

1. Download the file "fermi_lat_grbs.dat" from today's exercise page on the website. It contains 1000 gamma-ray bursts observed by the Fermi satellite between 2010 and 2018. 2. Use the `astropy.io.ascii.read()` routine to read the file into a table. You need to specify the `header_start`, `data_start`, and `delimiter` arguments. 3. Display the column headers using the `columns` or `colnames` method

```
In [101... import numpy as np
import astropy
from astropy import units as u
from astropy.table import Table

data = astropy.io.ascii.read(r"C:\Users\eklav\OneDrive - University of Illinois - U

data.columns
```

```
Out[101... <TableColumns names=('col0', 'name', 'ra', 'dec', 'trigger_time', 't90', 't90_error', 't9
0_start', 'fluence', 'fluence_error', 'flux_1024', 'flux_1024_error', 'flux_1024_tim
e', 'flux_64', 'flux_64_error', 'flnc_band_ampl', 'flnc_band_ampl_pos_err', 'flnc_band_
ampl_neg_err', 'flnc_band_epeak', 'flnc_band_epeak_pos_err', 'flnc_band_epeak_neg_er
r', 'flnc_band_alpha', 'flnc_band_alpha_pos_err', 'flnc_band_alpha_neg_err', 'flnc_ban
d_beta', 'flnc_band_beta_pos_err', 'flnc_band_beta_neg_err', 'flnc_spectrum_start', 'f
lnc_spectrum_stop', 'pflx_best_fitting_model', 'pflx_best_model_redchisq', 'flnc_best
_fitting_model', 'flnc_best_model_redchisq', '_1')>
```

Notice that RA and DEC are in sexagesimal format, represented as strings. Use `astropy.coordinates.Angle` to convert them to decimal degrees. Print the table

```
In [102... from astropy import coordinates
from astropy import units as u
x = data['ra']
y = data['dec']

# ra_units = [coordinates.Angle(i, unit=u.hour) for i in x]
# dec_units = [coordinates.Angle(i, unit=u.deg) for i in x]
data['ra'] = coordinates.Angle(data['ra'], unit = u.hour)

data['dec'] = coordinates.Angle(data['dec'], unit = u.deg)

data['ra'] = (data['ra']).to(u.deg)
data['dec'] = (data['dec']).to(u.deg)

print(data['ra'])
print(data['dec'])
# print(y)
```

```

      ra
    deg
-----
337.86999999999995
356.95916666666665
211.21999999999997
      209.14
2.8999999999999995
      9.479999999999999
244.03999999999996
20.722916666666666
      311.3299999999999
160.42999999999998
      ...
      308.56
235.66999999999996
159.91083333333333
      52.3
202.26999999999995
      304.9599999999999
246.66083333333327
      59.72
184.44499999999996
284.67999999999995
Length = 1000 rows
      dec
    deg
-----
-80.03999999999999
      -79.905
      -79.69
      -79.1
      -79.01
      -78.2
      -77.86
-77.78388888888888
      -74.5
      -74.28
      ...
      -39.36
      -39.34
-39.32805555555556
      -39.21
      -39.2
      -39.18
-39.15888888888889
      -37.2
-37.08611111111116
      -37.03
Length = 1000 rows

```

Do some brief analysis: a) Construct a boolean mask array that selects only those bursts with durations (t_{90}) < 2 seconds and relative errors in the duration (t_{90_error}/t_{90}) < 50%. b) Construct a second boolean mask array that selects $t_{90} > 2$ seconds with the same relative

error. c) Compare the median peak energy (flnc_band_epeak) of the two samples. Use the NumPy median function to compute the values

```
In [83]: mask_short_bursts = (data['t90'] < 2) & ((data['t90_error'] / data['t90']) < 0.5)
mask_long_bursts = (data['t90'] > 2) & ((data['t90_error'] / data['t90']) < 0.5)
short_bursts_epeak = data['flnc_band_epeak'][mask_short_bursts]
long_bursts_epeak = data['flnc_band_epeak'][mask_long_bursts]
median_peak_energy_short = np.median(short_bursts_epeak)
median_peak_energy_long = np.median(long_bursts_epeak)
median_peak_energy_short, median_peak_energy_long
```

```
Out[83]: (359.3643, 135.48020000000002)
```

