

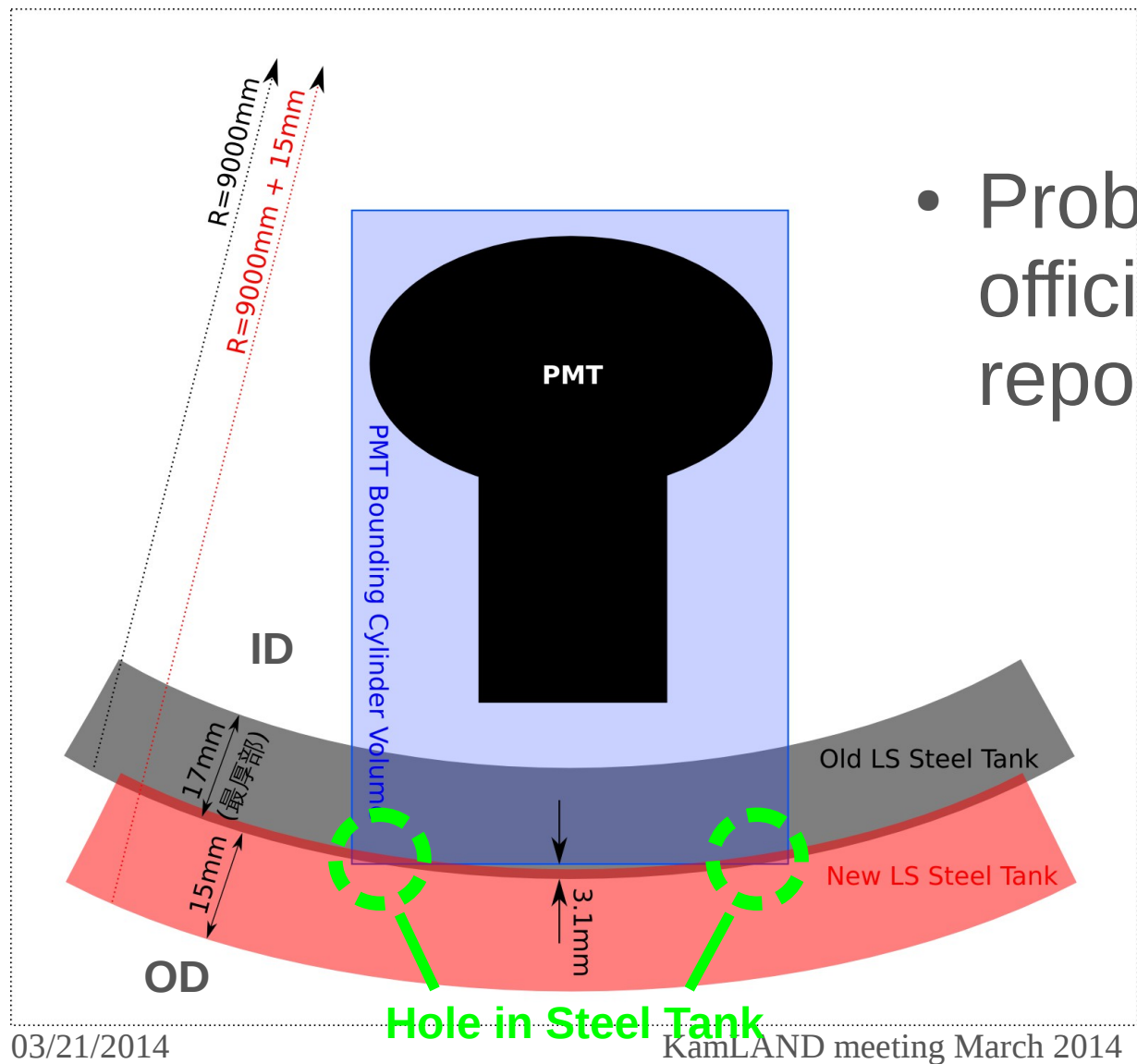
High Energy Event Reconstruction

Michinari Sakai

Known Geometry Problem with KLG4 for High Energy Analysis

- High energy event produces many photons
→ photons “warp” from ID to OD
- This is a problem when placing OD cuts to find fully contained events
- Cause due to PMT geometry overlap in KLG4

Known Geometry Problem with KLG4 for High Energy Analysis

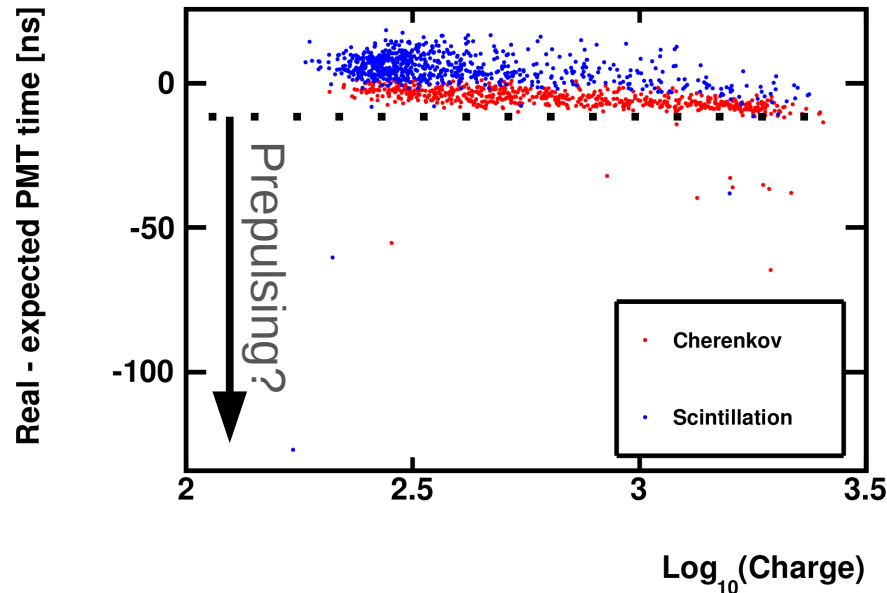


- Problem will be officially fixed in repository.

