Hamilfonian Mechanics: Basic Theory

o-) Hamilton's Canonical equations

- · Canonical momentum: Pi= OL
- Configuration space = "Photo of the system" (q') dim n ; n dof
- · Phase space: Describes physical states

(qi, Pi) - dim 2n; 2n dof, 2n coordinates burdle)

(. Vebcity phase space

(qi qi) ... dim 2Ns 2N dof = 2N coordinates (tangent bundle)

· Hamiltonian governs the dynamics

H(91Pt)=P,d'-L, g=q(91Pit)

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Theorem: The system of n (E-L) is equivalent to the system of 2n

first order Flamilton's equations

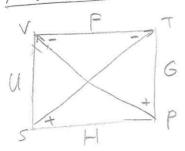
Proof: Variational) principle

- Sopi q'+ Pi oq' - opi oq' - opi opi) dt = 0 55=0= 5/cdt = 50(piqi-H)dt

= \ \delta \ \text{Pi} \left(\frac{\delta i - \delta H}{\delta pi} \right) \ \dt + \left(\frac{\delta i - \delta H}{\delta qi} \right) \ \dt = 0

Pemark: dH = OH q+ OH p+ OH OP p+ OH OP OP OP OP OP

H H # H(t) => H: conserved energy





$$\frac{\partial V}{\partial S} p = \frac{2 + 1}{\partial P} S$$

$$\frac{\partial V}{\partial S} p = -\frac{0 + 1}{\partial P} T$$

Proof:

$$dH = \frac{1}{2} \frac{1}{2}$$

b-) Poisson Brackets

Definition: For two phase space functions f and g. Their canonical poisson bracket is a new phase-space function. (fig) = 2f 29 - 2f 29 091

Properties

- Integral of motion
$$\frac{dI}{dt} = 0 = \{I_1H\}$$

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Nok that having two integrals of motion, In and I2
$$I_3 = \{I_1I_2\}$$
 is also an integral of motion.

C-) Canonical Transformations Definition: A caronical transformation $Q^{j} = Q^{j}(q^{i}, P_{i}, t)$, $\mathcal{L}_{j}^{i} = P_{j}(q^{i}, P_{i}, t)$ is such transformation so that It preserves the form of the Hamilton's equations. That is, one can find of new Hamiltonian" H'= H'(Qî, P; ,t). OH' = Q+ , OH' = P+ Ex: Not every transformation is canonical $H = \frac{P^2}{2m}$, $Q = \sqrt{q}$, P = P is this canonical? Free particle giessing H'= P2 $\frac{\partial H}{\partial P} = \frac{P}{M} = \frac{\partial H}{\partial P} = \frac{\dot{q}}{\dot{q}} = 2\dot{Q}\dot{Q} \neq \dot{Q}$ · Canonical transformations are generated by generating functions of the type Since old E(91P) and new E(Q1P) we have 4 possibilities (related

Ex. $F = F(q_1Q_1t)$ Adding a fotal derivative to the cagrangian does not charge the eom. $S = \int dt (pq - H - df) = \int f(q_1p - q_2p) dt$ $S = \int dt (pq - H'(f_1Q_1t)) dt$

 $\frac{\partial P}{\partial q} = -\frac{\partial P}{\partial \Omega}$

Remarks:

1-) Solve the (E-L) equations for 9, expressing it in terms of the boundary value In order to obtain this function

ii-) Once this solution is known, we plug it back to L, and integrate L over to get

ince
$$Ldt = P_i dq^i - Hdt$$

 $dS = L_f - L_i = (P_i dq^i - Hdt)_{final} - (P_i dq^i - Hdt)_{initio})$ (3.17)

ds=
$$\frac{05}{061}$$
 dt + $\frac{05}{091}$ dq + $\frac{05}{06}$ dt + $\frac{05}{09}$ dq dq + $\frac{05}{09}$ dq

leads us to Hamilton - Jacobs equation

$$\frac{\partial S}{\partial q_1} = Pi$$
, $\frac{\partial S}{\partial t} = -H$

Note also that upon 90=91, to=ti

Solution: Time evolution can be written as coordinate transformation $q(t) = q(q_0|P_0,t)$ $p(t) = p(q_0|P_0,t)$ where $(q_0|P_0)$ are the "old coordinates" and (q_1P) are the new ones".

Integrating (3.14) $\int (pdq - HoH) = \int (P_0 dq_0 - Hodt) - d(-5) dt$ $\int (pdq - HoH) = \int (P_0 dq_0 - Hodt) - d(-5) dt$

F=-5 generates time evolution

Liouville's Equation: Volume of the phase space remains invariant under canonical fransformations. In particular, time evolution preserves the phase space volume.



