RADIOACTIVITY

NAME: CLASS:

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1 Hydrogen						The Po	Periodic Tab	Table									+ He lium 2
7 Li Lithium 3	9 Be Beryllium											11 Boron 5	12 Carbon 6	14 Nitrogen 7	16 Oxygen 8	19 Flourine 9	20 Neon 10
Na Sodium 11	24 Mg Magnesium 12											27 A1 Aluminium 13		31 Phosphorus 15	32 S Sulphur 16	35.5 Cl Chlorine 17	40 Argon 18
39 K Potassium 19	40 Ca Calcium 20	45 Sc Scandium 21	48 Ti Titanium 22	51 V Vanadium 23	52 Cr Chromium 24	55 Mn Manganese 25	56 Fe Iron 26	Co Cobalt 27	59 Nickel 28	64 Cu Copper 29	65 Zn Zinc 30	70 Ga Gallium 31	73 Ge Germanium 32	75 As Arsenic 33	79 Se Selenium 34	80 Br Bromine 35	84 Kr Krypton 36
86 Rb Rubidium 37	St Strontium 38	89 Yttrium 39	91 Zr Zirconium 40	93 Niobium 41	96 Mo Molybdenum 42	96 Tc Technetium	Ru Ruthenium 44	Rh Rhodium 45	106 Pd Palladium 46	A B Silver 47	Cd Cadmium 48	114	Sn Tin 50	Sb Antimony 51	Te Tellurium 52	127 	131 Xe Xenon 54
CS Caesium 55	137 Ba Barium 56	La-Lu Lanthanide 57-71	179 Hf Hafnium 72	181	184 W	186 R.e. Rhenium 75	OS Osmium 76	192 Ir Irdium Irdium 77	195 Pt Platinum 78	197 Au Gold 79	Hg Mercury 80	204 T I Thallium 81	207 P b Lead 82	209 Bi Bismuth	209 Po Polonium 84	210 At Astatine 85	Radon 86
223 Fr Francium 87	226 Ra Radium 88	Ac-Lr Actinide 89-103	261 Reff Rutherfordium		266 	264 圆削 Bohrium 107	277 His	268 Mit Meitnerium 109	281 Ds Darmstadtium 110	Roentgenium 111	285 Uulb Ununbium 112	284 : Uudt Ununtrium 113	289 Uudg Ununquadium .	288 Uup Ununpentium	293 புமிர Ununhexium 116	292 Ulus Ununseptium	294 Ulu© Ununoctium 118

139		141		145	150	152	157	159	163	165	167	169	173	175
La		Ą		Pm	Sm	Eu	PS	Tb	Dv	Но	щ	E L	Yb	Lu
Lanthanum 57	Cerium 58	Praseodymium 59	Neodymium 60	Promethium 61	Samarium. 62	Europium 63	Gadolinium 64	Terbium 65	Dysprosium 66	Holmium	Erbium 68	Thulium 69	Ytterbium 70	Lutetium 71
 227	232	231	238	237	244	243	247	247	251	. 252	257	258	259	262
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 Actinium	Thorium	Protactinium	Uranium	Neptunium	Plutonium	Americium	Curium	Berkelium	Californium	Einsteinium	Fermium	Mendelevium	Nobelium	Lawrencium
 68	06	91	92	93	94	92	96	- 26	86	66	100	101	102	103

Symbol Name

Radioactive Decay

Radioactive decays occur when an atom of an isotope emits a radioactive particle. There are many types of radioactive decay, though the three main types of decay you will already be familiar with are: alpha (α) decay, beta (β) decay, and gamma (γ) decay. For A-level, there is also positron decay (also known as **beta+** (β ⁺) **decay**).

Alpha Decay

Alpha decays occur when an atom emits an alpha particle.

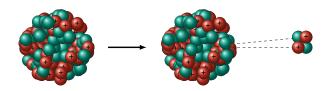


Figure 1: A representation of alpha decay. The atom emits 2 protons and 2 neutrons in the form of an alpha particle. The atom's mass number decreases by 4 and its atomic number decreases

Since an alpha particle contains 2 protons and 2 neutrons, the atom that emits the alpha particle loses 2 protons and 2 neutrons. This emission means that the atom actually changes into a different element! This process of turning one element into another through radioactive decay is called transmutation. Because we know exactly how many protons and neutrons the isotope loses, we can work out what the new element must be.

Let's use the example of Uranium-238. The atomic symbol for Uranium-238 is $^{238}_{92}$ U and we can use this to work out that it has 92 protons and 238-92 = 146 neutrons.

Uranium-238 undergoes alpha decay, emitting 2 protons and 2 neutrons. After the decay, the atom therefore has 90 protons and 150 neutrons, which means its atomic number is 90 and its mass number is 234.

The element in the periodic table with 90 protons is **Thorium**. Our new isotope is therefore $^{234}_{90}$ Th. In other words, Uranium-238 is turning into Thorium-234 by emitting an alpha particle!

We can summarise this radioactive decay in a decay equation:

$$^{238}_{92}U \longrightarrow ^{234}_{90}Th + ^{4}_{2}\alpha$$

Practice Questions 1-1 - Complete the full alpha decay equations.

$$^{238}\text{U} \longrightarrow ^{234}\text{Th} + ^{4}\text{2}\alpha$$

Figure 2: The mass numbers and the atomic numbers on the right hand side must equal the mass number on the left hand side.

$${}_{Z}^{A}X \longrightarrow {}_{Z-2}^{A-4}Y + {}_{2}^{4}\alpha$$

Figure 3: You may come across this more formal, abstract notation for alpha decay.

Practice Questions 1-2 - Which of these are valid alpha decay equations?

1.
$$^{177}_{80}$$
Hg \longrightarrow $^{4}_{2}\alpha$ + $^{171}_{78}$ Pt 4. $^{209}_{87}$ Fr \longrightarrow $^{4}_{2}\alpha$ + $^{203}_{84}$ Po 7. $^{258}_{101}$ Md \longrightarrow $^{4}_{2}\alpha$ + $^{255}_{97}$ Bk

$$2. \quad {}^{191}_{84} Po \quad \longrightarrow \quad {}^{4}_{2}\alpha \quad + \, {}^{187}_{82} Pb \qquad 5. \quad {}^{209}_{85} At \quad \longrightarrow \quad {}^{4}_{2}\alpha \quad + \, {}^{206}_{83} Bi \qquad 8. \quad {}^{238}_{96} Cm \quad \longrightarrow \quad {}^{4}_{2}\alpha \quad + \, {}^{234}_{94} Pu$$

$$3. \ \ ^{211}_{86} Rn \ \longrightarrow \ \ ^{4}_{2}\alpha \ + \ ^{208}_{83} Bi \ \ \ 6. \ \ ^{175}_{79} Au \ \longrightarrow \ \ ^{4}_{2}\alpha \ + \ ^{170}_{78} Pt \ \ \ 9. \ \ ^{202}_{82} Pb \ \longrightarrow \ \ ^{4}_{2}\alpha \ + \ ^{196}_{78} Pt$$

Beta Decay

1.2

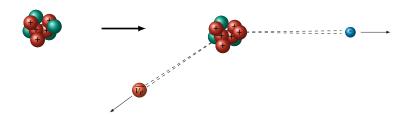


Figure 4: A representation of beta decay. In the nucleus, a neutron (green) turns into a proton (red) and an electron, and the atom emits the electron from the atom. When the neutron turns into a proton, an anti-electron neutrino is created, conserving Lepton number.

Beta particles are electrons. During beta decay, a neutron decays into a proton, an electron, and an anti-electron neutrino. The proton stays in the nucleus but the electron and anti-electron neutrino are ejected ejected from the nucleus.

Let's use the example of Carbon-14 with atomic symbol ${}^{14}_{6}$ C. This isotope of carbon has a mass number of 14 and an atomic number of 6, and therefore has 6 protons and 14-6=8 neutrons.

During beta decay, one of the neutrons turns into a proton. This means that after beta decay the atom has 6+1 = 7 protons and 8-1=7neutrons. Crucially, since both protons and neutrons have a relative atomic mass of 1, the mass number of the new atom stays the same i.e. the mass number is still 14 in this case.

Since the number of protons in the nucleus has increased to 7 the atom is no longer a Carbon atom and has transmuted into a new element. The element in the periodic table with atomic number 7 is Nitrogen. Therefore Carbon-14 turns into Nitrogen-14 when it undergoes beta decay.

$${}^{14}_{6}C \longrightarrow {}^{14}\underset{7}{\overset{1_{4+0}=14}{\longleftarrow}}{}^{0}_{1}\beta + \overline{\nu_{e}}$$

Figure 5: The mass numbers and the atomic numbers on the right hand side must equal the mass number on the left hand side.

