

# INTERFERENCE BY DIVISION OF AMPLITUDE

## Introduction

In the last section we dealt with interference produced by division of a wavefront. Here we are going to deal with the formation of interference by division of amplitude. If a plane wave falls on thin films of transparent material like soap bubbles or of a drop of oil spread on the surface of water, the wave reflected from the upper surface and that reflected from the lower surface get superposed to get interference pattern. These studies have many practical applications and also explain phenomenon of brilliant colours produced by a soap film exposed to white light (sunlight).

## Interference by a plane film

When a plane wave is incident on a thin film of uniform thickness (like a film of oil spread on water surface, or soap film), then the waves reflected from the upper surface interfere with the waves reflected from the lower surface. Depending upon the path difference between the two interfering waves, brightness or darkness can be seen.

**Note :** A film of thickness in the range  $0.5\mu\text{m}$  to  $10\mu\text{m}$  is considered as a thin film.

## Normal incidence of light on a plane film

Consider a thin transparent film of uniform thickness  $t$ , having refractive index  $\mu$  is kept in air. A plane wave from a monochromatic light of wavelength  $\lambda$  is falling normally on the film after reflection from a mirror  $G$  placed at an angle  $45^\circ$  with respect to the film. Then the waves reflected from the upper surface interfere with the waves reflected from the lower surface.

Clearly the wave reflected from the lower surface travels an additional optical path of  $2\mu t$  to join the wave reflected from the upper surface. Since the film is placed in air, then the wave reflected from the upper surface of the film will undergo a sudden change in phase of  $\pi$ , which is equivalent to a path difference of  $\frac{\lambda}{2}$ .

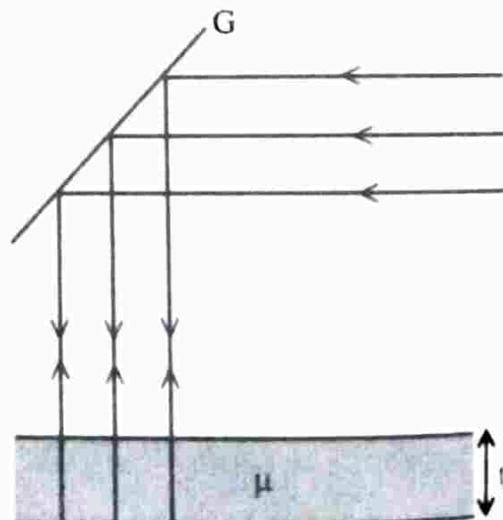


Figure 3.1

So the effective path difference between the two rays reflected from upper and lower surfaces of the film is

$$\Delta = 2\mu t - \frac{\lambda}{2}$$

If this path difference is an integral multiple of  $\lambda$ , the film appears to be bright. So the condition for brightness becomes

$$2\mu t - \frac{\lambda}{2} = n\lambda, \text{ where } n = 0, 1, 2, \dots$$

or  $2\mu t = \left(n + \frac{1}{2}\right)\lambda$

If the path difference is an odd integral multiple of  $\frac{\lambda}{2}$ , the film appears to be dark so the condition for darkness becomes

$$2\mu t - \frac{\lambda}{2} = (2n - 1)\frac{\lambda}{2}$$

or  $2\mu t = n\lambda \text{ where } n = 1, 2, 3, \dots$

### Oblique incidence of light on a plane film

Consider a thin transparent film of uniform thickness  $t$ , having refractive index  $\mu$  placed in air. A plane wave from a monochromatic light of wavelength  $\lambda$  is falling at an angle of incidence  $i$ . Part of a ray AB is reflected along BC, and part of it is transmitted into the film along BF. The transmitted ray BF makes an angle  $r$  with the normal to the surface at point G. The ray BF is in turn partially reflected back into the film along FD while a major part refracts into air along FK. Part of the reflected ray FD is transmitted at the upper surface and travels along DE. Since the surfaces of the film are parallel, the reflected rays BC and DE will be parallel to each other. The waves travelling along the paths BC and BFDE are derived from a single incident wave AB. Therefore they are coherent and can produce interference when they are made to superimpose.

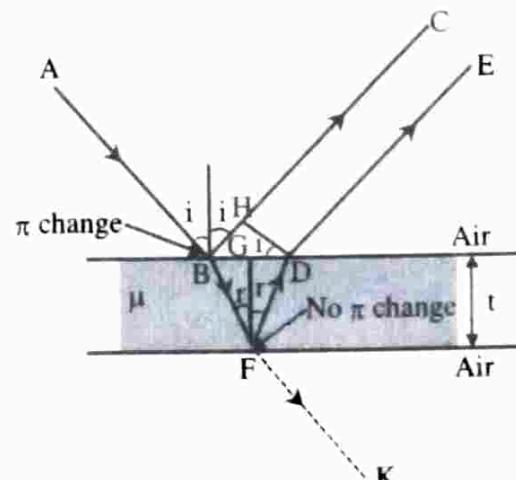


Figure 3.2

Let DH be the normal to BC. From H and D onwards, the rays HC and DE travel equal path.

∴ The optical path difference between the two reflected rays,  $\Delta = \mu(BF + FD) - BH$

$$\text{From the } \Delta BFG, \quad \frac{FG}{BF} = \cos r$$

$$\text{or} \quad BF = \frac{FG}{\cos r} = \frac{t}{\cos r}$$

$$\therefore BF + FD = 2BF \quad (\because BF = FD)$$

$$\text{or} \quad BF + FD = \frac{2t}{\cos r} \quad \dots\dots(2)$$

$$\text{From the } \Delta BHD, \frac{BH}{BD} = \sin i$$

$$\text{or} \quad BH = BD \sin i = 2BG \sin i \quad (\because BD = 2BG)$$

$$\text{From the } \Delta BFG, \frac{BG}{BF} = \sin r$$

$$\text{or} \quad BG = BF \sin r = \frac{t}{\cos r} \sin r = t \tan r$$

$$\therefore BH = 2t \tan r \sin i$$

$$\text{Using } \frac{\sin i}{\sin r} = \mu, BH = 2t \tan r \mu \sin r$$

$$BH = 2\mu t \frac{\sin^2 r}{\cos r} \quad \dots\dots(3)$$

Using equations 2 and 3 in equation (1), we get

$$\Delta = \mu \frac{2t}{\cos r} - \frac{2\mu t \sin^2 r}{\cos r}$$

$$\Delta = \frac{2\mu t}{\cos r} (1 - \sin^2 r) = 2\mu t \cos r$$

Since the film is placed in air, then the wave reflected from the upper surface of the film will undergo a sudden change in phase of  $\pi$ , which is equivalent to a path difference of  $\frac{\lambda}{2}$ . There is no path difference due to transmission at D. So the effective path difference between the two rays from upper and lower surfaces of the film is

$$\Delta = 2\mu t \cos r - \frac{\lambda}{2}$$

If this path difference is an integral multiple of  $\lambda$ , the film appears to be bright, so the condition for brightness becomes

$$2\mu t \cos r - \frac{\lambda}{2} = n\lambda$$

$$2\mu t \cos r = \left(n + \frac{1}{2}\right)\lambda \text{ where } n = 0, 1, 2, 3, \dots$$

If the path difference is an odd integral multiple of  $\frac{\lambda}{2}$ , the film appears to be dark. So the condition for darkness becomes

$$2\mu t \cos r - \frac{\lambda}{2} = (2n-1)\frac{\lambda}{2}$$

$$2\mu t \cos r = n\lambda \text{ where } n = 1, 2, 3, \dots$$

## The Cosine law

The effective path difference between the wave reflected from the lower surface and the upper surface of the film is given by

$$\Delta = 2\mu t \cos r \quad \dots\dots(4)$$

This is known as the cosine law.

### Example 1

A soap film is illuminated by white light incident at angle of  $30^\circ$ . The reflected light is examined by a spectroscope in which dark band corresponding to the wavelength  $6 \times 10^{-7} \text{ m}$  is found. Calculate the smallest thickness of the film.  $\mu = 1.33$

### Solution

$$i = 30^\circ, \lambda = 6 \times 10^{-7} \text{ m} \text{ and } \mu = 1.33$$

Dark band for reflected light is seen when

$$2\mu t \cos r = n\lambda$$

For minimum thickness  $n = 1$

$$\therefore 2\mu t \cos r = \lambda$$

$$t = \frac{\lambda}{2\mu \cos r} = \frac{6 \times 10^{-7}}{2 \times 1.33 \times \cos r}$$

$$\text{But } \cos r = \sqrt{1 - \sin^2 r} = \sqrt{1 - \frac{\sin i^2}{\mu^2}} \quad \left( \because \frac{\sin i}{\sin r} = \mu \right)$$

$$= \sqrt{1 - \frac{\sin^2 30}{1.33^2}} = 0.927$$

$$\therefore t = \frac{6 \times 10^{-5}}{2 \times 1.33 \times 0.927} = 2.433 \times 10^{-7} \text{ m.}$$

### Example 2

A soap film  $5 \times 10^{-5} \text{ cm}$  thick is viewed at an angle of  $35^\circ$  to the normal. Find the wavelength of light in the visible spectrum which will be absent from the reflected light  $\mu = 1.33$ .

