





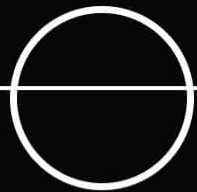
Rotation Curve of Milky Way

Zachary Cohen,
Dechong Wang, John
Wright



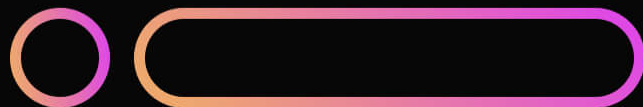
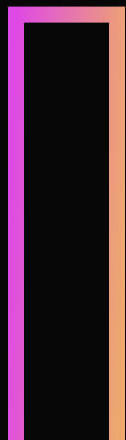
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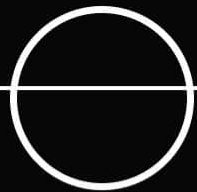
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Motivations



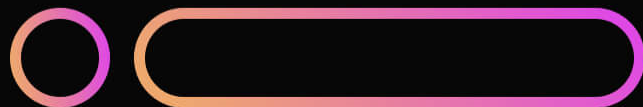
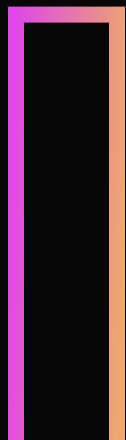
Motivations

- Estimate galactic mass distribution
- Understand relationship between mass distribution and rotation curve
- Understand the components of a galaxy



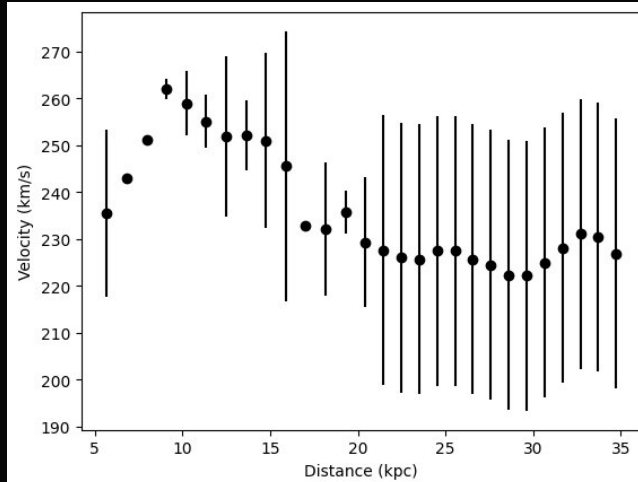
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Methods



Methods

- Given: rotation curve of galaxy [1]
- Want to infer mass distribution
- Model: velocity is a function of enclosed mass, which is a function of radius
- How does velocity depend on mass?
- How does mass depend on radius?
- Three components:
 - Galactic bulge (always enclosed, constant)
 - Galactic disc (thin cylinder, so increases quadratically up to disc radius)
 - Galactic halo (sphere, increases cubically)