Download the files "mcxc.dat" and "mcxc.readme" from today's exercise page on the web site. These files contain a catalog of X-ray-detected clusters of galaxies from Piffaretti et al. (2011) obtained through the Vizier service at the University of Strasbourg. 1. Use `astropy.io.ascii.read()` to read the table and its metadata into Python. This table is in "CDS" format, and you specify the metadata file using the `readme` argument. 2. Extract the log of L500 (luminosity in units of 10^{44} erg/s), log of M500 (mass in 10^{14} solar masses), and z (redshift) columns from the data into 1D arrays. Create a mask array selecting those clusters with redshift < 0.1 . 3. Now construct the array containing r . The X-ray luminosity and mass of galaxy clusters are correlated roughly such that, so the range in r is small. $r = \log L - 1.64 \log M$ $L \propto M^{5/3}$

```
In [93]: data = astropy.io.ascii.read(r"C:\Users\eklav\OneDrive - University of Illinois - U

log_L500 = data['L500'].data
log_M500 = data['M500'].data
z = data['z'].data
mask = z < 0.1
filtered_log_L500 = log_L500[mask]
filtered_log_M500 = log_M500[mask]
filtered_z = z[mask]
r = filtered_log_L500 - 1.64 * filtered_log_M500
```

```
In [94]: r
```

```
Out[94]: array([-1.012892, -0.48837 , -1.528762, -1.997207, -1.614681, -1.169538,  
-2.294202, -2.34182 , -1.294671, -1.260939, -1.572374, -1.259829,  
-1.815282, -1.163526, -2.483263, -2.112338, -1.616338, -2.366497,  
-2.387405, -0.699704, -1.025122, -1.309052, -0.592733, -2.145387,  
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-1.379327, -1.92645 , -0.648112, -0.673878, -2.154409, -0.482401,  
-1.09285 , -3.369212, -3.503389, -2.219902, -1.219815, -0.930073,  
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```

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-2.053815, -1.2781 , -2.323064, -3.090323, -1.226374, -3.065216,
-0.46924 , -0.692919, -2.517618, -2.725523, -2.670822, -2.110539,
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-0.465179, -2.078877, -1.753874, -2.615996, -2.260918, -1.755869,
-3.300709, -1.317805, -2.962304, -1.476309, -1.658059, -1.231028,
-1.609297, -2.092455, -1.551878, -2.77563 , -2.167596, -0.369748,
-2.313311, -2.208298, -2.201311, -2.461703, -0.566465, -1.980936,
-1.373977, -2.74229 , -2.179097])

```
In [98]: from astropy.table import Table, MaskedColumn
r = data['L500'] - 1.64 * data['M500']
mask = data['z'] >= 0.1
r_column = MaskedColumn(r, mask=mask, name='r', description='r = log L - 1.64 log M')
data.add_column(r_column)
print(data)
data.write(r'C:\Users\eklav\OneDrive - University of Illinois - Urbana\astro_310\la
```

MCXC	OName	AName	...	L500r4	r
			...		1e+37 W
-----	-----	-----	...	-----	-----
J0000.1+0816	RXC J0000.1+0816	UGC 12890	...	--	-1.012892
J0000.4-0237	RXC J0000.4-0237	--	...	--	-0.4883699999999997
J0001.6-1540	RXC J0001.6-1540	--	...	--	--
J0001.9+1204	RXC J0001.9+1204	A2692	...	--	--
J0003.1-0605	RXCJ0003.1-0605	A2697	...	--	--
J0003.2-3555	RXCJ0003.2-3555	A2717	...	--	-1.528762
J0003.8+0203	RXCJ0003.8+0203	A2700	...	--	-1.9972069999999997
J0004.9+1142	RXC J0004.9+1142	UGC 00032	...	--	-1.614681
J0005.3+1612	RXC J0005.3+1612	A2703	...	--	--
J0006.0-3443	RXCJ0006.0-3443	A2721	...	--	--
...
J2354.2-1024	RXCJ2354.2-1024	A2670	...	--	-2.4617029999999999
J2355.1-1500	BVH2007 242	VMF98 223	...	--	-0.5664649999999999
J2355.6+1120	RXC J2355.6+1120	A2675	...	--	-1.9809359999999998
J2355.7+1138	A2678	A2678	...	--	-1.3739769999999998
J2355.8+3423	RXC J2355.8+3423	A2677	...	--	--
J2357.0-3445	RXCJ2357.0-3445	A4059	...	--	-2.74229
J2359.3-6042	RXCJ2359.3-6042	A4067	...	--	-2.1790969999999996
J2359.4-3418	MS2356.9-3434	--	...	--	--
J2359.5-3211	RX J2359.5-3211 BSe RXJ2359.5-3211	--	--
J2359.9-3928	RXCJ2359.9-3928	A4068	...	--	--
Length = 1743 rows					