### Solution

$$t = 5 \times 10^{-5} \text{ cm} = 5 \times 10^{-7} \text{ m}, \quad i = 35^\circ, \text{ and } \mu = 1.33$$

using  $\frac{\sin i}{\sin r} = \mu$  we have  $\sin r = \frac{\sin i}{\mu} = \frac{\sin 35}{1.33}$   
 $\sin r = 0.431$

$$\therefore \cos r = \sqrt{1 - \sin^2 r} = \sqrt{1 - (0.431)^2} = 0.90$$

The condition for destructive interference is

$$2\mu t \cos r = n\lambda$$

$$\therefore \lambda = \frac{2\mu t \cos r}{n}$$

$$\text{for } n = 1 \quad \lambda_1 = 2\mu t \cos r = 2 \times 1.33 \times 5 \times 10^{-7} \times 0.90$$

$$\lambda_1 = 11.97 \times 10^{-7} \text{ m}$$

Similarly

$$\text{for } n = 2, \quad \lambda_2 = 5.985 \times 10^{-7} \text{ m}$$

$$\text{for } n = 3, \quad \lambda_3 = 3.99 \times 10^{-7} \text{ m}$$

$$\text{for } n = 4, \quad \lambda_4 = 2.99 \times 10^{-7} \text{ m.}$$

$\lambda_2$  and  $\lambda_3$  lie in the visible region, therefore the two wave lengths are absent in the reflected light.

### Example 3

A parallel beam of sodium light ( $\lambda = 5890 \text{ \AA}$ ) strikes a film of oil floating on water. When viewed at angle of  $30^\circ$  from the normal  $8^{\text{th}}$  dark band is seen. Determine the thickness of the film. ( $\mu$  of the oil = 1.5).

**Solution**

$$\lambda = 5890 \text{ Å} = 5890 \times 10^{-10} \text{ m}$$

$$i = 30^\circ, \mu = 1.5, n = 8$$

using

$$\mu = \frac{\sin i}{\sin r}$$

$$\sin r = \frac{\sin i}{\mu} = \frac{\sin 30}{1.5} = \frac{1}{2 \times \frac{3}{2}} = \frac{1}{3}$$

$$\begin{aligned}\cos r &= \sqrt{1 - \sin^2 r} = \sqrt{1 - \frac{1}{9}} = \sqrt{\frac{8}{9}} \\ &= \frac{2\sqrt{2}}{3}\end{aligned}$$

For the dark band, we have

$$2\mu t \cos r = n\lambda$$

$$t = \frac{n\lambda}{2\mu \cos r} = \frac{8 \times 5890 \times 10^{-10}}{2 \times \frac{3}{2} \cdot \frac{2\sqrt{2}}{3}}$$

$$= 1.67 \times 10^{-4} \text{ mm}$$

**Example 4**

When white light is incident on a soap film at an angle of  $\sin^{-1}(4/5)$  and the reflected light on examination by a spectroscope shows dark bands. Two consecutive dark bands correspond to wavelengths  $6.1 \times 10^{-7} \text{ m}$  and  $6 \times 10^{-7} \text{ m}$ . If  $\mu$  of the film be  $\frac{4}{3}$ , calculate its thickness.

**Solution**

The path difference for dark band in reflected light is given by

$$2\mu t \cos r = n\lambda$$

$$\text{For } \lambda_1 \text{ we have } 2\mu t \cos r = n\lambda_1 \quad \dots\dots(1)$$

$$\text{For } \lambda_2 \text{ we have } 2\mu t \cos r = (n+1)\lambda_2 \quad \dots\dots(2)$$

Comparing eqns (1) and (2), we have

$$n\lambda_1 = (n+1)\lambda_2$$

or

$$n = \frac{\lambda_2}{\lambda_1 - \lambda_2} = \frac{6 \times 10^{-7}}{6.1 \times 10^{-7} - 6 \times 10^{-7}} = 60$$

Putting the value of  $n$  in eqn (1), we get

$$2\mu t \cos r = 60 \times 6.1 \times 10^{-7}$$

$$\therefore t = \frac{60 \times 6.1 \times 10^{-7}}{2\mu \cos r} = \frac{60 \times 6.1 \times 10^{-7}}{2 \times \frac{4}{3} \times \cos r}$$

But  $i = \sin^{-1} \frac{4}{5}$  given

or  $\sin i = \frac{4}{5}$

using  $\mu = \frac{\sin i}{\sin r}$

or  $\sin r = \frac{\sin i}{\mu} = \frac{4}{5 \times \frac{4}{3}} = \frac{3}{5}$

$\therefore \cos r = \sqrt{1 - \sin^2 r} = \sqrt{1 - \frac{9}{25}} = \frac{4}{5}$

Then  $t = \frac{60 \times 6.1 \times 10^{-7}}{2 \times \frac{4}{3} \times \frac{4}{5}} = 1.716 \times 10^{-5} \text{ m}$

### Reflection and transmission coefficient

For simplicity we assume that light (electromagnetic wave) is a plane wave travelling in the  $z$ -direction. The field vectors  $\vec{E}_1$  and  $\vec{B}_1$  characterise the incident wave travelling in the positive direction,  $\vec{E}_R$  and  $\vec{B}_R$  the reflected wave travelling in the  $-z$  direction,  $\vec{E}_T$  and  $\vec{B}_T$  the wave transmitted into the second medium of refractive index  $\mu_2$ . The interface between the dielectric media is taken at  $z = 0$ .

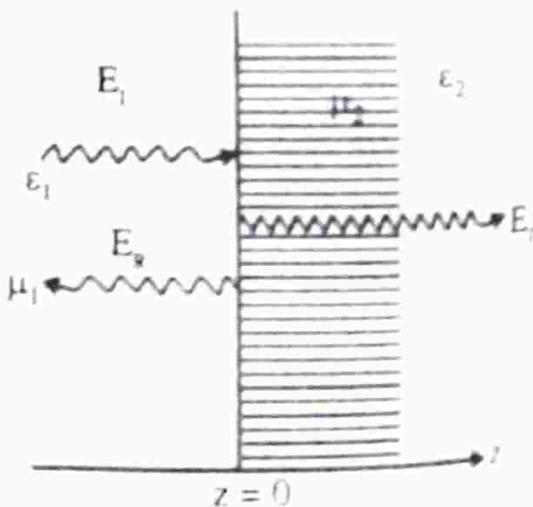


Figure 3.3

The incident, reflected and transmitted waves are given by

$$\text{Incident} \quad \vec{E}_I = \hat{i} E_{Io} e^{i(k_1 z - \omega t)} \quad \dots\dots(5)$$

where  $\hat{i}$  is the direction of the electric field vector,  $E_{Io}$  is the amplitude of the field and  $k_1$  is wave vector in medium 1.

$$\vec{B}_I = \hat{j} B_{Io} e^{i(k_1 z - \omega t)} \quad \dots\dots(6)$$

$\hat{j}$  is the direction of the magnetic field vector and  $B_{Io}$  is the amplitude of the wave.

$$\text{Reflected} \quad \vec{E}_R = \hat{i} E_{Ro} e^{-i(k_1 z + \omega t)} \quad \dots\dots(7)$$

$$\text{and} \quad \vec{B}_R = -\hat{j} B_{Ro} e^{-i(k_1 z + \omega t)} \quad \dots\dots(8)$$

$$\text{Transmitted} \quad \vec{E}_T = \hat{i} E_{To} e^{i(k_2 z - \omega t)} \quad \dots\dots(9)$$

$$\text{and} \quad \vec{B}_T = \hat{j} B_{To} e^{i(k_2 z - \omega t)} \quad \dots\dots(10)$$

where  $E_{To}$  and  $B_{To}$  are the amplitudes of the field vectors and  $k_2$  is the wave vector in the second medium.

The boundary conditions on  $\vec{E}$  and  $\vec{B}$  at  $z = 0$  are the continuity of their tangent components

$$E_{It} + E_{Rt} = E_{Tt} \quad (\text{From eqns 5, 6, 7, 8, 9, 10})$$

$$\text{or} \quad E_{Io} + E_{Ro} = E_{To} \quad \dots\dots(11)$$

$$\text{and} \quad \frac{B_{Io}}{\mu_1} - \frac{B_{Ro}}{\mu_1} = \frac{B_{To}}{\mu_2} \quad \dots\dots(12)$$

$$\text{Using } B_{Io} = \frac{E_{Io}}{v_1}, B_{Ro} = \frac{E_{Ro}}{v_1} \text{ and } B_{To} = \frac{E_{To}}{v_2}$$

$\therefore$  Eqn 12 becomes

$$\sqrt{\mu_1 \epsilon_1} \frac{E_{Io}}{\mu_1} - \sqrt{\mu_1 \epsilon_1} \frac{E_{Ro}}{\mu_1} = \frac{\sqrt{\mu_2 \epsilon_2}}{\mu_2} E_{To}$$

we have used  $v = \frac{1}{\sqrt{\mu\epsilon}}$  where  $\mu$  is the permeability and  $\epsilon$  is the permittivity

$$\text{or } \sqrt{\frac{\epsilon_1}{\mu_1}}(E_{lo} - E_{Ro}) = \sqrt{\frac{\epsilon_2}{\mu_2}} E_{To} \quad \dots\dots(13)$$

solving eqns 11 and 13 we get

$$E_{Ro} = \frac{\sqrt{\frac{\epsilon_1}{\mu_1}} - \sqrt{\frac{\epsilon_2}{\mu_2}}}{\sqrt{\frac{\epsilon_1}{\mu_1}} + \sqrt{\frac{\epsilon_2}{\mu_2}}} E_{lo} \quad \dots\dots(14)$$

and

$$E_{To} = \frac{2\sqrt{\frac{\epsilon_1}{\mu_1}}}{\sqrt{\frac{\epsilon_1}{\mu_1}} + \sqrt{\frac{\epsilon_2}{\mu_2}}} E_{lo} \quad \dots\dots(15)$$

$$\text{The refractive index } n = \frac{c}{v} = \sqrt{\frac{\mu\epsilon}{\mu_0\epsilon_0}}$$

Here  $\mu_1 = \mu_2 = \mu_0$  for both the medium.

$$\therefore n = \frac{\epsilon}{\epsilon_0} \quad \therefore n_1 \propto \sqrt{\epsilon_1} \quad \text{and} \quad n_2 \propto \sqrt{\epsilon_2}$$

Re-writing eqns 14 and 15 in terms of  $n_1$  and  $n_2$  we get

$$E_{Ro} = \frac{n_1 - n_2}{n_1 + n_2} E_{lo} \quad \dots\dots(16)$$

when  $n_2 > n_1$ ,  $E_{Ro}$  is negative showing that when a reflection occurs at a denser medium a phase change of  $\pi$  occurs.

$$E_{To} = \frac{2n_1}{n_1 + n_2} E_{lo} \quad \dots\dots(17)$$

$\frac{E_{Ro}}{E_{lo}}$  is called the amplitude reflection coefficient  $r$  and  $\frac{E_{To}}{E_{lo}}$  is called the amplitude transmission coefficient.

i.e.

$$r = \frac{n_1 - n_2}{n_1 + n_2} \quad \dots\dots(18)$$

$$t' = \frac{2n_1}{n_1 + n_2} \quad \dots\dots(19)$$

**Note :** We have been using the symbol  $\mu$  for refractive index. Here only  $n$  is used not to confuse with permeability  $\mu$ .

