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## Contents

Introduction of Fractional Velocity Dispersion Analysis (FVDA)

Research papers based on FVDA

A Python Package of FVDA -- pyFVDA

## FVDA contains two approaches

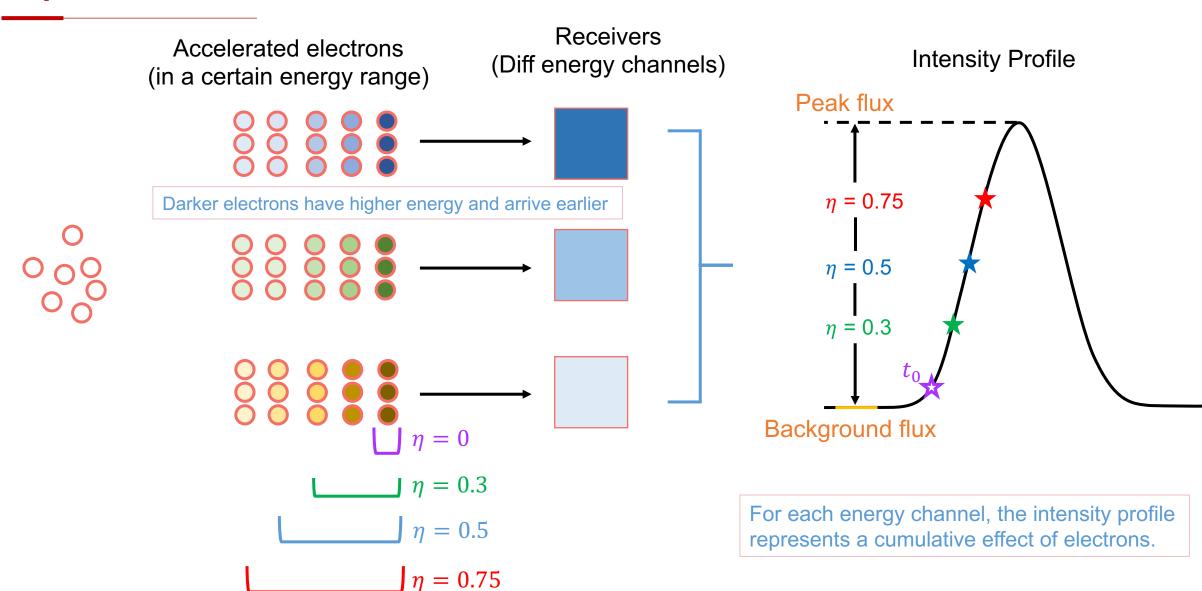
Original paper of the FVDA

Lulu Zhao, Gang Li, Ming Zhang, et al. <u>Statistical Analysis of Interplanetary Magnetic Field Path Lengths from Solar Energetic Electron Events Observed by WIND</u>. The Astrophysical Journal, 878:107 (9pp), 2019 June 20.

Summary of both approaches can be found:

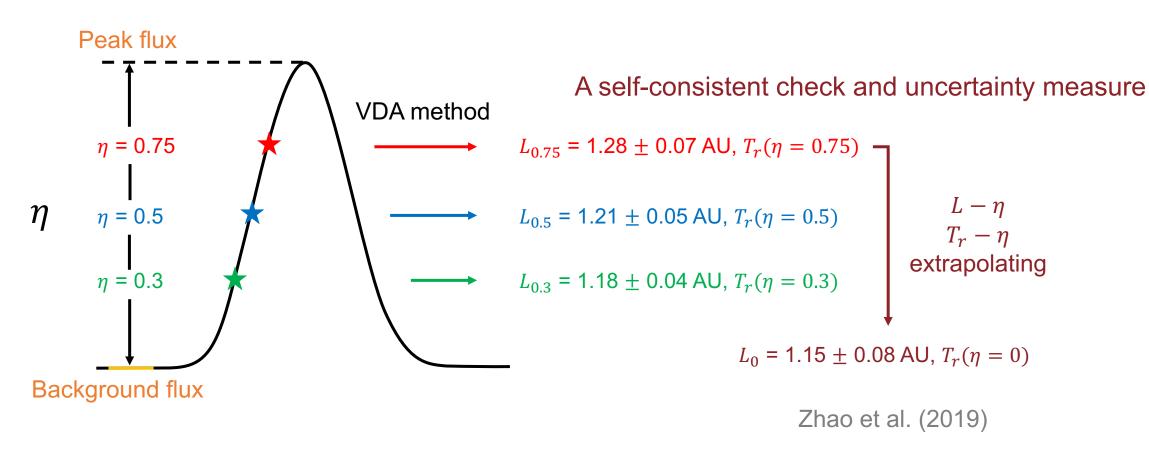
Xiangyu Wu, Gang Li, Lulu Zhao, et al. <u>Statistical Study of Release Time and Its Energy Dependence of In</u> <u>Situ Energetic Electrons in Impulsive Solar Flares</u>. Journal of Geophysical Research: Space Physics, 128, e2022JA030939.

