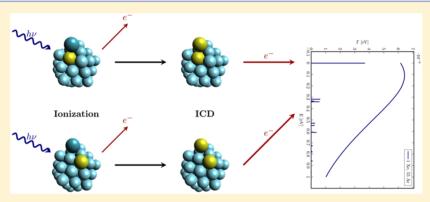
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# Long-Range Interatomic Coulombic Decay in ArXe Clusters: <sub>2</sub> Experiment and Theory

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  - Supporting Information



ABSTRACT: We report autoionization channels of Ar inner valence ionized states in mixed ArXe clusters and compare our experimental data obtained by electron-electron coincidence spectroscopy to our theoretical simulations for representative cluster structures. The combined experimental and theoretical data show that the autoionization of Ar 3s<sup>-1</sup> in ArXe is dominated by interatomic coulombic decay (ICD) to Xe atoms in the second and higher coordination shells of the originally excited atom. Clusters with a range of sizes, compositions, and structures were probed. The Xe content in the clusters was varied between 10% and 53%. Besides ICD, also electron transfer mediated decay (ETMD(3)) was found important in many of the calculated spectra, although it is seen with less intensity in the experimental spectra. From the calculations, we identify structural motifs in which the ETMD rate is minimized vs the ICD rate and suggest that these are preferentially realized in our experiment, in which clusters are formed by supersonic expansion of an Ar-Xe mixture. Suggested cluster structures either feature a clear segregation between Ar and Xe fractions, e.g., Xe core-Ar shell systems, or contain a few Xe atoms singled out at surface sites on an Ar cluster. These structures differ significantly from the majority of calculated minimum energy structures for ArXe systems of 38 atoms, which might show that the latter, annealed structures are not realized in our experiment. We show experimentally that the relaxation of Ar inner valence states by ICD and ETMD together has an efficiency of unity, within the experimental accuracy, for all clusters probed, except those with the lowest Xe content. The outer valence photoelectron spectra of ArXe are discussed also.

# INTRODUCTION

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27 Electron spectroscopy can make important contributions to the 28 research on composition and structure of free nanoparticles. 29 Besides photoionization, free electrons emerging from nano-30 particles can also be produced by the relaxation of electronically 31 excited states. One such process, which is of particular 32 relevance in weakly bonded systems, is the interatomic or 33 intermolecular coulombic decay (ICD).<sup>2</sup> In ICD, an electronic 34 excitation decays by energy transfer to one of its neighboring 35 atoms or molecules, thus releasing a free electron from the 36 latter site. ICD is an important relaxation channel, e.g., for

inner-valence holes in atoms, and also for core level vacancies, 37 e.g., in H<sub>2</sub>O, where it competes with Auger decay.<sup>3</sup>

By definition, ICD is particularly sensitive to the chemical 39 environment of the atom or molecule in which the primary 40 excitation has taken place. We suggest that this property of ICD 41 decay spectra can be used to derive information on a 42 nanoparticle or a solvation system. The term "excitation" is 43

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44 used here in a broader sense—in the systems we consider in 45 this work, ICD occurs from an excited, singly charged Ar state 46 created by photoionization. In other systems, ICD was also 47 observed from neutral excited states and states with more than 48 one vacancy.

We report here studies in which we produced heterogeneous so clusters of the noble gases Ar and Xe with different sizes, and with compositions ranging from a few Xe dopant atoms in an Ar matrix to clusters containing an equal amount of both species. By the use of electron—electron coincidence spectroscopy and of theoretical simulations on prototypical systems, we show how the radiationless decay spectrum of Ar inner valence (3s) ionized states connects to the structure of the clusters.

Intermolecular coulombic decay initially was predicted from 58 theoretical considerations of the energy levels in singly vs 59 doubly ionized, and doubly vs triply ionized, hydrogen-bonded 60 clusters. The first experimental work some years later used Ne 61 clusters, 5 but quickly was followed by demonstrations of ICD 62 in a diverse range of other systems. Experimental and 63 theoretical progress has been reviewed. 6-8

Soon after, related autoionization processes were discovered in which the initial vacancy is filled by an electron from a 66 neighboring unit. These were termed electron transfer mediated decay (ETMD).  $^{9-12}$  If the excess energy is 68 transferred to a third unit which is subsequently ionized, the 69 process is called ETMD(3); if the electron donor is ionized 70 once more, the process is called ETMD(2); and if the excess 71 energy is used to ionize the initially ionized unit, the process is 72 called "exchange ICD".  $^{13,14}$ 

We find that both energy and charge transfer induced autoionization play roles in Ar—Xe, and we detail the relevant processes below.

Rare gas clusters are suitable prototype systems for studies of ICD, as they can readily be produced by supersonic expansion through a cooled nozzle. The size of the clusters can be changed by varying the expansion parameters. The formation of heterogeneous rare gas clusters by coexpansion of a gas mixture cannot *a priori* be taken for granted, but it was experimentally shown for most combinations of two rare gases when suitable mixing ratios and expansion parameters were used (see references throughout this article).

The first experiments on ICD in heterogeneous systems used Ne—Ar clusters. Those were followed by studies of ICD-like inner-shell decays in aqueous solution. Pioneering work also showed the potential for studies of the interface between a substrate and an adsorbate by ICD. More detailed work on Ne—Ar clusters was recently presented by some of the current authors, showing that an analysis of the ICD spectra allowed to decide between structural alternatives for which the photoelectron data were indiscriminate. For this system, a detailed account of the photoelectron spectra with respect to structural features of the mixed clusters is also available. Showing that mixed clusters is also available.

Several experimental techniques have been used for the study of mixed Ar—Xe clusters, most notably fluorescence spectrossecopy, electron diffraction, and photoelectron spectroscopy. Early work focused on the fluorescence of a single Xe atom embedded in an Ar cluster. Increasing the Xe concentration in the expansion led to the formation of Xe<sub>2</sub> and larger Xe complexes inside the clusters. With increasing Xe content, the formation of Xe core—Ar shell systems with a sharp interface between the two species was shown in photoionization experiments. This finding was confirmed by electron diffraction. Always, the observed Xe content in the clusters

is higher than the Xe content in the expanding gas mixture, as 107 Xe can be condensed much more easily than Ar. 25,27 More 108 recently, photoelectron spectra of small Ar—Xe clusters were 109 analyzed, adding to the findings in earlier work. We will 110 discuss that paper in connection with our results below. For 111 completeness, we mention that Ar—Xe complexes can also be 112 produced by passing clusters from a neat Ar expansion through 113 a zone filled with Xe gas ("pick-up"). Clusters produced in 114 such a way can be structurally quite different from those 115 produced by coexpansion. 116

A number of theoretical works on Ar—Xe clusters aimed at 117 the prediction of minimum energy structures. Most recently, 118 these converged to structures which have the Xe atoms mainly 119 in the interior. <sup>31</sup> Xe was also found to diffuse into the interior 120 of Ar clusters in molecular dynamics simulations after pick- 121 up. <sup>32</sup> The electronic energy levels of very small Ar—Xe clusters 122 were calculated in ref 33, and the secondary electron spectra 123 were simulated for model structures of one argon atom placed 124 on a xenon surface. <sup>34</sup> Interestingly, it was found that Ar 3s<sup>-1</sup> is 125 stable against autoionization in an ArXe dimer, but is 126 destabilized by addition of further Xe atoms or an Ar solvation 127 shell to the system, thereby allowing for combinations of argon 128 and xenon atoms with larger interatomic distances. Again, we 129 will detail these results in conjunction with our current work. 130

The outline of our paper is as follows: We first delineate our 131 experimental and theoretical methods. Calculations were done 132 for a number of model structures which were systematically 133 varied, and are described in the following section. After that, we 134 describe and discuss our theoretical results for the auto- 135 ionization spectra of the model clusters. Experimental results 136 are then shown, and possible conclusions from comparing them 137 to the calculations are given. Details about the experimental 138 technique and calculated spectra for a wider range of structures 139 are given as Supporting Information.

# ■ EXPERIMENTAL SECTION

The apparatus used for the experiments consists of a supersonic 142 molecular jet with a cooled nozzle and a magnetic bottle 143 spectrometer, which detects photoelectrons and secondary 144 electrons produced after ionization with synchrotron radia- 145 tion. The focus on details specific for the current experiment. 147 Commercially obtained Ar and Xe gas was filled into separate 148 cylinders up to pressures suitable for producing a certain mixing 149 ratio. The gases were then allowed to mix before the expansion. 150 Expansion parameters are given in Table 1. Similar expansion 151 t1 conditions are further referred to by the labels "S"—"XL", as 152 given in the table.

