Exercise 1: reading a table and converting coordinates/times (10 points)

- 1. Download the file fermi_lat_grbs.dat from today's exercise page on the website. It contains 1000 gamma-ray bursts observed by the Fermi satellite between 2010 and 2018.
- 2. Use the astropy.io.ascii.read() routine to read the file into a table. You need to specify the header_start, data_start, and delimiter arguments.
- 3. Display the column headers using the columns or colnames method.
- 4. Notice that RA and DEC are in sexagesimal format, represented as strings. Use astropy.coordinates.Angle to convert them to decimal degrees. Print the table.
- 5. Do some brief analysis:
 - a) Construct a boolean mask array that selects only those bursts with durations $t_{90} < 2$ seconds and relative errors in the duration $\frac{t_{90_error}}{t_{90}} < 50\%$.
 - b) Construct a second boolean mask array that selects $t_{90}>2$ seconds with the same relative error.
 - c) Compare the median peak energy ($f_{lnc_band_epeak}$) of the two samples. Use the NumPy median function to compute the values.

```
In [2]: #Code Here
        import numpy as np
        import astropy
        from astropy import units as units
        from astropy.table import Table
        from astropy import coordinates
        from astropy import units as u
        # Q.1,2
        data = astropy.io.ascii.read(r"C:\Users\eklav\OneDrive - University of Illinois - U
        # Q.3
        print(data.columns)
        # Q.4
        data['ra'] = coordinates.Angle(data['ra'],unit = u.hour)
        data['dec'] = coordinates.Angle(data['dec'],unit = u.deg)
        data['ra'] = (data['ra']).to(u.deg)
        data['dec'] = (data['dec']).to(u.deg)
        print(data['ra'])
        print(data['dec'])
        mask_short_bursts = (data['t90'] < 2) & ((data['t90_error'] / data['t90']) < 0.5)
```

```
mask_long_bursts = (data['t90'] > 2) & ((data['t90_error'] / data['t90']) < 0.5)
short_bursts_epeak = data['flnc_band_epeak'][mask_short_bursts]
long_bursts_epeak = data['flnc_band_epeak'][mask_long_bursts]
median_peak_energy_short = np.median(short_bursts_epeak)
median_peak_energy_long = np.median(long_bursts_epeak)</pre>
```

<TableColumns names=('col0','name','ra','dec','trigger_time','t90','t90_error','t90_ start', 'fluence', 'fluence_error', 'flux_1024', 'flux_1024_error', 'flux_1024_time', 'flu x_64','flux_64_error','flnc_band_ampl','flnc_band_ampl_pos_err','flnc_band_ampl_neg_ err', 'flnc_band_epeak', 'flnc_band_epeak_pos_err', 'flnc_band_epeak_neg_err', 'flnc_band_epeak_neg_epeak_neg_err', 'flnc_band_epeak_neg_epeak_neg_epeak_neg_epeak_neg_epeak_neg_epeak_neg_epeak_neg_epeak_neg_epeak_neg_epeak_ep d_alpha','flnc_band_alpha_pos_err','flnc_band_alpha_neg_err','flnc_band_beta','flnc_ band_beta_pos_err','flnc_band_beta_neg_err','flnc_spectrum_start','flnc_spectrum_sto p','pflx_best_fitting_model','pflx_best_model_redchisq','flnc_best_fitting_model','f lnc_best_model_redchisq','_1')>

deg 337.8699999999995 356.95916666666665 211.21999999999997 209.14 2.899999999999995 9.47999999999999 244.0399999999996 20.722916666666666 311.3299999999999 160.4299999999998 308.56 235.66999999999996

159.91083333333333

52.3

202.2699999999995

304.959999999999 246.66083333333327

59.72

184.44499999999996

284.6799999999995

Length = 1000 rows

dec

deg

-80.0399999999999

-79.905

-79.69

-79.1

-79.01

-78.2

-77.86

-77.78388888888888

-74.5

-74.28

-39.36

-39.34

-39.3280555555556

-39.21

-39.2

-39.18 -39.15888888888889

-37.08611111111116

```
Length = 1000 rows
```

C:\Users\eklav\AppData\Local\Packages\PythonSoftwareFoundation.Python.3.11_qbz5n2kfr
a8p0\LocalCache\local-packages\Python311\site-packages\numpy\core\fromnumeric.py:77
1: UserWarning: Warning: 'partition' will ignore the 'mask' of the MaskedColumn.
 a.partition(kth, axis=axis, kind=kind, order=order)

```
In [3]: median_peak_energy_short, median_peak_energy_long
Out[3]: (359.3643, 135.48020000000002)
```

Exercise 2: working with and modifying a table (10 points)

Download the files "mcxc.dat" and "mcxc.readme" from today's exercise page on the website. These files contain a catalog of X-ray-detected clusters of galaxies from Piffaretti et al. (2011) obtained through the VizieR service at the University of Strasbourg.

