



8-P6SP18IRQ6JOSHC



9864208

FÖRSÄTTSLAD TENTAMEN/ EXAMINATION COVER

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IFYLLES AV STUDENT OCH TENTAMENSVAKT/

TO BE FILLED IN BY THE STUDENT AND THE INVIGILATOR:

KURSKOD / COURSE CODE				EFTERNAMN / FAMILY NAME															
A G 1 8 1 7				OLIVIER															
KURSNAMN / COURSE NAME				FÖRNAMN / FIRST NAME															
Kartprojektioner och referenssystem				NILS															
PROVKOD / TEST CODE				NAMNTECKNING / YOUR SIGNATURE															
T E N A																			
TENTAMENSdatum / EXAMINATION DATE				PERSONNUMMER / PERSONAL NUMBER															
Y/Y/Y/Y M/M D/D				Y/Y/Y/Y/M/M/D/D															
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PROGRAMKOD / PROGRAM CODE:				INLÄMNINGSTID / TIME SUBMITTED:				SIGNATUR TENTAMENSVAKT / SIGNATURE INVIGILATOR:				ANTAL BLAD / NO OF SHEETS:							
GIT-3				9.15								0 4							
MARKERA BEHANDLADE UPPGIFTER MED "X" OCH EJ BEHANDLADE UPPGIFTER MED "-" / MARK WITH "X" PROBLEMS SOLVED. MARK WITH "-" PROBLEMS NOT ATTEMPTED																			
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IFYLLES AV INSTITUTIONEN / TO BE FILLED IN BY THE DEPARTMENT:

BEDÖMNING / ASSESSMENT																			
1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20

BONUSPOÄNG/
BONUS POINTS:

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SLUTSUMMA /
FINAL POINTS:

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BETYG/
GRADE:

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9864208

Godkänns av examinator /
approved by Examiner.....

Family name, first name

OLIVIER, MILS

Personal Registration Number

20010610-1751

Programme

GIT-3

Sheet no.

01

Problem no.

01

a. $e = \frac{\sqrt{a^2 - b^2}}{a}$, $e' = \frac{\sqrt{a^2 - b^2}}{b}$

$$e^2 = \frac{a^2 - b^2}{a^2} \Leftrightarrow e^2 a^2 = a^2 - b^2 \Leftrightarrow b^2 = a^2 - e^2 a^2$$

$$e'^2 = \frac{a^2 - b^2}{b^2} \Rightarrow e'^2 = \frac{a^2 - (a^2 - e^2 a^2)}{a^2 - e^2 a^2} = \frac{a^2 e^2}{a^2 (1 - e^2)}$$

$$= \frac{e^2}{1 - e^2} \Leftrightarrow e' = \frac{\sqrt{e^2}}{\sqrt{1 - e^2}} = \frac{e}{\sqrt{1 - e^2}}$$

b. geodetic problems are about describing the earth on a reference ellipsoid, such as:

- calculating coordinates and shapes for objects.
- creating reference frames

c. Many triangulation-based reference systems need locally fitted reference ellipsoids. These ellipsoid have better accuracy locally but are not geocentric.

- d) i) no iv) yes
 ii) no v) no
 iii) yes vi) no
 iv)

Notes

not to be scanned!

$$h = \frac{\sqrt{e}}{R}$$

$$h = \frac{\sqrt{g}}{R \cos \varphi}$$

$$\cos \theta' = \frac{f}{\sqrt{e g}}$$

$$\xi = \frac{\sqrt{e g - f^2}}{R^2 \cos \varphi}$$

$$e = \left(\frac{\partial x}{\partial \varphi} \right)^2 + \left(\frac{\partial y}{\partial \varphi} \right)^2$$

$$f = \left(\frac{\partial x}{\partial \varphi} \right) \left(\frac{\partial x}{\partial \lambda} \right) + \left(\frac{\partial y}{\partial \varphi} \right) \left(\frac{\partial y}{\partial \lambda} \right)$$

$$g = \left(\frac{\partial x}{\partial \lambda} \right)^2 + \left(\frac{\partial y}{\partial \lambda} \right)^2$$

$h = \kappa \Rightarrow$ conformal

$h, \kappa = 1 \Rightarrow$ equi-distant / equivalent

2. a) Universal Transversal Mercator.