Prediction of the sizes of clusters from a supersonic 154 expansion mainly rests on empirical scaling laws,  $^{35}$  which 155 have been the subject of some discussion. Moreover, such 156 scaling laws originally were derived only for pure gas 157 expansions. A recent investigation of cluster sizes in a mixed 158 Ar—Kr expansion yielded a revised scaling law, which essentially 159 amounts to the use of a weighted average over the atom-specific 160 parameters of the two gases. This although the analogy to our case 161 is not complete, as phase segregation occurs in Ar—Kr to a 162 much lesser extent, we expect scaling law cluster sizes 163 between the values calculated for Ar and Xe. Recent analyses of 164 photoionization spectra suggest that mean sizes arrived at by 165 scaling laws are smaller than the actual size, at least for rare gas 166 clusters in the range  $\langle N \rangle < 1000.^{39-41}$  For the expansion 167

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Table 1. Expansion Parameters Used for Cluster  $Production^a$ 

	size label	Xe <sub>in</sub> (%)	T (K)	p (bar)	$\langle N_{ m Ar}  angle$	$\langle N_{ m Xe}  angle$
Ar	(1)	m (, , )	96.5	0.35	44	\ 'Ae/
	, ,					
Ar	(2)		96.5	0.67	200	
ArXe	"S"	1.2	174	0.32	4	74
ArXe	"M"	1.2	174	0.49	11	199
ArXe	"L"	1.2	174	0.68	23	431
ArXe	"S"	3.0	172	0.28	3	57
ArXe	"M"	3.0	172	0.51	12	233
ArXe	"L"	3.0	172	0.68	24	458
ArXe	"S"	5.0	167	0.37	7	129
ArXe	"M"	5.0	171	0.51	13	240
ArXe	"L"	5.0	167	0.68	28	537
ArXe	"XL"	2.5	154	2.12	995	18800
Xe		100	183.5	0.68		615

"Here,  $Xe_{in}$  is the molar fraction of Xe in the gas mixture before the expansion, T is the nozzle temperature, and p is the stagnation pressure. Experiments were performed with  $d=80~\mu m$  (size "S"—"L") and  $d=100~\mu m$  (size "XL") conical nozzles of 15° half opening angle. We basically have an Ar-seeded expansion of Xe gas, as the freezing point of Ar is much lower.  $\langle N_{Ar} \rangle$  and  $\langle N_{Xe} \rangle$  refer to cluster sizes for a pure Ar or pure Xe expansion, respectively, at the given conditions, calculated from a scaling law. Inaccuracies in the calculation of  $\langle N \rangle$  due to fluctuations of the input parameters are less than 6%; this figure does not include systematic errors of the empirical model.

168 conditions labeled "S"-"L", we therefore expect that the scaling 169 law estimates are a lower boundary for the actual cluster size. The expansion chamber for the supersonic jet is separated 171 from the interaction chamber by a non-magnetic, conical 172 skimmer with an opening of 1 mm in diameter (Beam 173 Dynamics). At few cm distance behind the skimmer, the cluster 174 jet was crossed by synchrotron radiation from the BESSY 175 electron storage ring at Helmholtz-Zentrum Berlin. Electrons 176 were detected by a short "magnetic bottle" time-of-flight 177 spectrometer that has been described earlier. 42 Due to its large 178 collection angle, this instrument is particularly suited for 179 experiments that employ electron-electron coincidence 180 detection. Data were recorded in two different beamtimes at 181 the UE112-PGM-1 beamline (sizes "S"-"L" in Table 1) and at 182 the TGM-4 beamline (size "XL" in Table 1). Linearly, 183 horizontally polarized radiation was used. The storage ring 184 was operated in single bunch conditions. Spectra of pure Ar 185 clusters shown for comparison are from ref 41, and a spectrum 186 of pure Xe clusters was recorded with the setup described in ref 187 40.

The entrance aperture and drift tube of the magnetic bottle spectrometer can be independently biased to influence the electron flight times. For the data shown here, slightly positive bias voltages were used (+1.8 to +2.2 V) in order to have all electrons arriving at the detector within one BESSY single bunch period (800 ns). The detection efficiency was determined as 0.6, with the method outlined in ref 42. Data were converted from flight times to kinetic energy by measuring calibration data for atomic photolines of known binding energy. The systematical uncertainty of this procedure

may lead to a common shift of all experimental kinetic energy 198 data shown in this Article of up to 30 meV.

The high detection efficiency of the instrument is a pre- 200 requisite for measuring two-electron emission processes, such 201 as photoelectron emission followed by ICD or ETMD, by using 202 electron—electron coincidence detection. This allows us to filter 203 the full electron spectrum for contributions arising after 204 emission of an Ar 3s photoelectron. ICD/ETMD spectra 205 presented here are based on this method, which has been 206 explained earlier. Technical details and full electron— 207 electron coincidence spectra are given in the Supporting 208 Information.

#### ■ THEORETICAL APPROACH

Experimental coincidence spectra are given by the multitude of 211 kinetic energies  $E_{\rm sec}$  of the secondary electron and the 212 intensities of the peaks. The latter are proportional to the 213 probability of the decay  $P=\Gamma t/\hbar$  as well as the theoretically 214 determinable decay width  $\Gamma=\hbar/\tau$ , which is inversely 215 proportional to the lifetime  $\tau$  of the decay process. The 216 manifold of the secondary energies and decay widths will 217 therefore compose the electron—electron coincidence spec- 218 trum. In the case of noble gas clusters with given geometrical 219 structures, three aspects have to be taken into account: different 220 decay mechanisms, the possibility to decay with multiple 221 interaction partners, and different decay channels  $\beta$  within each 222 decay mechanism. At the current stage of development, we 223 need to neglect the nuclear dynamics which might additionally 224 play an important role.

This means, for the evaluation of the decay width, we are 226 dealing with a sum over three different indices. For a given 227 structure, it is most convenient to start with a decomposition of 228 the system into interaction partners of the different relevant 229 decay mechanisms, i.e., pairs (combination of two atoms 230 regardless of their internuclear distance) for the ICD and triples 231 (combinations of three atoms) for the ETMD(3). These atoms 232 do not necessarily need to form bonds between each other or 233 even be close, but they are characterized according to fixed 234 internal coordinates. Each pair and each triple leads to a decay 235 spectrum that can be described by the energy of the secondary 236 electron and the decay width, which is, to a first order of 237 approximation, independent of other atoms present in the 238 cluster. In the following, this approach will be called the model 239 of pairs and triples.

In order to determine whether a channel  $\beta$  for a given 241 geometry and decay mechanism is open, i.e., whether it is in 242 accordance with energy conservation, knowledge of the 243 energies of the initial and the final states,  $E_{\rm in}$  and  $E_{\rm fin}$ , of the 244 corresponding processes is necessary. If the channel is open, the 245 excess energy is carried away by the emitted electron in the 246 form of its kinetic energy,  $E_{\rm sec}$ . These energies in the model of 247 pairs and triples can be approximated by

$$E_{\rm in} = SIP(X_{\rm in}) \tag{1}$$

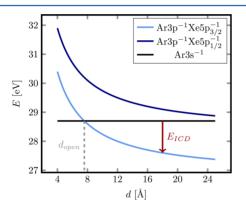
$$E_{\text{fin}}^{\beta} = \text{SIP}(X_{\text{D}}^{\beta}) + \text{SIP}(X_{\text{E}}^{\beta}) + \frac{1}{d}$$
 (2) <sub>250</sub>

$$E_{\text{sec}}^{\beta} = E_{\text{in}}^{\beta} - E_{\text{fin}}^{\beta} \tag{3}$$

where  $X_{\rm in}$  denotes the initially ionized atom and  $X_{\rm D}$  and  $X_{\rm E}$  252 describe the electron-donating atom and electron-emitting 253 atom, respectively.  $\beta$  denotes the decay channel, characterized 254 by the quantum numbers of the ionized atoms in the pairs and 255

256 triples,  $SIP(X^{\beta})$  the single ionization potential of vacancy  $\beta$  in 257 atom X, and d the interatomic distance between the atoms  $X_D$  258 and  $X_E$ . Atomic units are used. The initially ionized atom  $X_{in}$  259 can coincide with one or both of the final state atoms  $X_D$  and 260  $X_E$ . The distribution of the vacancies over the different atoms 261 determines the kind of electronic decay process at hand. Hence, 262 in an Auger process all three atoms coincide, for an ICD 263 process  $X_{in}$  coincides with  $X_D$ , for an ETMD(2) process  $X_D$  and 264  $X_E$  coincide, and for an ETMD(3) process all ionized states are 265 located on different atoms.

The electron emitted by autoionization has the kinetic 267 energy  $E_{\rm sec}$ . If the calculated value of  $E_{\rm sec}$  is negative, then the 268 final state energy is higher than the initial state energy, and the 269 process is energetically not accessible. Hence, the correspond-270 ing channel is closed, as shown in Figure 1 for the Ar  $3p^{-1}$ Xe 271  $5p_{1/2}^{-1}$  final state. Related to the opening of each channel is an 272 internuclear distance which we will call *channel opening distance* 273 in the following.