An anti-electron neutrino $(\overline{\nu_e})$ is produced in the decay in order to conserve lepton number, and the anti-neutrino must be of the electron flavour because the flavour of the lepton number must be conserved. Anti-electron neutrinos have tiny mass and no charge so they do not change any of the mass or atomic numbers in the decay equation.

$$_{Z}^{A}X \longrightarrow _{Z+1}^{A}Y + _{-1}^{o}\beta + \overline{\nu_{e}}$$

We can summarise this with the following decay equation:

$$^{14}_{6}C \longrightarrow ^{14}_{7}N + {^{0}_{-1}}\beta + \overline{\nu_{e}}$$

Figure 6: You may come across this more formal, abstract notation for beta decay.

Practice Questions 1-3 - Complete the full beta decay equations

Practice Questions 1-4 - Are these valid beta decay equations?

1.
$$^{115}_{49}\text{In} \longrightarrow ^{\circ}_{-1}\beta + ^{113}_{48}\text{Cd} + \nu_{\mu}$$

5.
$$^{^{1}54}_{63}Eu \longrightarrow ^{^{0}}_{-1}\beta + ^{^{1}54}_{65}Tb + \nu_e$$

$$2. \quad ^{214}_{82} Pb \quad \longrightarrow \quad \ ^{0}_{-1} \beta \quad + \quad ^{215}_{81} Tl \quad + \quad \overline{\nu_{\ell}}$$

6.
$$^{116}_{46}Pd \longrightarrow ^{0}_{-1}\beta + ^{117}_{46}Pd + \overline{\nu_e}$$

3.
$$^{91}_{36}\text{Kr} \longrightarrow ^{0}_{-1}\beta + ^{91}_{35}\text{Br} + \overline{\nu_e}$$

$$\stackrel{91}{_{36}}\mathrm{Kr} \longrightarrow \stackrel{^{0}}{_{-1}}\!\beta + \stackrel{91}{_{35}}\mathrm{Br} + \overline{\nu_{e}} \qquad \qquad 7. \quad \stackrel{63}{_{28}}\mathrm{Ni} \longrightarrow \stackrel{^{0}}{_{-1}}\!\beta + \stackrel{63}{_{27}}\mathrm{Co} + \overline{\nu_{e}}$$

4.
$$^{33}_{15}P$$
 \longrightarrow $^{0}_{-1}\beta$ + $^{31}_{15}P$ + $\overline{\nu_e}$

8.
$$^{98}_{37}\text{Rb}$$
 \longrightarrow $^{\circ}_{-1}\beta$ + $^{98}_{37}\text{Rb}$ + $\overline{\nu_e}$

Positron Decay

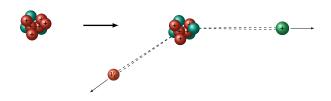


Figure 7: A representation of positron In the nucleus, a proton (red) turns into a neutron (green) and a positron, and the atom emits the positron.

Beta+ particles are positrons (anti-electrons). During beta+ decay, a proton decays into a neutron, a positron, and an electron neutrino. The neutron stays in the nucleus but the positron and electron neutrino are ejected from the nucleus.

Let's use the example of Fluorine-18 with atomic symbol ¹⁸₉F. This isotope of Fluorine has a mass number of 18 and an atomic number of 9, and therefore has 9 protons and 18-9 = 9 neutrons.

During beta+ decay, one of the protons turns into a neutron. This means that after beta+ decay the atom has 9-1 = 8 protons and 9+1= 10 neutrons. Crucially, since both protons and neutrons have a relative atomic mass of 1, the mass number of the new atom stays the same i.e. the mass number is still 18 in this case.

Since the number of protons in the nucleus has decreased to 8, the atom is no longer a Fluorine atom and has transmuted into a new element. The element in the periodic table with atomic number 8 is Oxygen. Therefore Carbon-18 turns into Oxygen-18 when it undergoes beta+ decay.

An electron neutrino (ν_e) is produced in the decay in order to conserve lepton number, and the neutrino must be of the electron flavour because the flavour of the lepton number must be conserved.

Like electron anti-neutrinos, electron neutrinos have tiny mass and no charge so they do not change any of the mass or atomic numbers in the decay equation.

We can summarise this with the following decay equation:

$$^{18}_{9}F \longrightarrow ^{18}_{8}O + ^{0}_{+1}\beta + \nu_{e}$$

$$^{18}_{9}F \longrightarrow ^{18}\underbrace{0 + 0 + 18}_{8+1=9} \beta + \nu_{\epsilon}$$

Figure 8: The mass numbers and the atomic numbers on the right hand side must equal the mass number on the left hand side.

$${}_{Z}^{A}X \longrightarrow {}_{Z-1}^{A}Y + {}_{+1}^{O}\beta + \nu_{e}$$

Figure 9: You may come across this more formal, abstract notation for beta+ decay.

Practice Questions 1-5 - Complete the full positron decay equations

Practice Questions 1-6 - Are these valid positron decay equations?

$$2. \quad {}^{240}_{92}\mathrm{U} \quad \longrightarrow \quad {}^{0}_{+1}\beta \quad + \quad {}^{238}_{91}\mathrm{Pa} \quad + \quad \overline{\nu_{e}} \qquad 6. \quad {}^{126}_{53}\mathrm{I} \qquad \longrightarrow \quad {}^{0}_{+1}\beta \quad + \quad {}^{124}_{54}\mathrm{Xe} \quad + \quad \nu_{e}$$

3.
$$^{81}_{32} \text{Ge} \longrightarrow {}^{\circ}_{+1} \beta + {}^{79}_{33} \text{As} + \nu_e$$
 7. $^{80}_{32} \text{Ge} \longrightarrow {}^{\circ}_{+1} \beta + {}^{81}_{32} \text{Ge} + \nu_{\tau}$

$$4. \quad {}^{136}_{52} \text{Te} \quad \longrightarrow \quad {}^{\circ}_{+1} \beta \quad + \quad {}^{134}_{52} \text{Te} \quad + \quad \nu_e \qquad 8. \quad {}^{106}_{45} \text{Rh} \quad \longrightarrow \quad {}^{\circ}_{+1} \beta \quad + \quad {}^{106}_{45} \text{Rh} \quad + \quad \nu_e$$

Gamma Decay

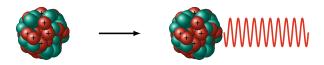


Figure 10: A representation of gamma decay. An atom emits an electromagnetic wave, lowering the overall energy of the atom.

Gamma decay is the simplest of the three decays. During gamma decay, an atom emits an electromagnetic wave in the gamma frequency part of the electromagnetic spectrum. Since gamma waves have no charge and no mass, the proton number and atomic number of the atom stays the same; the atom simply has less energy overall than it did before it emitted the gamma wave.

Let's look at Dysprosium-152 with atomic symbol $^{152}_{66}$ Dy. Dysprosium has **66** protons and **152-66 = 86** neutrons. After emitting a gamma wave, it still has **66** protons and **86** neutrons. Therefore, its decay equation can be written:

$$^{152}_{66}$$
Dy $\longrightarrow ^{152}_{66}$ Dy $+ ^{0}_{0}\gamma$

 $^{152}_{66}$ Dy $\longrightarrow ^{52}_{66}$ Dy $+ ^{0}_{0}$

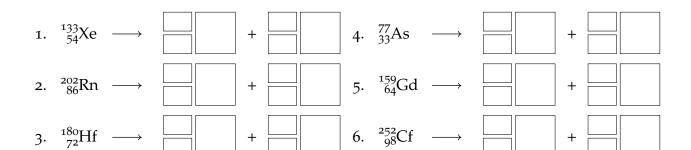
Figure 11: The mass numbers and the atomic numbers on the right hand side must equal the mass number on the left hand side.

$${}_{Z}^{A}X \longrightarrow {}_{Z}^{A}Y + {}_{o}^{o}\gamma$$

Figure 12: You may come across this more formal, abstract notation for gamma decay.

Only α and β decays result in **transmutation** i.e. the turning of one element into another.

Practice Questions 1-7 - Complete the full gamma decay equations



Practice Questions 1-8 - Are these valid gamma decay equations?