### **Explanation of the idea of FVDA**



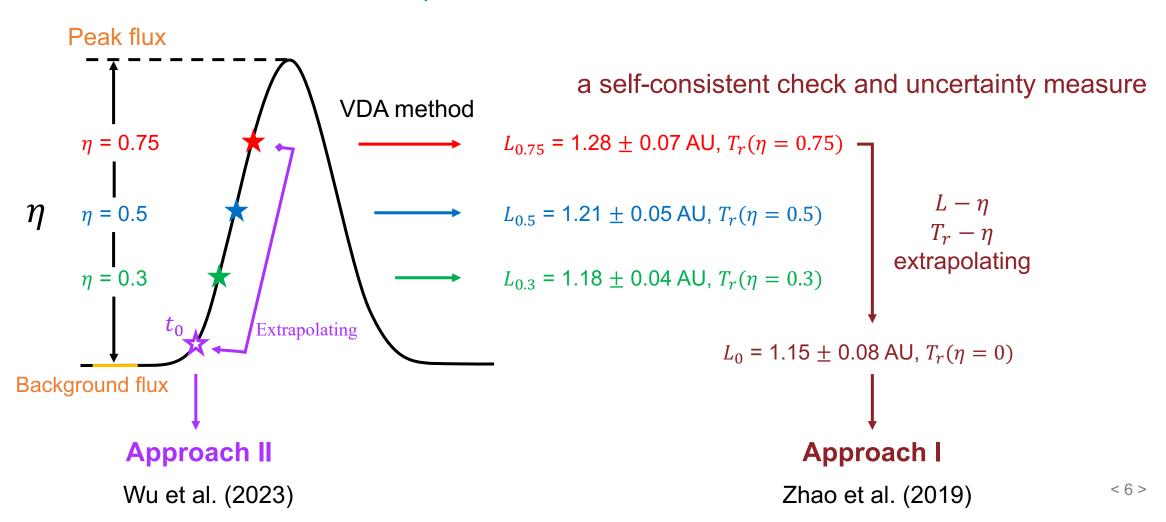
Simultaneous release

- Scatter-free propagation
- Less affected determine onset/peak time, contamination effect

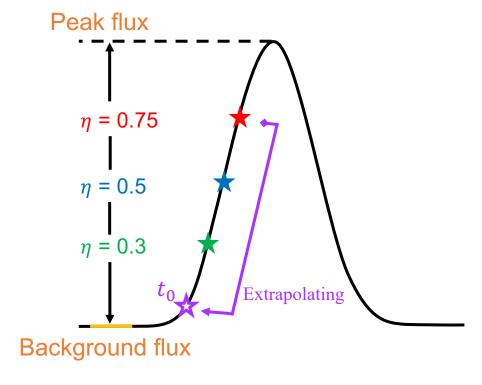


Simultaneous release

- Scatter-free propagation
- Less affected determine onset/peak time, contamination effect



### Approach II



$$t_i(\eta) = a\eta^2 + b\eta + t_i^0(1)$$

$$T_r(E_i) = t_i^0 - \frac{L}{v_i}$$
 (2)

For each energy channel:

- (1) Obtain onset time  $t_i^0$  using Eq.(1)
- (2) Compute the release time  $T_r$  from Eq.(2)