PMT Prepulsing at High Energies

- PMT prepulsing occurs when many photons ($>\sim 100$)
- Prepulsing is most probable when photons hit dynode along dynode axis.
- Prepulsing can give wrong bias to fitters that use PMT hit timing

PMT Prepulsing at High Energies



Sample muon #1

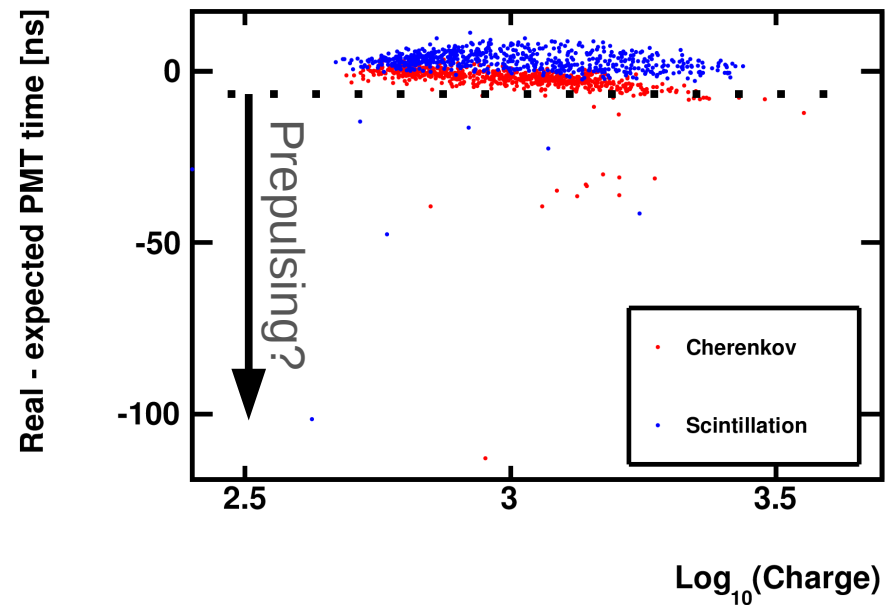
Run 5000

Event 25663

Recon Energy 2512.59 MeV

Badness 13.706

Impact Param 396.618 cm



Sample muon #2

Run 5000