 $1. \quad ^{151}_{62} Sm \quad \longrightarrow \quad ^{0}_{0} \gamma \ \, + \quad ^{151}_{62} Sm \quad 4. \quad ^{102}_{43} Tc \quad \longrightarrow \quad ^{0}_{0} \gamma \ \, + \quad ^{102}_{42} Mo \quad 7. \quad ^{39}_{17} Cl \quad \longrightarrow \quad ^{0}_{0} \gamma \ \, + \quad ^{38}_{17} Cl$

 $\longrightarrow \quad {}^0_0\gamma \ + \quad {}^{132}_{53}I \qquad 5. \quad {}^{135}_{52}Te \quad \longrightarrow \quad {}^0_0\gamma \ + \quad {}^{134}_{51}Sb \qquad 8. \quad {}^{208}_{81}Tl \quad \longrightarrow \quad {}^0_0\gamma \ + \quad {}^{207}_{80}Hg$

 $3. \quad \overset{84}{34} Se \quad \longrightarrow \quad {}^{\scriptscriptstyle 0}{}_{\scriptscriptstyle 0}\gamma \ + \quad \overset{84}{33} As \quad \ 6. \quad \overset{^{\scriptscriptstyle 231}}{89} Ac \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 230}}{89} Ac \quad \ 9. \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle 103}}{43} Tc \quad \longrightarrow \quad \ \, \overset{^{\scriptscriptstyle 0}}{\scriptscriptstyle 0}\gamma \ + \quad \overset{^{\scriptscriptstyle$

Practice Questions 1-9 - Review: Mixed Decay Equations

 $^{172}_{69}\text{Tm} \longrightarrow ^{0}_{-1}\beta +$ 1. +

 $^{204}_{85}$ At \longrightarrow $^{4}_{2}\alpha$ + 2. +

 $^{246}_{98}Cf \longrightarrow ^{4}_{2}\alpha +$ 3. +

 $_{30}^{71}$ Zn \longrightarrow $_{+1}^{0}\beta$ + 4. +

 $^{160}_{74}W \longrightarrow ^{0}_{0}\gamma +$ 5. +

 $^{28}_{12}\text{Mg} \longrightarrow ^{0}_{-1}\beta +$ 6. +

 $^{175}_{80}\text{Hg}$ \longrightarrow $^{0}_{+1}\beta$ + 7. +

 $^{36}_{17}\text{Cl} \longrightarrow ^{0}_{0}\gamma +$ 8.

 $^{213}_{87} Fr \longrightarrow ^{4}_{2} \alpha +$ 9.

10. $^{40}_{17}\text{Cl} \longrightarrow ^{0}_{-1}\beta +$

1.
$$^{127}_{50}\text{Sn} \longrightarrow ^{127}_{49}\text{In} + \boxed{ }$$

$$2. \quad {}^{228}_{89}\text{Ac} \quad \longrightarrow \quad {}^{228}_{90}\text{Th} \quad + \qquad \qquad + \qquad \qquad + \qquad \qquad \qquad$$

$$4. \quad {}^{211}_{83}\text{Bi} \quad \longrightarrow \quad {}^{207}_{81}\text{Tl} \; + \boxed{ } \boxed{ }$$

$$5. \quad {}^{101}_{40}\text{Zr} \quad \longrightarrow \quad {}^{101}_{41}\text{Nb} \; + \boxed{ } \boxed{ } \boxed{ } + \boxed{ }$$

$$6. \quad {}^{170}_{69}\text{Tm} \longrightarrow \quad {}^{170}_{68}\text{Er} + \boxed{ } \boxed{ } \boxed{ } + \boxed{ }$$

$$7. \quad {}^{244}_{94}Pu \quad \longrightarrow \quad {}^{240}_{92}U \ + \boxed{ } \boxed{ } \boxed{ } + \boxed{ }$$

$$8. \quad {}^{219}_{85}\text{At} \quad \longrightarrow \quad {}^{215}_{83}\text{Bi} \; + \boxed{ } \boxed{ } \boxed{ } + \boxed{ }$$

$$9. \quad {}^{187}_{81}\text{Tl} \quad \longrightarrow \quad {}^{183}_{79}\text{Au} \; + \boxed{ } \boxed{ } \boxed{ } \boxed{ } + \boxed{ } \boxed{ }$$

2 Ionisation

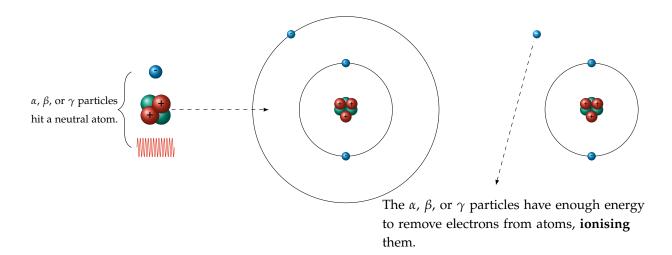
When a radioactive particle hits a neutral atom, it can provide enough energy to **remove an electron** from the atom. When the atom loses an electron, it has a net positive charge, and is therefore called an **ion**. **Ionisation** is the name given to this process.

The danger of radioactive particles is closely related to their ionising ability (i.e. their ability to knock electrons off of neutral atoms). The order of ionising ability from most to least ionising amongst the radioactive particles is as follows:

- 1. Alpha (α) Particles.
- 2. Beta (β) Particles.
- 3. Gamma (γ) Waves.

 α **particles** are most ionising because they have a +2 overall charge: they **pull** electrons off other atoms very easily, ionising them.

 $\beta+$ particles (positrons) are a more nuanced case that we will discuss in a moment



 β - particles are the next most ionising because they have a high velocity: they carry large amounts of kinetic energy that can be transferred to other electrons, giving these electrons enough energy to escape from their host atoms

 γ rays are weakly ionising and will pass through most materials without ionising them. If conditions are just right, their energy can be transferred to electrons in the neutral atom causing the electrons to escape from their host atom.

 β + particles are more complicated. Since β + particles are identical to electrons except they have opposite charge, in theory they should have the exact same ionising ability. However, as soon as a β + particle meets an electron it annihilates to form 2 γ rays. There is a huge asymmetry of matter to anti-matter in the universe - i.e. there are far more electrons than positrons - and so positrons usually do not manage to travel very far before meeting an electron and annihilating.

Penetration

The penetrative ability of radioactive particles is strongly related to their ionisation ability: the most highly ionising particles do not penetrate materials very far because they ionise atoms very easily! The order of penetrative ability from most to least penetrating amongst the radioactive particles is as follows:

- 1. Gamma (γ) waves (photons).
- 2. Beta- (β^-) particles (electrons).
- 3. Alpha (α) particles (Helium nuclei).
- 4. Beta+ (β^+) particles (positrons).

Notice that this is the reverse of the ionisation ability list, as expected.

Alpha particles are stopped very easily by almost any material; remember, they only travel a few centimetres in air before ionising and so their penetrative ability in anything more dense than air will be even less. Outside of the body, alpha particles pose little risk to humans: the layer of dead skin on the surface of our bodies is thick

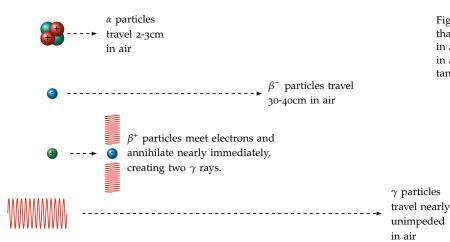
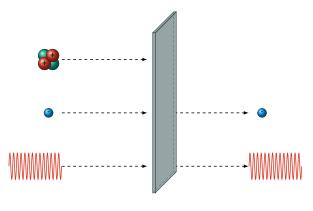


Figure 13: α particles are so ionising that they only travel a few centimetres in air. β particles can travel up to 40cm in air, and γ waves can travel large distances in air before ionising an atom.

enough to ionising the alpha particles and stop them travelling further.

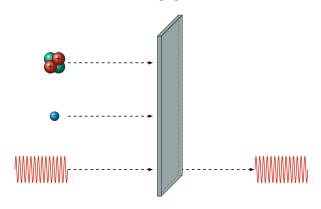
Beta particles penetrate 10x further than alpha particles in air, and therefore can penetrate small thicknesses of some materials e.g. a few mm of aluminium foil. Beta particles can therefore be harmful outside of the body provided you are close enough to the source.

Only thick lead (2 inches or more) or thick concrete (several metres) stops gamma waves by any significant amount. Because they're electromagnetic waves with no mass, no charge, and a tiny wavelength, they pass through most things without ionising them. Strictly speaking it's impossible to fully stop gamma waves. Thick lead and concrete just reduce the chances of a gamma wave getting through, but some gamma waves will still get through. We can only reduce the intensity of the gamma rays i.e. the number arriving per second in a given area. Gamma rays can therefore be harmful if the source is outside the body, but their probability of hitting an atom and ionising it is quite low.



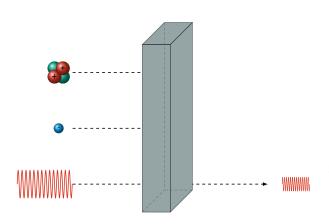
 α particles are so ionising that even a thin sheet of paper is enough to stop them.

Thin paper



Aluminium stops α particles easily and can also stop $\beta^$ particles if the thickness is more than 5mm.