**Figure 1.** Illustration of channel opening: initial (Ar  $3s^{-1}$ ) and final state (Ar  $3p^{-1}$  Xe  $5p_{3/2}^{-1}$  and Ar  $3p^{-1}$  Xe  $5p_{1/2}^{-1}$ ) energies for ICD in ArXe pairs as a function of internuclear distance, using eq 2 and the experimental ionization potentials given in Table 3. When the final state energy is lower than the initial state energy, the corresponding channel is open. We call the distance at which the energies are equal the "channel opening distance". If the curves do not cross, the corresponding channel is closed.

This ad hoc approach allows us to easily correct for energetic shifts of ionization potentials as observed in larger clusters by adding experimentally or theoretically determined energy shifts  $\Delta E(X_I^\beta)$  for a given vacancy I (= in, D, E) and channel  $\beta$ . This yields the working expression for the kinetic energy of the secondary electron:

$$E_{\text{sec}}^{\beta} = \text{SIP}(X_{\text{in}}) + \Delta E(X_{\text{in}}) - \text{SIP}(X_{\text{D}}^{\beta}) - \Delta E(X_{\text{D}}^{\beta})$$
$$- \text{SIP}(X_{\text{E}}^{\beta}) - \Delta E(X_{\text{E}}^{\beta}) - \frac{1}{d}$$
(4)

The total decay width  $\Gamma$  is given by the sum over the decay widths of all decay mechanisms. In the case of the ArXe 283 clusters, we only consider ICD and ETMD(3), because the 284 ETMD(2) pathway is energetically not accessible:

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$$\Gamma = \Gamma_{\rm ICD} + \Gamma_{\rm ETMD} \tag{5}$$

286 These are given by the sum over the decay widths of all 287 geometrically different pairs i for the ICD and all geometrically 288 different triples j for the ETMD(3), respectively, and over all 289 channels  $\beta$ :

$$\Gamma_{\rm ICD} = \sum_{i,\beta} N_{{\rm ICD},i} \Gamma_{{\rm ICD},i,\beta}$$
(6) <sub>290</sub>

$$\Gamma_{\text{ETMD}} = \sum_{j,\beta} N_{\text{ETMD},j} \Gamma_{\text{ETMD},j,\beta}$$
(7) 29

Here,  $N_{\text{ICD},i}$  and  $N_{\text{ETMD},j}$  denote the numbers of geometrically 292 equal pairs and triples in a given cluster structure. The total 293 number of pairs reads  $N_{\text{ICD}} = N_{\text{in}} N_{\text{D/E}} = N_{\text{Ar}} N_{\text{Xe}} = \sum_i N_{\text{ICD},i}$ , and 294 the number of triples is  $N_{\text{ETMD}} = N_{\text{in}} N_{\text{D/E}} (N_{\text{D/E}} - 1) = 295$   $N_{\text{Ar}} N_{\text{Xe}} (N_{\text{Xe}} - 1) = \sum_j N_{\text{ETMD},j}$ . The numbers of geometrically 296 equivalent pairs  $N_{\text{ICD},i}$  and triples  $N_{\text{ETMD},j}$  strongly depend on 297 the structure of the cluster. From these relationships, it can be 298 seen that a higher xenon content in the cluster statistically 299 favors ETMD(3) over ICD.

In previous work,  $^{34,44}$  we have shown that the decay width 301 for a single pair or triple for a distinct channel  $\beta$  following 302 Wentzel,  $^{45}$  Feshbach,  $^{46,47}$  and Fano  $^{48}$  reads 303

$$\Gamma_{\beta}(E_{\rm res}) = 2\pi \left| \left\langle \Phi_{\rm in} | \hat{V} | \chi_{\beta} \right\rangle \right|^2 \tag{8}$$

where  $|\Phi_{\rm in}\rangle$  and  $|\chi_{\beta}\rangle$  are the initial and final states, respectively, 305 and  $\hat{V}$  is the interaction operator between the two, which can be 306 approximated by its asymptotic behavior. For the ICD the 307 approximation reads

$$\Gamma_{\text{ICD},i,\beta} = (2J_{\text{in}} + 1) \times \frac{3c^4}{8\pi} \sum_{M'_{\text{in}}} \left| \begin{pmatrix} J'_{\text{A}} & 1 & J_{\text{A}} \\ -M'_{\text{A}} & M'_{\text{A}} - M_{\text{A}} & M_{\text{A}} \end{pmatrix} \right|^2 \frac{\sigma^{(X_{\text{E}})}(\omega_{\text{vp},\beta})}{R_i^6 \omega_{\text{vp},\beta}^4 \tau_{\text{in},\beta}}$$
(9) 300

where  $J_{\rm A}$ ,  $J_{\rm A}'$  and  $M_{\rm A}$ ,  $M_{\rm A}'$  denote the total angular momentum of 310 the initially ionized atom (which is the same as the electron 311 donor atom) in the initial and final states, respectively, which 312 otherwise depend on intrinsic atomic properties which can be 313 determined experimentally, like the ionization cross sections 314  $\sigma^{(X_{\rm E})}(\omega_{\rm vp,\beta})$ , the radiative lifetimes of the initially ionized state 315  $\tau_{\rm in,\beta}$ , and the excess energy transferred to the emitting atom 316 (the energy of the virtual photon)  $\omega_{\rm vp,\beta}$ .

For the ETMD(3) the approximation reads

 $\Gamma_{\text{ETMD},i,\beta} =$ 

$$\frac{c}{\pi} \sum_{m,M_{\text{in,D'}}} \frac{\Theta_{m,k}(\alpha_j) \sigma^{(X_{\text{E}})}(\omega_{\text{vp},\beta}) \left| \langle \tilde{D}_{m,j,\beta}(M_{\text{in,D'}}) \rangle \right|^2}{R_j^6 \omega_{\text{vp},\beta}}$$
(10) 319

$$=\frac{c}{\pi R_j^6} \sum_{M_{\text{in},D'}} \left[ \left| \left\langle \tilde{D}_{x,j,\beta}(M_{in,D'}) \right\rangle \right|^2 (2 + \sin^2 \alpha_j) + \right.$$

$$\left| \left\langle \tilde{D}_{z,j,\beta}(M_{\text{in},D'}) \right\rangle \right|^2 (1 + \cos^2 \alpha_j) \right] \frac{c\sigma^{(X_E)}(\omega_{\text{vp},\beta})}{\omega_{\text{vp},\beta}} \tag{11)}$$

where  $\tilde{D}_{m,j,\beta}$  are calculated transition dipole moments and 321  $\Theta_m(\alpha_j)$  is a function depending on the angle  $\alpha_j$  of the triple and 322 the direction of the dipole transition moment m (compare ref 323 34).

The decay width for the total system is, in the end, 325 normalized to the decay width per initially ionized atom since 326 we are interested in the events after one ionization which could 327 be located at any of the atoms of the corresponding atom type. 328 We thereby yield an averaged decay width for the cluster 329 structure. This step allows the comparison of the total average 330

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Table 2. Structural and Calculated Properties of Some Model Clusters<sup>a</sup>

#	no. of Xe	no. of Ar	$\%~{\rm Xe_{cl}}$	structure class	position	relative position	ICD	ETMD	% ICD	related Figure
1	1	12	7.7	cs			0.000	0.000		
2	1	54	1.8	cs			0.000	0.000		7
3	2	53	3.6	cs			0.033	0.167	16.7	7
4	1415	642	68.8	cs			7.039	2.573	73.2	SI
5	561	3310	14.5	cs, icosahedral			1.654	0.256	86.6	8
6	561	3310	14.5	cs, cuboctahedral			1.723	0.355	82.9	8
7	1	13	7.1	cs	surface		0.792	0.000	100.0	4
8	1	13	7.1	cs	edge		0.013	0.000	100.0	
9	1	13	7.1	cs	vertex		0.050	0.000	100.0	
10	2	13	13.3	cs		closest	0.101	0.151	40.1	5
11	2	13	13.3	cs		middle	0.159	0.008	95.4	5
12	2	13	13.3	cs		farthest	0.159	0.002	98.9	5
13	1	55	1.8	cs	surface		0.064	0.000	100.0	4
14	1	55	1.8	cs	edge		0.048	0.000	100.0	
15	1	55	1.8	cs	vertex		0.027	0.000	100.0	
16	1	54	1.8	in	edge		0.077	0.000	100.0	
17	1	54	1.8	in	vertex		0.053	0.000	100.0	
18	2	53	3.6	in			0.117	0.055	68.1	
19	6	49	10.9	in			0.375	0.405	48.1	6
20	6	49	10.9	mixed			0.292	0.674	30.3	6
21	2	36	5.3	cs			0.117	0.151	43.6	SI
22	4	34	10.5	mixed			0.248	0.587	29.7	SI
23	5	33	13.2	mixed			0.335	0.844	28.4	SI
24	13	25	34.2	mixed			0.828	5.414	13.3	SI
25	24	14	63.2	mixed			1.488	16.718	8.2	SI
26	25	13	65.8	in			1.432	4.995	22.3	SI
27	35	3	92.1	in			2.017	7.051	22.2	SI

"The cluster structures are characterized by their respective numbers of Ar and Xe atoms, and are arranged in three groups according to gross structural properties. The first group (#1-6) contains Xe core—Ar shell systems of various sizes and relative Xe contents. In the second group (#7-20), model structures for small clusters are displayed, in which 1-2 Xe atoms are added at systematically varied locations to an icosahedral Ar cluster consisting of one or two full shells, or in which 1-6 Ar atoms of such a cluster were replaced by Xe atoms. The last group (#21-27) contains data for calculated minimal energy structures of mixed ArXe clusters. The cluster structures are classified as core—shell structures (cs), structures with segregated xenon atoms (in), and completely mixed structures. The columns "ICD" and "ETMD" are the respective partial decay widths in  $10^{-4}$  eV, and "% ICD" gives the percentage of ICD in the total decay width. See text for details.