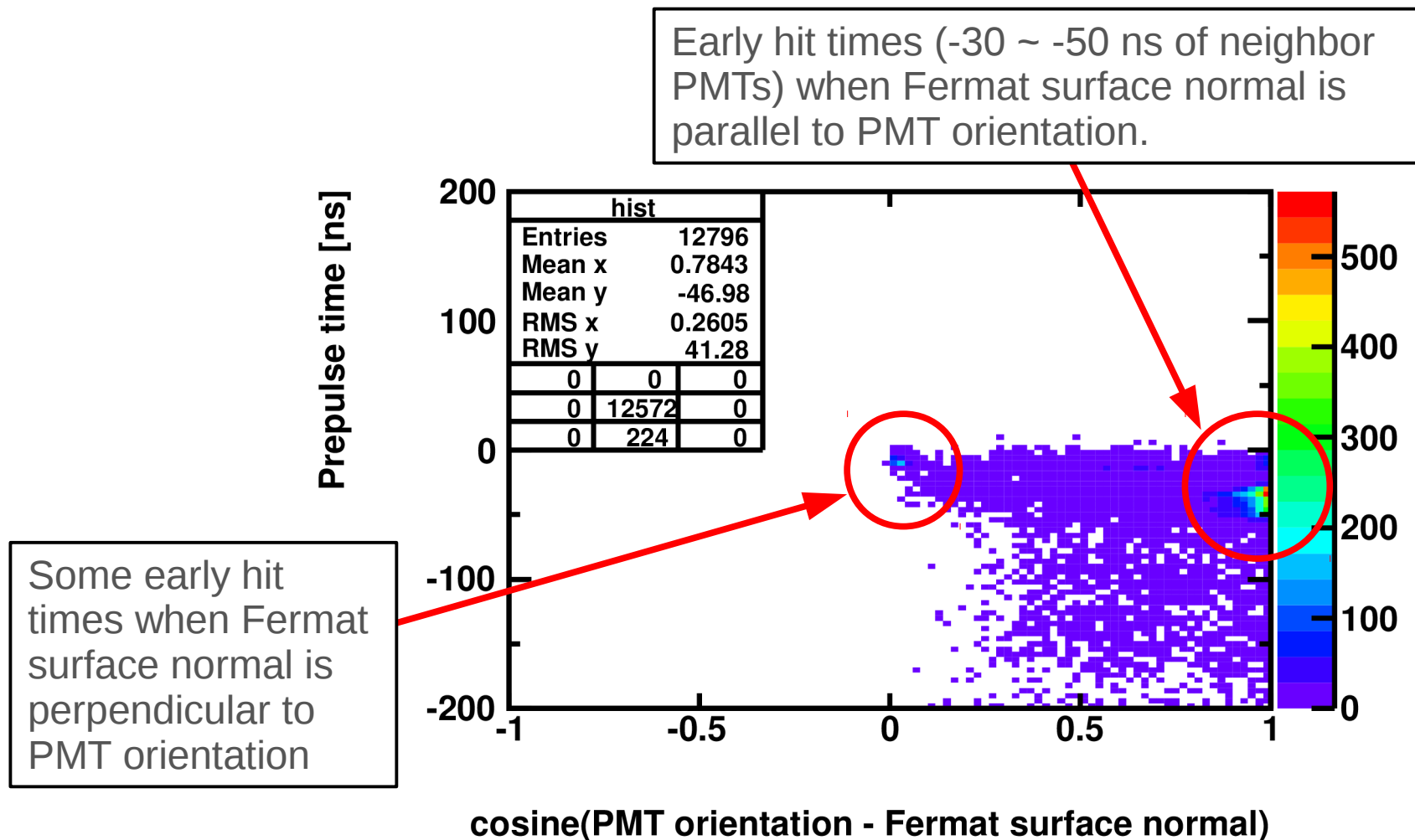
Event 37715

Recon Energy 5031.46 MeV

Badness 16.4625

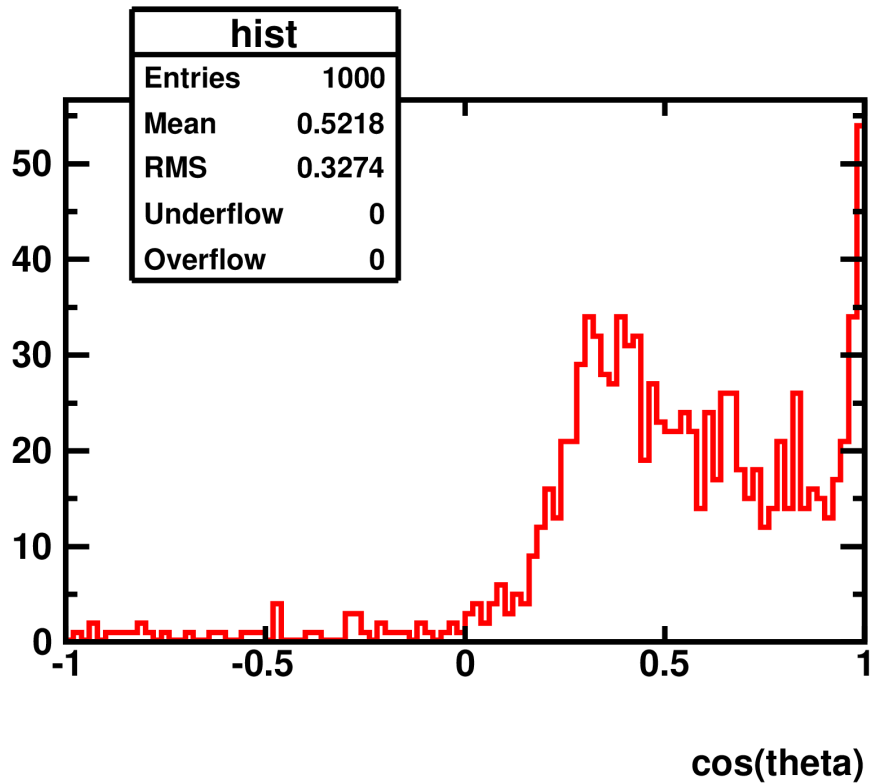
Impact Param 187.912 cm

PMT Prepulsing at High Energies (1000 muons from RTQ run 5000~)

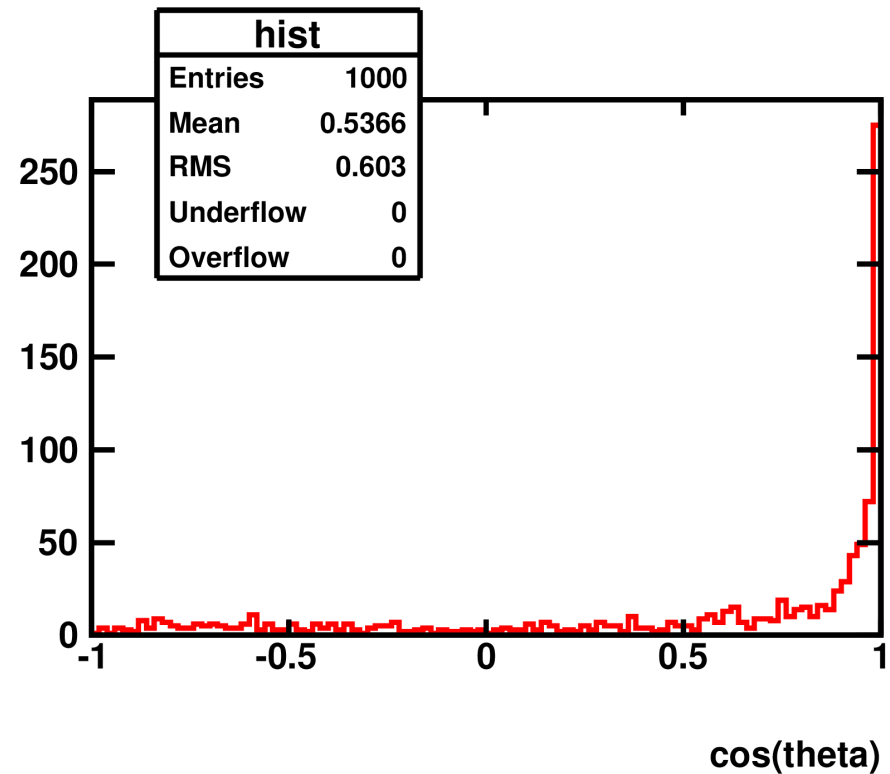


Agreement of Michi Fitter / Muon Fitter (1000 muons, RTQ run 5000~)

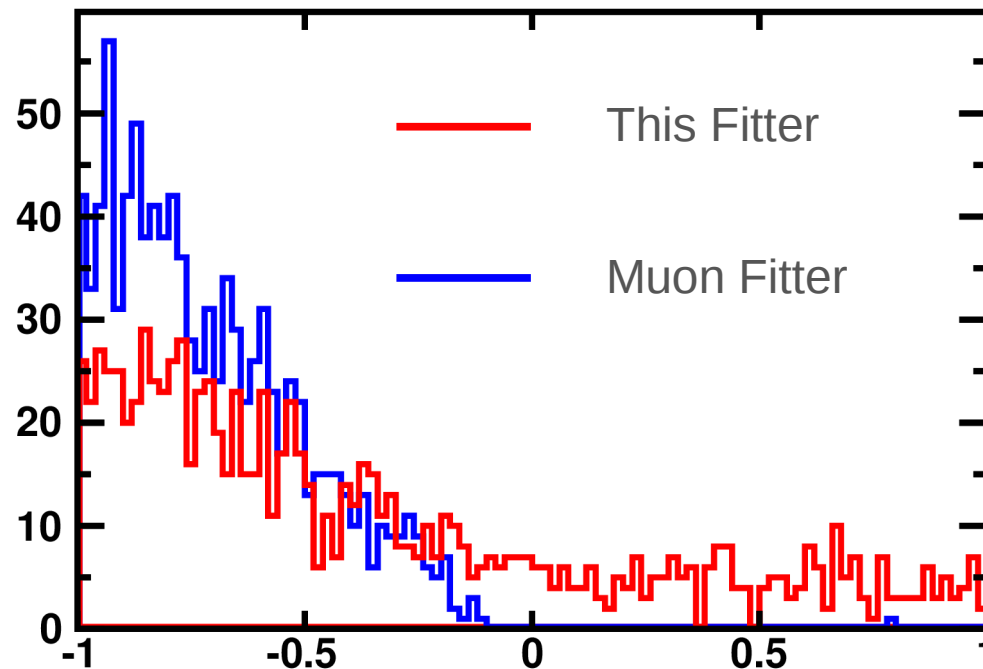
No cut



Prepulse cut



Zenith Angle of Both Fitters (1000 muons, RTQ run 5000~)



