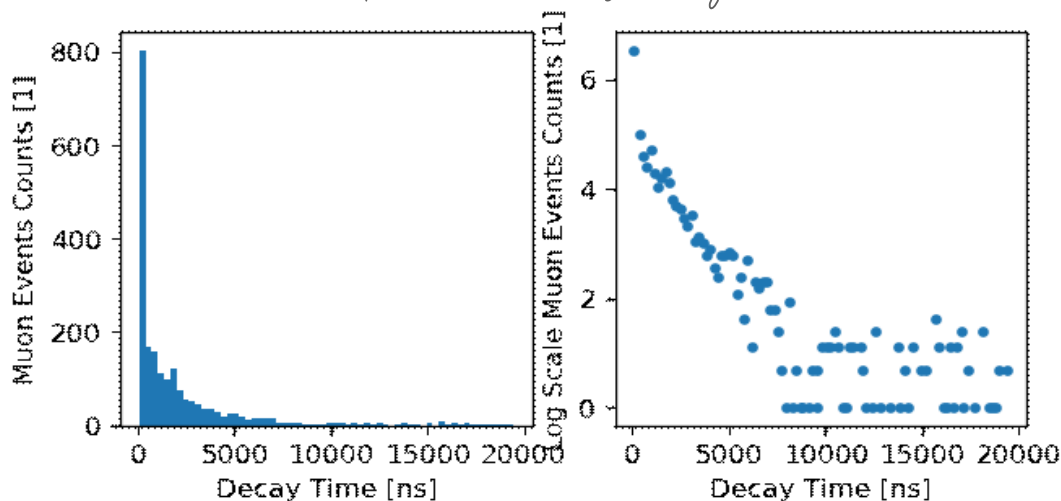
muon events: 171, total events: 47484

most events are filtered out by the overestimated threshold

the remained data is something like noise or other kind of cosmic particles

Solution: reset threshold voltage, set to 110~120mV

restart data record process, turn high voltage to 1200V (max)



A. the figure shows clearly that the impact of noise is more significant at larger time zone, where the effective count of muons is relatively smaller. The long-tail data make no sense because the minimum count of events is 1 and the probability of muon decay is very low. Thus even one noise event will result in high fluctuation.

B. The peak at short time zone should be cut off. The anomaly here is caused by possible noise and the limited precision of apparatus. Since the resolution of measurement is 20ns, all the muons decay in 20ns are counted in the same bins. And that's why the first bin is much higher than other part.

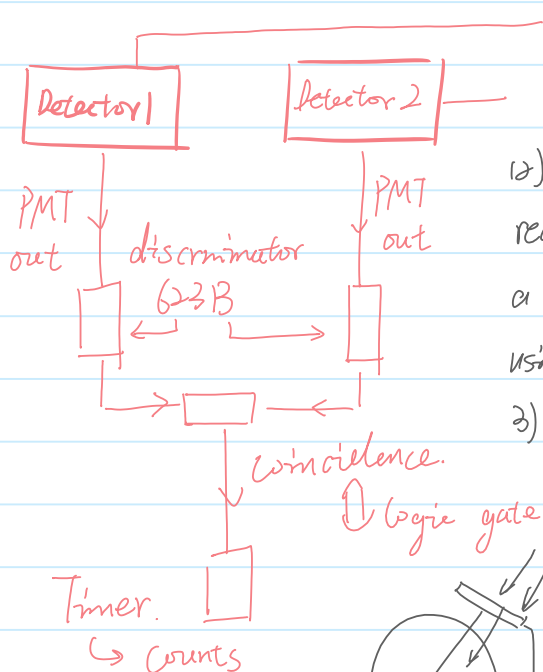
To correctly fit the exponential decay curve, the long tail and the first bin should be abandoned.

# Flux Measurement Setup

Wednesday, May 22, 2019

4:23 PM

Measure muon flux using two detectors array.