331 decay widths between cluster structures of different sizes and 332 compositions.

In this work we evaluate the secondary energies and decay widths with the program HARDRoC<sup>20,49</sup> using the expersis imental ionization energies of Table 3 (below) and data from the literature given in Tables II and III of ref 34.

## 337 CLUSTER STRUCTURES

338 From the experimental results, we can obtain two hints about 339 the cluster structure: the minimum size and the xenon content. 340 In the following we will use these to construct representative 341 cluster structures, for which we will simulate the ICD and 342 ETMD spectra. At this point we put in a variety of structures in 343 order to learn about general structure—spectrum relationships, 344 therefore making it possible to exclude some of them by 345 comparison of the ICD and ETMD spectra to the experimental 346 ones.

In our previous study,<sup>34</sup> the ArXe clusters were assumed to to be large, having more than 1500 atoms. The model structures assumed an fcc structure and not an icosahedral structure, which is energetically favored for small clusters. Most estimated cluster sizes given in Table 1 are much smaller. We use the setimated number of atoms for a pure Ar expansion as the

lower limit for the actual cluster size and start our search for 353 model structures from there. For the smaller clusters measured, 354 these expectation values range from 3 to 21 Ar atoms. The 355 cluster sizes for expansion of a mixed gas are unknown, but for 356 xenon as the second component, a larger size is expected 357 compared to a pure argon expansion, due to the higher 358 polarizability and hence larger cohesive energy of xenon. All 359 clusters contributing to the coincidence spectrum contain at 360 least one xenon atom, as otherwise decay by ICD or ETMD is 361 energetically not allowed.

In a gas expansion, a range of cluster sizes with different 363 structures is produced. However, for small rare gas clusters 364 consisting of less than approximately 1000 atoms, so-called 365 islands of stability were found. <sup>50</sup> In mass spectra, clusters with 366 mass numbers corresponding to icosahedral shapes had an 367 abundance that was (slightly) enhanced over other sizes. We 368 therefore started from icosahedral structures and constructed 369 mixed cluster structures by substituting individual atoms with 370 one or the other kind (Ar or Xe), or adding individual atoms 371 close to the desired position within a given structure. After that, 372 the cluster structures were optimized using the unified force field implemented in the Avogadro program (version 1.1.0) <sup>\$1,52</sup> 374 to give local minimum structures. The method was chosen due 375 to its low computational cost, the possibility to find the next

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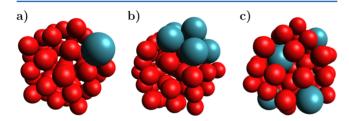
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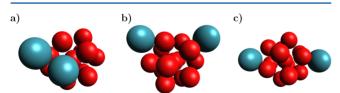
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As discussed in the Introduction, the small clusters will likely 383 deviate from a core—shell structure with a xenon core 384 surrounded by argon shells, as observed for large clusters. 24,32 385 We therefore start from argon clusters of 13 or 55 atoms, and 386 modify them in one of the following ways:

- add one xenon atom on top of a surface/edge/vertex of the argon core, with an example shown in Figure 2a (one additional xenon atom on top of a surface of a 55-atom argon core) (Table 2, #7–9, #13–15)
- substitute one argon atom in a surface/edge/vertex position (Table 2, #16 and #17)
- substitute one or two argon atoms in the core of the argon cluster by xenon atom(s) for the simulation of the core—shell structures (Table 2, #1-3)
- add two xenon atoms on top of surfaces of a 13-atom argon core in different relative positions (see Figure 3), in order to achieve a xenon content in the cluster close to the measured one (see below) (Table 2, #10–12)
- substitute six argon atoms of a 55-atom argon cluster by xenon atoms, either grouping them or distributing them throughout the cluster (see Figure 2b,c), in order to achieve a xenon content in the cluster close to the measured one (Table 2, #19 and #20)



**Figure 2.** Cluster structures derived from a 55-atom icosahedral structure: (a) one xenon atom on the surface, (b) six xenon atoms inside a 55-atom cluster, grouped on one side, and (c) six xenon atoms distributed randomly in a 55-atom cluster. They refer to #13, #19, and #20 in Table 2.



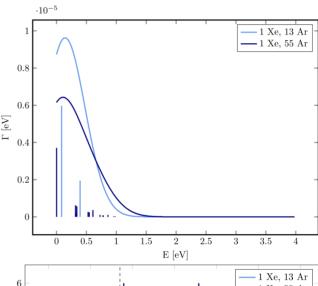
**Figure 3.** Cluster structures composed of 13 argon atoms and two additional xenon atoms on different surfaces of the argon icosahedron, varying in Xe–Xe distance: (a) closest possible, (b) middle, and (c) farthest away. They refer to #10–12 in Table 2.

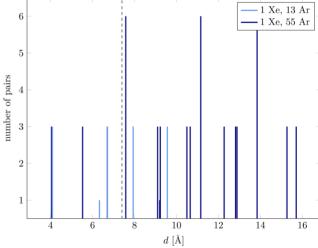
These structures were then optimized using force field  $_{406}$  methods. Furthermore, for comparison with the largest  $_{407}$  measured ArXe clusters, we constructed also larger core—shell  $_{408}$  structures (Table 2, #4–6). These were idealized icosahedral  $_{409}$  and cuboctahedral structures based on the van der Waals radii  $_{410}$  of argon and xenon,  $_{747}$  = 1.88 Å<sup>53</sup> and  $_{738}$  = 2.16 Å.<sup>54</sup>

## THEORETICAL RESULTS

The geometrical properties as well as the total ICD and ETMD 412 decay widths of the investigated structures are shown in 413 Table 2. It is important to remember that the first ICD channel 414 opens at a channel opening distance of 7.58 Å. In the smallest 415 clusters, with only one argon layer around each xenon atom, the 416 Ar—Xe pair distances are below that value; therefore, no ICD 417 can take place. Also, by definition, two Xe atoms are required 418 for ETMD(3); hence, it cannot occur in clusters with only one 419 xenon atom.

We will now discuss features of several structures, using 421 selected examples. First we consider clusters having an argon 422 core of 13 or 55 atoms plus a single xenon atom on one of the 423 surfaces (see Figure 2a for the structure of clusters with the 55- 424 atom core). Secondary electron spectra simulated for these 425 systems are shown in Figure 4, top panel.

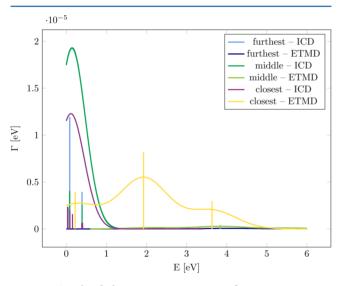




**Figure 4.** Top panel: Simulated electron emission spectra of an Ar 3s vacancy in argon clusters consisting of 13 and 55 atoms, with an additional xenon atom residing on top of one of the argon surfaces. Bottom panel: Distribution of Ar—Xe distances d in the studied ArXe clusters. The ICD channels to final states involving Xe  $5p_{3/2}^{-1}$  and  $5p_{1/2}^{-1}$  vacancies open at 7.58 and 36.00 Å, respectively; the former distance is marked by a vertical dotted line. Only pairs at longer interatomic distances contribute to the electron emission spectra. Larger Ar—Xe distances, as found in the larger cluster, correspond to features at a higher kinetic energy of the ICD electron (top panel).