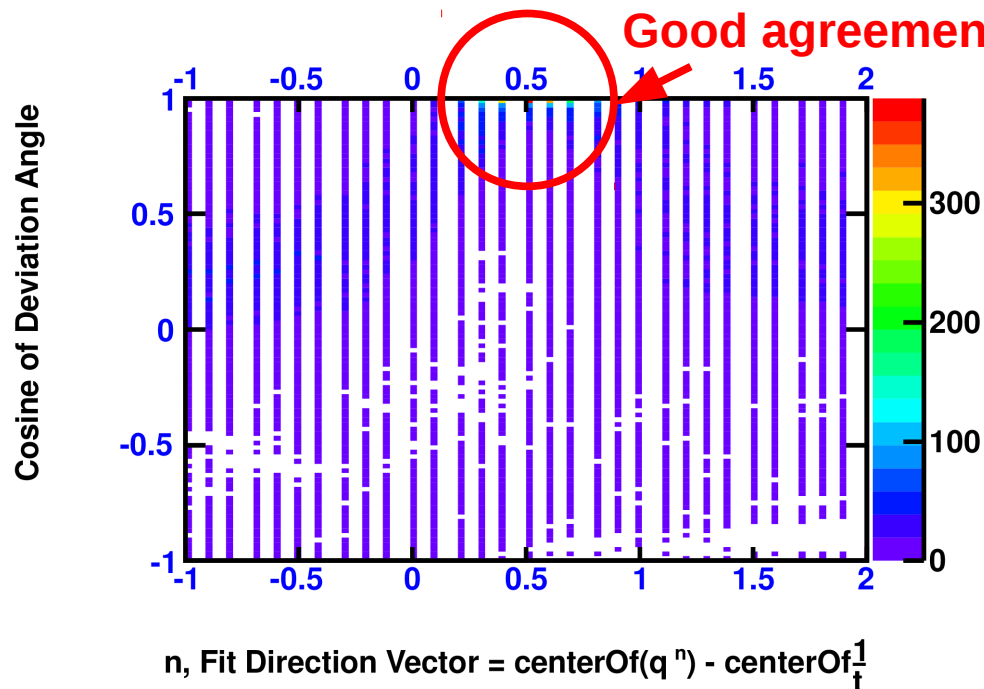
- Muon Fitter is fitting through-going muons well.
- New fitter is less accurate for through going muons..

Center of time/charge vs. muon fitter

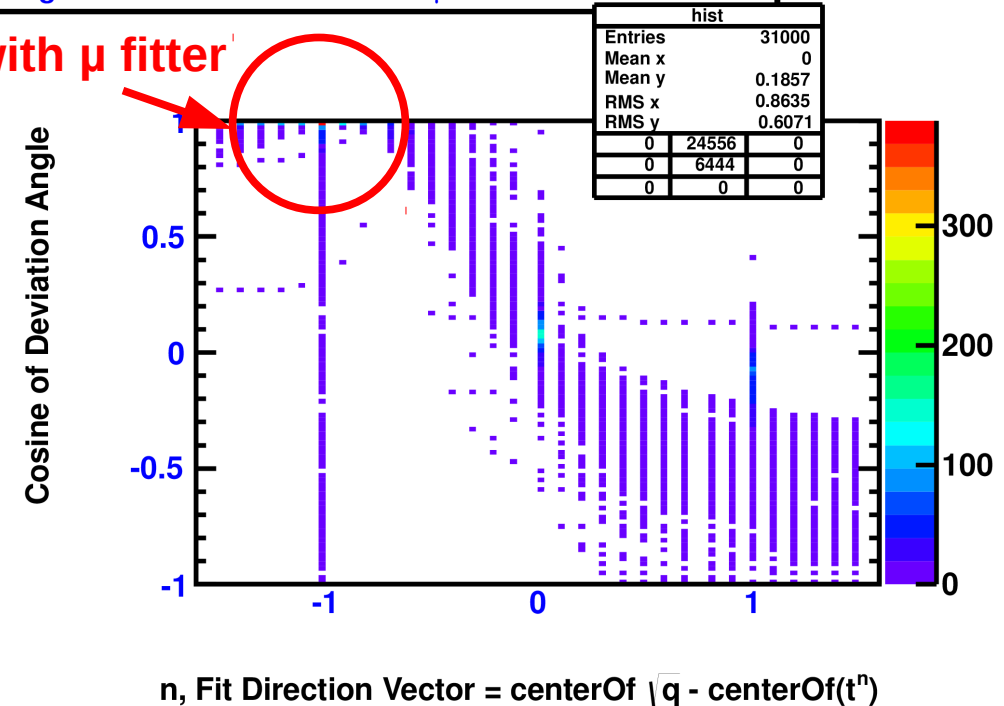
(1000 muons, RTQ run 5000~)

- Exponent modifier m, n for center of (time^m) and (charge^n) was varied in steps of 0.1
- Vector pointing from center of time^{-1} to center of $\text{charge}^{0.5}$ gave best result

Agreement between KamLAND μ Fitter and Mich Fitter



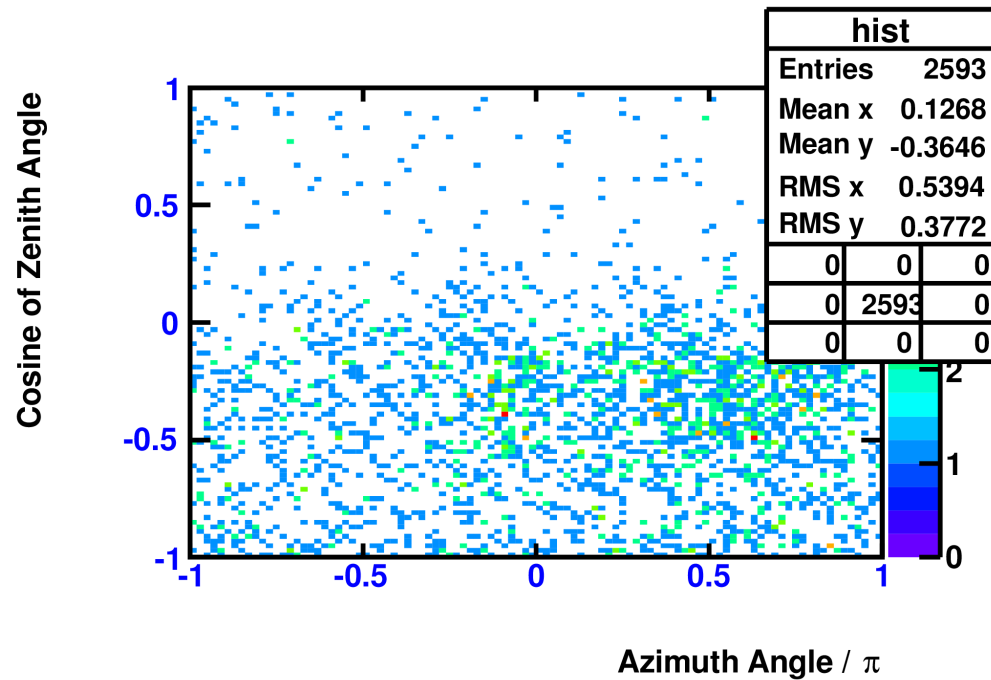
Agreement between KamLAND μ Fitter and Mich Fitter