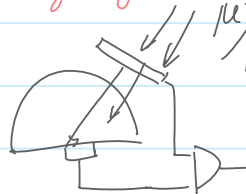


High Voltage  
DC Power.

1) the detectors are powered by high voltage DC power

2) the PMT send signal to discriminators. when the signals reach the threshold voltage, the discriminator will send a square wave pulse. width can be adjusted using the skew driver.

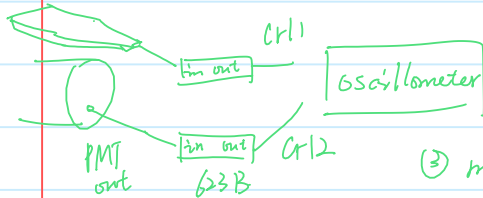
3) the coincidence is actually a logical AND gate, that means only when both two detectors sending a signal above threshold will the final counter receive a counting signal.



the PMT is triggered when particles excited a blue flash the logical AND makes sure that only particles crossing

both detector can be counted, which filter out most of the random noise happened in the detectors (consider the probability that two irrelevant random noise happen simultaneously is very small)

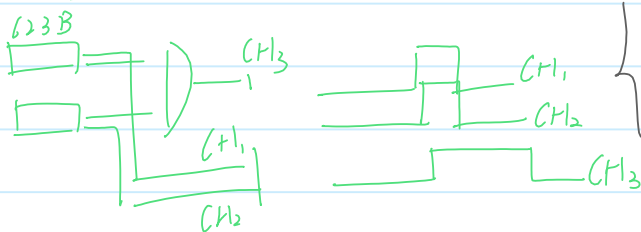
- check discriminator. ① set the threshold at minimum.



② PMT out  $\rightarrow$  discriminator  $\rightarrow$  oscilloscope. check CH1 and CH2. edge trigger. DC coupling. result like this

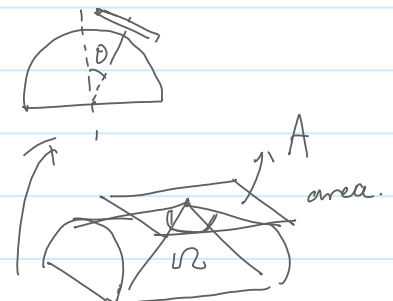
③ measure minimum threshold, ground and the hole in panel, both 29.3mV

- check coincidence



angle.  $\theta_1, \theta_2, \theta_3$   
count  $n_1, n_2, n_3$ .  
time  $t_1, t_2, t_3$ .

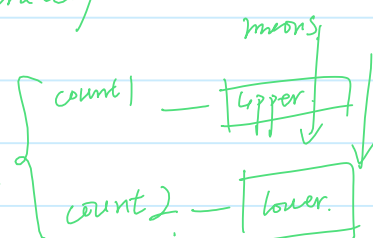
$$\frac{n_i/t_i \cdot A \cdot \Omega}{\text{flux}}$$



- check counts separately

count 2 > count 1

larger detection area.



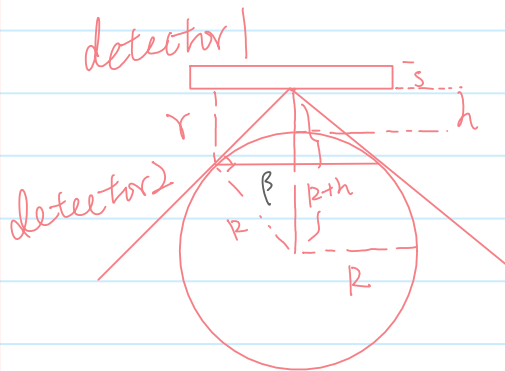
Set PMT I and PMT II parallel at zenith angle  $\theta$ . change location of the PMT II to measure flux at different direction