Because the structures contain only one xenon atom, 428 ETMD(3) is not possible; only ICD is expected. The spectrum 429 pertaining to the clusters with larger (55 atoms) Ar core shows 430 signals at higher energies of the secondary electron. These 431 features can be explained by investigating the Ar-Xe distance 432 distributions in the clusters, shown in the lower panel of 433 Figure 4. Every different value of the atom pair distance results 434 in a different energy of the secondary electron, and only pairs 435 with a distance larger than the channel opening distance 436 (dashed gray line) will contribute to the spectrum. In our 437 model, there is a one-to-one correspondence between 438 interatomic distance d and kinetic energy of the ICD electron: 439 the larger the interatomic distance, the higher is the kinetic 440 energy of the secondary electron for a given channel. Therefore, the spectrum of the cluster with a 55-atom core extends to 442 higher energies. This relation between ICD energy and cluster size holds true for all cluster structures. 443

For clusters with more than one xenon atom, ETMD(3) is 445 also possible. Depending on the positions of the xenon atoms 446 relative to each other and to the Ar core, the spectra are 447 expected to be different. As an example, we discuss clusters with 448 a core of 13 argon atoms and two xenon atoms on different 449 surfaces. Their xenon content is 13.3%, and thus close to the 450 experimental value of 10-12% for the lowest xenon admixture. 451 The cluster structures were illustrated in Figure 3, and 452 simulated spectra are shown in Figure 5. In all three cases,

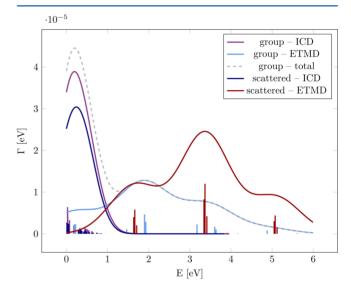


**Figure 5.** Simulated electron emission spectra of an Ar 3s vacancy in argon clusters consisting of 13 argon atoms with two additional xenon atoms on top of two different argon surfaces. ICD and ETMD spectra are shown for three different relative positions of the two xenon atoms. The ICD spectra for the farthest and middle relative position coincide. If the xenon atoms are close to each other, an ETMD process is clearly visible. In the other two cases, the spectra are dominated by the ICD spectrum.

453 both ICD and ETMD(3) are energetically allowed. The ICD 454 spectra do not change qualitatively when the second xenon is 455 added. However, the ETMD(3) spectra are very sensitive to the 456 relative positions of the two xenon atoms. Two aspects have to 457 be taken into account: an energy shift of the peaks due to 458 different charge distances d in the final state (i.e., the 459 interatomic Xe–Xe distance) and the decrease of the decay 460 width with  $R^{-6}$ , R being the distance between the electron 461 transfer unit and the xenon atom ionized in the final state. The

larger the distance between the xenon atoms, the higher are the  $^{462}$  energies of the secondary electrons. For the case of xenon  $^{463}$  atoms on argon surfaces, this at the same time means a larger  $^{464}$  distance between the Ar–Xe pair involved in the electron  $^{465}$  transfer and the electron emitting xenon atom; hence, the decay  $^{466}$  widths are smaller. Therefore, a significant contribution from  $^{467}$  ETMD(3) compared to ICD is seen only if the two xenon  $^{468}$  atoms reside on two adjacent surfaces. The three energetically  $^{469}$  separated peaks in the ETMD(3) spectrum pertain to different  $^{470}$  combinations of the Xe  $^{5}$ p $_{j}^{-1}$  vacancy states, which have a  $^{471}$  substantial fine-structure splitting (compare to Figure 6).

For the 55-atom clusters, a xenon content of 10-12% 473 corresponds to six xenon atoms out of a total of 55 atoms. 474 These can form a connected group in one part of the cluster or 475 can be evenly distributed, as shown in Figure 2b,c. The ICD 476 and ETMD(3) spectra for these structures are shown in 477 f6 Figure 6.



**Figure 6.** Same as **Figure 5**, for clusters consisting of 55 atoms, six of which are xenon atoms. These are either grouped on one side of the cluster (label "group") or evenly distributed in the cluster interior (label "scattered").

For both cases, the overall ICD spectra are very similar and 479 show a maximum at a kinetic energy of approximately 0.3 eV. 480 In the case of grouped xenon atoms, the ICD spectrum extends 481 to slightly higher energies, corresponding to ArXe pairs at 482 slightly larger spatial distances. A more significant difference is 483 observed in the ETMD(3) spectra. The interatomic xenon 484 distance is shorter in clusters, in which the xenon atoms form a 485 connected group. Hence, the triple structure of the ETMD(3) 486 spectrum is seen at lower kinetic energies. Additionally, the 487 probability for electron transfer from a xenon to an argon atom 488 decreases exponentially with the interatomic distance. There- 489 fore, only those xenon atoms which have direct argon atom 490 neighbors make a significant contribution to the total decay 491 width, and the total ETMD(3) decay width is lower for the 492 structure with grouped xenon atoms than for the structure with 493 evenly distributed xenon atoms (see Table 2).

For comparison with the previous assumptions of a xenon 495 core surrounded by argon atoms, we investigated clusters of 13 496 and 55 atoms with one or two xenon atoms in the core. Some 497 of the corresponding spectra are shown in Figure 7.

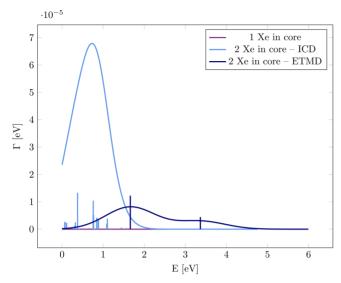


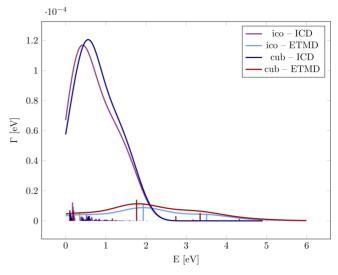
Figure 7. Same as Figure 5, for argon clusters with one or two xenon atoms in the core of the cluster (55 atoms total). All ICD and ETMD channels are closed if the cluster contains only a single xenon atom. For clusters with two xenon atoms in the core, the first ICD channel is open for some pairs, and additionally ETMD is possible for multiple triples. Due to a different distance distribution compared to the clusters with an argon core (Figure 5), the ICD peak is shifted to higher kinetic energies.

For a core consisting of two xenon atoms, both an ICD spectrum and an ETMD(3) spectrum can be seen. Compared to the cluster structures involving an argon core, the ICD peak is shifted to higher ICD electron energies. Additionally, the ICD and ETMD(3) spectra overlap. For ETMD(3), only three of the four different decay channels are energetically allowed.

In the first experiments, very large argon—xenon clusters of 506 ~4380 atoms with a xenon content of 19% were measured, 557 which we approximate in one example by a cluster consisting of 508 561 xenon atoms surrounded by five layers of argon atoms 509 (3871 atoms in total). In order to take the better stabilization of 510 charges in larger clusters into account, the ionization energies 511 of the XL clusters (see Table 3) were used. The spectra are 512 shown in Figure 8.

Compared to the small clusters with a xenon core in Figure 7, the ICD peak is broadened and shifted to higher ICD s1s electron energies, which is expected due to the larger interatomic Ar–Xe distances in the large cluster and the better stabilization of charges, which is relevant mostly for the final state, but also for the initial state. The ETMD(3) peak is shifted to slightly higher energies as well. This allows opening of the fourth decay channel, pertaining to Xe  $5p_{1/2}^{-1}$  Ar  $3p_{1/2}^{-1}$  final states.

We also note that the percentage of decay via ICD is practically independent of the Xe content for large clusters with Xe a Xe core—Ar shell structure. This can be expected, as both decay processes are most pronounced for the atoms in the interface region of the xenon core and the argon shell. The ETMD(3), depending on the electron transfer from one xenon atom to the initially ionized argon atom, and hence on the corresponding orbital overlap, is purely an interface effect that involves the argon atoms with direct xenon neighbors and the xenon layers closest to the argon shell. The ICD has a stronger long-range character, due to which the description of decays with partners of size at least twice the smallest channel opening distance is necessary for clusters. However, the bigger



**Figure 8.** Same as Figure 5, for a large cluster consisting of a xenon core of 561 atoms surrounded by five complete layers of argon atoms (Xe content, 14.5%) for both an icosahedral and a cuboctahedral basic structure. The ICD and ETMD peaks overlap such that the ETMD peak is unlikely to be visible.

the clusters are, the more homogeneous the interface region 535 becomes for clusters of different sizes. This manifests itself in 536 the distribution of relevant Ar—Xe pair distances (compare 537 Figure 4, lower panel). As a consequence, the ICD spectra of 538 two large clusters of different sizes hardly differ.