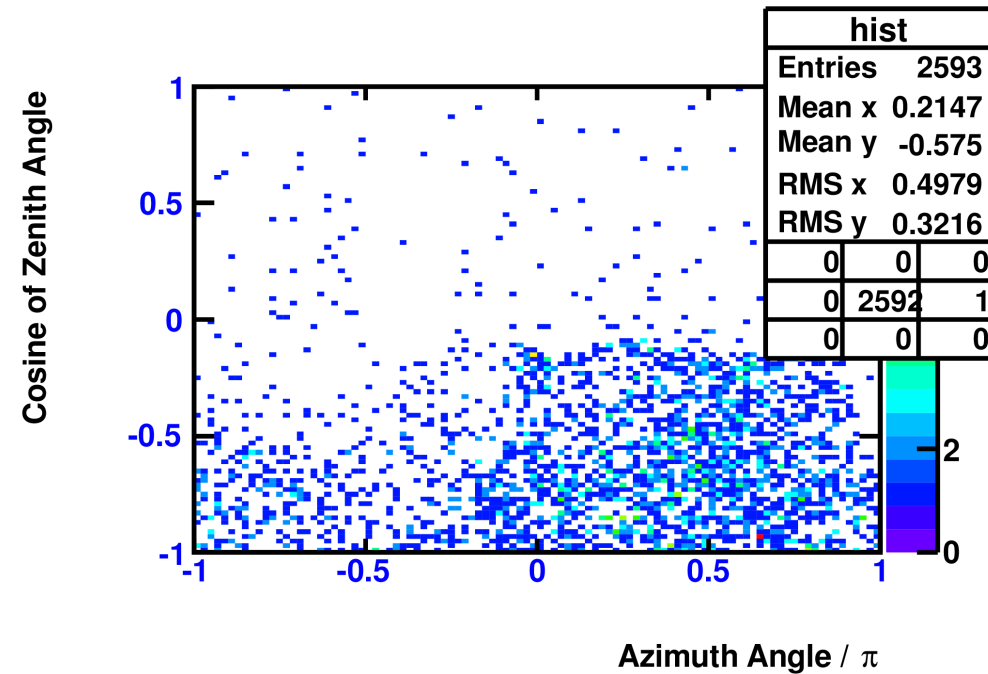
Atmospheric Neutrino

Events from Tamae, Shimizu

Michi ν Fitter

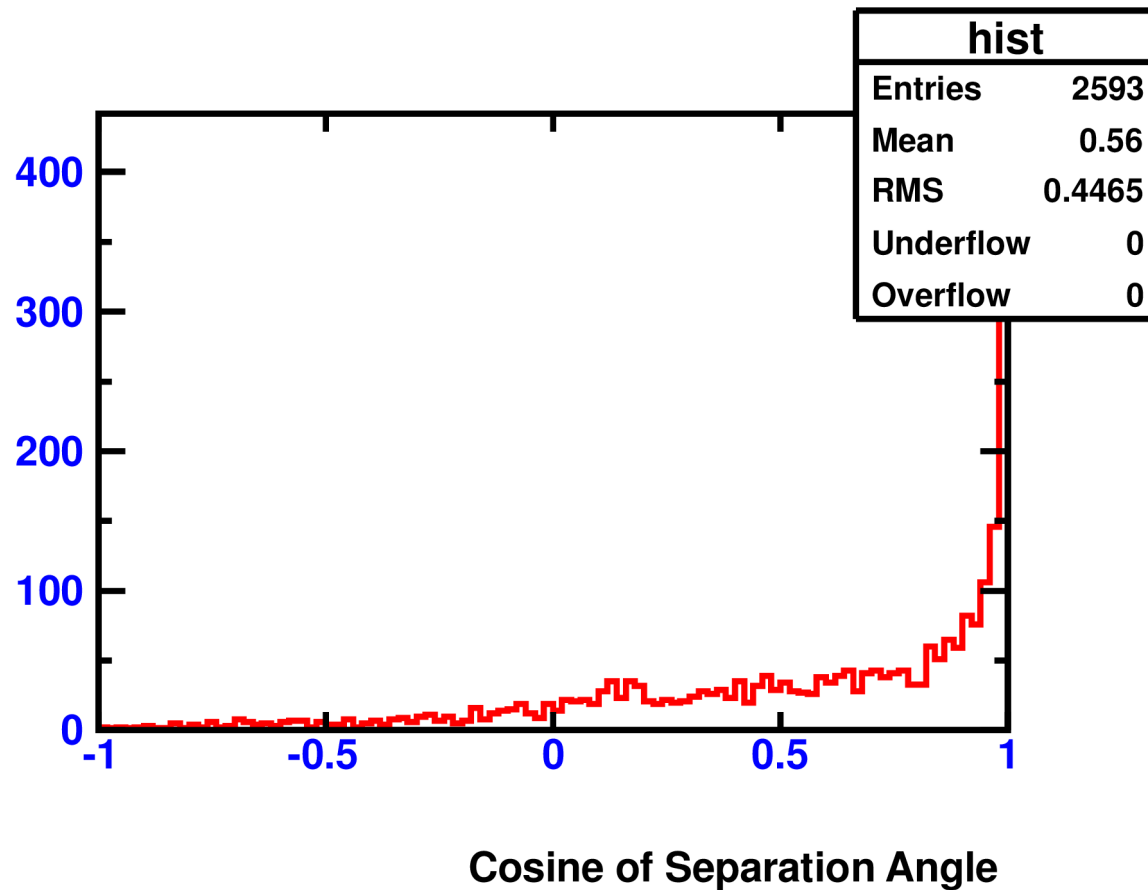


KamLAND μ fitter



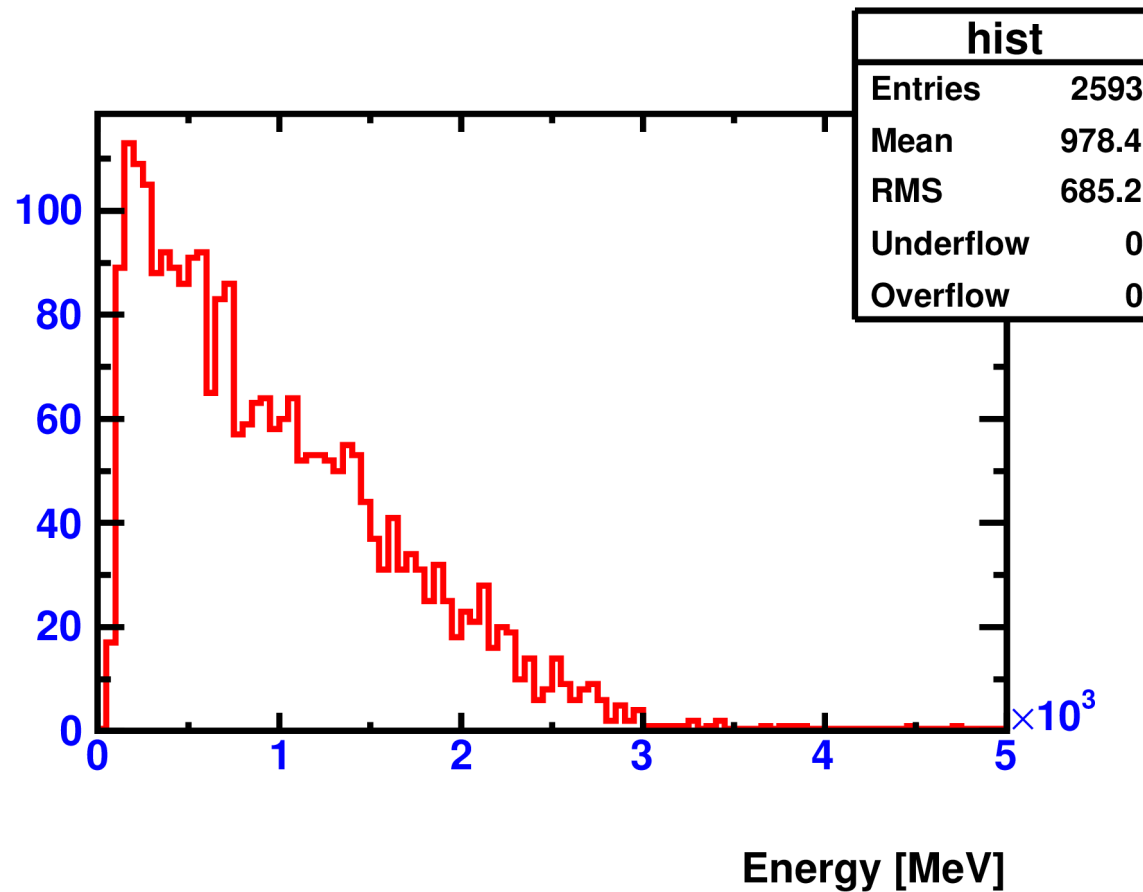
Atmospheric Neutrino

Atmospheric Neutrino Reconstructed Direction wrt Muon Fitter



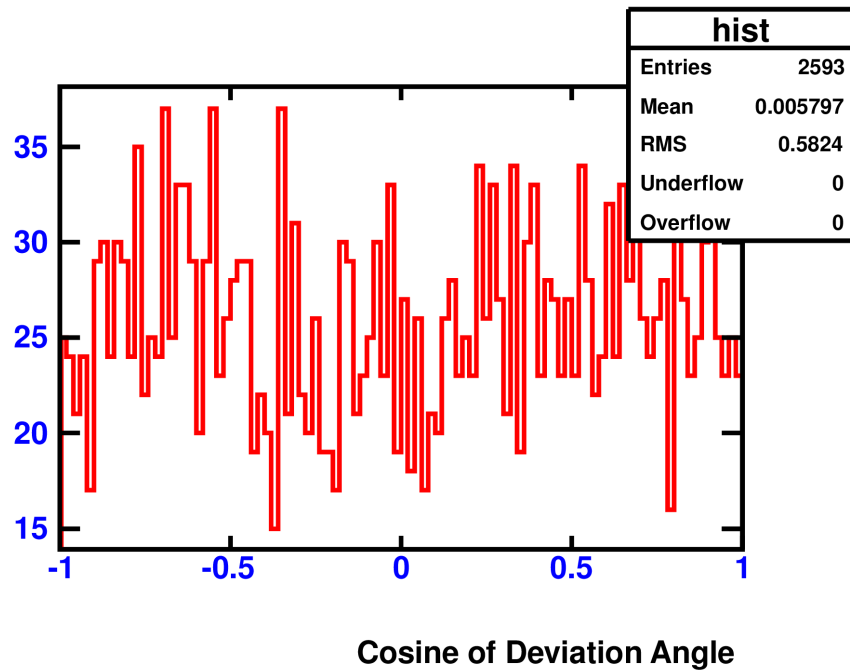
Atmospheric Neutrino

Atmospheric Neutrino Reconstructed Energy