- 1. Use astropy.io.ascii.read() to read the table and its metadata into Python. This table is in "CDS" format, and you specify the metadata file using the readme argument.
- 2. Extract the log of L500 (luminosity in units of $10^{44} {\rm erg/s}$), log of M500 (mass in 10^{14} solar masses), and z (redshift) columns from the data into 1D arrays. Create a mask array selecting those clusters with redshift < 0.1.
- 3. Now construct the array r containing $r=\log L-1.64\log M$. The X-ray luminosity and mass of galaxy clusters are correlated roughly such that $L\propto M^{5/3}$, so the range in r is small.
- 4. Modify the table to add a masked column (<code>MaskedColumn</code> object) for r, and use the description " $r = \log L 1.64 \log M$ " for that column. Use the mask array you created to mask those clusters you don't want to store r values for. Print the table to check whether your column was added correctly. You should see "--" for masked-out rows.
- 5. Write the modified table to a file named "mcxc_new.csv" in comma-separated value (CSV) format.

```
In [5]: #Code Here
    from astropy.table import Table, MaskedColumn
    data = astropy.io.ascii.read(r"C:\Users\eklav\OneDrive - University of Illinois - U

1500 = data['L500'].data
    M500 = data['M500'].data
    z = data['z'].data
    mask = z < 0.1
    filter_1500 = 1500[mask]
    filter_M500 = M500[mask]
    filtered_z = z[mask]</pre>
```

```
r = filter_1500 - 1.64 * filter_M500
print(r)
r = data['L500'] - 1.64 * data['M500']
mask = data['z'] >= 0.1
r_column = MaskedColumn(r, mask=mask, name='r', description='r = log L - 1.64 log M
```

```
[-1.012892 -0.48837 -1.528762 -1.997207 -1.614681 -1.169538 -2.294202
-2.34182 -1.294671 -1.260939 -1.572374 -1.259829 -1.815282 -1.163526
-2.483263 -2.112338 -1.616338 -2.366497 -2.387405 -0.699704 -1.025122
-1.309052 -0.592733 -2.145387 -3.618647 -2.754706 -1.25405 -2.210964
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-0.183307 -1.104538 -0.645966 -1.256982 -0.876203 -1.530939 -0.069357
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-1.526178 -1.195995 -2.232193 -2.052756 -1.769778 -3.321541 -2.053815
        -2.323064 -3.090323 -1.226374 -3.065216 -0.46924 -0.692919
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-2.352053 -1.668644 -1.346296 -1.338956 -0.86651 -2.596195 -0.465179
-2.078877 -1.753874 -2.615996 -2.260918 -1.755869 -3.300709 -1.317805
-2.962304 -1.476309 -1.658059 -1.231028 -1.609297 -2.092455 -1.551878
-2.77563 -2.167596 -0.369748 -2.313311 -2.208298 -2.201311 -2.461703
-0.566465 -1.980936 -1.373977 -2.74229 -2.179097]
```

```
In [6]: data.add_column(r_column)
    print(data)
    data.write(r'C:\Users\eklav\OneDrive - University of Illinois - Urbana\astro_310\la
```

MCVC	ON	A N I = =		1.500-4	_
MCXC	OName	AName	• • •	L500r4	r
			• • •		1e+37 W
			• • •		
J0000.1+0816	RXC J0000.1+0816	UGC 12890	• • •		-1.012892
J0000.4-0237	RXC J0000.4-0237				-0.48836999999999997
J0001.6-1540	RXC J0001.6-1540				
J0001.9+1204	RXC J0001.9+1204	A2692			
J0003.1-0605	RXCJ0003.1-0605	A2697			
J0003.2-3555	RXCJ0003.2-3555	A2717			-1.528762
J0003.8+0203	RXCJ0003.8+0203	A2700			-1.9972069999999997
J0004.9+1142	RXC J0004.9+1142	UGC 00032			-1.614681
J0005.3+1612	RXC J0005.3+1612	A2703			
J0006.0-3443	RXCJ0006.0-3443	A2721			
J2354.2-1024	RXCJ2354.2-1024	A2670			-2.461702999999999
J2355.1-1500	BVH2007 242	VMF98 223	• • •		-0.5664649999999999
	RXC J2355.6+1120	A2675	• • •		-1.980935999999999
			• • •		
J2355.7+1138	A2678		• • •		-1.3739769999999998
	RXC J2355.8+3423	A2677	• • •		
J2357.0-3445	RXCJ2357.0-3445	A4059	• • •		-2.74229
J2359.3-6042	RXCJ2359.3-6042	A4067	• • •		-2.1790969999999996
J2359.4-3418	MS2356.9-3434				
J2359.5-3211	RX J2359.5-3211	BSe RXJ2359.5-3211			
J2359.9-3928	RXCJ2359.9-3928	A4068			
Length = 1743 rows					