Thin Aluminium



Thick lead or concrete stops α and β^- particles easily. It can also stop gamma rays and reduces their intensity significantly.

Thick Concrete/Lead

Half-Life

A radioactive isotope is an isotope that is **unstable** and one that will emit a radioactive particle. Unfortunately, radioactive particles decay randomly and we have no way of knowing when an individual nucleus of a radioactive isotope will decay and emit a radioactive particle. Radioactive decays are also spontaneous, meaning they cannot be triggered by e.g. temperature, pressure, or other physical conditions.

• Unstable: unstable nuclei are likely to randomly, and spontaneously decay because they have too much energy and/or too many nucleons.

- Random: it is not possible to predict when an individual nucleus will decay, we can only predict the probability or fraction of nuclei that will decay in a given time.
- Spontaneous: radioactive decay cannot be controlled by external factors such as temperature or pressure.

However, physicists discovered that if they had enough nuclei together in one place they could predict how long it would take a certain proportion of them to decay. In particular, they discovered that the time it takes for 50% of the radioactive nuclei to decay was always the same, and they called this the half-life.

The Half-Life of an isotope is the average time taken for 50% of the radioactive atoms in a sample to decay.

Let's say we have 100 atoms of the isotope Rubidium-84. After 1 half life (HL), half of the atoms will decay, leaving 50 atoms. After another half-life, half of those 50 atoms will have decayed, leaving 25 atoms. After another half-life, half of those 25 atoms will have decayed, leaving 12.5 atoms. Since we can't have half an atom, in reality there will be either 12 or 13 atoms remaining, on average. After another half-life, we expect 50% of the 12 or 13 atoms to have decayed, leaving 6 or 7 atoms. We can summarise this using the following decay chain:

$$100 \text{ Atoms} \xrightarrow{1 \text{ HL}} 50 \text{ Atoms} \xrightarrow{2 \text{ HLs}} 25 \text{ Atoms} \xrightarrow{3 \text{ HLs}} 13 \text{ Atoms} \xrightarrow{4 \text{ HLs}} 7 \text{ Atoms}$$

It is very important to understand that since radioactive decay is completely random, the actual number of atoms remaining fluctuates a bit around the expected values; if we flip 50 coins, we do not expect exactly 25 to turn up heads, but rather 25 on average. The same is true of radioactive decay. Figure 14 below shows a sample of atoms from a radioactive isotope decaying at random.

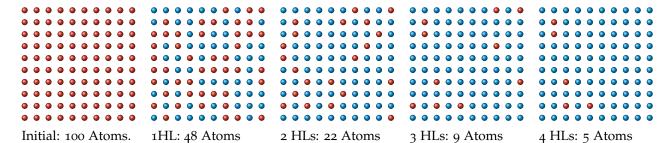


Figure 14: A visual representation of Half-Life. After each half-life, 50% of the remaining atoms decay.

Practice Questions 4-1 - How many atoms remain?

All half-lives are with respect to the initial number of atoms, not the previous box.

1.
$$462 \text{ Atoms} \xrightarrow{1 \text{ HL}} \boxed{\qquad} \xrightarrow{2 \text{ HL}} \boxed{\qquad} \xrightarrow{3 \text{ HL}} \boxed{\qquad} \xrightarrow{4 \text{ HL}} \boxed{\qquad}$$

2.
$$812 \text{ Atoms} \xrightarrow{1 \text{ HL}} \boxed{\qquad} \xrightarrow{2 \text{ HL}} \boxed{\qquad} \xrightarrow{3 \text{ HL}} \boxed{\qquad} \xrightarrow{4 \text{ HL}}$$

3. 198 Atoms
$$\xrightarrow{1 \text{HL}}$$
 $\xrightarrow{2 \text{HL}}$ $\xrightarrow{3 \text{HL}}$ $\xrightarrow{4 \text{HL}}$

Practice Questions 4-2 - How many atoms remain after the given half-lives?

Half-Lives No. of Atoms	Half-Li	ves No. of Atoms	Half-Lives	No. of Atoms
0 26	О	429	О	258
1	2		5	
2	4		7	
5	7		8	
8	9		9	
10	10		10	

So far we have talked about how atoms decay after a given number of half lives. But we haven't discussed the time element much. Remember our definition of half-life: the time-taken for 50% of the atoms in a sample to decay.

Each unstable isotope has its own characteristic half-life that tells us how long it takes for 50% of its atoms to decay. Let's use the example of Rubidium-84, which has a half-life of 33 days.

This means it takes 33 days for half of the atoms in a sample of Rubidium-84 to decay. If we start with a sample of 80 Rubidium-84 atoms, there are 40 atoms of Rubidium-84 remaining after 33 days (1 Half-Life). After a further 33 days (66 days in total, 66/33 = 2Half-Lives), there are 20 atoms of Rubidium-84 left. After another 33 days (99 days in total, 99/33 = 3 Half-Lives), there are 10 atoms of Rubidium-84 remaining, and so on.

In order to understand how much of an isotope remains after a given time, we need to calculate many half-lives have occurred in that time. For example, if we wanted to know how much Rubidium-84 remained after 90 days, we would need to first realise that 3 half-lives (90/30) had occurred in that time.

You can think of a radioactive element's half-life as its own fingerprint: no two radioactive isotopes share the same half-life.

Isotope	Half-Life
Na-22	2.6 years
H-3	12 years
Se-75	120 days
Ru-105	4.44 hours
Ag-110	25 S
U-238	4.5 billion years

Figure 15: Some sample half-lives of radioactive isotopes

Practice Questions 4-3 - Calculate how many half-lives have occurred

Isotope Half-	Life = 147 years	Isotope Half-Life = 644 seconds	Isotope Half-Life = 731 hours	
Time (years)	Half-Lives	Time (seconds) Half-Lives	Time (hours) Half-Lives	
0	0	0	0	
147	1	644	1462	
441	3	1288	2193	
735		3220	5848	
1029		5152	6579	
1470		5796	7310	

Practice Questions 4-4 - Calculate the number of half-lives occurred to find out how many atoms remain

Isotop	e Half-Life =	952 days	Isotope H	lalf-Life = 271	seconds
Time (days)	Half-Lives	No. of Atoms	Time (seconds)	Half-Lives	No. of Atoms
0	0	7866	0	0	2314
1904			542		
2856			1084		
3808			1355		
4760			1626		
8568			2168		

4.1 Activity and Becquerels

So far we have talked about how many atoms decay and how many remain after a given number of half-lives. But there are other units that we can use to measure radioactive decay.

We can measure the Activity of an isotope, which gives us a measure of how many radioactive decays per second occur: the higher the activity, the more radioactive decays per second.

The unit used to measure activity is the Becquerel (Bq).

The Becquerel (Bq) measures how many disintegrations/decays per second occur.

1 Bq = 1 radioactive disintegration/decay per second.

The activity is quite closely related to the number of atoms: if we have more atoms decaying, there will be more radioactive decays per second, and therefore a higher activity.

Similarly, if we have a higher mass of radioactive isotope, this will contain more atoms, and therefore the sample will have a higher activity.

Number of Atoms, Activity, and Mass all follow the same behaviour and rules when it comes to radioactive decay.

Background Radiation and Corrected Count Rate

If you turn on a Geiger counter even far away from any radioactive sources it will still click and measure an activity. This is because of background radiation. There are many sources of background radiation:

- Air: there is a small amount of radioactive radon gas in the air.
- Buildings and materials: the materials that make up the buildings and some of the materials around us have small amounts of radioisotopes in them.
- Food and other organic matter: Living things that contain Carbon contain a small amount of radioactive Carbon-14. Some of the foods we eat also contain radioisotopes e.g. Potassium in bananas.
- Cosmic rays: cosmic rays produced in outer space hurtle down to Earth and react with atoms in the upper atmosphere and produce radioisotopes in the process.
- Industrial and medical: hospitals and some industry make use of radioisotopes.

When measuring the activity of a source, we must subtract the background radiation count rate from the measured activity. This calculated activity is called the **corrected count rate**.

Cosmic rays are produced in highly energetic astrophysical phenomena such as when two black holes collide.

Worked Example 4-4 - Corrected count rate

Q: The activity of a sample of radioisotope is 460 Bq. 30 minutes later, the activity of the sample has fallen to 240 Bq. The background radiation is measured to have an activity of 20 Bq. Calculate the half-life of the radioisotope.

A: Before doing anything we must correct the activity for the background count rate. This gives a corrected initial activity of 460 - 20 = 440 Bq and a corrected final activity of 240 - 20 = 220 Bq.

We can see that the corrected activity has halved in 30 minutes, and therefore the half-life is 30 minutes.

Notice that if we did not correct for the count rate, we would have overestimated the half life. This problem gets worse the lower the count rate measured.