### Non-reflecting films

Here we will see how to reduce the reflectivity of thin film surfaces. The technique of reducing reflectivity is known as blooming. Consider a light beam propagating in a medium of refractive index  $\mu_1$  is incident normally on a dielectric of refractive index  $\mu_2$ . Let  $r$  and  $t$  be the amplitude reflection and transmission coefficients.

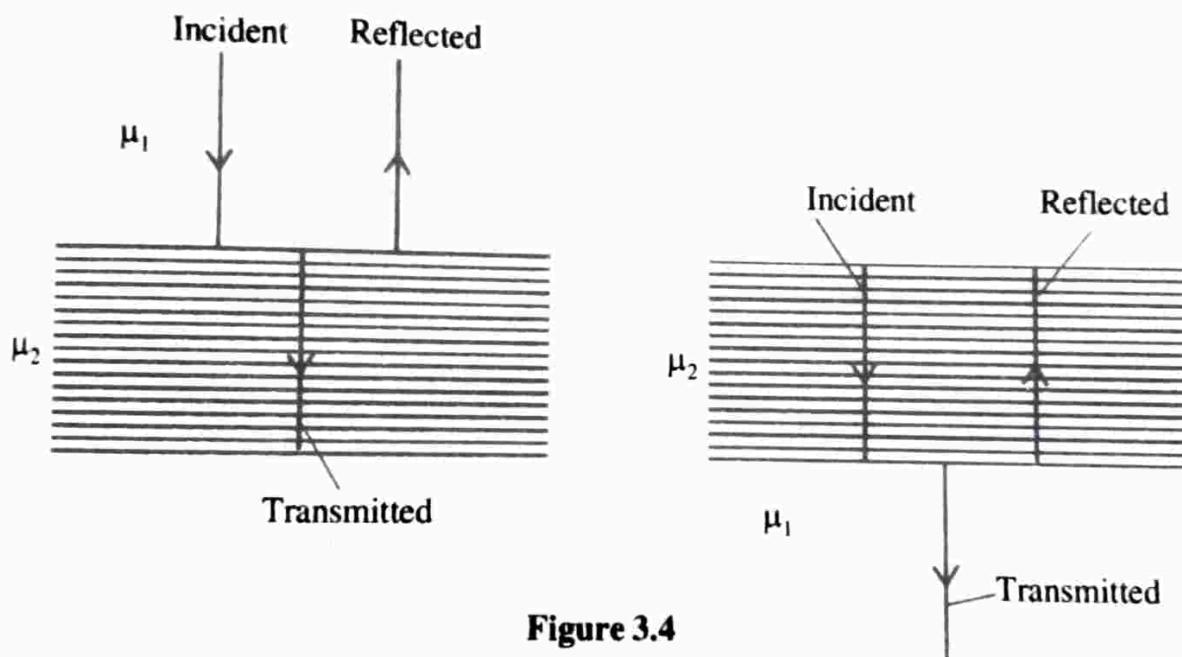


Figure 3.4

Then

$$r = \frac{\mu_1 - \mu_2}{\mu_1 + \mu_2} \quad \dots\dots(20)$$

and

$$t = \frac{2\mu_1}{\mu_1 + \mu_2} \quad \dots\dots(21)$$

If the light beam propagating in a medium of refractive index  $\mu_2$  is incident on a medium of refractive index  $\mu_1$ , then the amplitude reflection coefficient  $r'$  and amplitude transmission coefficient  $t'$  are given by

$$r' = \frac{\mu_2 - \mu_1}{\mu_2 + \mu_1} \quad \dots\dots(22)$$

and

$$t' = \frac{2\mu_2}{\mu_2 + \mu_1} \quad \dots\dots(23)$$

From eqn 22, we get

$$r' = -\frac{(\mu_1 - \mu_2)}{(\mu_1 + \mu_2)} = -r \quad \dots\dots(24)$$

We know that the intensity of light is directly proportional to the square of the amplitude. Hence the square of the amplitude reflection coefficient gives the fraction of intensity of light get reflected. It is called reflection coefficient or reflectivity and is denoted by R.

i.e.

$$r^2 = R = \left( \frac{\mu_1 - \mu_2}{\mu_1 + \mu_2} \right)^2 \quad \dots\dots(25)$$

### Reduction of reflectivity of lens surfaces

As an application of the thin film interference phenomenon, here we will see how to reduce the reflectivity of lens surfaces. We know that in optional instruments there are many lens surfaces. Each surface reflects light and the loss of intensity of light will be large. For example the reflectivity for normal incidence is given by

$$R = \left( \frac{\mu_1 - \mu_2}{\mu_1 + \mu_2} \right)^2$$

Suppose light falls on a crown glass surface from air then  $\mu_1 = 1$  and  $\mu_2 = 1.5$

$$\therefore R = \left( \frac{1-1.5}{1+1.5} \right)^2 = 0.04$$

It shows 4% light is reflected from a crown glass from each surface.

If the glass is flint  $\mu_2 = 1.67$

Then

$$R = \left( \frac{1-1.67}{1+1.67} \right)^2 = 0.06$$

i.e. 6% light is reflected from a flint glass from each surface. If there are large number of surfaces the loss will be severe. In order to reduce these losses lens surfaces are coated with a thick non-reflecting film whose refractive index is less than that of glass. Suppose a thin film of a transparent film is coated on the surface of a glass. The material of the thin film is so selected that its refractive index is equal to the

square root of the product of refractive indices of air and of glass. If  $\mu_f$  is the refractive index of film and  $\mu_g$  and  $\mu_a$  are respectively the refractive indices of glass and air then (see example 5)

$$\mu_f \approx \sqrt{\mu_f \mu_a}$$

since  $\mu_a = 1$

$$\mu_f = \sqrt{\mu_g} \quad \dots\dots(26)$$

For crown glass  $\mu_g = 1.5$   $\therefore \mu_f = \sqrt{1.5} = 1.22$

For flint glass  $\mu_g = 1.67$   $\therefore \mu_f = \sqrt{1.67} = 1.29$

Such a value of  $\mu_f$  results into reflection of equal quantities of light from the outer surface of the film and the boundary surface between the film and the glass. Since in both reflections the light travels from a rarer medium to a denser medium, then same phase change occurs in each reflections. If the thickness of the coating is

$\frac{\lambda}{4\mu_f}$  (or  $\frac{\lambda'}{4}$  where  $\lambda'$  is the wave length of light in the film) then the light reflected

from these two surfaces (upper and lower) will have a path dif-

ference of  $\frac{\lambda'}{2}$  or phase differ-  
ence  $\pi$ . Then the conditions for  
complete destructive inter-  
ference takes place.

It should be remembered that  
the thickness of the film can be  
one quarter wavelength for one  
particular wavelength only. This  
is usually chosen in the yellow-  
green portion ( $5500\text{\AA}$ ) of the  
spectrum, because our eye is  
most sensitive for this region of

spectrum. Thus when incident light is white some reflections take place for longer (red) and shorter (violet) wave lengths and the reflected light has a purple blue.

This method has been found useful in reducing reflection from a lens or a prism.  
such a coating is therefore termed as non-reflecting film and the technique em-  
**ployed is called blooming.**

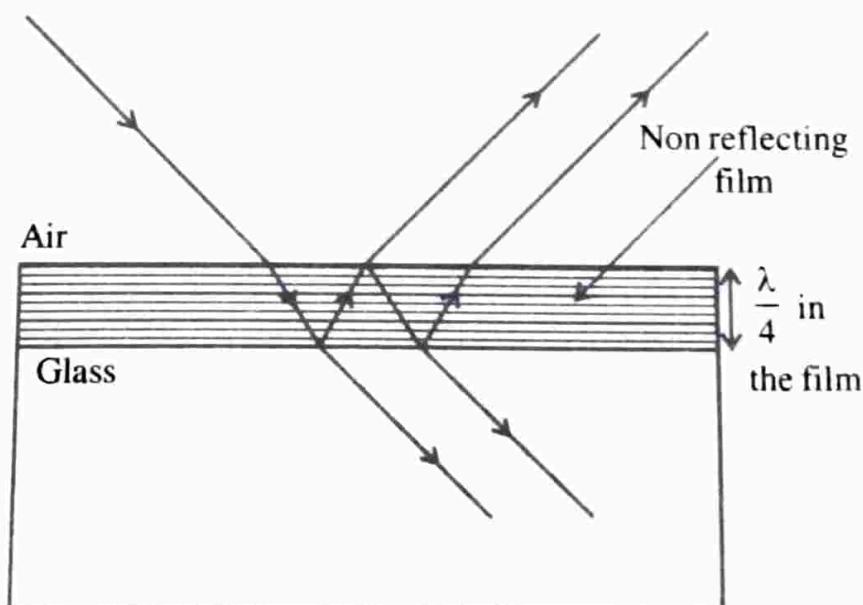


Figure 3.5

**Note :** The best materials suitable for using as non-reflecting films are magnesium fluoride ( $MgF_2$ ) with refractive index 1.38 and cryolite with refractive index 1.36.