### **Advantages of Approach II**

- Avoid contamination effect, less uncertainty than VDA
- More accurately determine the onset time → Energy-dependent Tr of electrons

## Comparison of VDA and FVDA

#### **VDA**

$$T_r = t_i^0 - \frac{L}{v_i}$$

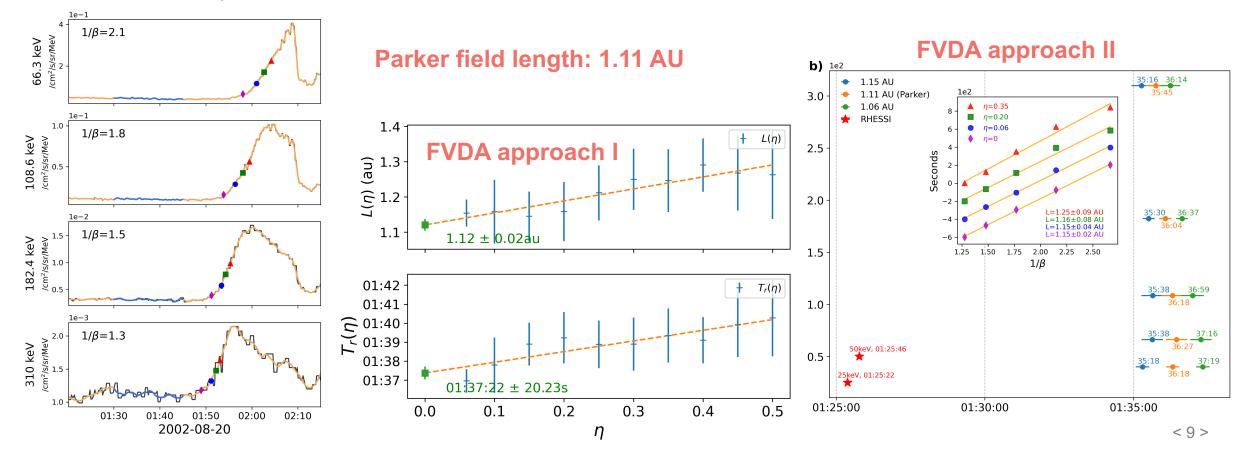
- Non-physical path length
  - (1) Energy-dependent Tr of SEE
  - (2) Incorrect determination of onset

#### **FVDA**

- Two approaches
- Self-consistent check
  - Further analysis on the cause

#### **Self-consistent check**

- For the cases with simultaneous release,  $T_r$  from **Approach I** and **II** should be close.
- An example event 2002-08-20

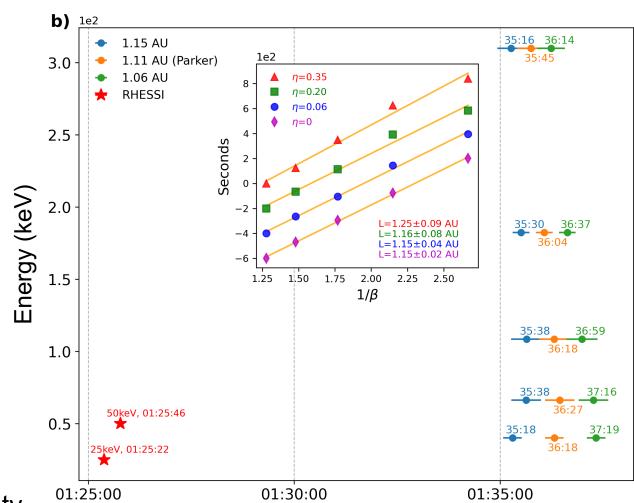


#### **Self-consistent check**

- Parker field length: 1.11 au
- Results from FVDA approach I
  - $1.12 \pm 0.02$  au
  - $01:37:22 \pm 20 \text{ s}$
- Results from FVDA approach II

$$T_r(E_i) = t_i^0 - \frac{L}{v_i}$$
,  $L = [1.06, 1.15]$  au

- 01:35 ~ 01:37
- Max dt < 2 mins, within the uncertainty.</li>



### **Summary**

- An improved method Release time & Path length of SEE
- Two approaches
- If electrons (different energy) are released simultaneously
  - I & II yield similar Tr
- If not
  - Path length from Approach I could be non-physical
  - Approach II should be used to obtain Tr of electrons

### 2 - Research based on FVDA

- Identification of two electron populations in an impulsive Solar Energetic Electron (SEE)
  event. G. Li et al 2020 ApJL 900 L16.
- Constraints on the electron acceleration process in solar flare. G. Li et al, Geophysical Research Letters, 48, e2021GL095138.
- Statistical study of release time and its energy dependence of energetic electrons in SEE events. X. Wu et al, Journal of Geophysical Research: Space Physics, 128, e2022JA030939.