Parents and Daughters - What are the radioactive atoms decaying into?

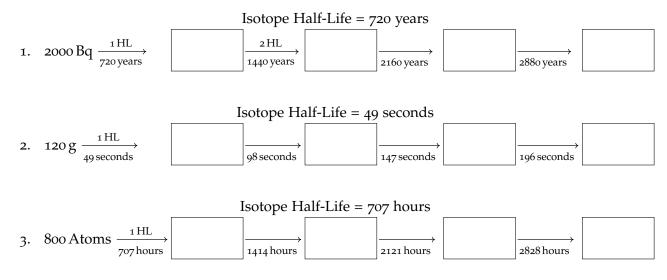
In Figure 14 we showed the red, undecayed particles turning into the blue, decayed particles. What do the red and blue particles represent? It depends on the kind of decay the atoms are undergoing. Let's use the example of Rubidium-84, which undergoes beta- decay.

$$^{84}_{37}\text{Rb} \longrightarrow ^{84}_{38}\text{Sr} + ^{0}_{-1}\beta + \overline{\nu_e}$$

This means that the new, blue atoms are actually Strontium atoms. So while the number of Rubidium atoms is decreasing with each decay, the number of Strontium atoms is increasing. We call the initial isotope the parent nucleus, and the new isotope that is formed the daughter nucleus. In this case the Rubidium is the parent isotope and Strontium the daughter isotope.

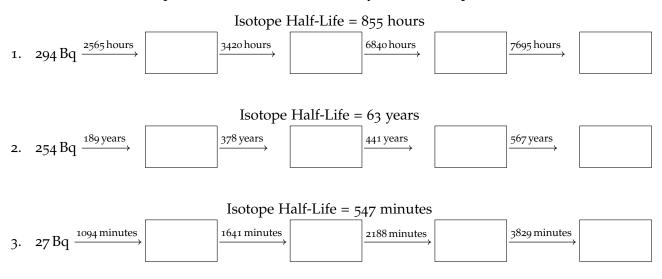
Practice Questions 4-5 - Calculate the number of atoms remaining

All arrows are with respect to the initial activity. Write the number of half-lives that have occurred since the initial value above the arrows.



Practice Questions 4-6 - How many atoms remain after the given time?

Write the number of half-lives that have occurred and use this to calculate the activity in each of the boxes. The half-lives are with respect to the initial time and activity, not from the previous box.



Practice Questions 4-7 - What's the half-life of the isotope?

Write the number of half-lives that have occurred and use this to calculate the activity in each of the boxes. The half-lives are with respect to the initial time and activity, not from the previous box.

Half Life Graphs

Let's conduct our radioactive decay experiment again, this time starting with **80** atoms of an isotope with a half life of **50** years. We know that after 50 years (1 half-life) the number of atoms remaining of the original isotope should be roughly **40**. After 100 years (2 half-lives), it should be roughly **20**, after 150 years (3 half-lives) there should be around **10** atoms remaining, and after 200 years (4 half-lives), roughly **5** atoms should remain.

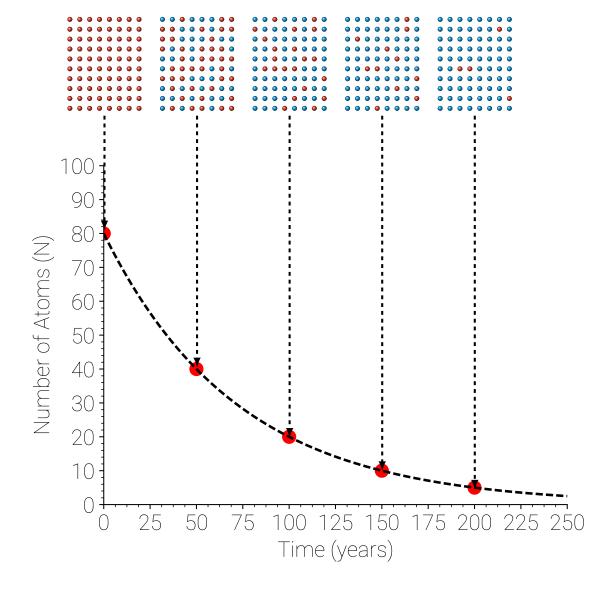
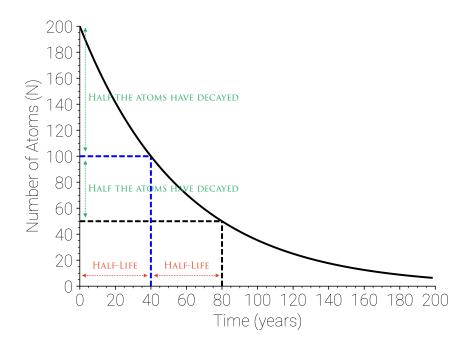


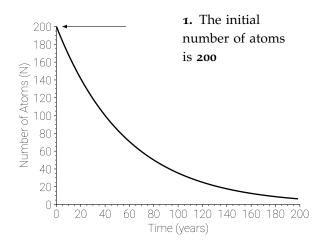
Figure 16: The classic **exponential decay** curves of radioactive decay. After each half-life, 50% of the remaining atoms decay.

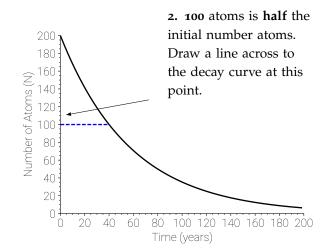


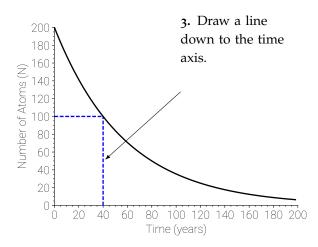
If we plot the number of atoms vs time, we get a graph of that looks like the one in Figure 16. This type of curve is called an exponential decay curve. Notice that the curve gets less steep with time: this is because there are fewer radioactive particles left that are able to decay, and so the number decaying in each period of time decreases. However, the proportion/percentage of particles that decay in a given time period is always the same.

The amazing thing about these kinds of exponential decay curves is that we can find out the half-life of an isotope using them. Figure 18 shows the basic procedure for finding the half-life of a radioisotope using its decay graph. Remember that our y-axis units could equally well be activity/mass.

Figure 17: We find the half-live by picking an initial number of atoms/activity/mass, and finding how long it takes to get to half of that number of atoms/activity/mass.







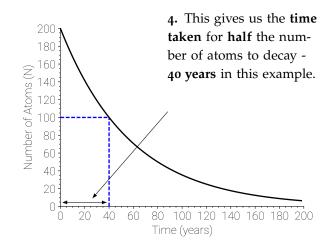


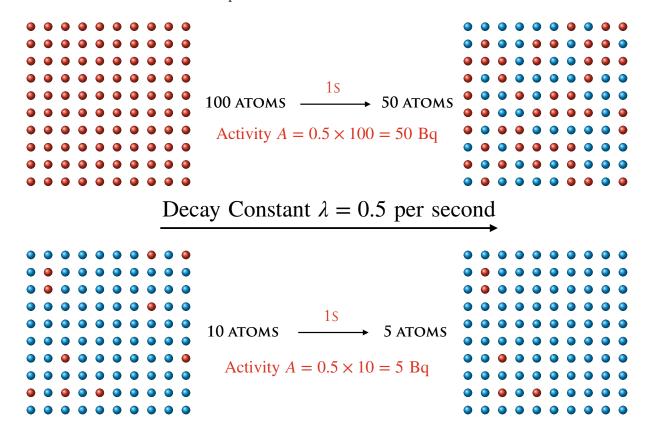
Figure 18: How to find the half life from a radioactive decay curve. The procedure for finding the half life from a radioactive decay curve. This also works if "Activity" or "Mass" is on the y-axis.

Half-Life II

So far we have focussed on situations where we calculate the number of half-lives remaining after a whole number of half-lives have occurred. For example, if an isotope has 100 atoms and a half-life of 25 seconds, the number of atoms remaining after 50 seconds would be 25 atoms. But what if we wanted to know the number atoms remaining after 40 seconds? Similarly, if we started with 100 atoms and find after 60 seconds that we have 20 atoms remaining, can we calculate the half-life of the isotope from that knowledge?

The answer is yes, but we need to introduce some new mathematical machinery. This mathematical machinery is the main difference between GCSE radioactivity and A-level radioactivity, and it will be the focus of this section.

The first observation we need to make is that the number of radioactive particles emitted per second by an isotope (the activity) depends on the number of atoms of the isotope.



Therefore, we can say:

$$A \propto N$$

where A is the activity and N is the number of particles of the isotope.

Using our usual rules of proportionality:

$$A \propto N$$

$$\implies A = \lambda N$$

where λ is a constant known as the **decay constant**.

