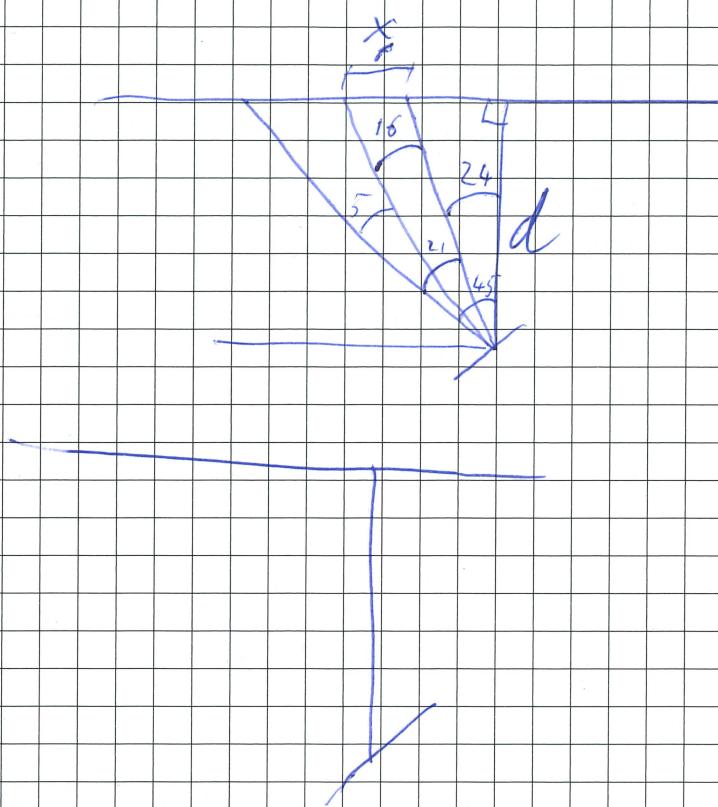


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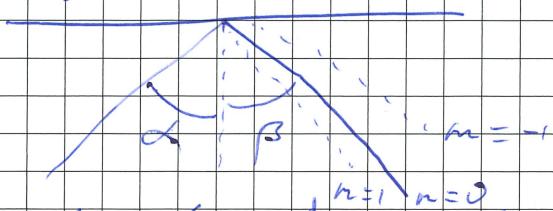


$$x = d \tan(45^\circ + 24^\circ) - d \tan(24^\circ)$$

Equation of the grating is,

$$m\lambda = d_0 (\sin(\alpha) + \sin(\beta))$$

with  $\alpha$  and  $\beta$  defined as



We have measured  $\alpha$  and  $\beta$  with  $\alpha = 45^\circ \pm 2^\circ$  and  $\beta = 10^\circ \pm 5^\circ$ . This gave us a  $d_0 = 7.49 \times 10^{-2} \text{ m}$  for the grating separation.

We have then found a formula for  $\beta$  as function of  $\alpha$  and  $\lambda$  as

$$\beta = \arcsin\left(\frac{m\lambda}{d_0} - d_0 \sin\left(\frac{\alpha}{2}\right)\right)$$

Up for the calculated the length ~~of~~ between the zeroth order spot and a spot of order  $m$  and wavelength  $\lambda$  on the screen is,

$$l = \text{distance screen} \cdot \tan\left(\frac{\beta_1}{2} + \frac{60^\circ}{45^\circ - \beta}\right)$$

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This gave a distance of 6 cm of the wavelength wavelengths 400-380 nm - 600 nm of order 1, which matches our findings.

We then made a Littrow configuration which sends the first order light emitted by the laser back into the laser itself via reflection by the grating. We then measured the power of the laser with and without the feedback and without the grating and found:

- Without the grating the power reached was the highest at around 30 mA with a setting of  $\approx 20$  mW
- With the grating the power was reduced due to the separation of the laser beam also ~~reducing~~ reducing the power off the 2nd order beam. The threshold of lasing was around 20 mA
- With feedback the lasing threshold got reduced to 25 mA and the power is higher than without feedback but with the grating.

We ran into issues with our power measurements due to the power meter fluctuating a lot at 20 mA. The fluctuations were between 10 - 3250 µW. We measured again with another power meter confirming the above results which ~~it~~ also includes graphics on github.