A global map projection with:

- 60 longitude zones, 6° each
- scale factor: 0.9996
- False northing: 10 000 km if southern hemisphere
- false easting: 500 km
- 20 latitude bands with some variations.

b) a/ $e = \left(\frac{\partial x}{\partial \bar{\phi}}\right)^2 + \left(\frac{\partial y}{\partial \bar{\phi}}\right)^2 = (R \cos \lambda)^2 + (-R \sin \lambda)^2 = R^2 (\cos^2 \lambda + \sin^2 \lambda) = R^2$

$f = \left(\frac{\partial x}{\partial \bar{\phi}}\right) \left(\frac{\partial x}{\partial \lambda}\right) + \left(\frac{\partial y}{\partial \bar{\phi}}\right) \left(\frac{\partial y}{\partial \lambda}\right) = (R \cos \lambda) (R \sin \lambda (\frac{\pi}{2} - \bar{\phi})) + (-R \sin \lambda) (R \cos \lambda (\frac{\pi}{2} - \bar{\phi}))$

$= R^2 \cos \lambda \sin \lambda (\frac{\pi}{2} - \bar{\phi}) - R^2 \sin \lambda \cos \lambda (\frac{\pi}{2} - \bar{\phi}) = 0$

$g = \left(\frac{\partial x}{\partial \lambda}\right)^2 + \left(\frac{\partial y}{\partial \lambda}\right)^2 = (R \sin \lambda (\frac{\pi}{2} - \bar{\phi}))^2 + (R \cos \lambda (\frac{\pi}{2} - \bar{\phi}))^2$

$= R^2 \sin^2 \lambda (\frac{\pi}{2} - \bar{\phi})^2 + R^2 \cos^2 \lambda (\frac{\pi}{2} - \bar{\phi})^2$

$= R^2 (\frac{\pi}{2} - \bar{\phi})^2 (\sin^2 \lambda + \cos^2 \lambda) = R^2 (\frac{\pi}{2} - \bar{\phi})^2$

$$\begin{cases} e = R^2 \\ f = 0 \\ g = R^2 (\frac{\pi}{2} - \bar{\phi})^2 \end{cases}$$

b/ $h = \frac{\sqrt{e}}{R} = \frac{\sqrt{R^2}}{R} = \frac{R}{R} = 1$

$n = \frac{\sqrt{g}}{R \cos \bar{\phi}} = \frac{\sqrt{R^2 (\frac{\pi}{2} - \bar{\phi})^2}}{R \cos \bar{\phi}} = \frac{R (\frac{\pi}{2} - \bar{\phi})}{R \cos \bar{\phi}} = \frac{\frac{\pi}{2} - \bar{\phi}}{\cos \bar{\phi}}$

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Continue 2.

$$c./ \cos \theta' = \frac{f}{\sqrt{eg}} = \frac{0}{\sqrt{eg}} = 0 \Rightarrow \theta' = \arccos(0) = 90^\circ$$

$$d./ \xi = \frac{\sqrt{eg-f^2}}{R^2 \cos \bar{\theta}} = \frac{\sqrt{eg}}{R^2 \cos \bar{\theta}} = \frac{\sqrt{R^2 R^2 (\frac{\pi}{2} - \bar{\theta})^2}}{R^2 \cos \bar{\theta}} = \frac{R^2 (\frac{\pi}{2} - \bar{\theta})}{R^2 \cos \bar{\theta}}$$

$$= \frac{(\frac{\pi}{2} - \bar{\theta})}{\cos \bar{\theta}}$$

$$e. 1 \neq \frac{\pi}{2} - \bar{\theta} \Rightarrow h \neq k \Rightarrow \underline{\text{not conformal}}$$

$$1 \cdot \frac{(\frac{\pi}{2} - \bar{\theta})}{\cos \bar{\theta}} \neq 1 \Rightarrow h \cdot k \neq 1 \Rightarrow \underline{\text{not equivalent}}$$

answer: neither.

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a. Conventional international origin

The mean pole position in 1900-1905.
used as a fixed reference point for terrestrial
reference systems. The (TRS) need a fixed pole to
define the rotation axis and the poles are constantly
moving.

b. J2000.0 is a time epoch of 1 January 1st,
ut1 = 12:00, year 2000 in the Julian calendar.
Since "everything" is changing with time and reference
systems need fixed references. This time epoch
is used to define celestial reference systems.

c.

$(\phi, \lambda, h)_{\text{SWEREF99}} \longrightarrow (x, y, z)_{\text{SWEREF99}}$ GRS 80 ellipsoid

$(x, y, z)_{\text{SWEREF99}} \longrightarrow (x, y, z)_{\text{RT90}}$ Helmert's transformation

$(x, y, z)_{\text{RT90}} \longrightarrow (\phi, \lambda, h)_{\text{RT90}}$ Bessel's 1841
ellipsoid

$(\phi, \lambda, h)_{\text{RT90}} \longrightarrow (x, y)_{\text{RT90}}$ Bessel's 1841
ellipsoid.