### Interference by a film with two non-parallel reflecting surfaces – wedge shaped film

Consider a wedge shaped thin air film enclosed between two plane surfaces OA and OB inclined at angle  $\theta$ . As shown in figure the thickness of the film goes on increasing from O onwards.

When a monochromatic parallel beam of light falls on such a film the rays reflected from upper and lower surface of the film interfere and produce fringes. These fringes are straight and are formed parallel to the edge of the film. (See footnote). The interfering rays do not enter the eye parallel to the edge but appear to diverge from a point near to the film. The effect is best observed when the angle of incidence is very small.

We know that the path difference between the reflected rays from a thin film is given by  $2\mu t \cos r - \frac{\lambda}{2}$ . If we are observing the reflected ray only in vertical direction then  $i$  will be small and consequently  $r$  will also be small and we can take  $\cos r \approx 1$ .

∴ The path difference between the reflected rays =  $2\mu t - \frac{\lambda}{2}$ .

For constructive interference

$$2\mu t - \frac{\lambda}{2} = n\lambda \quad (n = 0, 1, 2, \dots)$$

or  $2\mu t = \left(n + \frac{1}{2}\right)\lambda \quad \dots(27)$

We observe bright fringes at all places which satisfy the above condition.  
For destructive interference

$$2\mu t - \frac{\lambda}{2} = (2n - 1) \frac{\lambda}{2}$$

$$2\mu t = n\lambda \quad \dots(28)$$

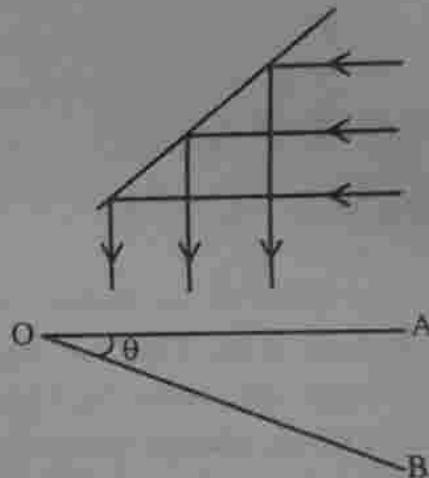


Figure 3.6

We observe dark fringes at all places which satisfy the above condition.

### Expression for fringe width in wedge shaped film

Suppose  $n^{\text{th}}$  dark fringe is formed at a distance  $x_n$  from O.

$$\text{From the figure } \frac{t}{x_n} = \tan \theta$$

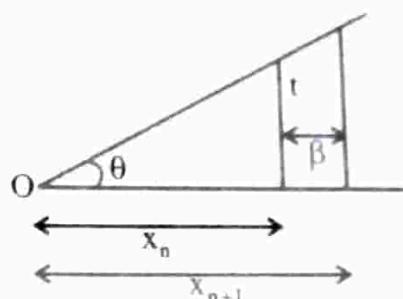


Figure 3.7

or

$$t = x_n \tan \theta$$

Using the condition for destructive interference (eqn 28), we get

$$2\mu x_n \tan \theta = n\lambda \quad \dots \dots (29)$$

Similarly for  $(n+1)^{\text{th}}$  dark fringe, we get

$$2\mu x_{n+1} \tan \theta = (n+1)\lambda \quad \dots \dots (30)$$

eqn 30 – eqn 29 gives

$$2\mu \tan \theta (x_{n+1} - x_n) = \lambda$$

But  $x_{n+1} - x_n = \beta$ , the fringe width

$$\text{Then } 2\mu \tan \theta \beta = \lambda$$

or

$$\beta = \frac{\lambda}{2\mu \tan \theta}$$

If  $\theta$  is small  $\tan \theta \approx \theta$

$$\therefore \beta = \frac{\lambda}{2\mu \theta} \quad \dots \dots (31)$$

Since the fringe width does not depend on  $n$ , all the fringes are of equal thickness. Fringe width increases in direct proportion to  $\lambda$  and in inverse proportion to wedge angle  $\theta$ . If we determine  $\beta$  for any given wavelength  $\lambda$ , we can determine the angle of the wedge.

**Note :** This can be made use of for the determination of the wavelength of a given source of light and also can be used to determine the thickness of fine wires. (see example 7 and 9)

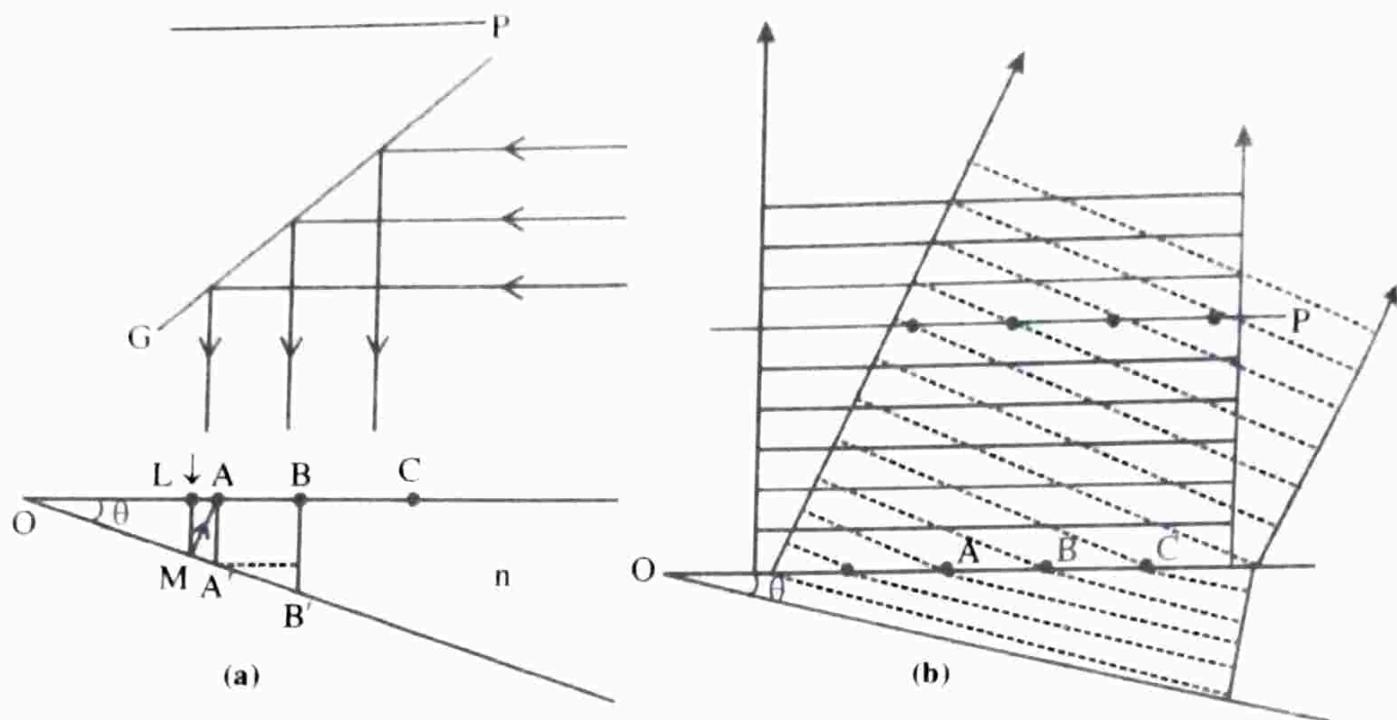


Figure 3.8

Suppose a parallel beam of light is incident normally on the upper surface of the film. In the successive positions of crests (at a particular instant of time) reflected from the upper surface and lower surface of the film are shown by solid and dashed lines respectively. Obviously a photographic plate  $P$  will record straight line interference fringes which will be parallel to the edge of the wedge. (The edge is the line passing through the point  $O$  and perpendicular to the plane of the paper).

### Fringes produced by extended source of light

If the wedge is illuminated by extended source of light (white light) then at the edge of the film appears dark, and near to the edge a few coloured fringes are seen in the order of violet to red colour. Then there is overlapping of colours. And further for large thickness, the film becomes uniformly white. i.e. no fringe pattern will be observed.

### Colours of thin films

It is a commonly observed phenomenon that a thin film of oil on water surface, oil spreads on roads, soap bubbles etc. when viewed under an extended source of white light (sun) appears to be beautifully coloured. It is due to the phenomenon of interference. When an observer views different regions of the film at different angles, depending on the angle and thickness of the film a particular wavelength in white light may satisfy the condition for constructive or destructive interference. The colours which satisfies the condition for constructive interference is present in the region. The colour which satisfies the condition for destructive interference is absent in the region and the corresponding complementary colour is present in the region. So

different points on the film appear in different colours depending on the angle of reflected light coming to the observer's eye from that point. As the observer changes his position the thickness and angle of reflected light vary from point to point, different colours are seen.

**Note :** No colours are visible with thick films. Increasing the thickness of the film, causes the reflected rays to get far apart from each other. As the pupil of the eye is small only one of the two rays that produce interference can enter the eye at a time and interference cannot occur. To see the interference, the film should not be more than few wavelength thick.

### Example 5

A non-reflecting film of refractive index  $\mu_f$  is coated on a glass having refractive index  $\mu_g$ . Show that  $\mu_f = \sqrt{\mu_g \times \mu_a}$  where  $\mu_a$  is the refractive index of air.