We can write $A = -\frac{dN}{dt}$ since this simply says that the activity is the rate of change of the number of particles of the parent isotope. The negative sign is because the number of parent isotope particles decreases during the decay, while the number of daughter particles increases. Since a change from parent to daughter only occurs when a radioactive particle is emitted, it makes sense that the activity (which measures the number of radioactive decays per second) is related to the rate of change of the number of particles.

Combining these equations leads to:

$$A = -\frac{dN}{dt} \propto N$$

$$\implies A = -\frac{dN}{dt} = \lambda N$$

This kind of equation is called a first order separable differential equation, which you learn to solve in further maths or your degree. The basic idea is that we move all N terms to one side and all t terms to one side and integrate each side of the equation with respect to those variables:

$$-\frac{dN}{dt} = \lambda N$$

$$\Rightarrow -\frac{1}{N} dN = \lambda dt$$

$$\Rightarrow \int \frac{1}{N} dN = -\int \lambda dt$$

$$\Rightarrow \ln N = -\lambda t + C$$

$$\Rightarrow N = e^{-\lambda t + C}$$

$$\Rightarrow N = N_0 e^{-\lambda t}$$

It is not necessary to understand or know the derivation of this equation, and so if you can't understand the derivation simply sear the following equation into your brain:

$$N = N_0 e^{-\lambda t}$$

and since we know that $A = \lambda N \Longrightarrow N = \frac{A}{\lambda}$ we can substitute that into the above equation and obtain:

$$A = A_0 e^{-\lambda t}$$

in terms of activity.

In the case of γ decay, the "daughter" nuclei is the same as the "parent" nuclei except in a lower-energy state.

If we set t = 0 in $N(t) = e^{C}e^{-\lambda t}$ we find $N(0) = N_0 = e^c$. In other words, the constant term is the initial number of particles in the sample.

The **decay constant** λ tells us the number of radioactive decays per second, and the **half-life** $t_{1/2}$ tells us how long it takes for 50% of the atoms to decay. Intuitively, it feels as if there could be a link between the two quantities, and there is, which we can find by solving the above equation for the time it takes for the activity or number of particles to halve:

$$N = N_0 e^{-\lambda t}$$

$$\Rightarrow \frac{N_0}{2} = N_0 e^{-\lambda t_{1/2}}$$

$$\Rightarrow \frac{1}{2} = e^{-\lambda t_{1/2}}$$

$$\Rightarrow \ln(\frac{1}{2}) = -\lambda t_{1/2}$$

$$\Rightarrow \ln 1 - \ln 2 = -\lambda t_{1/2}$$

$$\Rightarrow -\ln 2 = -\lambda t_{1/2}$$

$$\Rightarrow \frac{\ln 2}{\lambda} = t_{1/2}$$

$$t_{1/2} = \frac{\ln 2}{\lambda} = \frac{0.69}{\lambda}$$

6.1 A more intuitive mathematical understanding of half-life

These equations look intimidating, and they are, and there is in fact a much easier way of working out how many atoms remain after a given time period.

Consider a situation in which we have 120 atoms with a half-life of 30 minutes. Using the notation from the equation above, we would say $N_0=120$ and $t_{1/2}=30$ mins.

Let's calculate how many atoms remain after a given number of halflives:

$$N_0 imes rac{1}{2} = N_0 imes \left(rac{1}{2}
ight)^1 = 60$$
 atoms. After 1 half-life.
$$\left(N_0 imes rac{1}{2}
ight) imes rac{1}{2} = N_0 imes \left(rac{1}{2}
ight)^2 = 30 ext{ atoms. After 2 half-lives.}$$
 $\left(\left(N_0 imes rac{1}{2}
ight) imes rac{1}{2} = N_0 imes \left(rac{1}{2}
ight)^3 = 15 ext{ atoms. After 3 half-lives.}$

Clearly we can see a pattern: to work out how many atoms remain after a certain number of half-lives, you just multiply the initial number of atoms N_0 by $\left(\frac{1}{2}\right)^n$ where n is the number of half-lives:

$$N = N_0 \left(\frac{1}{2}\right)^n$$

To work out how many atoms remain when a non-integer number of half-lives occur, we use the exact same formula. For example, with a half-life of 30 minutes, if we wanted to know the number of atoms remaining after 45 minutes, that corresponds to 45 mins/30 mins = 1.5 half-lives. The number of atoms remaining would be $N=120 imes \left(\frac{1}{2}\right)^{1.5}=42$ atoms. Therefore, we can just use the same techniques as before and note that this equation could be written:

$$N=N_0\left(\frac{1}{2}\right)^{\frac{t}{t_{1/2}}}$$

Practice Questions 6-1 - Calculate the number of half-lives occurred to find out how many atoms remain

Isotope Ha	alf-Life	e = 884 years	Isotope H	lalf-Life	e = 421 days
Time (years)	$\frac{t}{t_{1/2}}$	No. of Atoms	Time (days)	$\frac{t}{t_{1/2}}$	No. of Atoms
0	О	1136	0	О	7877
1594			603		
2182			924		
3666			1332		
4475			3370		
6658			4365		

Practice Questions 6-2 - Calculate the number of half-lives occurred to find out how many atoms remain

- 1. A sample of niobium-92 has a half-life of 34.7 million years. Its initial activity is 614 Bq. What is the activity of the element after 115 million years?
- 2. A sample of bismuth-208 has a half-life of 368 thousand years. Its initial activity is 789 Bq. What is the activity of the element after 700 thousand years?
- 3. A sample of caesium-137 has a half-life of 30.17 years. Its initial activity is 896 Bq. What is the activity of the element after 80 years?

Practice Questions 6-3 - Calculate the number of half-lives occurred to find out how many atoms remain

- 1. The isotope iodine-129 has a half-life of 15.7 million years. Its current activity is 809 Bq. What was its activity 87.0 million years ago?
- 2. The isotope niobium-94 has a half-life of 20300 years. Its current activity is 665 Bq. What was its activity 43917.0 years ago?
- 3. The isotope cadmium-109 has a half-life of 1.267 years. Its current activity is 273 Bq. What was its activity 4.0 years ago?

Worked Example 6-3 - Carbon dating - calculating the age of a sample.

Q: The sample of the isotope technetium-97 has a half-life of 2.6 million years. Its current activity is 176 Bq, and its initial activity was 3562 Bq. What is the age of the sample?

A: Identifying the variables, we find $t_{1/2}=2.6$ million years, $A_0=3562$ Bq, A=176 Bq. Using $A=A_0\left(\frac{1}{2}\right)^{\frac{t}{t_{1/2}}}$ and solving for t:

$$A = A_0 \left(\frac{1}{2}\right)^{\frac{t}{t_{1/2}}}$$

$$\Rightarrow 176 = 3562 \left(\frac{1}{2}\right)^{\frac{t}{t_{1/2}}}$$

$$\Rightarrow 0.049 = \left(\frac{1}{2}\right)^{\frac{t}{t_{1/2}}}$$

$$\Rightarrow \ln 0.049 = \frac{t}{t_{1/2}} \ln \left(\frac{1}{2}\right)$$

$$\Rightarrow \frac{-3.01}{\ln 0.5} = \frac{t}{t_{1/2}}$$

$$\Rightarrow 4.338 = \frac{t}{t_{1/2}}$$

$$\Rightarrow t = 4.338 \times 2.6 \text{ million years}$$

$$\Rightarrow t = 11.2 \text{ million years}$$

Practice Questions 6-4 - Calculate the number of half-lives occurred to find out how many atoms remain

- 1. The sample of the isotope samarium-147 has a half-life of 106 billion years. Its current activity is 616 Bq, and its initial activity was 1585 Bq. What is the age of the sample?
- 2. The sample of the isotope berkelium-247 has a half-life of 1380 years. Its current activity is 1555 Bq, and its initial activity was 4940 Bq. What is the age of the sample?
- 3. The sample of the isotope silicon-32 has a half-life of 170 years. Its current activity is 2285 Bq, and its initial activity was 3323 Bq. What is the age of the sample?

Worked Example 6-4 - Carbon dating - calculating the age of a sample.

Q: After a period of 6563 years, the activity of a sample of plutonium-240 decreases from 3884 Bq to 3005 Bq. What is the half-life of the isotope?