All these results neglect the effect of nuclear dynamics, which 540 might affect the spectra. Possible reasons are distance- 541 dependent channel closings and openings, a different 542 distribution of interatomic distances due to vibrations, and 543 their different impact on the ICD and the ETMD(3). For 544 example, for the ICD in the neon dimer, the treatment of 545 nuclear dynamics increases the decay width by approximately a 546 factor of 2.56 However, the atomic displacement can be 547 expected to be far less severe in clusters than in dimers or 548 trimers. Additionally, the elongation of one interatomic 549 distance in a cluster automatically results in the shortening of 550 a different interatomic distance, which will reduce the overall 551 effect. Therefore, there will be a broader spectrum of final state 552 charge distances d, resulting in a broadening of the peak. This 553 was taken into account by folding the stick spectrum with 554 Gaussians with a width of 300 and 600 meV for the ICD and 555 ETMD(3), respectively. While the width of the ICD peaks was 556 estimated from the potential energy surface of the ArXe dimer, 557 the situation is more complex in case of the ETMD(3), with 558 three degrees of freedom involved. It was therefore guessed to 559 be twice the width of the ICD peak.

The channel closing due to shortening of the interatomic 561 distance is not relevant for the ETMD(3) process in the ArXe 562 clusters treated in this work, because this would require 563 unnaturally short interatomic distances. However, for the ICD, 564 it might result in a lower number of pairs with open decay 565 channels and hence a decrease of the decay width. Since the 566 peaks in the convoluted spectrum contain contributions from 567 several atom pairs at different distances, this would result in a 568 small peak shift to higher energies of the combined peak at 569 lowest energy. For a more detailed discussion of nuclear 570 dynamics for the ICD in dimers vs clusters, see ref 57.

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Table 3. Properties of the Outer Valence Spectra of Mixed ArXe Clusters<sup>a</sup>

			Ar 3p		Xe 5p <sub>1/2</sub>	$x_{p_{1/2}}$ Xe $x_{p_{3/2}}$		
	size label	Xe <sub>in</sub> (%)	$E_{\rm b}~({\rm eV})$	w (eV)	$E_{\rm b}$ (eV)	$E_{\rm b}~({\rm eV})$	w (eV)	Xe <sub>cl</sub> (%)
Ar	(1)		15.3	1.1				
Ar	(2)		15.1	1.3				
ArXe	"S"	1.2	15.36	0.9	13.07	11.75	0.85	12
ArXe	"M"	1.2	15.30	1.0	12.97	11.61	0.85	11
ArXe	"L"	1.2	15.25	1.1	12.91	11.52	1.08	10
ArXe	"S"	3.0	15.39	0.8	13.06	11.68	1.02	29
ArXe	"M"	5.0	15.31	0.6	12.96	11.44	1.24	53
Xe					12.76	11.19	1.18	100
ArXe	"XL"	2.5	15.15	1.3	12.59	11.01	1.24	19
Ar, Xe	(th.)		15.3		13.0	11.5		
(atomic) <sup>59,6</sup>	50		15.76   15.94		13.43	12.13		

"Spectra of pure Ar and Xe clusters are included for reference. All spectra were recorded at  $h\nu = 17$  eV, except for the pure Xe clusters ( $h\nu = 60$  eV). Binding energies  $E_b$  were determined as the center of gravity of the respective feature, while band widths w are the fwhm of a Gaussian fit. The average Xe content of the cluster ensemble, Xe<sub>cl</sub>, was determined from the areas of the respective photo lines, corrected by their atomic photoionization cross sections (Ar, 33.0 Mb; Xe, 51.3 Mb). Uncertainties are estimated as  $\pm 3\%$  for the Xe content and 0.05 eV for the binding energies. The row labeled "(th.)" gives the binding energies used in the simulations of this paper, and the last row shows their atomic counterparts. Values for Ar 3p in the last row refer to the fine structure states.

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#### EXPERIMENTAL RESULTS

**Outer Valence Spectra.** This section describes and discusses the outer valence spectra of the mixed Ar–Xe clusters. The main information we obtain from these spectra within the context of our theoretical considerations are the single ionization potentials and the relative Xe contents in the mixed clusters (see Table 3).

Figure 9 shows the outer valence spectra of small and large 580 ArXe clusters, compared to clusters of the pure gases. 581 Qualitatively similar spectra have been published without 582 detailed discussion in ref 28. Differences are seen in particular 583 for the Ar component. In pure Ar clusters, outer valence 584 photoionization leads to a broad band, caused by both spin-585 orbit coupling and crystal field splitting. A distinction between 586 the Ar  $3p_{1/2}$  and the Ar  $3p_{3/2}$  derived cluster bands therefore cannot be made. The relative importance of these two mechanisms remains under debate. For larger clusters (e.g.,  $\langle N \rangle$  = 190, black trace), the maximum at low binding energies becomes more pronounced and is identified with emission from the cluster interior. 40,61 At the particular photon energy selected here, atop of this band a sharp feature is also visible (in Figure 9 at a kinetic energy of about 1.8 eV). For 594 larger clusters (above  $\langle N \rangle \approx 100$ ), over a range of photon energies of about 3 eV, its appararent binding energy changes in a way which is reminiscent to photoemission of crystalline bulk matter ("dispersion").41,62

Ar 3p spectra of the smaller mixed clusters (Figure 9c and 599 Figure 10) show neither of these traits. Rather, the Ar band is 600 symmetric, less wide than in the pure clusters, and at a higher 601 binding energy. Even for very large clusters (Figure 9b), the 602 asymmetry of the 3p band and the "dispersing feature" do not 603 appear. We characterize the spectral shape by giving a single 604 value for the Ar 3p binding energy and the full width at half-605 maximum (fwhm) of the feature. Values for different expansion 606 parameters, and for both Ar and Xe valence lines, are collected 607 in Table 3.

We find experimentally that the clusters we produced have a 609 Xe content, Xe<sub>cl</sub>, between 10% and 50%. A slight decrease of 610 the binding energy with cluster size is seen for all outer valence

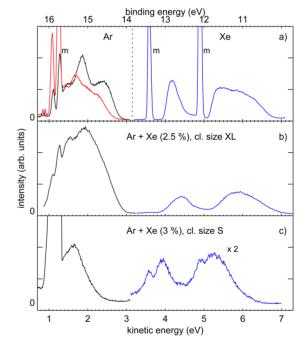
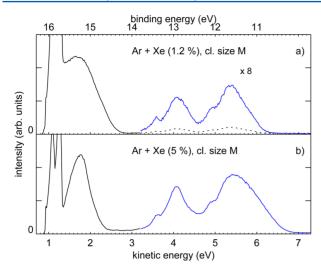


Figure 9. Outer valence photoelectron spectra of mixed Ar–Xe clusters, in comparison to the pure species. Panel (a) shows the outer valence region of homogeneous Ar(1) (red), Ar(2) (black), and Xe clusters (see text and Table 1 for details). Panels (b) and (c) show spectra of the mixed species with different mean sizes. Sharp lines marked "m" (for "monomer") result from photoionization of uncondensed atoms into the Ar  $3p_{1/2,3/2}$  and Xe  $5p_{1/2,3/2}$  final states. Labels in (b) and (c) give the Xe content in the expanding gas mixture (Xe<sub>in</sub>), which is lower than the Xe content observed in the heterogeneous clusters (Xe<sub>d</sub>). The photon energy was 17 eV, in contrast to the pure Xe cluster spectrum (60 eV).

lines, and is attributed to a larger final state polarization energy 611 in larger clusters.

For further interpretation of the shape of the Ar peaks, we 613 refer to photoemission spectra of condensed Ar monolayers, 614 measured in several settings. 63,64 Spectra were reported for 615



**Figure 10.** Outer valence photoelectron spectra of mixed Ar—Xe clusters from gas mixtures with different Xe concentrations. For better visibility, a scaling factor is applied to the Xe part of the spectrum in panel (a). The dotted line shows the unscaled spectrum in this region. See Figure 9 and text for details.