Proof:
$$(n=1)$$
 $V_t = \int dQ dP = \int facobion | dq dP$
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 $J = \begin{vmatrix} QQ & QQ & QP \\ QQ & QP \end{vmatrix} = \frac{QQ}{QQ} \frac{QP}{QP} - \frac{QQ}{QQ} \frac{QQ}{QQ} = \frac{QQ}{QQ} = \frac{QQ}{QQ} \frac{QQ}{QQ} = \frac{QQ}{QQ} = \frac{QQ}{QQ} = \frac{QQ}{QQ} = \frac$

Idea: Find a special generating function 5, 50 that new Hamiltonian d-) Hamilton - Jacobi Theory $\partial \hat{J} = \frac{\partial H}{\partial P} = 0$ = $\partial \hat{J} = \partial \hat{J} = constant$ $\hat{P}_{j} = -\frac{\partial H'}{\partial \Omega^{j}} \partial \rightarrow \hat{P}_{j} - \beta \hat{\tau}$ constant $P_{j} = -\beta = -\frac{\partial S(q_{i}Q_{i})}{\partial Q_{j}} = \frac{\partial S(q_{i}Q_{i}t)}{\partial Q_{j}}$ $P_{j} = \frac{\partial S(q_{i}Q_{i}t)}{\partial Q_{j}}$ This is an implicit solution, the explicit one is found by inversion $H' = H + \frac{\partial f}{\partial t} = 0 \implies \frac{\partial s}{\partial t} + H(q', \frac{\partial s}{\partial q'}, t') = 0 \qquad \text{one-PDE for } S(q, t')$ $Pj = \frac{\partial s}{\partial q'}$. Solve (H-J) equations interested in the solutions such that Type of solutions which depends on a integration constants are omplete integral . A trick for finding this is to use additive seperation of variables S(qit) = S(q1) + S2(q2) - Sn(qn) + St(t) If H+H(t), when 5 put into the HJ H+H(x) S= XCx +So x-independent Ex: Freefall in homogeneous gravitational field $H = \frac{p^2}{2m} - mgX$, V(x) = -mgX $\frac{1}{2m}\left(\frac{\partial S}{\partial X}\right)^2 + V(x) + \frac{\partial}{\partial t}(S_0 - Et) = 0$ 1 (25)2 +V(x) -E+ 350=0

$$\frac{(QS)^{2}-2ME-V)}{So} = So = \int dx | 2m (E-V)'$$

$$So = \int dx \sqrt{2m(E+mgx)}, performing the integral$$

$$S = \frac{1}{3gm} \left(2m (E+mgx) \right)^{3/2} - E+$$

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Taking now the K-30 limit

 $\frac{1}{2m}\left(\frac{35}{0x}\right)^2 + V + \frac{35}{02} = 0 \quad \text{which is the (H-J) equation}$

e-) Integrable systems

Definition: The dynamical system with a dof (2n-dimensional phase space) is completely (Liouville) integrable, if it posses in independent conserved quantities Fi(qip)=fi, {H,Fi3=0, that are in involution: {Fi,Fj3=0 for all ij

· Liouville's Theorem: The solution of eom of a completely integrable system can be obtained by quadrature, that is finite numbers of integrations and algebraic operations

Proof is constructive and tells how knowing of Fig, I can find the solution by

. Interesting observation is that I need only n (not 2n). Each integral of motion

similar steatement in field theory; In any gauge theory Number of True = Number of apparent - 2 (Number of dof)

EX: EM described in terms of (4,A)

Posy-Oth => EIB are invariant where 1 is arbitrary function 产一年1

True = 4-2=2 polarizacition

of photon 1: gauge function and has one dof(scalar) 4: apparent (4) B3

1: gauge function

Ex: Oravity

gur 4x4 symmetric modifix
4x4 symmetric modifix

gen -> gen + Vu Syt Vv Syn Signinger function with 4 dof

True = 10-8=2 > polarization of graviton

f-) constraints

· Let not all q', P; be independent, that is at any instant of time we have a anstraint

- Remark. Any phase space function of can be understood as "Hamiltonian" Time evolution of I = 0, soit is

dq' = (q',f) dp = (p,f) Time evolution generated by I leaves the Hamiltonian invariant.

