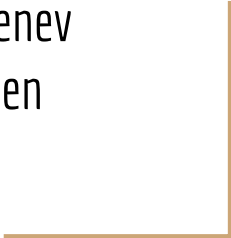


# Measurement of Subpixel Structure

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By: Benjamin Oommen



# Foreword

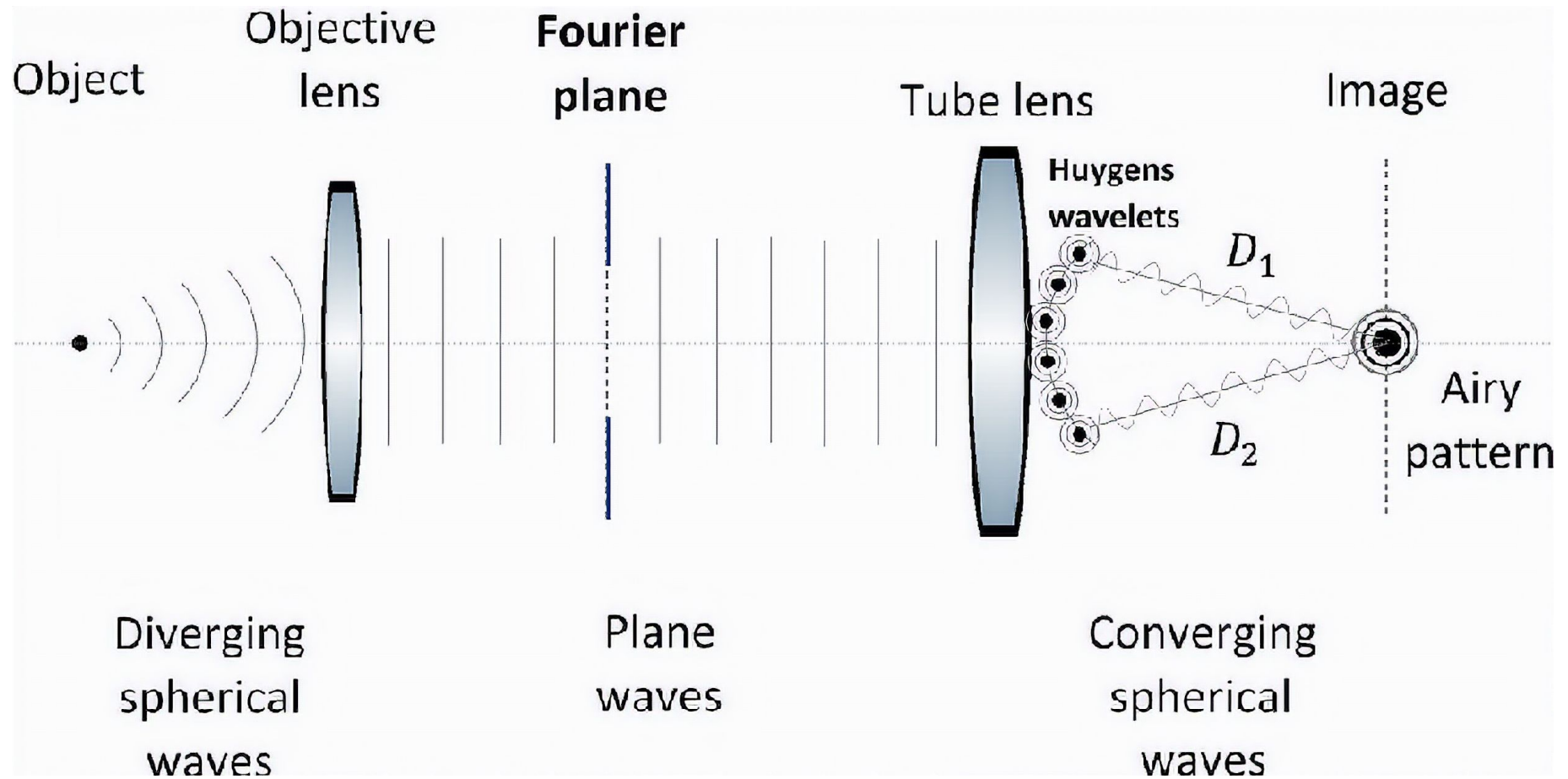
I would like to express my thanks to Dr. Kaloyan Penev who graciously set aside time every day to help me with this project and explain new concepts to me.

