

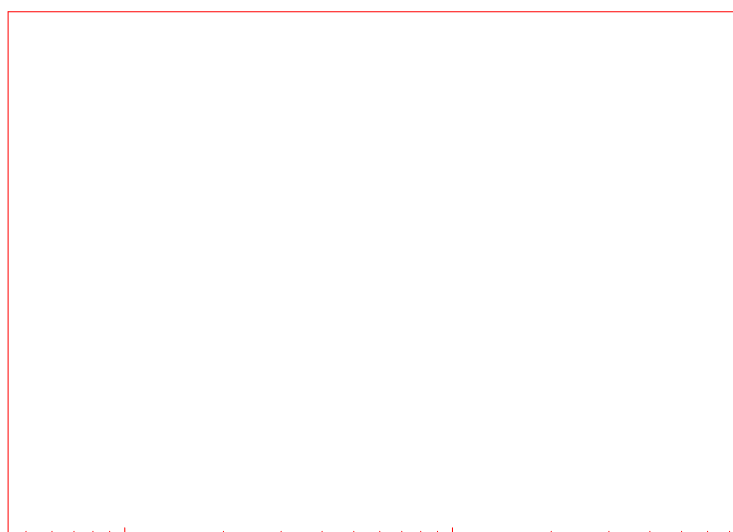
LOFAR calibration strategy

J.E. Moorham, STRO

14th November 2005

Abstract

This is the state of affairs at the LOFAR Critical Design Review,



2 Development strategy

Reduction of the SRT observation that include the experimental LOFAR/ HAT station. This exercise is the modelling of LOFAR HBA voltage beam, and the combination of (highly) differential station. This is the first example of a project that cannot be handled by any of the existing package.

Ex1.7: 117 MHz (n SRT LFFE) exercise the Minimum Ionospheric Model (MIN) concept, and other low-frequency issue that are highly relevant for LOFAR. Optionally another LFFE field without a bright source could be attempted, to exercise solving for phase gradient only (after cross-calibration with a calibrator source). It could also be used to exercise redundant spacing calibration.

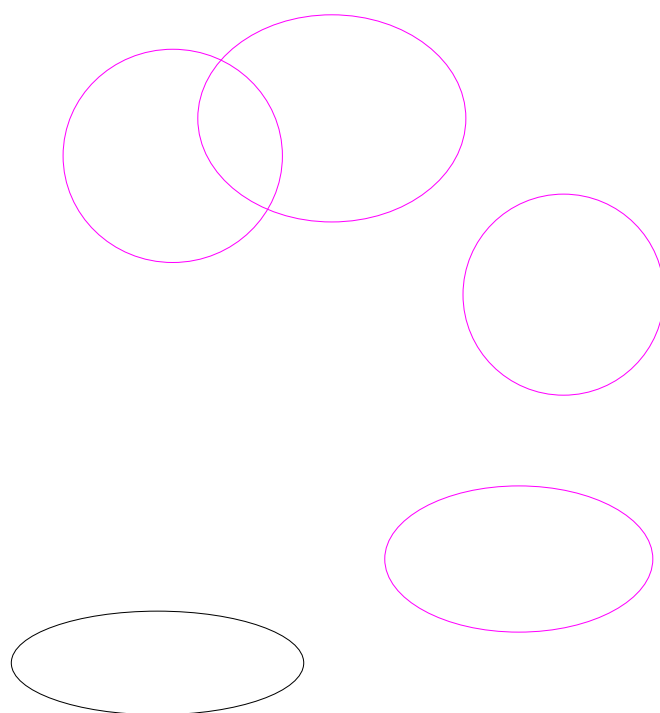
Ex...: This field contains a substantial number of bright point source in the field. It exercises the solving for individual station beam shape (which include pointing error).

See also the appendix.

3. The subtraction = source subtraction

As stated before, calibration can be equated with the subtraction of foreground source. Because of image-plane effect, they must be subtracted from the uv-data as much as possible. Unfortunately, in order to preserve the field size, subtraction must be done at full (ft) resolution, which is expensive. We distinguish three categories of source:

1. **Cat I sources** are the 20- 0 source per FOV that are so bright that they



Solving for MLE parameters (Meta Params)

Arguably the most important function of the calibration system is solving for M.

7 Local Sky Model (LSM)

The Local Sky Model (see fig 5) represent the subset of the GSM that is relevant to a particular observation. It contains the following main component :

The **source list** refer to the source that are included in the LSM. The majority will be parametrised source component (P

field-of-view

Some of the effects have been analysed to some extent (cite BSR), but it remains unclear what their precise impact on the image will be in practice. In addition, there are bound to be other ones, that we have not yet thought of. In some cases, the artefact may be impossible to get rid of entirely, so the best we can do is make sure that it cannot be confused with a scientifically interesting feature like the EoR. In all cases, the only way ahead is to use adequate tools to patiently chip away at the problem.

phers (MIM)