### Notes of programming the CCD:

We found that the microcontroller that we were using was not setting its own clock correctly (it was set to 8 MHz but we found it to be more around 16 MHz). We adjusted the values of the timers so the oscilloscope shows the correct values. We found for the Master-clock (f<sub>osc</sub>) the values:

$$ARR = 13 - 1$$

$$CCA = 6 - 1$$

$$ICG: ARR = 201600 - 1$$

$$CCR = 2448 - 1$$

0

$$SH: ARR = \frac{85}{8} - 1$$

$$CCR = 1078 - 1$$

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This gives an integration time of .311 μs

We found that to get a correct delay between the rising edge of ICG and the rising edge of the SH pulse we need to set the timers up in a master-slave condition. This would allow the timers to start with the correct delay between them and keep the delay.

We chose TIM2 for the ICG and TIM4 for the SH both operating in PWM mode channel 1. The master timer (TIM2) is set to enable MSM pin and trigger event is Update event. The slave timer (TIM4) is set to trigger mode with source ITR1 because of the tables

Slave TIM		ITR0/ITR1	ITR2/ITR3	
TIM2	TIM1	Tim0	Tim3	TIM4
TIM3	TIM1	Tim2	11 5	
TIM4	TIM1	11 2	11 3	11 4
TIM5	TIM2	11 3	11 4	11 8

Here for example if TIM3 is set as slave (Row) then setting the trigger source to ITR2 would make TIM5 the master. We set in the slave configuration panel the MSM bit to Enable and Trigger condition to Rising and Polarity to Low.

In the main.c file we wrote

```
HAL_TIM_PPLWM_SetStart(&htim3, TIM_CHANNEL1);
  11
  ( htim4,
    11
  )
  11
  ( htim5,
    11
  )
  11
  ( htim2,
    11
  )
-- HAL_TIM_SET_COUNTER(&htim4, -10);
```

The last line causes the SH timer to start 10 clock cycles later than the ICG and this results in the desired delay between the ICG and SH.

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We noticed that there was streaking of the CCD output and attributed it to a faulty cable we replaced the cables but could not solve the issue we then saw the effect of placing the special board vertically and started to measure more. we measured if the clocks that micro-inverter produced and noticed they were more sinusoidal than stuck. we

12 - 06 - 2023

We are trying to see if the master clock timing and duty cycle have effect on the noise. To research this we are trying to have a master clock setting of  $ARR = 12 - 1$ ,  $CCR = 6 - 1$  and scaling the other clocks accordingly. This unfortunately has not helped in fixing the noise.

13 - 06 - 2023

We are trying to see if some wire is causing the noise. We did not find it Maximillian then suggested to put the outputs impedance of the ODSOilloscope to  $100\text{ M}\Omega$  instead of  $50\text{ }\Omega$  because the voltage is higher at  $100\text{ M}\Omega$  impedance so the noise would be less. This also did not help. We then tried to let the microcontroller and the scope have the same ground by connecting it to the same group of outlets but this also did not help. This has then turned the lights off so the measurement of the laser and suddenly the address light went away so we think the fluorescent lights caused the problem we also tried to measure the output of the CCD directly instead of putting it through a transistor we do not know if this was helping in reducing the noise.

14 - 06 - 2023

See gif hub

15 - 06 - 2023

We are having difficulty calibrating the spectrometer with the moon light due to the peaks were always not having truly peaks but more hills. This makes the peak guessing much more difficult. We now chose four peaks ~~position~~ which we calibrate  $585, 25$  nm at pixel  $\approx 2360, 02$  ~~approx~~  $+ 10$  px  $1736, 72$

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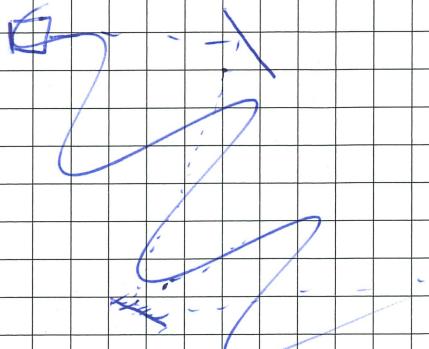
Then a  $640 \text{ nm} \pm 2 \text{ nm}$  peak at pixel  $275,45 \pm 20 \text{ pixels}$ . Then another peak at  $650,65 \text{ nm}$  at  $2960,9 \text{ pixels} \pm 20 \text{ pixels}$ . And then a peak at  $614,31 \text{ nm} \pm 2 \text{ nm}$  at  $2198,79 \text{ pixels} \pm 20 \text{ pixels}$ . This gives  $S = 222.033$ ,  $A_1 = 0.32820$ ,  $A_2 = 9000.35$ ,  $A_3 = 1,26870 \cdot 10^{-3}$  with no errors as I have not calculated them yet. Robin is working on it.