### Solution

Let  $\mu_a$ ,  $\mu_f$  and  $\mu_g$  be the refractive indices of air, non-reflecting film and glass respectively. The amplitude reflection and transmission coefficients  $r$  and  $t$  are given by

$$r = \frac{\mu_1 - \mu_2}{\mu_1 + \mu_2} \quad \text{and}$$

$$t = \frac{2\mu_1}{\mu_1 + \mu_2}$$

Let  $a$  be the amplitude of the incident wave

$\therefore$  The amplitude of the reflected wave =  $a_r$

$$= \frac{(\mu_a - \mu_f)}{(\mu_a + \mu_f)} a \quad \dots(1)$$

The amplitude of the transmitted wave in the film =  $a_t$

$$= \frac{2\mu_a}{\mu_a + \mu_f} a \quad \dots(2)$$

The amplitude of the reflected wave from the glass surface =  $a_{rt}$

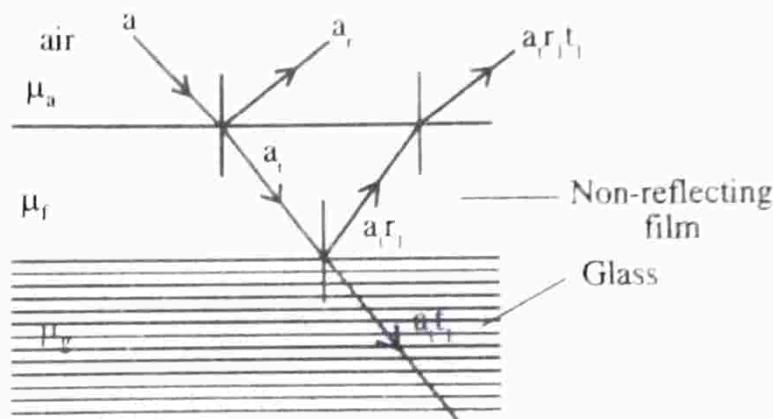


Figure 3.9

$$= \frac{2\mu_a}{\mu_a + \mu_f} \cdot \frac{(\mu_f - \mu_g)}{(\mu_f + \mu_g)} \cdot a \quad \dots\dots(3)$$

The amplitude of the transmitted wave from film to air =  $a_t r_1 t_1$

$$= \frac{2\mu_a}{\mu_a + \mu_f} \frac{(\mu_f - \mu_g)}{(\mu_f + \mu_g)} \frac{2\mu_f}{\mu_f + \mu_a} a \quad \dots\dots(4)$$

For complete destructive interference the amplitudes of the two waves in air (one reflected and the other transmitted) must be the same.

i.e. eqn (1) = eqn (4)

We get

$$\frac{\mu_a - \mu_f}{\mu_a + \mu_f} = \frac{4\mu_a \mu_f (\mu_f - \mu_g)}{(\mu_a + \mu_f)^2 (\mu_f + \mu_g)}$$

Since  $\frac{4\mu_a \mu_f}{(\mu_a + \mu_f)^2} \approx 1$ , we have

$$\frac{\mu_a - \mu_f}{\mu_a + \mu_f} = \frac{\mu_f - \mu_g}{\mu_f + \mu_g}$$

Cross multiplying we get

$$\mu_a \mu_f - \mu_f^2 + \mu_a \mu_g - \mu_f \mu_g = \mu_a \mu_f - \mu_a \mu_g + \mu_f^2 - \mu_f \mu_g$$

or

$$2\mu_f^2 = 2\mu_a \mu_g$$

or

$$\mu_f^2 = \mu_a \mu_g$$

$$\therefore \mu_f = \sqrt{\mu_a \times \mu_g}$$

**Note :** It may be noted that the calculation is done for normal incidence but figure drawn is for oblique incidence. This is done for clarity.

### Example 6

A light of wavelength  $6000\text{\AA}$  is falling on a glass plate ( $\mu = 1.5$ ). What is the refractive index and minimum coating thickness for minimum reflectivity.

### Solution

$$\lambda = 6000 \times 10^{-10} \text{ m}$$

$$\mu_g = 1.5$$

$$\mu_f = \sqrt{\mu_g} = \sqrt{1.5} = 1.225$$

minimum thickness of coating.

$$t = \frac{\lambda}{4\mu_f} = \frac{6000 \times 10^{-10}}{4 \times 1.225}$$

$$= 1.2245 \times 10^{-7} \text{ m}$$

### Example 7

In a wedge shaped film the distance between the successive fringes is measured to be 1.25 mm. The angle of the wedge is 40 seconds. Calculate the wavelength of light used.  $\mu = 1.4$ .

### Solution

$$\beta = 1.25 \times 10^{-3} \text{ m}, \theta = 40 \text{ s} = \frac{40}{3600} \text{ degree}$$

$$\theta = \frac{40}{3600} \times \frac{\pi}{180} \text{ radian} = 0.194 \times 10^{-3} \text{ radian}$$

$$\mu = 1.4$$

$$\text{We have } \beta = \frac{\lambda}{2\mu\theta} \text{ or } \lambda = 2\beta\mu\theta$$

$$\lambda = 2 \times 1.25 \times 10^{-3} \times 1.4 \times 0.194 \times 10^{-3}$$

$$\lambda = 6790 \times 10^{-10} \text{ m}$$

### Example 8

A wedge shaped air film is illuminated with a light of wavelength 4000 Å. There are 15 fringes per cm. What is the angle of the wedge. Assume normal incidence.

### Solution

$$\lambda = 4000 \times 10^{-10} \text{ m}$$

$$\text{Number of fringes per cm} = 15$$

$$\text{Number of fringes per m} = 15 \times 10^2$$

$$\therefore \text{Fringe width, } \beta = \frac{1}{15 \times 10^2} = \frac{10^{-2}}{15} \text{ m}$$

$$\text{Using } \beta = \frac{\lambda}{2\mu\theta}$$

$$\therefore \theta = \frac{\lambda}{2\mu\beta} = \frac{4000 \times 10^{-10} \times 15}{2 \times 1 \times 10^{-2}}$$

$$= 3 \times 10^{-4} \text{ radian}$$

**Example 9**

Two plane glass surfaces in contact along one edge are separated at the opposite edge by a thin wire. If 20 fringes are observed between these edges in sodium light of normal incidence. What is the thickness of the wire. ( $\lambda = 5890\text{\AA}$ )

**Solution**

If the thickness of the wire is  $t$ , then

$$2t = n\lambda$$

or  $t = \frac{n\lambda}{2} = \frac{20 \times 5890 \times 10^{-10}}{2} = 5.89 \times 10^{-6} \text{ m}$

**Example 10**

The thickness of a wedge shaped square cellophane film ( $\mu = 1.5$ ) at two ends is  $t_1$  and  $t_2$  respectively. If the number of fringes seen on the film by the light of wavelength  $6000\text{\AA}$  is 10. Find the value of  $t_2 - t_1$ .

**Solution**

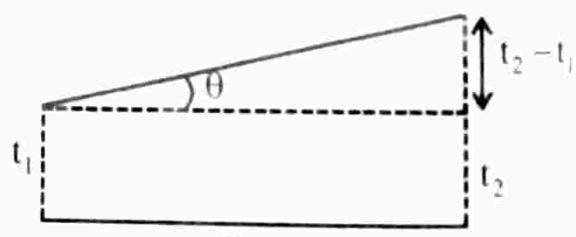
$$\mu = 1.5, \quad \lambda = 6000 \times 10^{-10} \text{ m} \quad n = 10$$

If the side of the square film is  $x$ . then from the figure

$$\theta = \frac{t_2 - t_1}{x}$$

or

$$t_2 - t_1 = x\theta$$



But

$$\beta = \frac{x}{10} = \frac{\lambda}{2\mu\theta}$$

**Figure 3.10**

or

$$\theta = \frac{10\lambda}{2\mu x}$$

$$\therefore t_2 - t_1 = \frac{10\lambda}{2\mu} = \frac{5\lambda}{\mu} = \frac{5 \times 6000 \times 10^{-10}}{1.5}$$

$$= 2 \times 10^{-6} \text{ m.}$$

## Newton's rings

When we place a planoconvex lens on a plane glass plate, a thin film of air is formed between the lens and the glass plate. The thickness of the film at the point of contact is zero and increases as one moves away from the centre. When monochromatic light is allowed to fall on this arrangement normally (as shown in figure) interference takes place between the light reflected from the upper and lower surfaces of the air film and interference pattern can be observed using microscope. The fringes are in the form of concentric rings since the thickness of air film is constant over a circle having the point of lens-glass plate contact as the centre. These rings are called Newton's rings.