# Problem

- Detection of exoplanets yields lots of noise
- Inconsistencies on outer edges of pixels on detector plane
- If we were able to deduce how sensitivity changes at some point on the pixel, we could mathematically correct for this phenomenon

# Experiment

- We emit spherical waves through a mask which are reflected off mirrors and converge to the detector plane
- The mask allows for the light to form a certain pattern on the detector plane
- We do this for a grid of spots to create more data points to interpolate on.

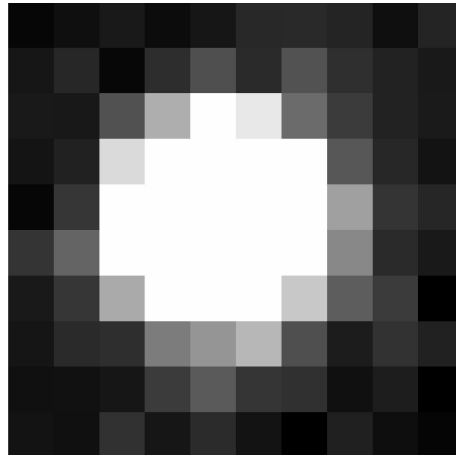


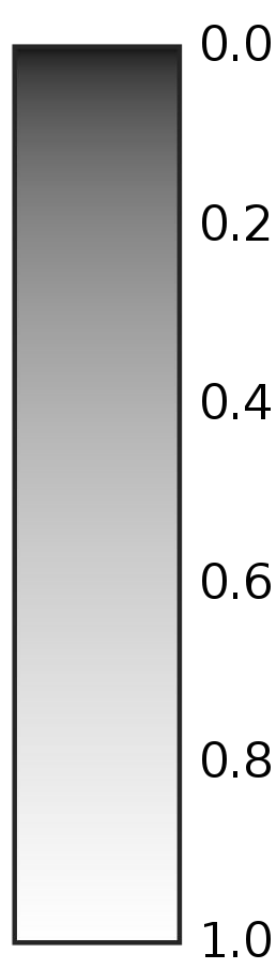
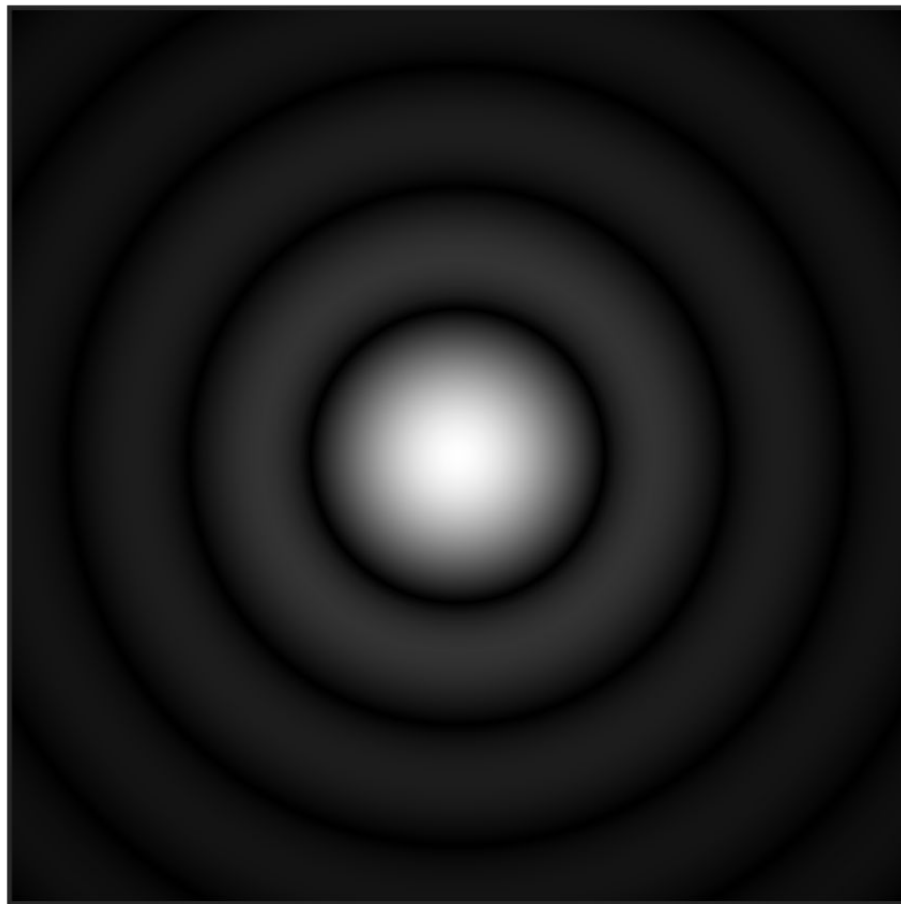
Source: [Wikipedia](#)