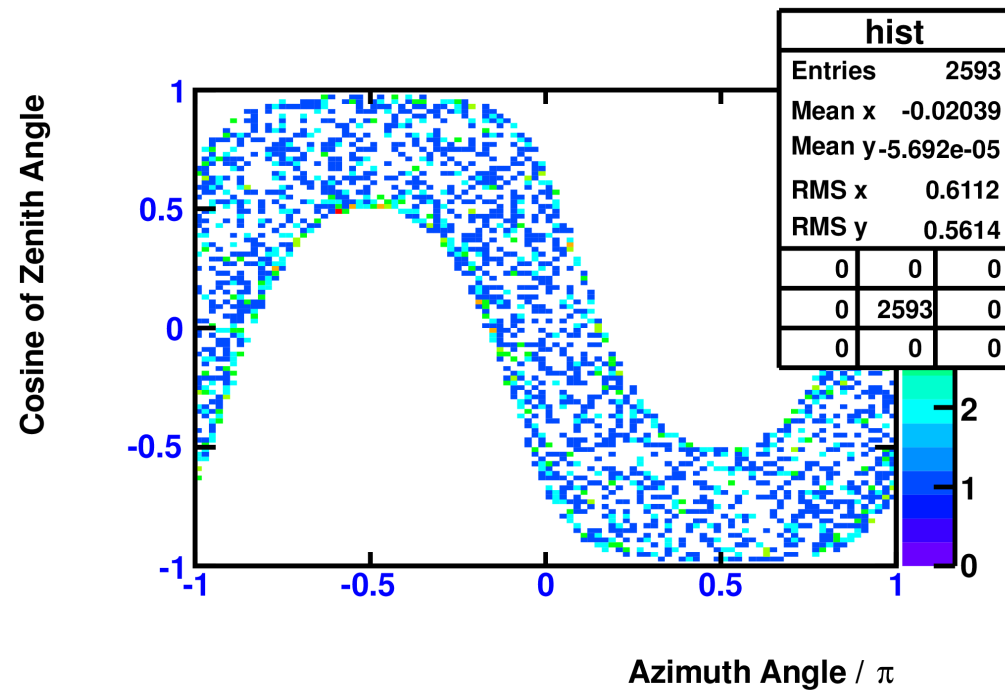


Atmospheric Neutrino

Angle Difference Between Solar Position and ν Direction

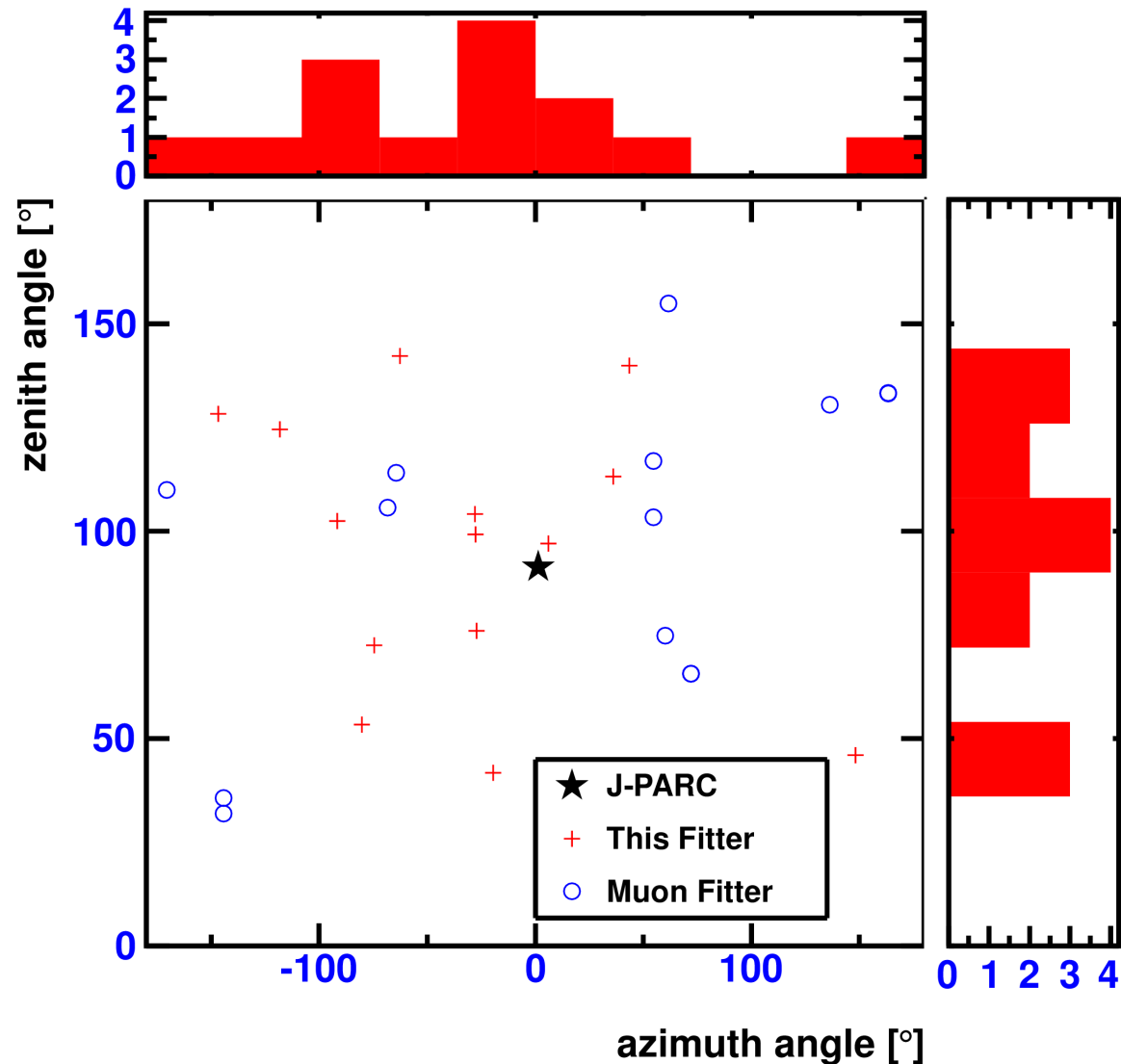


Solar Zenith and Azimuth Angle



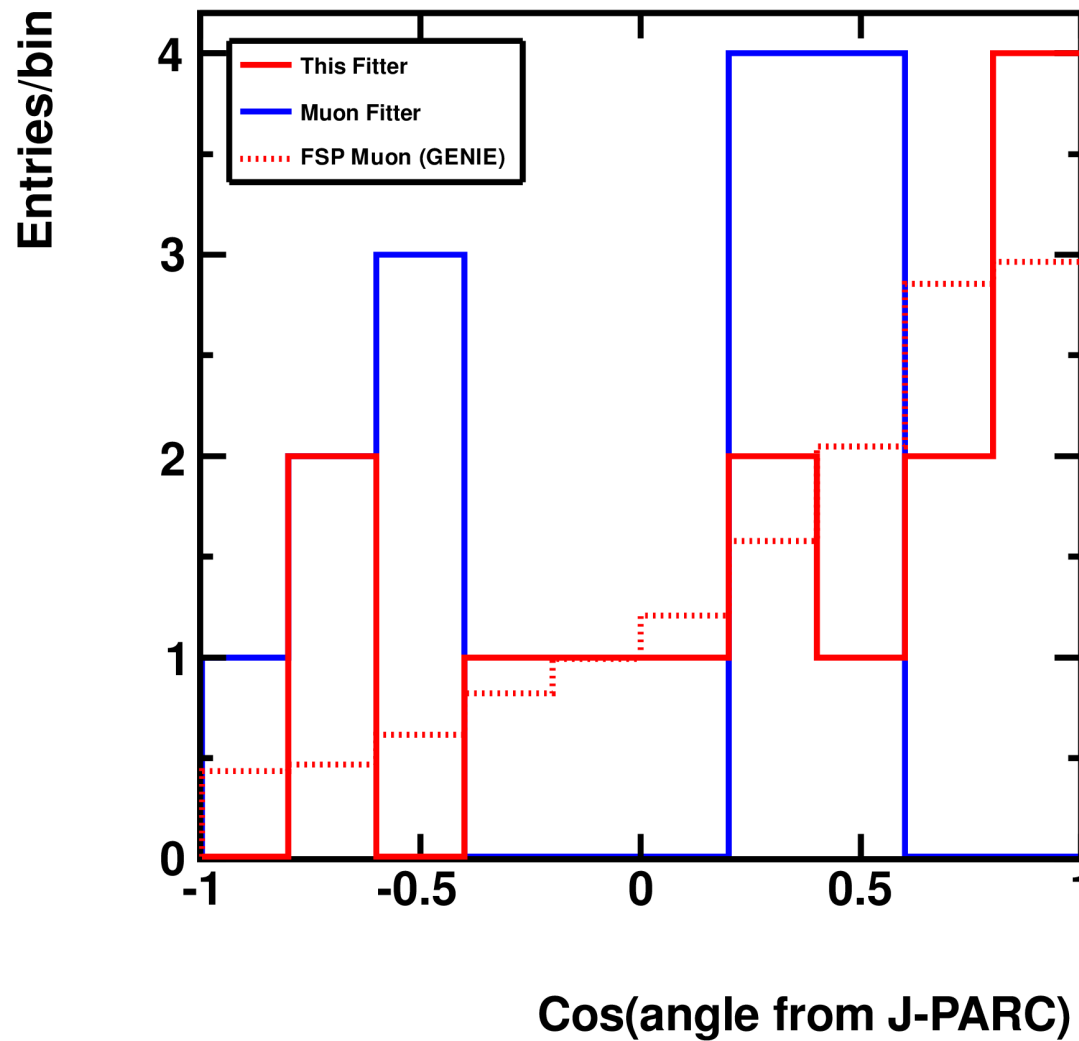
Reconstructed T2K Event Direction

Events from Ikeda Haruo



Reconstructed T2K Event Direction

Events from Ikeda Haruo



Summary