616 physisorption of Ar on two different metal single-crystal 617 substrates, and for Ar atop of a Xe spacer layer adsorbed on 618 the metal. While the binding energy depends on the substrate, 619 the spectral shape is very similar in all cases. Most spectra were 620 recorded for emission along the surface normal and show a 621 double peak split by about 0.5 eV, leading to a structure with a 622 fwhm (w) of about 1 eV. Although the splitting is larger than 623 the gas phase fine structure split of 0.18 eV, the associated 624 states have been assigned to Ar  $3p_{1/2}$  and  $3p_{3/2}$ . A crystal field 625 splitting of the 3p<sub>3/2</sub> state has been assumed to be also present, 626 but with a smaller value of approximately 0.1-0.2 eV.64 The 627 spectrum clearly changes when going to an emission angle of 628 40° with respect to the surface normal (Figure 2 in ref 64). The 629 higher binding energy peak significantly loses intensity, and the 630 spectrum is now dominated by a single peak comprising both 631 crystal field split substates of Ar  $3p_{3/2}$ , with a fwhm (w) of only 632 0.4 eV. As our measured spectrum is comprised of 633 contributions recorded under all emission angles with respect 634 to the cluster surface, grazing emission will be the rule. We 635 therefore believe that the arguments given above make it 636 plausible that only a single peak is observed in all our spectra. Emission from the Xe 5p state shows a much larger fine 638 structure splitting than that from the Ar 3p state; it dominates 639 the spectrum even in clusters (where other broadening 640 mechanisms are also present). The shift of the cluster bands 641 with respect to the monomer lines in the largest mixed clusters 642 we have probed (Figure 9b) are similar to the ones in the pure Xe spectrum, the scaling law size of which is at somewhat more 644 than five filled shells. This indicates an approximate upper limit 645 for the size of the Xe core in our mixed clusters, as their binding 646 energy shift might be somewhat higher than that in pure clusters, due to polarization of covering Ar layers. Xe cluster 648 features in the spectra of mixed clusters appear less asymmetric 649 than in the pure Xe spectrum, and also less asymmetric than in 650 some of the ArXe spectra shown in ref 28. This might be caused 651 by the difference in photon energy ( $h\nu = 90$  eV in ref 28 and 652 for the pure Xe spectrum,  $h\nu = 32$  eV for the mixed ArXe 653 spectra in this work).

In Figure 10, we compare the outer valence spectra of conditions with similar expansion conditions but different

composition of the gas mixture. A Xe-rich mixture obviously 656 leads to clusters with more intense Xe photolines, but besides 657 that the main change is a narrowing of the Ar band. A 658 comparison with the Ar monomer features (clipped in the 659 figure) also shows an increased degree of condensation for the 660 Ar gas in the expansion.

For the smallest clusters we produced, a possible model is  $^{662}$  provided by calculated minimum energy structures of  $^{663}$  Ar $_{\rm N}$ Xe $_{\rm 38-N}$  (Figure 7 in ref 31). We there see that, for a Xe  $^{664}$  content of less than 60%, core—shell systems are *not* formed,  $^{665}$  and although the Xe atoms tend to connect, the degree of  $^{666}$  Ar $_{\rm X}$ Xe mixing is large. While for low Xe content (approx- $^{667}$  imately 10%) some Ar atoms have only Ar nearest neighbors,  $^{668}$  for larger Xe content all Ar atoms seem to see both Ar and Xe  $^{669}$  nearest neighbors. We believe this leads to the Ar 3p narrowing  $^{670}$  pointed out above. For Figure 10, possibly, some clusters of our  $^{671}$  ensemble already are in the core—shell regime, which would  $^{672}$  lead to an even lower 3p width.

For the largest clusters we have produced (Figure 9b), the 674 low binding energy of the Xe lines and the broadening of the Ar 675 feature without appearance of the "bulk Ar" maximum support 676 formation of a Xe core covered by at least two layers of Ar. If 677 the scaling law size for a pure Ar expansion is taken as the lower 678 limit for the cluster size, from the observed Xe content we 679 arrive at clusters composed of a Xe core with four layers, 680 covered by three layers of Ar (see Supporting Information). 681

Finally, comparing Figure 10 to Figure 9 reveals that much 682 larger changes in the valence emission spectra occur as a result 683 of changes in size than changes of the (relative) Xe content. 684 This finding is supported by the spectra shown in ref 28.

Inner Valence Spectra. We now focus on the Ar inner 686 valence (3s) vacancy states.

Figure 11 shows the Ar part of the inner valence spectrum for 688 f11 representative ArXe clusters. The total photoelectron signal 689 looks similar to literature spectra for pure Ar clusters. 66,67 A 690 low-energy tail seen for the Ar 3s monomer line results from 691 the transmission properties of the electron spectrometer, 692 together with the strongly positive angular distribution 693 parameter of this line. 67,68 The Ar 3s cluster line shows no 694 splitting into bulk and surface components, different from the 695 literature on pure clusters but in agreement with our discussion 696 of the outer valence spectra. The binding energy of the cluster 697 line has values between 28.85 and 28.70 eV for all cluster 698 ensembles with size labels "S"—"L", and 28.67 eV for the largest 699 clusters measured. A small but systematic decrease of the 700 binding energy is observed when the cluster size is increased. 701 The value assumed in the simulations is 28.7 eV.

If we produce a spectrum only from those electrons recorded 703 as part of a two-electron coincidence (electron pair with kinetic 704 energies  $(e_1,e_2)$ ) the apparent intensity drops (compare panel b 705 to panel a, and panel d to panel c in Figure 11). This can partly 706 be attributed to the finite detection efficiency of the 707 spectrometer. Moreover, the monomer part of the Ar 3s signal 708 completely disappears, as 3s photoionization of an uncon-709 densed Ar atom in the gas jet cannot lead to emission of a 710 second electron. Further features of these diagrams and the 711 least-squares fits shown in the Figure are discussed below.

ICD/ETMD Spectra. In Figure 12, we show the spectra of 713 ft2 ICD/ETMD electrons pertaining to emission of an Ar 3s 714 cluster photoelectron. Due to their low kinetic energy, without 715 use of a coincidence method they could hardly be separated 716 from the background of inelastically scattered photoelectrons. 69 717 This and the following Figure constitute the central 718

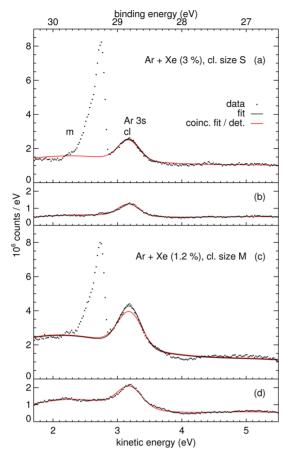


Figure 11. Electron spectra of mixed Ar-Xe clusters after photoionization in the inner valence region. Symbols show the Ar 3s photoline from clusters ("cl") and uncondensed Ar monomers ("m"), atop of a background resulting from inelastic intracluster scattering of outer valence photoelectrons ("excitonic satellites").65 The photon energy was  $h\nu = 32$  eV. Panel (a) shows all electrons accumulated. For the same conditions, in panel (b) we show only electrons that were detected in coincidence with a secondary electron of lower kinetic energy. The black solid trace in (a) and red solid trace in (b) result from a least-squares fit to the cluster part of the spectrum. The red solid trace in (a) is the fit shown in (b), divided by the detection efficiency of the spectrometer of 0.6. Panels (c) and (d) show the same type of data for smaller clusters with a lower Xe content. See text for details.

719 experimental result of the article. Spectra recorded at  $h\nu = 34 \text{ eV}$ 720 as a cross-check quantitatively agree to those shown here. This 721 underpins our assignment of this intensity to an autoionization process. Further details on the data acquisition and analysis 723 methodology, as well as the 34 eV spectra, are given in the Supporting Information. 724

Spectra for the three different cluster sizes in Figure 12 are 726 significantly different: Spectra for the larger clusters acquire 727 more intensity in the 1-1.5 eV region and have less intensity for energies below 0.5 eV. In contrast to that, the ICD/ETMD spectrum hardly varies when the composition of the expanding gas mixture, and with that the relative Xe content of the clusters, is changed (Figure 13). The full set of ICD/ETMD 732 spectra is shown as Supporting Information.

We now discuss which fraction of Ar 3s vacancy states relaxes 734 via either ICD or ETMD. A method to derive this information 735 has been established by some of us. <sup>70</sup> Briefly, we correct the 736 intensity of the Ar 3s photoline in the coincident spectra by the

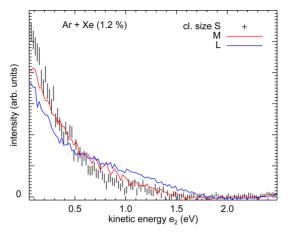


Figure 12. Energy spectrum of all coincident secondary (ICD or ETMD) electrons of kinetic energy  $e_2$  pertaining to primary electrons of kinetic energy  $e_1$  in the Ar 3s binding energy region. Spectra were recorded with a photon energy of  $h\nu = 32$  eV. Black error bars show the data points for the smallest clusters measured ("S"); two larger clusters sizes are shown by the red and blue traces. Error bars for the latter are smaller than the ones shown and have been omitted. For better comparison, all spectra are shown area-normalized.