16-06-2023

The setup had fallen off the stand so we are recalibrating the whole spectrometer again. While trying to setup the calibration spectrum after ~~calibration~~ I found that the weird highest peak of the moon Camp set which was not the  $585 \text{ nm}$  peak was also slightly present in the spectrum of the flash light.

Peaks:

Peak wavelength	Position (time oscilloscope)	Pixel
$585,25 \text{ nm}$	$3,00 \text{ ns} \pm 0,1 \text{ ns}$	$1593,52 \pm 50$
$640,00 \text{ nm}^{\text{new}}$	$4,92 \text{ ns} \pm 0,1 \text{ ns}$	$2613 \pm 50$
$650,65 \text{ nm}$	$3,30 \text{ ns} \pm 0,1 \text{ ns}$	$2814 \pm 50$
$614,31 \text{ nm} \pm 3 \text{ nm}$	$3,85 \pm 0,1 \text{ ns}$	$2044,26 \pm 50$



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TITLE Photo's of Oscilloscope

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5	Photo 34 beginning with measurement cases	
44	question knowing	
50	not use not	19-06-2023
57	Calibration 01	19-06-2023
58	Measurement flashlight 1	19-06-2023
59	current 25, no feedback integration time 100 ms	
60	current 28 (basing), no feedback	
61	current 28 (zoom), no feedback	
62	integration time 223 ms	
64	current 28, feedback	first peak. 5.00°
65	current 28, feedback	second peak. 5.03°
66	current 25, feedback	5.07°
67	current 28, feedback	5.14°
68	current 28, feedback	5.23°
69	current 28, feedback	5.35°
70	current 28, feedback	5.42°
71	current 28, feedback	5.45°
72	current 25, feedback	5.23°
67(68)		
45 (new)		

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19-06-2023 We measured the laser

5 Specular with feedback and without and  
laser and with laser and found the results  
as is plotted on figure. What was odd is that  
the laser emitted a peak clearly to the left  
of the  $\approx 6400$  nm peak from the laser lamp and  
but the laser itself was speckled at  $\approx 850$  nm.  
This discrepancy we put down to inaccuracy  
measurements and/or calibration. We subtract sound  
phot the peak of the laser moves to the left  
(smaller wavelength) as we increased the angle of  
the grating. ~~at different times we also tried measuring~~  
~~at a different light and the same but the signal~~  
we got was too weak for the CCD to detect.   
we will try again tomorrow and also make the poster.

20 - 06 - 2023

I calculated the Full width half max from  
of the peaks at  $28\text{ m}\lambda$  and  $20\text{ m}\lambda$  no feedback or  
in case and sound the FWHM of  $25\text{ m}\lambda$  is  
 $\approx 34$  nm and FWHM of  $20\text{ m}\lambda$  is  $\approx 10$  nm.

I also recalibrated the CCD photo G7 in Sylabs.

We found that while adding the error on  
the calibration that it would vary wildly

21 - 06 - 2023

We processed more of the data of the laser  
measurements by calculating the maximum  
peak of the laser at  $25\text{ m}\lambda$  ~~and no~~ feedback and  
then making all the weaker events less relative  
to that peak by setting the peak at 1 and  
then dividing every ~~re~~ datapoint by that peak.

We also tried to add the errors ~~on~~ on  
the ~~new~~ photos but we found them not  
represent the data well enough and also it was  
cluttered with it. Therefore we decided to have  
the errors be present in the x-axis label. The  
y-axis showing the relative intensity ~~plus~~ has the  
errors included.

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For instructions on the CD program a IT  
is on github.

5

22-06-2023

10

Tom and Tristan made the poster last  
we encountered some difficulties but managed to  
make it work in the end. I also needed to adjust  
some graphics for better display on the poster.

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