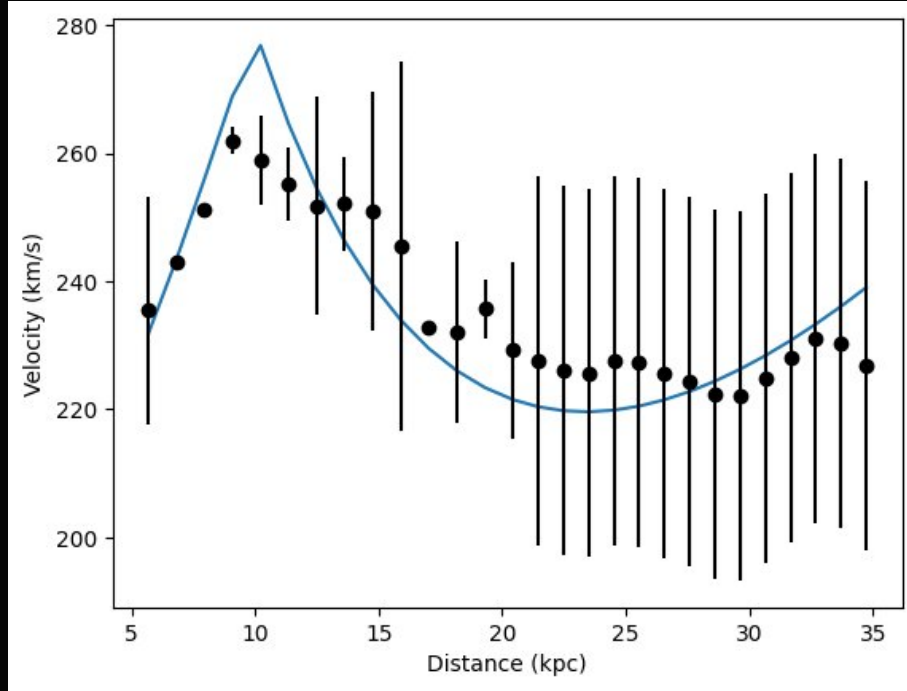


$$v(r) = \sqrt{\frac{m(r)G}{r}}$$

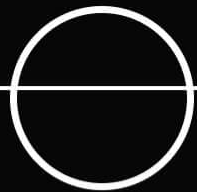
$$\frac{v^2}{G} = m_b \frac{1}{r} + \rho_d \frac{\pi \min(r^2, r_d^2)}{r} + \rho_h \frac{4}{3} \pi r^2$$
$$m(r) = m_b + \rho_d \pi \min(r^2, r_d^2) + \rho_h \frac{4}{3} \pi r^3$$

$$[1.9702122 \times 10^{10}, 1.5506895 \times 10^{11}, 2.9332181 \times 10^{11}] M_{\odot}$$

- Estimation: if we rewrite the previous equations, we get
- And we can apply standard linear regression to get easy mass estimates

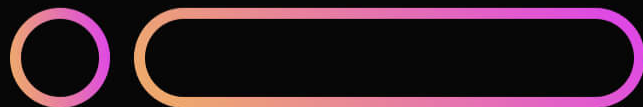
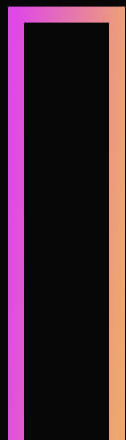


$$\frac{v^2}{G} = m_b \frac{1}{r} + \rho_d \frac{\pi \min(r^2, r_d^2)}{r} + \rho_h \frac{4}{3} \pi r^2$$



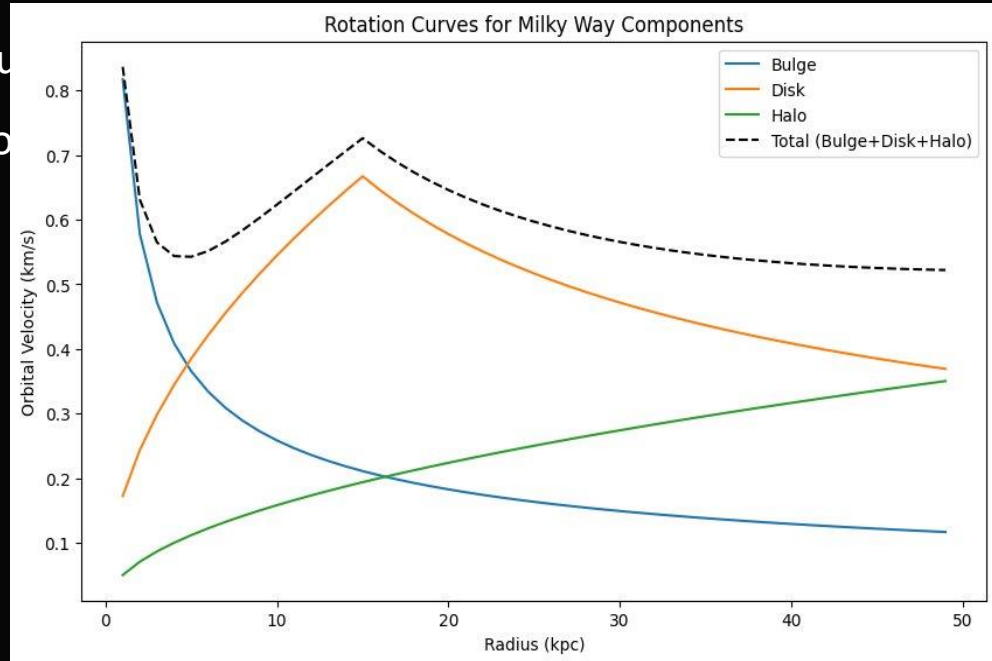
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Results