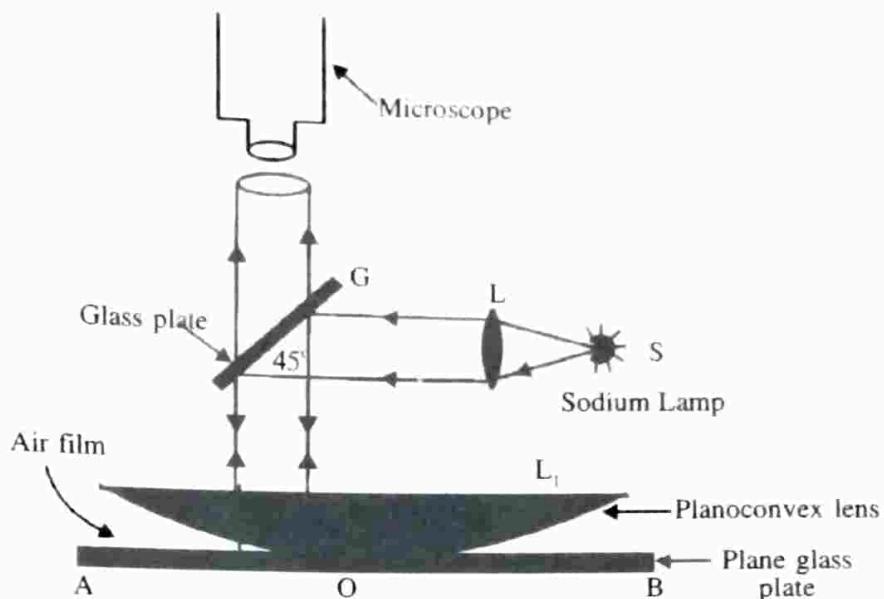


Figure 3.11

## Theory

Let  $R$  be the radius of curvature of the lens. Let a dark fringe be located at  $Q$ . Let the thickness of the air film at  $Q$  be  $PQ = t$ , and radius of the circular fringe at  $Q$  be  $OQ = r_n$ . Using Pythagoras theorem we have

$$PM^2 = PN^2 + MN^2$$

i.e.

$$R^2 = r_n^2 + (R - t)^2$$

$$R^2 = r_n^2 + R^2 - 2Rt + t^2$$

or

$$r_n^2 = 2Rt - t^2$$

since  $R \gg t$ ,  $2Rt \gg t^2$

hence  $t^2$  can be neglected

$$\therefore r_n^2 = 2Rt \quad \dots\dots(1)$$

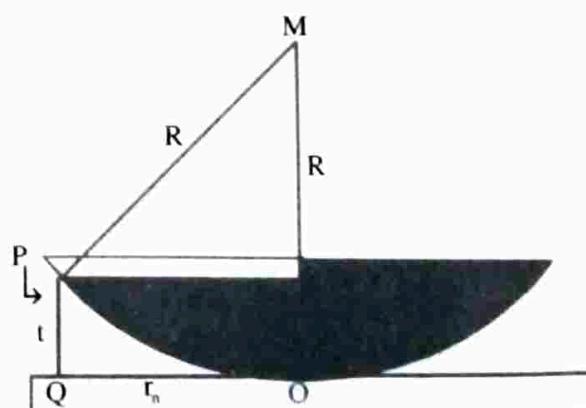


Figure 3.12

using the condition for darkness

$$2\mu t \cos r = n\lambda$$

$\mu = 1$  and  $r = 0$  (since incidence is normal)

$$\therefore 2t = n\lambda$$

$\therefore$  equation 1, becomes

$$r_n^2 = n\lambda R \quad \dots\dots(2)$$

or  $r_n = \sqrt{n\lambda R}$  where  $n = 1, 2, 3, \dots$  it shows that the radii of dark rings are proportional to the square root of natural numbers.

$\therefore$  The diameter of the ring  $D_n = 2r_n$

$$\text{or } D_n = 2\sqrt{2Rt} = 2\sqrt{n\lambda R}, \quad n = 1, 2, 3, \dots$$

Similarly the diameter of the  $(n+1)^{\text{th}}$  bright ring will be

$$D_{n+1} = 2\sqrt{(n + \frac{1}{2})\lambda R}$$

$$\text{or } D_n = 2\sqrt{(n - \frac{1}{2})\lambda R} \text{ for } n^{\text{th}} \text{ ring.}$$

Since thickness of the air film at the centre is zero, the path difference at the centre is zero and a dark spot is surrounded by the first bright ring.

Let us consider the  $(n+k)^{\text{th}}$  dark ring. The radius of the  $(n+k)^{\text{th}}$  ring be  $r_{n+k}$ .

Then

$$r_n^2 = n\lambda R \quad \dots\dots(3)$$

$$D_n^2 = 4n\lambda R$$

$$D_{n+k}^2 = 4(n+k)\lambda R \quad \dots\dots(4)$$

$$D_{n+k}^2 - D_n^2 = 4k\lambda R$$

So the difference of the squares of diameters of  $(n+k)^{\text{th}}$  ring and  $n^{\text{th}}$  ring is independent of  $n$  and only depends on  $k$ .

$$\lambda = \frac{D_{n+k}^2 - D_n^2}{4kR}$$

In this way the wave length of monochromatic light can be found using Newton's

**Note:** If the medium other than air, say liquid, forms the film  $\lambda$ , can be calculated by

$$\lambda = \mu \frac{D_{n+k}^2 - D_n^2}{4kR} \text{ where } \mu \text{ is the refractive index of the liquid.}$$

### Determination of wavelength of monochromatic light

The experimental arrangement (see figure 1.11) for observing Newton's rings consists of a monochromatic light S whose wavelength  $\lambda$  to be determined, placed at the focus of a lens L. The rays emerging out through the lens are parallel. These horizontal parallel rays fall on a glass plate G kept at  $45^\circ$  and are partially reflected from it. These reflected rays fall normally on the planoconvex L<sub>1</sub> placed on the glass plate AB. It is due to the superposition of two reflected rays one from the top and the other from the bottom surfaces of the glass plate interference rings are formed. These rings are observed through a microscope arranged vertically above the glass plate G. The microscope is focussed well so that the rings are clearly seen. Then by working the tangential screw, the point of intersection of the cross wire is kept at the central dark. Then the microscope is moved to the left and to the right in order to ensure that about 30 dark rings are clearly seen.

Starting from the central spot the microscope is moved to the left by working the tangential screw. The tangential screw is slowly adjusted so that the cross wire is tangential to the 28<sup>th</sup> dark ring. The microscope reading on the horizontal scale is taken. Then by working the tangential screw the cross wire is kept tangential to the 26<sup>th</sup>, 24<sup>th</sup>, 22<sup>nd</sup> etc. dark rings up to the second ring on the left and the reading to each ring is taken. Then by working the tangential screw, the microscope is moved in the same direction until the cross wire is tangential to the second dark ring on the right. The corresponding reading is taken. Similarly the cross wire is kept tangential to the 4<sup>th</sup>, 6<sup>th</sup>, 8<sup>th</sup> etc. dark rings up to the 28<sup>th</sup> dark ring on the right. Then reading corresponding to each ring is taken. (The tangential screw should be worked only in one direction from the position of the 28<sup>th</sup> ring on the left to the position of the 28<sup>th</sup> ring on the right. This is to avoid back lash error).

The difference in readings on the left and right at each ring gives its diameters. The value of  $D^2$  is calculated. The values of  $D_{n+k}^2 - D_n^2$  are calculated for a value of  $k = 10$ . Then the mean value of  $D_{n+k}^2 - D_n^2$  is found.

The focal length of the lens is determined by the plane mirror method as it is a long focus lens. For this the convex lens is placed in front of an illuminated wire gauze with a plane mirror held behind the lens. The distance of the lens is adjusted to get a clear image of the wire gauze side by side with it. The distance between the

lens and wire gauze is measured. This gives the focal length of the lens. Repeat this three or four times take the average value. The radius of curvature of the lower surface of the lens is found by Boy's method. For this convex lens is placed in front of an illuminated wire gauze with a piece of black paper held behind the lens. The distance of the lens is adjusted to get a clear image at the wire gauze side by side with it. The distance  $d$  between the lens and wire gauze is measured. The experiment is repeated and the average value of  $d$  is found. Then the radius of curvature of the convex lens is calculated using the formula (Boy's method)

$$R = \frac{fd}{f-d}$$

The wavelength of the given light is hence calculated using the formula

$$\lambda = \frac{D_{n+k}^2 - D_n^2}{4kR}, \text{ } k \text{ may be taken to be 10.}$$

### Determination of refractive index of liquid by Newtons rings

In the experimental arrangement shown above one or two drops of liquid whose refractive index is to be determined is put between lens and glass plate.

When the film is of refractive index  $\mu$ , the diameter of the  $n^{\text{th}}$  dark ring is

$$D_n'^2 = \frac{4n\lambda R}{\mu} \quad \dots\dots(1)$$

For the  $(n+k)^{\text{th}}$  dark ring

$$D_{n+k}'^2 = \frac{4(n+k)}{\mu} \lambda R \quad \dots\dots(2)$$

$$D_{n+k}'^2 - D_n'^2 = \frac{4k\lambda R}{\mu} \quad \dots\dots(3)$$

For air

$$D_{n+k}^2 - D_n^2 = 4k\lambda R \quad \dots\dots(4)$$

$\frac{\text{eqn 4}}{\text{eqn 3}}$  gives

$$\mu = \frac{D_{n+k}^2 - D_n^2}{D_{n+k}'^2 - D_n'^2} \quad \dots\dots(5)$$

First the diameters of  $n^{\text{th}}$  and  $(n+k)^{\text{th}}$  dark rings are measured when the film is air. And the diameters of same rings are measured when the film is of liquid. Then by using eqn (5),  $\mu$  can be determined.

### Testing of optical flatness

An important application of the principle involved in the Newton's rings exper-

ment lies in the determination of the optical flatness of a glass plate. When a beam of parallel monochromatic light is allowed to fall on a film with flat surfaces equal thickness fringes will be observed. In other words if the fringes obtained are not of equal thickness it means that the surfaces are not optically flat.

To test whether a given surface is optically flat or not, take an optically plane surface OA and surface OB to be tested. A wedge shaped air film of varying thickness is formed in between two surfaces. The fringes so obtained are observed through a microscope and if the fringes are not of equal thickness, the surface OB is not flat. Polishing the surface and the process is repeated till fringes of equal thickness are observed. This is how manufacturers of optical flats detect even minute deviations from true optical flatness.

### Example 11

Newton's rings are observed in reflected light of  $\lambda = 5.9 \times 10^{-7}$  m. The diameter of the 10<sup>th</sup> dark ring is 0.5 cm. Find the radius of curvature of the lens and the thickness of air film.