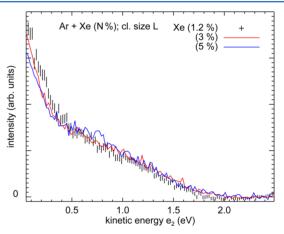


Figure 13. Energy spectrum of all coincident secondary (ICD or ETMD) electrons, kinetic energy  $e_2$ , for clusters of the same size, but from gas mixtures with different Xe concentration. See Figure 12 for details.

detector efficiency, and divide it by the total (coincident and 737 non-coincident) Ar 3s intensity. The result gives the branching 738 ratio of decay via ICD/ETMD vs the sum of all channels, 739 including those not involving electron emission. (If all Ar 3s 740 vacancies decay by emission of another electron, the coincident 741 and non-coincident count rate differ only by the probability for 742 the spectrometer to actually detect the secondary electron (0.6 743 for this data set).) To arrive at quantitative results, we have 744 performed least-squares fits of the non-coincident and 745 coincident e1 spectra. The fits assumed one or two Gaussian 746 peaks, respectively, atop of a background modeled by two more 747 Gaussian curves with very large widths. The peak pertaining to 748 the Ar 3s cluster photoline, with the background added, is 749 shown by the black solid trace in Figure 11a and red solid trace 750 in Figure 11b. Moreover, the line labeled 'c.fit/det.' shows the 751 fit to the coincident events, corrected by the detection 752 efficiency, and plotted atop of the non-coincident intensity. 753 This virtually agrees with the fit to the total spectrum in 754 Figure 11a, but not in 11c. The corresponding figure for the 755

f14

756 ICD/ETMD efficiency comes out as one for 11a, and 757 significantly smaller than one for 11c. Figure 14 shows all

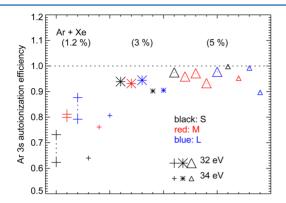


Figure 14. Efficiency of the decay of Ar 3s ionized states in ArXe clusters by emission of a secondary electron via ICD or ETMD. Values are arranged by Xe content of the initial gas mixture ("+" symbols, 1.2%; asterisks, 3%; triangles, 5%). Symbol sizes indicate the photon energy (large symbols, 32 eV; small symbols, 34 eV), and color indicates the cluster size (black symbols, S; red, M; blue, L). See text for details.

758 results of this analysis. The error in the autoionization efficiency 759 shown is mainly determined by uncertainties in the peak/760 background separation. For the gas mixture with a  $Xe_{\rm in}$  of 1.2%, 761 we therefore show results from fits using two different choices 762 of the background shape in Figure 14. Statistical errors are 763 much smaller than the differences between the pairs of values 764 arrived at such.

765 We found an Ar 3s autoionization efficiency which is 766 compatible with unity, within the accuracy of our experiment, 767 for all clusters except those with the lowest Xe content (Xe<sub>cl</sub> 768 approximately 10-12%). For the latter, also a decrease of the 769 efficiency with decreasing cluster size is seen. Two main reasons 770 can be identified that may cause a decrease in autoionization 771 efficiency. First, in small clusters with a low Xe content, for 772 some Ar atoms no Xe partners allowing either ICD or ETMD 773 might be available. A simple example are clusters containing 774 just a single Xe atom. This mechanism would gain in 775 importance for cluster ensembles with small  $\langle N \rangle$ . Second, if 776 only Xe partners at very large distances are available, ICD might 777 be outpaced by other decay processes, most importantly 778 fluorescence.

# 79 DISCUSSION

780 We now interpret our experimental findings in view of the 781 simulated results described in the Theoretical Results section. 782 Figure 14 shows that, for Ar 3s<sup>-1</sup> states, autoionization by either 783 ICD or ETMD is the only mode of relaxation, except for the 784 clusters with the smallest Xe content. We consider this a 785 remarkable result, because ICD of these states is only possible 786 with a Xe atom in the second coordination shell of the decaying 787 Ar vacancy (or shells even farther apart), which significantly 788 slows down the ICD process. Also, decay rates for ETMD were 789 found to be small in early studies of this effect. 9 Nevertheless, 790 even in this situation, the non-local autoionization channels 791 foreclose radiative decay.

792 Structurally, all experimental ICD/ETMD spectra have a 793 maximum at or near the lowest kinetic energy that could be 794 measured (50 meV), and they decrease in intensity to less than 795 half for kinetic energies already below 1 eV. Qualitatively, this shape fits to the ICD contributions in the autoionization 796 spectra discussed in the Theoretical Approach section, but not 797 to the ones from ETMD. The calculated ETMD spectra have a 798 typical three-fold structure with peaks between 3.5 and 5 eV, 799 1.5–2.5 eV, and around 0 eV kinetic energy. The former two 800 features seem not to be present in the experimental spectra. 801 Due to restrictions in the detector electronics, electron pairs 802 with very similar kinetic energy cannot be detected. The highest 803 kinetic energy for an autoionization electron that can be 804 detected in coincidence with a photoelectron of  $\sim$ 3.5 eV 805 ( $h\nu = 32$  eV) is around 2.5 eV. Very weak intensity is observed 806 at this energy, and also at higher  $e_2$  energies that can be probed 807 at  $h\nu = 34$  eV (Supporting Information).

As a cross-check for our theoretical methods, we have 809 calculated the radiationless decay spectrum of Ar 3s<sup>-1</sup> in ArKr, 810 which may arise *only* from ETMD(3), 12 and arrived at 811 qualitative agreement with the experimental data. 812

The size-dependent trend within the experimental data is in 813 qualitative agreement to the change in the calculated spectral 814 shape for ICD in systems of increased size (compare Figure 5 815 to Figures 7 and 8).

We are therefore led to the conclusion that the auto- 817 ionization spectra we incur are dominated by ICD to non-818 nearest-neighbor atoms, while the intensity of the ETMD 819 channel is low. We recall the factors that lead to a propensity 820 for decay via one mechanism or the other, which were 821 discussed in the Theoretical Results section. For ETMD to 822 occur with a measurable rate, two Xe atoms have to be present 823 in the vincinity of the decaying Ar atom (Figure 5). Should that 824 be the case for a large part of the Ar atoms in some cluster 825 structure, ETMD can even dominate over ICD (Figure 6 and 826 Supporting Information). Model structures shown in Figures 5 827 and 6 have a Xe content similar to the lowest one probed in our 828 experiment (~11%). For a higher Xe content, ETMD will 829 always dominate the spectra if the two species are given the 830 chance to mix, as, e.g., in the calculated minimum energy 831 structures of N = 38 ArXe clusters (Supporting Information). 832 On the other hand, in core-shell systems, decay via ETMD is 833 probable only for Ar atoms in the interface layer, as this 834 mechanism requires some wave function overlap between Ar 835 and one of the Xe atoms involved. For the core-shell systems 836 we have investigated, the likelihood of decay via ETMD stays 837 below 27% even for a Xe content above 50% (Table 2).

In summary, there are two structural motifs that can best be 839 reconciled with our experimental data: 840

- 1. small clusters containing only few Xe atoms that are  $_{841}$  spread out to distant positions on the cluster surface, and  $_{842}$
- 2. systems with a compact Xe core and Ar outer layers.

These structures are compatible with the findings from core- 844 level photoelectron spectroscopy by Lindblad et al., 28 who 845 proposed the former structure for the smallest clusters in their 846 study and the latter for larger ones. As even in these two cases 847 intensity for decay via ETMD would not vanish completely, we 848 have to leave it open, though, whether other mechanisms are 849 present which further suppress ETMD vs ICD.

While in early work on ICD, it has been almost a paradigm 851 that the decay involves two *neighboring* atoms or molecules, 6 852 more recently significant intensity for non-nearest-neighbor 853 ICD has been found also in other systems. One example is 854 Ne—Ar after Ne 2s ionization, where decay to Ar in the second 855 coordination shell is clearly visible, despite the fact that in this 856 system decay involving nearest neighbors is energetically 857

858 allowed.<sup>20</sup> A further theoretical delineation of the respective 859 factors has recently been presented by one of us.<sup>57</sup>

#### 60 SUMMARY

861 We have presented comprehensive experimental and theoretical 862 data for the autoionization of inner valence ionized states in 863 ArXe clusters. Both ICD and ETMD are allowed for most cases 864 we considered. Because of energetical reasons, ICD requires a 865 separation of the final state vacancies of 7.58 Å or more, which 866 is about two times the typical Ar—Xe distance. We found that 867 autoionization has an efficiency of 0.8—1.0, within the 868 experimental accuracy; that is, it dominates over other modes 869 of relaxation. By comparing our measured spectra to 870 calculations, we identified "long-range" ICD as the most 871 important decay mode and were able to determine the most 872 probable structural motifs to be a xenon core surrounded by 873 several argon layers for large clusters and argon clusters with 874 very few xenon atoms spread out over the surface.

## ASSOCIATED CONTENT

# 876 Supporting Information

877 The Supporting Information is available free of charge on the 878 ACS Publications website at DOI: 10.1021/acs.jpcc.6b06665.

Detailed experimental approach, results, and theoretical spectra for structures 21–27 (PDF)

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892 Notes

893 The authors declare no competing financial interest.

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