- Sample of spots on a grid and their respective point spread functions

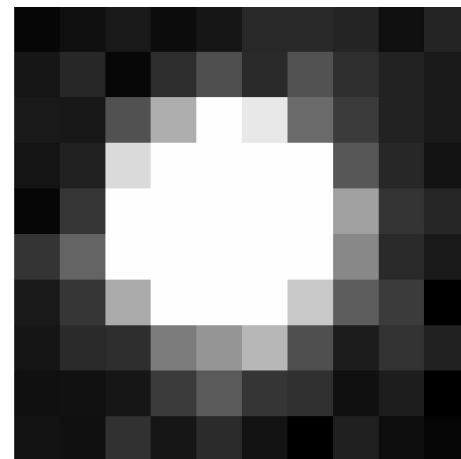
# Point Spread Function

- The waves converge to a spot, and the distribution of intensity is modeled by the point spread function.
- The point spread function is defined as the convolution between an Airy Disk Function and a 2D gaussian function.





Example of airy rings produced by the airy disk function, and its similarity to the response of the pixels of a spot





# Point Spread Function (Fourier Space)

## **Airy Disk**

Represents the intensity  
when given an x,y coordinate  
on the aperture

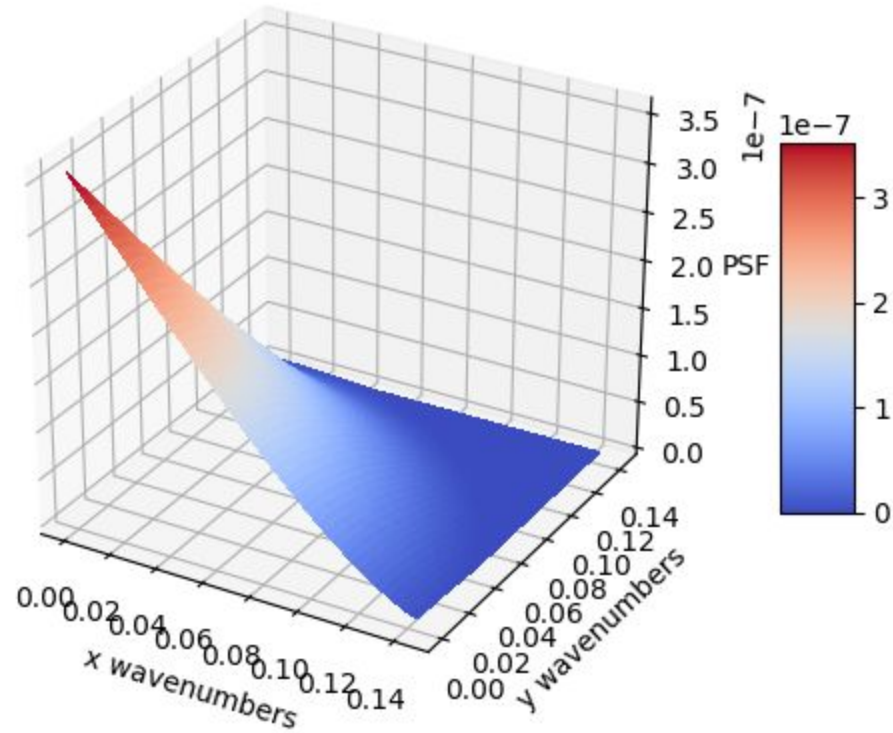
## **Gaussian**

Used to account for  
mechanical jitter

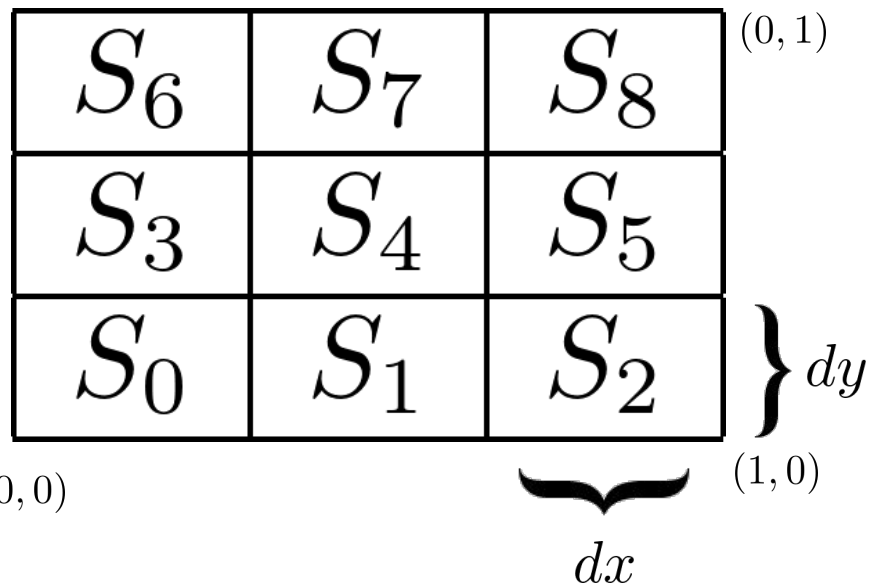
## **Fourier Point Spread Function (PSF)**

$$PSF(\lambda, N, k_x, k_y, \sigma_x, \sigma_y) = G(\sigma_x) \cdot G(\sigma_y) \cdot OTF$$

## PSF Values Graph (In the fourier space)



# Modeling Pixels



- To the left is the way we chose to model a pixel, which is a grid of subpixels
- In our actual model, each pixel was split into a 10x10 grid of subpixels

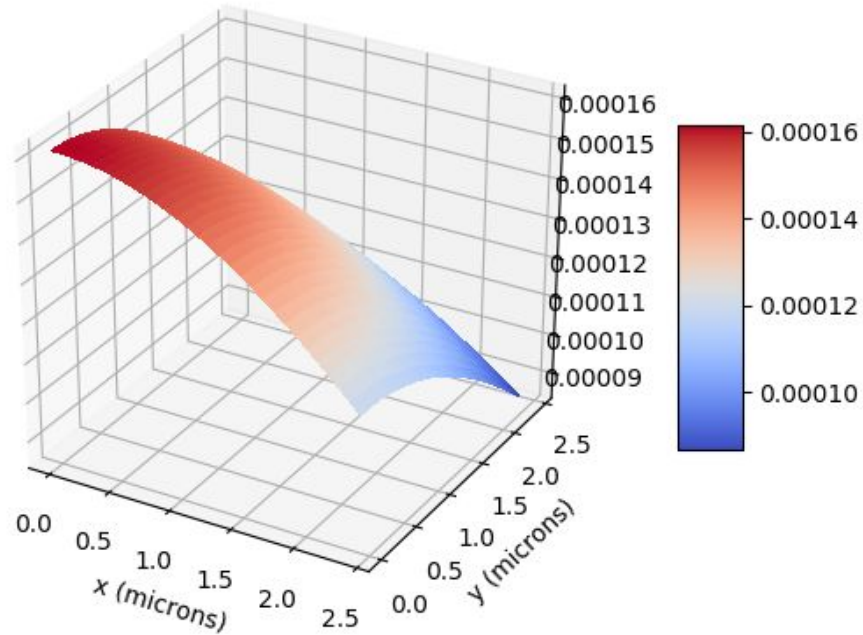
# Response (PRF)