A: Identifying the variables, we find t=6563 years, $A_0=3884$ Bq, A=3005 Bq. Using $A=A_0\left(\frac{1}{2}\right)^{\frac{1}{t_1/2}}$ and solving for $t_{1/2}$:

$$A = A_0 \left(\frac{1}{2}\right)^{\frac{t}{t_{1/2}}}$$

$$\implies 3005 = 3884 \left(\frac{1}{2}\right)^{\frac{t}{t_{1/2}}}$$

$$\implies 0.774 = \left(\frac{1}{2}\right)^{\frac{t}{t_{1/2}}}$$

$$\implies \ln 0.774 = \frac{t}{t_{1/2}} \ln \left(\frac{1}{2}\right)$$

$$\implies \frac{-0.257}{\ln 0.5} = \frac{t}{t_{1/2}}$$

$$\implies 0.370 = \frac{t}{t_{1/2}}$$

$$\implies t_{1/2} = \frac{6563}{0.370}$$

$$\implies t_{1/2} = 17,729 \text{ years}$$

Practice Questions 6-5 - Calculate the number of half-lives occurred to find out how many atoms remain

- 1. After a period of 717 thousand years, the activity of a sample of aluminium-26 decreases from 7004 Bq to 3561 Bq. What is the half-life of the isotope?
- 2. After a period of 1380 years, the activity of a sample of berkelium-247 decreases from 2728 Bg to 349 Bq. What is the half-life of the isotope?
- 3. After a period of 6.5 million years, the activity of a sample of palladium-107 decreases from 771 Bq to 269 Bq. What is the half-life of the isotope?

Worked Example 6-5 - Calculate the decay constant and the activity

Q: The isotope lanthanum-137 has a half-life of 60 thousand years. Calculate the decay constant λ , and hence the activity of a sample of $N = 2.3 \times 10^{15}$ atoms of the isotope

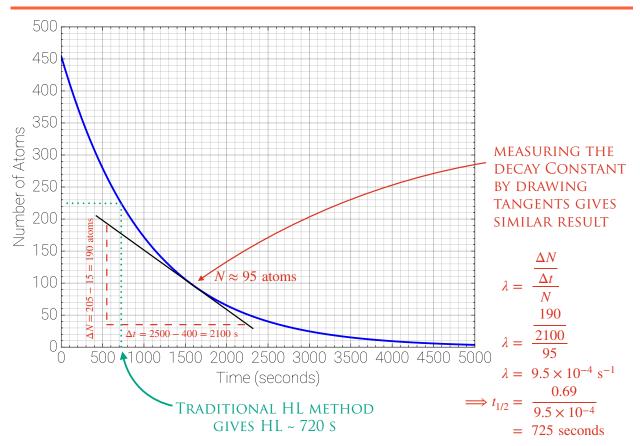
A: Using $\lambda = \frac{\ln 2}{t_{1/2}}$ we can find the decay constant straightforwardly $\lambda = \frac{0.69}{60000 \text{ years}} = 1.15 \times 10^{-5} \text{ year}^{-1}$. We can then find the activity of the sample using $A = \lambda N = 1.15 \times 10^{-5} \text{ year}^{-1} \times 2.3 \times 10^{15} = 2.3 \times 10^{10}$ decays per year, which can easily be converted into decays per second if required.

Practice Questions 6-6 - Calculate the decay constant and the activity

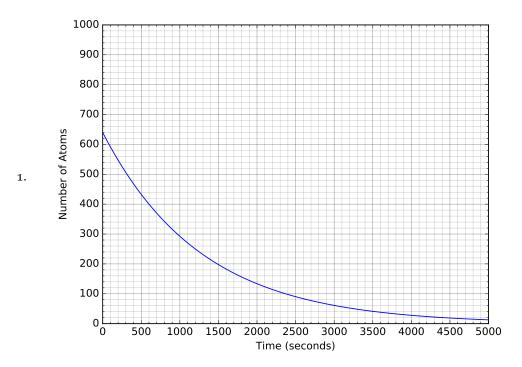
1. The isotope americium-241 has a half-life of 432.2 years. Calculate the decay constant λ , and hence the activity of a sample of $N = 6.4 \times 10^{10}$ atoms of the isotope

- 2. The isotope curium-245 has a half-life of 8500 years. Calculate the decay constant λ , and hence the activity of a sample of $N = 8.7 \times 10^{23}$ atoms of the isotope
- 3. The isotope curium-247 has a half-life of 15.6 million years. Calculate the decay constant λ , and hence the activity of a sample of $N = 3.1 \times 10^{18}$ atoms of the isotope

Worked Example 6-6 - Calculate the decay constant and half-life from the graph

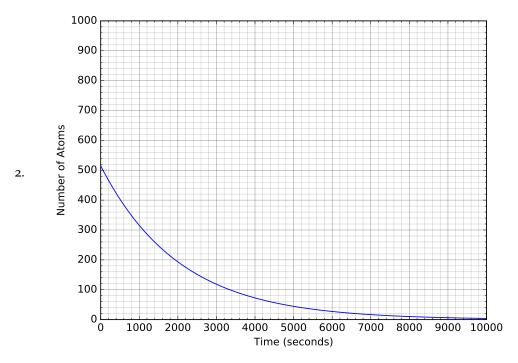


Practice Questions 6-7 - Find the decay constant and the half-life from the graph



Half-Life (from graph) Decay Constant (calculated)

Decay Constant (from graph)



Half-Life (from graph) Decay Constant (calculated)

Decay Constant (from graph)

Worked Example 6-7 - Carbon dating - calculating the age of a sample.

Q: The half-life of Carbon-14 is 5730 years. The ratio of Carbon-14 to Carbon-12 in a fossil is 60% the ratio found in living specimens. Calculate the age of the sample.

A: The ratio of C-14 to C-12 is constant in living organisms i.e. $\frac{C-14_{living}}{C-12_{living}} = const.$

In organisms that have died, the amount of C-14 decreases due to beta decay but the amount of C-12 remains the same.

Therefore, the statement "The ratio of Carbon-14 to Carbon-12 in a fossil is 60% the ratio found in living specimens" means we can write, mathematically:

$$\frac{C-14_{dead}}{C-12_{dead}} = 0.6 \times \frac{C-14_{living}}{C-12_{living}} \Longrightarrow C-14_{dead} = 0.6 \times C-14_{living} \text{ since } C-12_{living} = C-12_{dead}.$$

We now simply plug into our half-life equations $N=N_0\left(\frac{1}{2}\right)^{t/t_{1/2}}$ or $N=N_0e^{-\lambda t}$ to give:

$$C - 14_{\text{dead}} = C - 14_{\text{living}} e^{-\lambda t}$$

$$\implies 0.6 \times C - 14_{\text{living}} = C - 14_{\text{living}} e^{-\lambda t}$$

$$\implies 0.6 = e^{-\lambda t}$$

$$\implies \ln 0.6 = -\lambda t$$

$$\implies t = \frac{0.51}{0.69} \times 5730 \text{ years}$$

$$\implies t = 4220 \text{ years}$$

Practice Questions 6-8 - Carbon dating - calculating the age of a sample.

The half-life of Carbon-14 is 5730 years.

- 1. The ratio of Carbon-14 to Carbon-12 in a fossil is 60% the ratio found in living specimens. Calculate the age of the sample.
- 2. The ratio of Carbon-14 to Carbon-12 in a fossil is 25% the ratio found in living specimens. Calculate the age of the sample.
- 3. The ratio of Carbon-14 to Carbon-12 in a fossil is 7% the ratio found in living specimens. Calculate the age of the sample.

Worked Example 6-8 - Carbon dating - calculating the age of a sample.

Q: The half-life of Carbon-14 is 5730 years. A bone sample from an archaeological dig has an activity of 5.2 Bq per gram. A bone sample from a modern skeleton has an activity of 6.5 Bq per gram. Calculate the age of the sample.

A: The ratio of C-14 to C-12 is constant in living organisms i.e. $\frac{C-14_{\text{living}}}{C-12_{\text{living}}} = \text{const. Using } A = \lambda N$, this means the activity is proportional in the same way.

We can therefore say that the initial activity $A_0 = 6.5$ Bq and the current activity A = 5.2 Bq and plug this into our $A = A_0 \left(\frac{1}{2}\right)^{t/t_{1/2}}$ or $A = A_0 e^{-\lambda t}$ equations to give:

$$\implies 5.2 = 6.5 e^{-\lambda t}$$

$$\implies 0.8 = e^{-\lambda t}$$

$$\implies \ln 0.8 = -\lambda t$$

$$\implies t = \frac{0.223}{0.69} \times 5730 \text{ years}$$

$$\implies t = 1850 \text{ years}$$

Practice Questions 6-9 - Carbon dating - calculating the age of a sample.

The half-life of Carbon-14 is 5730 years. The activity of a bone sample from a modern skeleton has an activity of 6.5 Bq per gram.

- 1. A bone sample from an archaeological dig has an activity of 3.25 Bq per gram. Calculate the age of the sample.
- 2. A bone sample from an archaeological dig has an activity of 2.0 Bq per gram. Calculate the age of the sample.
- 3. A bone sample from an archaeological dig has an activity of 0.75 Bq per gram. Calculate the age of the sample.