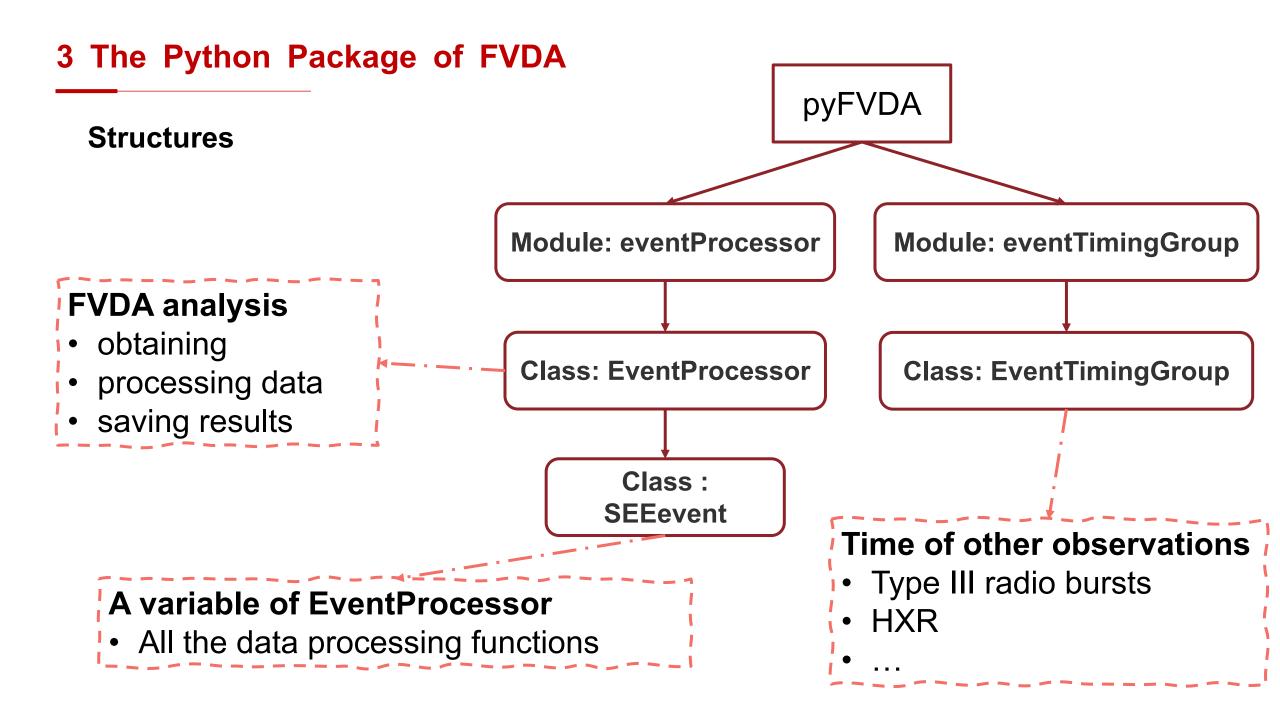
### 3 The Python Package of FVDA

### **Objectives & Features**

- Simple & Fast calculate the Tr & Path length Help analyze SEE events
- Include Approach I & II
- Obtaining data from the server, interactively processing data, and saving results

#### **Updates**

- An initial version is now available on Github
  - Approach I & II are included
  - Data from WIND & STEREO



## 3 The Python Package of FVDA

### **Next steps**

- Will be available on pypi soon
- Will include SolO

### **Github Page**

https://github.com/Xiangyu-W/pyFVDA