- The PRF can be evaluated as the point spread function integrated over each subpixel, and then multiplied by its corresponding subpixel sensitivity map coefficient.
- This represents the intensity of the subpixel. Summing these integrals results in the PRF value of the pixel

# Discrete Cosine Transform (DCT-II)

- Function for intensity is in the fourier space, so an inverse fourier transform is required to get back into the real space
- Results in a DCT-II (Discrete Cosine Transform type 2) due to the imaginary portions of our integral cancelling out
- We use FFTW to calculate this transform
- FFTW is written purely in C and results in excellent performance

# Intensities



# Solving for the sensitivity map

- After executing the DCT, we now have the values of the intensities for all subpixels of each pixel
- We can now use least squares to solve for the coefficients

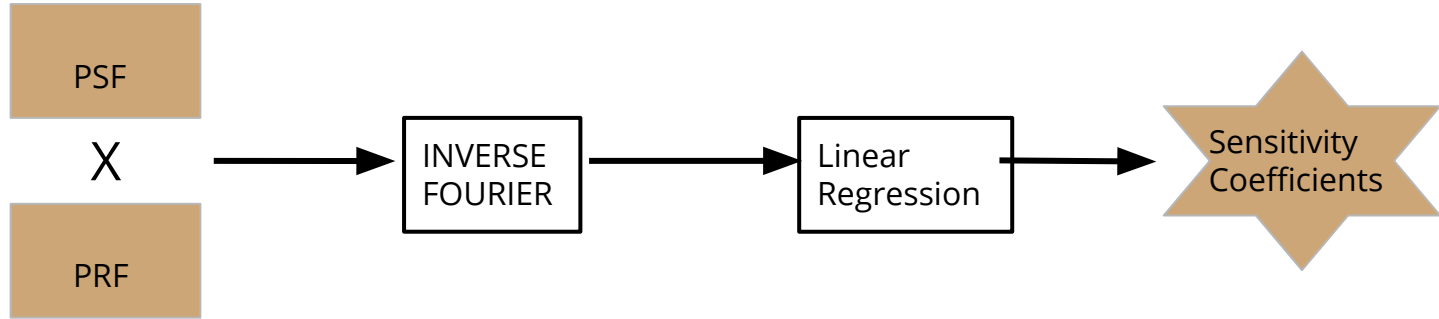
$$\begin{bmatrix} r_0 \\ r_1 \\ r_2 \\ \vdots \\ r_{p_y+p_x P_y} \end{bmatrix} = \begin{bmatrix} I_{0,0} & I_{0,1} & I_{0,2} & \cdots & I_{0,n_y+n_x N_y} \\ I_{1,0} & I_{1,1} & I_{1,2} & \cdots & I_{1,n_y+n_x N_y} \\ I_{2,0} & I_{2,1} & I_{2,2} & \cdots & I_{2,n_y+n_x N_y} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ I_{p_y+p_x P_y,0} & I_{p_y+p_x P_y,1} & I_{p_y+p_x P_y,0,2} & \cdots & I_{p_y+p_x P_y,n_y+n_x N_y} \end{bmatrix} \begin{bmatrix} S_0 \\ S_1 \\ S_2 \\ \vdots \\ S_{n_y+n_x N_y} \end{bmatrix}$$

# Optimizations

- PSF and PRF models were written in C++
- Convolutions were performed using fast fourier transforms rather than numeric integration



# Outcomes



# References

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- “Point Spread Function.” *Wikipedia*, Wikimedia Foundation, 27 July 2019, [en.wikipedia.org/wiki/Point\\_spread\\_function](https://en.wikipedia.org/wiki/Point_spread_function).

# Software Libraries Used

C/C++:

- FFTW (Fastest Fourier Transform in the West) - used to compute fast fourier transforms
- boost - used to compute the bessel integral

Python:

- numpy - used for various numerical methods involving matrices
- scipy - used for spline generation and interpolation
- matplotlib - used for the generation of graphs