Over the year, the idea for an ionospheric model for LOFAR have undergone

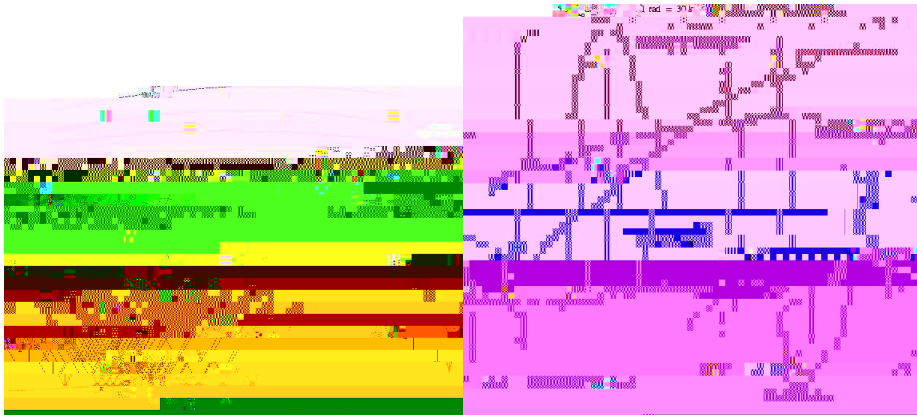


Figure 8: *The*

12 Station beam shapes

The LOFAR station primary beam will have much higher side-lobes than the parabolic dish that we are used to. In the words of Jaap Bregman: “LOFAR

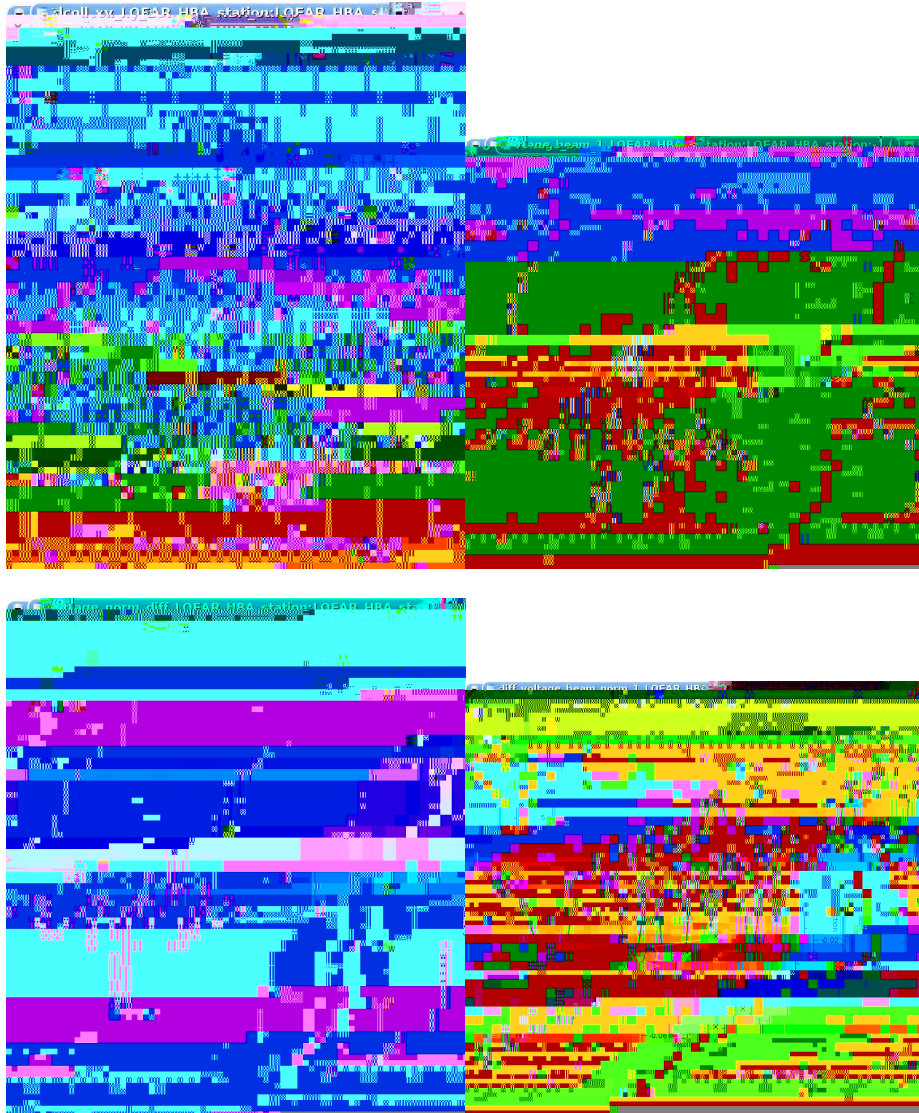


Figure 9: Simulation of LOFAR HBA station beams, using MeqTrees. The top-left (meqbrowser) image shows the HBA antenna configuration, while top-right shows one of the two voltage beams @ 150 MHz, when pointed at azimuth (ϕ) of 0 rad, and elevation of 1.0 rad. Bottom-left shows the normalised difference between two voltage beams, and bottom-right shows the normalised

at II source in tation beam idelobe

MIM validation

Minimi ation of peeling contamination

primary beam.

B



C.2 The MeqBrowser GUI

The MeqBrowser GUI (see fig 11) is a powerful tool for the generation, control, and inspection of MeqTrees. It has the following main features :

- Script editor. Once a TDL script is loaded, it may be edited. The result

It cannot be stressed enough that the functionality of the MegBrowser GUI enormously reduces the turnaround time for experiments and debugging. If anything, this speeds up development time proportionally, and reduces the chance of undetected errors.

C.3 Lowering the threshold

MG_cript (example, demonstration, experimentation, importable function) and 'official' TDL_cript.

D Glossary

Since the LOFAR calibration system necessarily contains a few innovations, the

olver/condeq: A olver collect condition equation from it condeq children, and accumulate them into a olution matrix. It invert te matrix,