

Plots re-done

Sheean Jolicoeur

Department of Physics & Astronomy, University of the Western Cape, Cape Town 7535,
South Africa

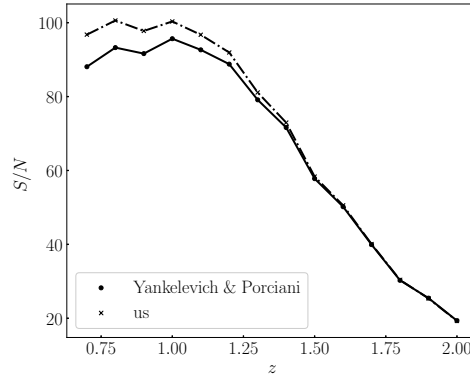


Figure 1. Reproduction of Yankelevich & Porciani Newtonian SNR.

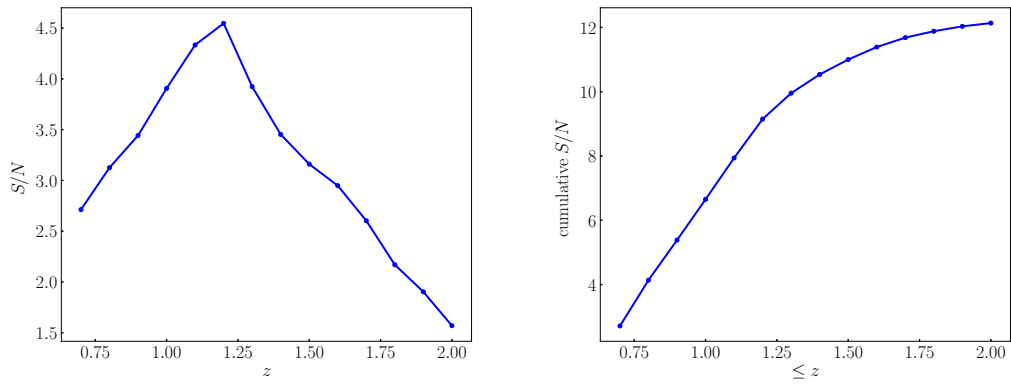


Figure 2. Relativistic SNR per z -bin (*left*) and cumulative (*right*) for a Stage IV $H\alpha$ survey

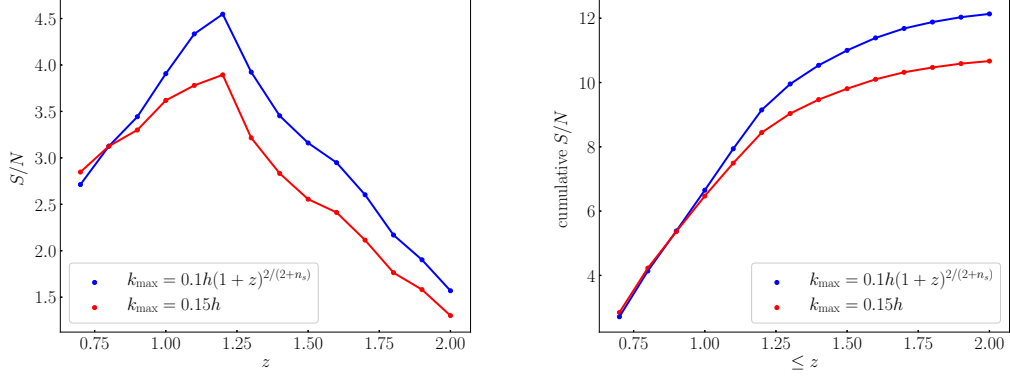


Figure 3. Effect of changing k_{max} on SNR z -per bin (*left*) and cumulative SNR (*right*).

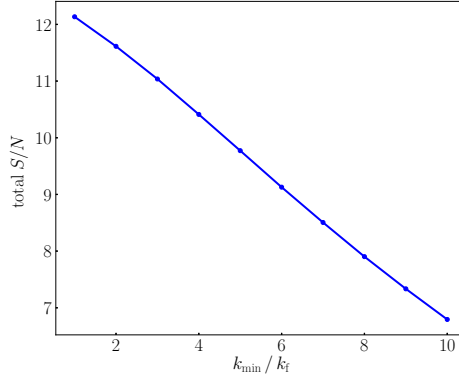


Figure 4. Effect of changing k_{min} on total relativistic SNR

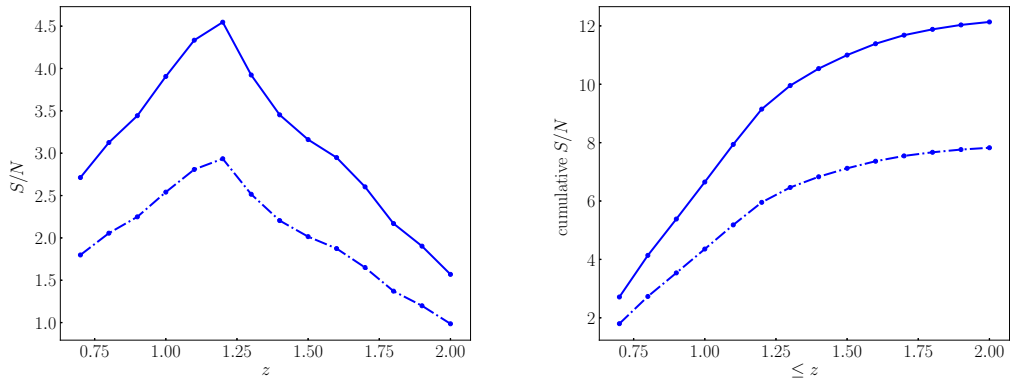


Figure 5. The effect of omitting the second-order relativistic contribution to the bispectrum (dot-dashed curves)

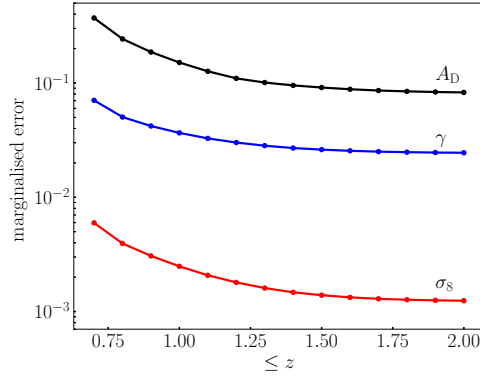


Figure 6. Marginal errors on the relativistic contribution, growth index and clustering amplitude, using the full bispectrum.