# Geometry and Solid Angle

Wednesday, May 22, 2019

3:39 PM

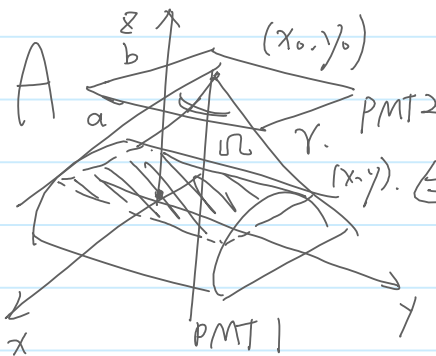
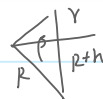
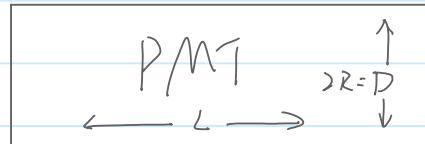
## Measurement of muon flux



$$\begin{cases} r = \frac{R}{R+h} \sqrt{(R+h)^2 - R^2} \\ \beta = \frac{(R+h)^2 - R^2}{R+h} \end{cases}$$

$$c^2 = \left(\frac{b-d}{2}\right)^2 = h^2$$

$$A = ab + \frac{(b+d)}{2} \sqrt{c^2 - \left(\frac{b-d}{2}\right)^2}$$

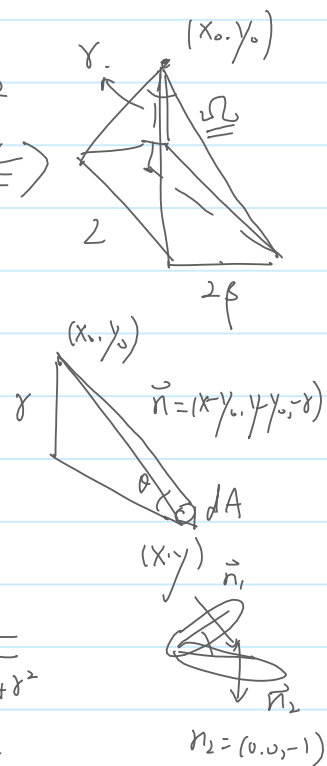


$$dA = dx dy$$

$$r = \sqrt{(x-x_0)^2 + (y-y_0)^2 + h^2}$$

$$dA_{eff} = dA \cos \theta$$

$$\cos \theta = \frac{|\vec{n}_1 \cdot \vec{n}_2|}{|\vec{n}_1| |\vec{n}_2|} = \frac{r}{\sqrt{(x-x_0)^2 + (y-y_0)^2 + h^2}}$$



R[cm]	16.5	16.5	16.5	16.4
h[cm]	6.5	6.4	6.4	6.5
s[cm]	2.7	2.8	2.8	3
a[cm]	10.0	10.1	10.0	
b[cm]	10.0	10.2	10.0	
c[cm]	7.5	7.2	7.3	
d[cm]	5.5	5.3	5.0	
L[cm]	36.2	35.9	35.9	36.0
m[cm]	5.0	4.8	4.7	

$$\text{thus. } A \cdot \Omega = \int_{-\frac{b}{2}}^{\frac{b}{2}} dx_0 \int_{\frac{L}{2}-m-a}^{\frac{L}{2}-m} dy_0 \int_{-\frac{b}{2}}^{\frac{b}{2}} dy \int_{-\frac{L}{2}}^{\frac{L}{2}} dx \frac{r}{[(x-x_0)^2 + (y-y_0)^2 + h^2]^{\frac{3}{2}}}$$

$$\text{Flux} = \frac{N_\mu}{\Delta t \cdot A \cdot \Omega}$$

$$\Rightarrow \underline{\underline{S^{-1} \text{ sr}^{-1} \cdot \text{cm}^{-2}}}$$

$$\text{Flux} \cdot 2\pi \Rightarrow \underline{\underline{5^{-1} \text{ cm}^{-2}}} \quad \Omega < 2\pi$$



that's why we need use solid angle to correct the result the separation between two detectors lead to less observed muon events. prefer a