Worked Example 6-9 - Carbon dating - calculating the age of a sample.

Q: When examining a small sample of an old skull, scientists found that 2.3×10^{11} % of the carbon was C-14, where as in recently deceased skulls the C-14 proportion is 1.0×10^{-10} %. Calculate the age of the skull.

A: The initial proportion of C-14 is $N_0 = 1.0 \times 10^{-12}$ and the current proportion is $N = 2.3 \times 10^{13}$ (where we have converted from percentages into decimals). Plugging into $N = N_0 \left(\frac{1}{2}\right)^{t/t_{1/2}}$ or $N = N_0 e^{-\lambda t}$ gives:

$$\implies 2.3 \times 10^{-13} = 1.0 \times 10^{-12} e^{-\lambda t}$$

$$\implies 0.23 = e^{-\lambda t}$$

$$\implies \ln 0.23 = -\lambda t$$

$$\implies t = \frac{1.47}{0.69} \times 5730 \text{ years}$$

$$\implies t = 12,207 \text{ years}$$

Practice Questions 6-10 - Carbon dating - calculating the age of a sample.

The proportion of C-14 in recently deceased skulls is 1.0×10^{-10} %. The half-life of C-14 is 5730 years.

- 1. When examining a small sample of an old skull, scientists found that 2.3×10^{11} % of the carbon was C-14. Calculate the age of the skull.
- 2. When examining a small sample of an old skull, scientists found that 8.5×10^{12} % of the carbon was C-14. Calculate the age of the skull.
- 3. When examining a small sample of an old skull, scientists found that 6.3×10^{11} % of the carbon was C-14. Calculate the age of the skull.

Worked Example 6-10 - Carbon dating - calculating the age of a sample.

Q: Carbon-14 has a half-life of 5730 years. A living organisms contains approximately 1 atom of C-14 for every 10¹⁸ atoms of C-12. A fossil contains 12 atoms of C-14 for every 10²⁰ atoms of C-12. Estimate its age.

A: The initial proportion of C-14 is
$$N_0 = \frac{1}{1 \times 10^{18}} = 1.0 \times 10^{-18}$$
 and the current proportion is $N = \frac{12}{1 \times 10^{20}} = 1.2 \times 10^{-19}$ using $N = N_0 \left(\frac{1}{2}\right)^{t/t_{1/2}}$ or $N = N_0 e^{-\lambda t}$ gives:

$$\implies 1.2 \times 10^{-19} = 1.0 \times 10^{-18} e^{-\lambda t}$$

$$\implies 0.12 = e^{-\lambda t}$$

$$\implies \ln 0.12 = -\lambda t$$

$$\implies t = \frac{2.12}{0.69} \times 5730 \text{ years}$$

$$\implies t = 17,607 \text{ years}$$

Practice Questions 6-11 - Carbon dating - calculating the age of a sample.

Carbon-14 has a half-life of 5730 years. A living organisms contains approximately 1 atom of C-14 for every 10^{18} atoms of C-12.

- 1. A fossil contains 32 atoms of C-14 for every 10^{20} atoms of C-12. Estimate its age.
- 2. A fossil contains 73 atoms of C-14 for every 10^{20} atoms of C-12. Estimate its age.
- 3. A fossil contains 2 atoms of C-14 for every 10^{20} atoms of C-12. Estimate its age.

Q: In a particular rock sample, 145 of the initial 256 Carbon-14 atoms have decayed. Calculate the age of the rock.

A: If 145/256 = 57% of the atoms have decayed, this means that the remaining proportion is N = (256 - 145)/256 = 111/256 = 0.43. Plugging into $N = N_0 \left(\frac{1}{2}\right)^{t/t_{1/2}}$ or $N = N_0 e^{-\lambda t}$ gives:

$$\implies 0.43N_0 = N_0 e^{-\lambda t}$$

$$\implies 0.43 = e^{-\lambda t}$$

$$\implies \ln 0.43 = -\lambda t$$

$$\implies t = \frac{0.844}{0.69} \times 5730 \text{ years}$$

$$\implies t = 7,008 \text{ years}$$

Practice Questions 6-12 - Carbon dating - calculating the age of a sample.

Carbon-14 has a half-life of 5730 years.

- 1. In a particular rock sample, 64% Carbon-14 atoms have decayed. Calculate the age of the rock.
- 2. In a particular rock sample, 14% Carbon-14 atoms have decayed. Calculate the age of the rock.
- 3. In a particular rock sample, 95% Carbon-14 atoms have decayed. Calculate the age of the rock.

Worked Example 6-12 - Carbon dating - calculating the age of a sample.

Q: Using Carbon-14 to date samples is not always possible. For dating rocks taken from the moon on the Apollo missions, scientists used the decay of potassium-40 to argon-40. The half-life of potassium-40 is 1.3×10^9 years. In one particular rock sample scientists found 0.84 μ g of argon-40 and 0.10 μ g of potassium-40. Calculate the age of the rock sample in years.

A: The K-40 decays into Ar-40 and therefore as the mass of K-40 decreases the mass of Ar-40 increases in exactly proportional amounts. The sample must have started with $0.84+0.10=0.94\mu g$ of K-40, but there is now only $0.1\mu g$ remaining. We then simply plug this into our $N=N_0\left(\frac{1}{2}\right)^{t/t_{1/2}}$ or $N=N_0e^{-\lambda t}$ equations to give:

$$\implies 0.1 = 0.94 e^{-\lambda t}$$

$$\implies 0.106 = e^{-\lambda t}$$

$$\implies \ln 0.106 = -\lambda t$$

$$\implies t = \frac{0.106}{0.69} \times 1.3 \times 10^9 \text{ years}$$

$$\implies t = 4.2 \times 10^9 \text{ years}$$

Practice Questions 6-13 - Alternatives to Carbon dating

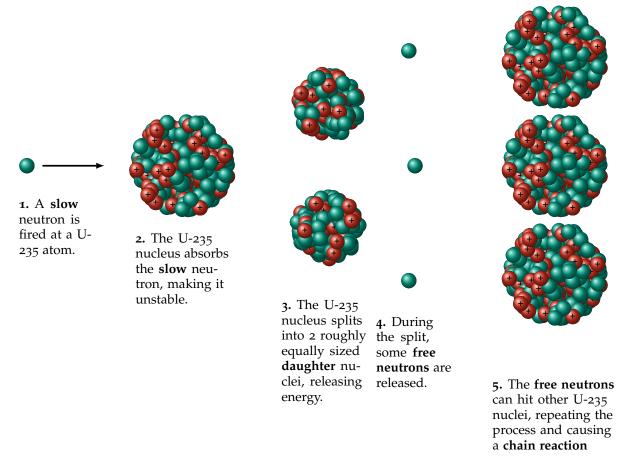
The half-life of potassium-40 is 1.3×10^9 years.

1. In one particular rock sample scientists found 0.54 μg of argon-40 and 0.20 μg of potassium-40. Calculate the age of the rock sample in years.

- 2. In one particular rock sample scientists found 1.4 μg of argon-40 and 0.74 μg of potassium-40. Calculate the age of the rock sample in years.
- 3. In one particular rock sample scientists found 0.22 µg of argon-40 and 1.38 µg of potassium-40. Calculate the age of the rock sample in years.

Nuclear Fission

Nuclear fission involves splitting atoms (usually Uranium) into smaller elements and releasing vast amounts of energy. Fission can either be controlled and used in nuclear reactors to produce useful energy or uncontrolled and used in nuclear bombs to produce explosions.



Nuclear Fission Equations

When the parent Uranium-236 nucleus splits into two daughter nuclei, it generates free neutrons. The number of free neutrons released depends on the fission products i.e. the exact two daughter nuclei the U-236 breaks into. Some reactions produce more free neutrons than others. Let's look at the example of Uranium-235 absorbing a neutron and fissioning into Krypton and Barium:

$$^{235}_{92}U + ^{1}_{0}n \longrightarrow ^{90}_{36}Kr + ^{143}_{56}Ba + 3 (^{1}_{0}n)$$

Figure 19: The mass numbers and the atomic numbers on the right hand side must equal the mass number on the left hand side.

Practice Questions 7-1 - How many neutrons are released in the following fission reactions?

1.
$${}^{238}_{92}U + {}^{1}_{0}n \longrightarrow {}^{181}_{71}Lu + {}^{55}_{21}Sc + {}^{1}_{0}n$$

2.
$${}^{23^2}_{90}\text{Th} + {}^{1}_{0}\text{n} \longrightarrow {}^{91}_{37}\text{Rb} + {}^{140}_{53}\text{I} + {}^{1}_{0}\text{n}$$

Practice Questions 7-2 - What is the second daughter nucleus?