### Solution

$$\lambda = 5.9 \times 10^{-7} \text{ m}, D_{10} = 0.5 \text{ cm} = 0.5 \times 10^{-2} \text{ m}, n = 10$$

We have

$$D_n^2 = 4nR\lambda$$

$$\therefore R = \frac{D_n^2}{4n\lambda} = \frac{0.5 \times 10^{-2}}{4 \times 10 \times 5.9 \times 10^{-7}} = 1.059 \text{ m}$$

Using  $2t = n\lambda$

$$\begin{aligned} t &= \frac{n\lambda}{2} = \frac{10 \times 5.9 \times 10^{-7}}{2} \\ &= 29.5 \times 10^{-7} \text{ m.} \end{aligned}$$

### Example 12

Newton's rings are formed with a monochromatic light between a flat glass plate and a convex lens are viewed normally. What will be the wavelength of light used if the diameter of the 20<sup>th</sup> bright ring is 0.56 cm,  $R = 1.2$  m.

### Solution

$$D_{20} = 0.56 \text{ cm} = 0.56 \times 10^{-2} \text{ m and } R = 1.2 \text{ m.}$$

The diameter of the  $n^{\text{th}}$  bright ring is given by

$$D_n = 2\sqrt{(n - \frac{1}{2})\lambda R}$$

Squaring and rearranging, we get

$$\lambda = \frac{D_n^2}{4(n - \frac{1}{2})R}$$

$$\lambda = \frac{(0.56 \times 10^{-2})^2}{4(20 - \frac{1}{2}) \times 1.2} = \frac{0.56^2 \times 10^{-4}}{4 \times 19.5 \times 1.2}$$

$$\lambda = 3.35 \times 10^{-7} \text{ m.}$$

### Example 13

A Newton's ring arrangement is used with a source emitting two wave lengths  $\lambda_1 = 6 \times 10^{-7} \text{ m}$  and  $\lambda_2 = 4.5 \times 10^{-7} \text{ m}$  is found that the  $n^{\text{th}}$  dark ring due to  $\lambda_1$  coincides with  $(n+1)^{\text{th}}$  dark ring due to  $\lambda_2$ . Find the diameter of the  $n^{\text{th}}$  dark ring for  $\lambda_1$ .  $R = 90 \text{ cm} = .90 \text{ m}$ .

### Solution

$$\lambda_1 = 6 \times 10^{-7} \text{ m}, \lambda_2 = 4.5 \times 10^{-7} \text{ m} \text{ and } R = 90 \text{ cm} = .90 \text{ m.}$$

$$\text{The diameter of } n^{\text{th}} \text{ dark ring } D_n = 2\sqrt{nR\lambda}$$

As the  $n^{\text{th}}$  dark ring due to  $\lambda_1$  coincides with  $(n+1)^{\text{th}}$  dark ring due to  $\lambda_2$ , the diameters are the same, then we have

$$2\sqrt{nR\lambda_1} = 2\sqrt{(n+1)R\lambda_2}$$

$$\therefore n\lambda_1 = (n+1)\lambda_2$$

$$n \times 6 \times 10^{-7} = (n+1) \times 4.5 \times 10^{-7}$$

$$6n = 4.5n + 4.5$$

or

$$1.5n = 4.5$$

$\therefore$

$$n = 3.$$

Then

$$D_3 = 2\sqrt{3 \times .90 \times 6 \times 10^{-7}} = 2.545 \times 10^{-3} \text{ m.}$$

### Example 14

In the Newton's rings arrangement, the radius of curvature of the curved surface is 50 cm. The radii of the 9<sup>th</sup> and 16<sup>th</sup> dark rings are 0.18 cm and 0.2235 cm respectively. Calculate the wavelength.

### Solution

$$R = 50 \text{ cm} = 0.5 \text{ m}$$

$$r_9 = 0.18 \text{ cm} = 0.18 \times 10^{-2} \text{ m}$$

$$D_9 = 2 \times r_9 = 0.36 \times 10^{-2} \text{ m}$$

$$r_{16} = 0.2235 \text{ cm} = 0.2235 \times 10^{-2} \text{ m}$$

$$D_{16} = 2 \times r_{16} = 0.4470 \times 10^{-2} \text{ m}$$

Using

$$\lambda = \frac{D_{n+k}^2 - D_n^2}{4kR}$$

$$\lambda = \frac{D_{16}^2 - D_9^2}{4 \times 7 \times R} = \frac{(0.4470 \times 10^{-2})^2 - (0.36 \times 10^{-2})^2}{4 \times 7 \times 0.5}$$

$$\lambda = \frac{(19.98 - 12.96) \times 10^{-6}}{14}$$

$$\lambda = 5.014 \times 10^{-7} \text{ m.}$$

### Example 15

In Newton's ring experiment the radius of curvature of a lens is 5m and its diameter is 2cm. Calculate the total number of rings formed. If the wavelength of incident monochromatic light is 5500Å. Also calculate the number of rings if the system is kept in water ( $\mu = 1.33$ )

### Solution

$$D_n = 2 \times 10^{-2} \text{ m}, R = 5 \text{ m}, \lambda = 5500 \times 10^{-10}$$

$$\mu = 1.33, r_n = 1 \times 10^{-2} \text{ m}$$

$$r_n = \sqrt{n\lambda R}$$

$$\text{or } n = \frac{r_n^2}{\lambda R} = \frac{(10^{-2})^2}{5500 \times 10^{-10} \times 5} = 36.36$$

Therefore 36 dark and 36 bright rings are formed

When the system is kept in water

$$r_n = \sqrt{\frac{n\lambda R}{\mu}}$$

$$\text{or } n = \frac{r_n^2 \mu}{\lambda R} = 36.36 \times 1.33 \\ = 48.35$$

Therefore in water 48 dark and 48 bright fringes are formed.

## Interferometers

Instruments based on the principle of interference of light are called as interferometers. These instruments use large path differences between the two interfering beams to obtain interference. They are used to study the fine structure of spectral lines and variation of refractive indices of glass and also of liquids at different conditions such as pressure, temperature and density. The interferometers used for the study of refractive indices are called as refractometers.

### Michelson's interferometer

Prof. A.A. Michelson (1852-1931) devised an interferometer based on the principle of interference called Michelson's interferometer. It consists of two highly polished front silvered plane mirrors  $M_1$  and  $M_2$  mounted mutually perpendicular to each other. The mirror  $M_1$  is mounted on a fine pitch screw (capable of reading up to  $10^{-5}$  cm accurately) by which  $M_1$  can be moved to and fro exactly parallel to itself. The mirror  $M_2$  is fixed and provided with levelling screws at its back for adjusting the plane of the mirror.  $G_1$  and  $G_2$  are optically plane and equally thick parallel plates of glass held parallel to each other and at  $45^\circ$  to mirrors  $M_1$  and  $M_2$ .  $G_1$  is semi-silvered at its back surface and  $G_2$  is plane. The plane of  $G_2$  can be slightly adjusted by means of fine screw. The plate  $G_2$  acts as a compensating plate and required essentially to work with white light source.

### Working

Light from a monochromatic source  $S$  after being rendered parallel by a convex lens falls on the glass plate  $G_1$ . At  $G_1$  the beam is divided into two parts, one part goes to  $M_1$  after reflection and the other part is transmitted and goes to  $M_2$ . Both the rays fall on the mirrors normally and retrace their paths. The ray coming from  $M_1$  is partially transmitted at  $G_1$  and proceeds along AT. The ray coming from  $M_2$  gets reflected at  $G_1$  and proceeds along AT. Thus the telescope T receives two sources. Since these two are derived from the same source by the process of division of amplitude, they are coherent. The two beams entering the telescope differ in their path lengths. This path difference between the two interfering rays arises due to the difference in distances of  $M_1$  and  $M_2$  from  $G_1$  and also due to the fact that the rays travelling from  $M_2$  to  $G_1$  suffer a phase change of  $\pi$  when it is reflected at  $G_1$ . The two beams incident on the telescope interfere with each other and fringes are observed in the field of view of telescope T.

### Function of compensating plate $G_2$

The two rays travel through glass unequal distances before reaching the telescope. The ray going towards  $M_1$  and reaching the telescope passes through the

glass plate thrice while the other ray passes the glass plate only once. To equalise this a compensating plate  $G_2$  is introduced between the plates.

### Adjustments of the setup

Michelson interferometer is to be adjusted before using it for doing experiments. The following adjustments are to be done to obtain fringes.

- The interferometer is illuminated with a broad source of monochromatic light. A sodium vapour lamp or mercury vapour lamp provide this. If the source is small a convex lens can be made use of to broaden the source.
- The distances of the mirrors  $M_1$  and  $M_2$  from the back surface of the plate  $G_1$  must be almost equal to within a millimetre.
- The mirror  $M_2$  is made perpendicular to the mirror  $M_1$ . To do this a pin is placed between the source  $S$  and the plate  $G_1$ . On looking through the telescope four images (two pairs) are seen. One pair is due to reflections from the silvered part of  $G_1$  and the other pair is due to the reflection from the unsilvered part of  $G_1$ . By adjusting the screws at the back of one of the mirrors ( $M_1$  or  $M_2$ ), the weaker images are made to coincide with the stronger ones. Now only two images are seen through the telescope. After this being done the pin is removed.

When the adjustments are over circular fringes can be seen through the telescope. Further the centre of the fringe system can be brought in the field view by slightly adjusting the screws provided at the back of the mirror.