# Data and Some Error Estimation

Wednesday, May 22, 2019 6:45 PM

	Angle[Deg]	Start Time	End Time	Delta_T	Delta_T	Count
1	90	3:26:13 PM	4:33:05 PM	1:06:52	4012.00	600(err)
2	70	4:36:30 PM	5:01:23 PM	0:24:53	1493.00	281
3	90	5:05:21 PM	5:19:18 PM	0:13:57	837.00	170(con)
4	90	5:19:18 PM	5:35:12 PM	0:15:54	954.00	389
5	60	5:37:20 PM	5:46:55 PM	0:09:35	575.00	104(con)
6	60	5:46:55 PM	5:56:45 PM	0:09:50	590.00	219
7	60	11:40:45 AM	11:57:20 AM	0:16:35	995.00	177
8	50	11:59:55 AM	12:14:15 PM	0:14:20	860.00	104(con)
9	50	12:14:15 PM	12:29:55 PM	0:15:40	940.00	204
10	40	12:32:19 PM	1:01:39 PM	0:29:20	1760.00	140
11	30	1:03:11 PM	1:31:43 PM	0:28:32	1712.00	92

$$\begin{cases} \text{count} = N \\ \delta \text{count} = \sqrt{N} \end{cases}$$

$$\delta t = 1 \text{ s.}$$

$$Flux = \frac{N}{\delta t A \Omega}$$

$$\delta Flux = Flux \cdot \sqrt{\left(\frac{\delta N}{N}\right)^2 + \left(\frac{\delta \delta t}{\delta t}\right)^2 + \left(\frac{\delta A}{A}\right)^2 + \left(\frac{\delta \Omega}{\Omega}\right)^2}$$

$\approx 0 \quad \downarrow$

$$I = \text{NIntegrate}\left[\frac{y}{((x - x_0)^2 + (y - y_0)^2 + z^2)^{3/2}}, \{x, -L/2, L/2\}, \{y, -\beta/2, \beta/2\}, \{x_0, L/2 - m - a, L/2 - m\}, \{y_0, -b/2, b/2\}\right]$$

83.0584

hard to do error propagation on Integral.

$$\rightarrow \left(\frac{\partial I}{\partial L}\delta L\right)^2 + \left(\frac{\partial I}{\partial a}\delta a\right)^2 + \left(\frac{\partial I}{\partial b}\delta b\right)^2 + \left(\frac{\partial I}{\partial R}\delta R\right)^2 + \left(\frac{\partial I}{\partial h}\delta h\right)^2$$

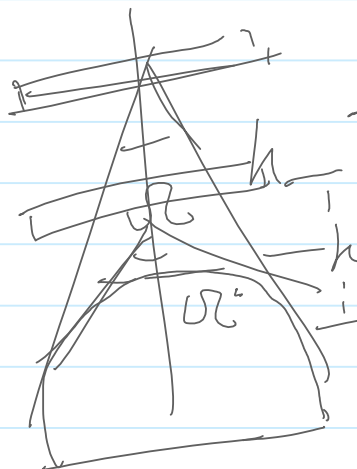
$$\begin{cases} \beta = \beta(R, h) \\ \delta = \delta(R, h) \end{cases} \quad \frac{\partial I}{\partial h} = ???$$

rough estimation: minimum precision  $\sim a/\text{cm}$ . typical deviation of length measurement  $\sim 0.3 \text{ m}$

I dimension:  $\text{cm}^2$  or  $S \cdot \frac{A}{r^2}$

maximum  $(h)$   
error percentage  $\leq 10\%$ . square  $\rightarrow \sqrt{2}$ .

$$\frac{\delta I}{I} \leq \sqrt{2} \cdot 0.1$$



$\odot$ .  
 $h$  should be treated carefully!!  
significantly affect the estimated value of solid angle  $\Omega$ .