Specifically, a conserved quantity obeys (I,H) = (H, I) = 0. This means that Hamiltonian is invariant under the flow generated by I -- symmetry

Constants are "conserved quantities" and imply the presence of symmetries in physical systems. constraint description

" clever description" (Occam) · Standart example

. n=1 dof

. 2n=2 phase space (4,P4)

. No constraint & Eliminate constraint

. N=2 dof X . phase space (x,y, Px, Py) · constraint x2+y2-l2=0

X= lcos4 y= esiny

· Advantages

- · Can calculate "force of constraint" (Tension of string T) · Extra symmetry
- . To deal with constraints one uses the method of Lagrange multipliers Namely in the presence of m number of constraints \$\phi_j\$, we define

total Hamiltonian

HT= H+ = xf (+1 P)

Attl: Lagrange multipliers

The eom are then derived with this total Hamiltonian

2 possibilities 1) Enlarge phase space even further Consider Aj(t) as new coordinate We have to vary wit 27 as well (2+M) E-L equations dimension 3 constraint Qj=0, f=1...m: on shell constraints + n modified (E-L) for gt L7= L+ NJPJ 85= 862 + 579 \$5 + x7 647 > = / (3L da + dt (3g dg) - dt (3L) ofg] + xi [30 69 + d (30 69) - St (30) 69) + \$5 5 xi $= \int \left[\frac{\partial (\zeta + \lambda \phi)}{\partial q} - \frac{\partial}{\partial t} \left(\frac{\partial (\zeta + \lambda \phi)}{\partial q} \right) \right] dq + \phi_j dA_j$ = \[\left(\frac{\partial LT}{\partial q} - \partial \left(\frac{\partial LT}{\partial q} \right) \right) \right) \rightarrow q + \partial p \rights \right) together with on-shell constraint We can use these to solve for gi as well as Aj. Note also that proceeding to the Hamiltonian picture, we find that PX5=0 and so HT 2) Keep Not) as arbitrary function. In this case we have n (E-L) for qu Flow generated by constraint

All the endpoints are physically equivalent (related by which depends on Aj. Time evolution is not unique. the symmetry)

2

· Remark. One source of constraints is a Legendre transformation as L->H

 $P_i = \frac{\partial L}{\partial q^i} \stackrel{\text{Invert}}{=} q^i = q^i (q^j, P_j, t)$ may not exist

Infact:

Air = Orl ogiogi ... if this matrix has seve eigenvalues (det A=0)

one can write qi = qi(qkpkixj,t) J = 1-m where m is the number of zero

Toy example: consider the motion of free particle

L= 1 m x² but using "lousy time"

T=T(t): inertial time and t is lousy time starting from the c written in terms of the lowsy time and performing the Standart Legendre transformention we find that corresponding Hamiltonian

VOLAIShes; H=0: does not generate dynamics

This is a characteristic of a totally constraint system

so there is no "time evolution" and the only flow is with the constraint. (Total Hamiltonian is non trivial and determines this flow. Despite the lack of time evolution, relational predictions are still possible for this type of a system.

(-) If Li, L5 are onserved, so is Lk/

2-)
2.11 a).
$$Cf(x)(g(x)) = Cf(x) + f'(x) \cdot x + f'(x) \cdot x^{2} ... / g(x) + g'(x) \cdot x + - 1 = 0$$

$$Cf(x)(g(x)) = \sum_{n=0}^{\infty} \frac{f(n)(n)}{n!} Cf(x)^{n}$$

$$= -\frac{\partial x}{\partial x} \sum_{n=0}^{\infty} \frac{1}{(n!)^{(n)}} \sum_{n=0}^{\infty} \frac{1}{(n!)^{(n)}} = -\sum_{n=0}^{\infty} \frac{1}{(n!)^{(n)}} \sum_{n=0}^{\infty} \frac{1}{(n!)^{(n)}}$$

$$= -\frac{\partial x}{\partial x} \sum_{n=0}^{\infty} \frac{1}{(n!)^{(n)}} \sum_{n=0}$$

$$\begin{aligned} & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x) \right] = - \frac{\partial a}{\partial a} f \\ & - \left[P_{a_1} F(x)$$

$$[X_{\alpha},V(x)]=0$$

$$(-) \quad \dot{p}_{\alpha} = [P_{\alpha},H] = [P_{\alpha},P_{\alpha}] + V(x)] = [P_{\alpha},V(x)] = -\frac{\partial V_{\alpha}}{\partial x^{\alpha}}$$