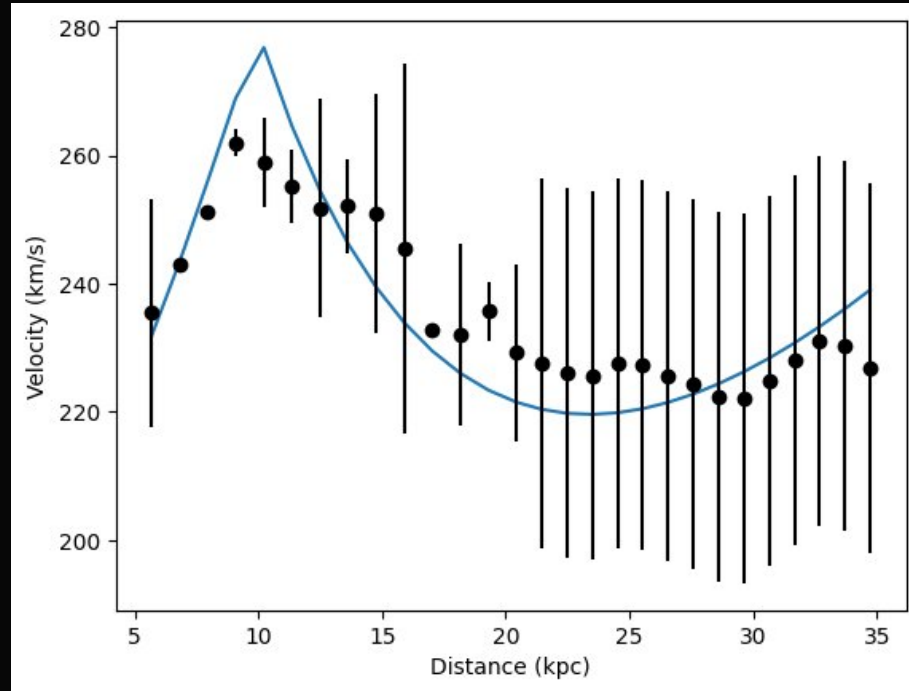


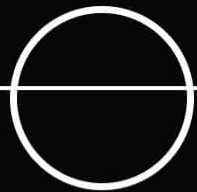
Results

- What does this model produce
- Comparison with observations
- Mass estimates
- Bulge, disc, halo



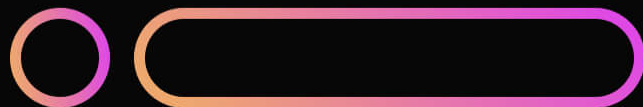
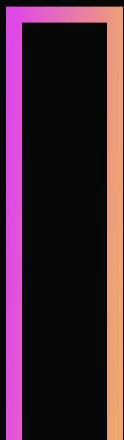
Comparison with observations






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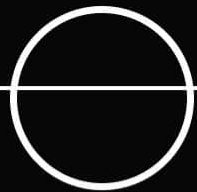
Conclusions





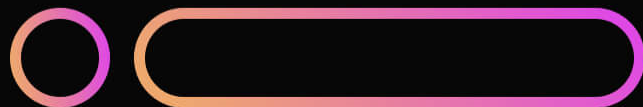
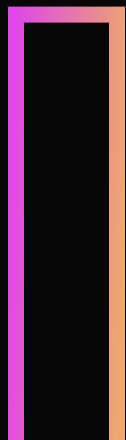
Simple model of mass
distribution works reasonably
well

Mass is distributed along the
components of the Milky Way,
but there is still unexplained
mass present



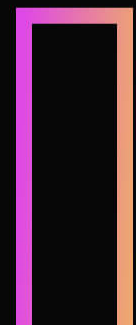
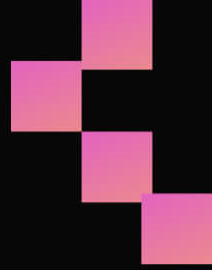
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References





References



[1] Carignan, Claude, et al. "The extended HI rotation curve and mass distribution of M31." The Astrophysical Journal 641.2 (2006): L109.



THE END
THANKS