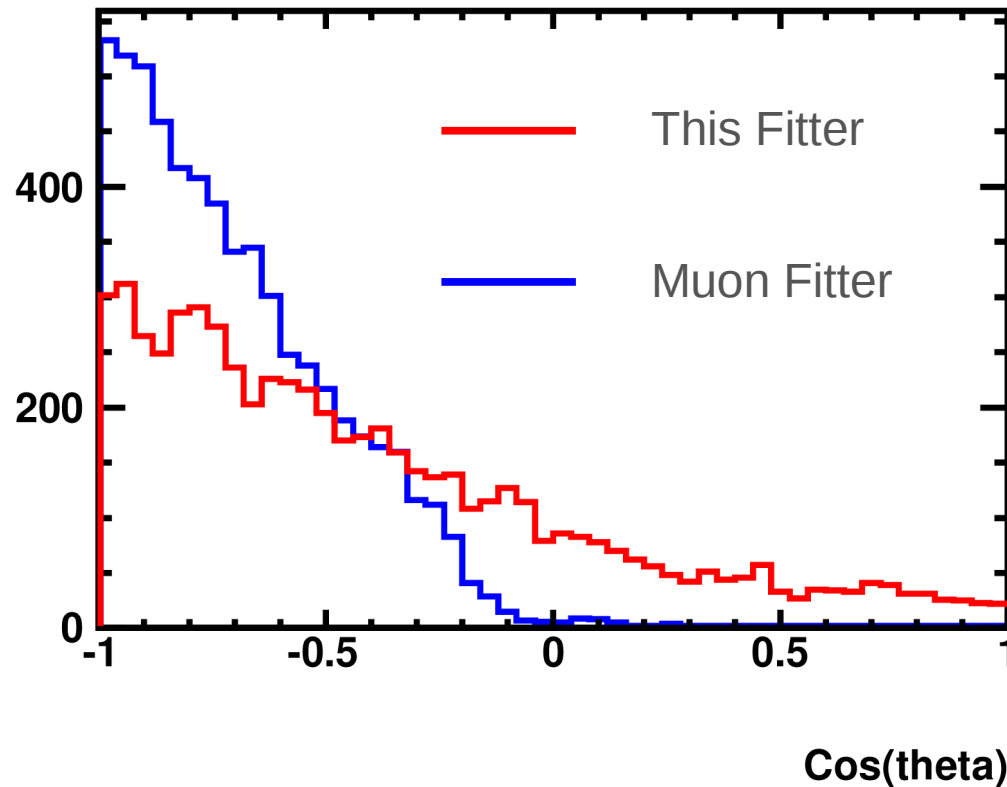
- Identified KLG4 geometry problem causing photon “warp” from ID to OD.
- Identified PMT prepulse problem → Solved by placing cut outside of mean $\pm 3\sigma$ of neighbor PMTs.
- Atmospheric/T2K event candidates direction fitted with μ fitter and new fitter.

To Do

- Explore flavor discrimination using Deep Neural Network
- Improve direction fitter by improving t_0 used in center of $1/\text{Time}$.
- Constrain Dark Matter mass (using WIMP Sim?).
- Also look at specifically very high energy events (~ 50 GeV?) b/c backgrounds are lower at higher energies.

Backup Slides

Zenith Angle of Both Fitters (muons from RTQ run 5000~5099)



- Muon Fitter is fitting through-going muons well.
- New fitter is less accurate for through going muons..