C)
$$\pi_a = Pa + Aa$$

$$= \left[(x_a | P_b) = \delta ab \right]$$

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$$d-\int \sum x_{a1}H^{3} = \hat{x}_{a} = \frac{\rho_{0}}{m} =$$

$$= \sum x_{a1}\frac{\rho^{3}}{2m} + V(x) = \sum x_{a1}(\frac{\pi - A}{2m})^{2} + V(x)$$

$$H = (\frac{\pi - A}{2m})^{2} + V(x)$$

$$e-\int \hat{\rho}_{a1} = (p_{a1}H)^{2} = \sum p_{a1}\frac{\rho_{b}\rho^{b}}{2m} + V(x) = \frac{1}{m} \int_{a}^{a} \int_{a}^{b} \frac{\rho^{b} - \frac{\partial V}{\partial x_{0}}}{\partial x_{0}}$$

$$\hat{\rho}_{a2} = \int_{a1}^{a1} \frac{1}{n} \int_{a1}^{a1} \frac{\rho_{b}\rho^{b}}{2m} + V(x) = \frac{1}{m} \int_{a1}^{a1} \frac{\rho^{b}\rho^{b} - \frac{\partial V}{\partial x_{0}}}{\partial x_{0}}$$

$$\hat{\rho}_{a2} = \int_{a1}^{a1} \frac{1}{n} \int_{a1}^{a1} \frac{\rho^{b}\rho^{b}}{2m} + V(x) = \frac{1}{m} \int_{a1}^{a1} \frac{\rho^{b}\rho^{b}}{2m} + V(x) = \frac{1}{m}$$

On the other hard this is not general case

$$\frac{dx}{dt} = \frac{\lambda px}{m}$$

$$= \int \frac{dx}{dt} = \frac{px}{m}$$

$$= \int \frac{dx}{dt} = \frac{px}{m} = \int \frac{x}{m} (T - T_0) + \frac{x}{m}$$

$$= \int \frac{dx}{dt} = \frac{x}{m}$$

That is me have recovered the Neutonian mechanics using that by going back to the inertial time T

$$S = \int dt \left(\frac{1}{2} m q^{2} \right)$$

$$S = \int dt \left(\frac{1}{2} m q^{2} \right)$$

$$Q = A \implies Q = A \implies Q$$

$$q(t) = At + B$$

$$q_1 - q_0 = A(t_1 - t_0)$$

$$A = \left(\frac{q_1 - q_0}{t_1 - t_0}\right)$$

$$B = q_0 t_1 - t_0 q_1$$

$$t_1 - t_0$$

$$t = t_1 \cdot q = q_1 \implies q_1 = At_1 + t_2$$

$$t = t_1 \cdot q = q_1 \implies q_1 = At_1 + t_2$$

$$t = t_1 \cdot q = q_1 \implies q_1 = At_1 + t_2$$

$$= \int dt \left(\frac{q_1 - q_0}{t_1 - t_0}\right)^2 \frac{M}{2}$$

$$= \left(t_1 - t_0\right) \frac{M}{2} \left(\frac{q_1 - q_0}{t_1 - t_0}\right)^2$$

6)
$$\frac{0.5}{0.91} = P_1 = M \frac{(9_1 - 9_0)}{t_1 - t_0} = P$$

$$\frac{0.5}{0.90} = -\frac{1}{0.00} =$$

$$\frac{0.5}{0.90} = -P_0 = -\frac{m(9_1 - 9_0)}{E_1 - E_2} = -\frac{1}{2} \frac{m(9_1 - 9_0)^2}{(E_1 - E_2)^2} = E = \frac{1}{2} \frac{m(9_1 - 9_0)^2}{(E_1 - E_2)^2}$$

S(901911/2)
$$MU\left(\frac{(9^2+9^2)\cos[\omega(t_1-t_2)]-29091}{2\sin[\omega(t_1+t_2)]}\right)$$

$$A = \frac{90 \sin(\omega t_1) - 91 \sin(\omega t_2)}{\sin(\omega t_1)\cos(\omega t_2) - \sin(\omega t_2)\cos(\omega t_2)}$$

$$R = \frac{900}{\sin(\omega t_1)\cos(\omega t_2) - \sin(\omega t_3)\cos(\omega t_3)}$$

$$R = \frac{900\cos(\omega t_1) + 91\cos(\omega t_3)}{\sin(\omega t_1)\cos(\omega t_3) - \sin(\omega t_3)\cos(\omega t_3)}$$

 $\frac{dX}{dt} = \lambda \frac{Px}{m} = \lambda \frac{Px}{m} = \lambda \frac{Px}{m} (t-6s) + \lambda s : \text{ Newtonian equation (almost)}$

6.1.

4.)
$$\frac{\partial S}{\partial t} + f(q^{2}, \frac{\partial S}{\partial q}, t) = 0$$

0.) $\frac{1}{\sqrt{2}} = \frac{1}{2} m(i^{2} + i^{2} +$