### Types of fringes

The shape of the fringes formed in Michelson's interferometer depends on the inclination of  $M_1$  and  $M_2$ . This can be easily understood from the figure below. Let

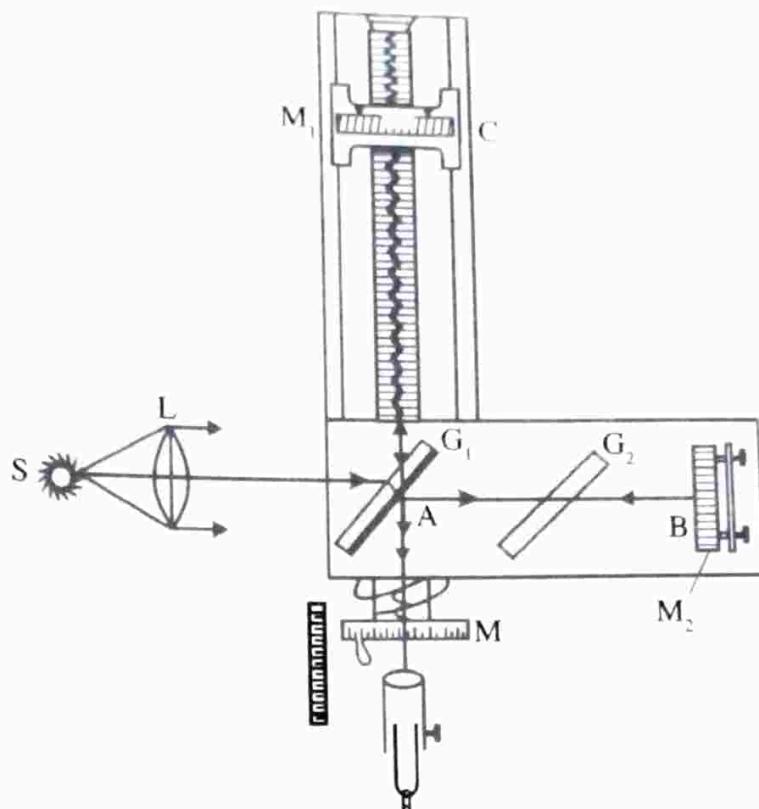


Figure 3.13

$M'_2$  be the virtual image of mirror  $M_2$  formed by the reflection in  $G_1$ , and  $S_o$  be the image of the source formed in  $G_1$ . The interference pattern may be regarded due to the rays of light reflected from the surface of  $M_1$  (real) and  $M'_2$  (virtual). Therefore the interference will be due to the air film enclosed between the two mirrors  $M_1$  and  $M'_2$ . If  $S_1$  and  $S_2$  are the images of the source in the mirror  $M_1$  and  $M'_2$ . Then we can consider that the whole arrangement produces two coherent sources by the process of division of amplitude.

### Circular fringes

When the mirrors  $M_1$  and  $M_2$  are exactly at right angles to each other, then the image  $M'_2$  is parallel to  $M_1$ . If the difference in the distances of mirror  $M_1$  and  $M_2$  from  $G_1$  is  $d$ . Then there will be a air film of constant thickness  $d$  between  $M_1$  and  $M_2$  and the virtual coherent sources  $S_1$  and  $S_2$  will also be parallel but separated by a distance  $2d$ . And the path difference between the rays coming from  $S_1$  and  $S_2$  and inclined at an angle  $\theta$  is  $2d \cos \theta$ .

If the two rays entering the telescope satisfy the condition

$$2d \cos \theta - \frac{\lambda}{2} = n\lambda \quad (\text{see foot note})$$

$$2d \cos \theta = \left(n + \frac{1}{2}\right)\lambda$$

Then the two rays are in a condition of constructive interference and therefore a bright circular fringe is seen at all those points which are at inclination  $\theta$ .

$$\text{If the path difference } 2d \cos \theta - \frac{\lambda}{2} = (2n - 1) \frac{\lambda}{2}$$

$$2d \cos \theta = n\lambda$$

Then the two rays are destructively interfere and there a dark circular fringe is seen at the inclination  $\theta$ .

When  $\theta = 0$ , then the path difference between the two rays is  $2d$ .

If  $2d = \left(n + \frac{1}{2}\right)\lambda$  the centre is bright and when  $2d = n\lambda$  the centre is dark.

**Note :** Since the beam splitter ( $G_1$ ) is just a simple glass plate, the beam reflected from the mirror  $M_2$  will undergo a phase change  $\pi$ , this will introduce an additional path difference  $\frac{\lambda}{2}$  in  $2d \cos \theta$ .

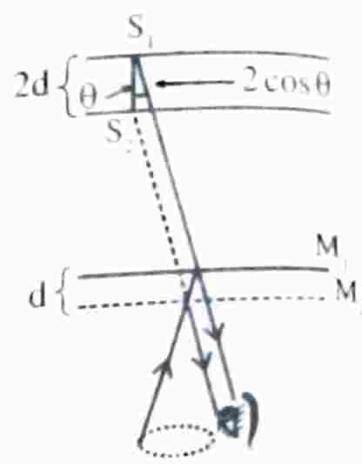


Figure 3.14

## Localised fringes

When the mirrors  $M_1$  and  $M_2$  are kept not exactly at right angles to each other, then  $M_1$  and  $M_2'$  will not be parallel, hence space between them is a wedge shaped air film. Thus we get equally spaced (fringes of equal thickness) localised fringes. The fringes are nearly straight but slightly curved.

## White light fringes

When white light is used instead of monochromatic light no fringes will be seen except for a path difference so small that it does not exceed a few wavelengths. When the path difference is very small the different colours overlap hence only first few coloured fringes are seen. Here the central fringe corresponding to zero path difference is dark.

## Uses of Michelson's interferometer

Michelson's interferometer is used for variety of purposes. Here we discuss only three of them.

### (i) Determination of the wavelength of monochromatic light

Make the necessary adjustments of the interferometer with monochromatic light whose wavelength is to be determined. By adjusting the mirror  $M_1$ , bright spot is obtained at the centre. Telescope is focused on the bright centre spot. The mirror  $M_1$  is moved slowly backward or forward, the number of bright fringes crossing at the centre is counted. The distance through which mirror moved is read from the micrometer screw.

If the difference between the distances of  $M_1$  and  $M_2$  from  $G_1$  is  $d$ , then for the dark spot at the centre, we have

$$2d = n\lambda$$

where  $n$  is the order of the fringe.

If now  $M_1$  is further moved away by a distance  $\frac{\lambda}{2}$ ,  $2d$  changes by  $\lambda$  and again we get dark spot at the centre. Now the dark spot at the centre will be  $(n + 1)$  instead of  $n$ . Thus each time when  $M_1$  is moved by a distance  $\frac{\lambda}{2}$  one fringe will cross the field of view.

Now moving the mirror  $M_1$  by a distance  $x$ , if the number of fringes which cross the field of view is  $N$ . Then

$$x = N \cdot \frac{\lambda}{2}$$

or

$$\lambda = \frac{2x}{N}$$

Knowing N and x the wavelength can be calculated.

## (ii) Determination of the difference between wavelengths of two close spectral lines

When the source of light emits composite light of wavelengths  $\lambda_1$  and  $\lambda_2$  are used to illuminate interferometer then each source produces its own fringes. As a result of superposition of these fringes the resultant fringes may or may not be clearly visible. When the bright fringe produced by one source falls on the bright fringe produced by the other source then the fringes observed are brightest and clear. This can be done by adjusting the mirror  $M_1$ . The displacement in the mirror  $M_1$  for two successive clear positions is obtained on the micrometer screw. Let this distance be x. Then the path difference be  $2x$ . If the number of fringes of wavelength  $\lambda_1$  in this path be n and the number of fringes corresponding to fringes of wavelength  $\lambda_2$  will be  $n + 1$ .

∴

$$2x = n\lambda_1 = (n+1)\lambda_2$$

or

$$n = \frac{\lambda_2}{\lambda_1 - \lambda_2}$$

or

$$2x = \frac{\lambda_1 \lambda_2}{\lambda_1 - \lambda_2} \approx \frac{\lambda^2}{\lambda_1 - \lambda_2} \quad (\because \lambda_1 \lambda_2 \approx \lambda^2)$$

or

$$\lambda_1 - \lambda_2 = \frac{\lambda^2}{2x}$$

Knowing  $\lambda$  (geometric mean of two wavelengths) and x, the difference in the wavelengths  $\lambda_1 - \lambda_2$  can be calculated.

**Note :** In the experiment the reading of ten successive positions of maxima are taken and mean value of x is found.

## (iii) Standardisation of the metre

One of the most outstanding uses of Michelson's interferometer is to measure length in terms of wavelengths very accurately. When the mirror  $M_1$  is moved slowly and one position to another counting the number of fringes in monochromatic light which cross the centre of the field of view will give a measure of the distance the mirror has moved in terms of  $\lambda$ . Let  $x_1$  be the position corresponding to the dark fringe of order  $n_1$  and  $x_2$  be the position corresponding to the dark fringe of order  $n_2$ . We have

$$2x_1 = n_1 \lambda$$

and

$$2x_2 = n_2 \lambda$$

or

$$x_1 - x_2 = (n_1 - n_2) \frac{\lambda}{2}$$

Hence the distance moved equals the number of fringes counted multiplied by a half wavelength. The distance measured need not correspond to an integral number of half wavelengths. Fractional parts of a whole fringe displacement can easily be estimated to one fiftieth of a fringe. This gives the distance to an accuracy of one hundredth wavelength.

The most important measurement made with the interferometer by using the above principle was the comparison of the standard metre, kept at the International Bureau of weights and Measures at Sèvres near Paris, with the wavelengths of intense red, green and blue lines of Cadmium. This was first performed by Michelson and Benoit in 1895. But it would be impossible to count directly the number of fringes for a displacement of the movable mirror from one end of the standard metre to the other. This is because for a path difference of more than 20 cm it is not possible to obtain fringes. To circumvent this difficulty Michelson and Benoit divided the metre into a number of substandards called etalons. Each etalon is twice the length of the preceding one. The length of the largest etalon was 10 cm and that of the shortest one was 0.039 cm. The smallest etalon was calibrated in terms of wavelengths by actually counting the number of fringes which cross the cross wire of the telescope. This etalon was then compared with the next larger etalon and so on until the longest etalon was reached. Finally the length of the longest etalon i.e. 10 cm was then compared with standard metre. With this Michelson and Benoit found that one standard metre consists of 1,553,164.13 wavelengths of red light of cadmium, 1,966,249.7 wavelengths of green light of cadmium and 2,083,372.7 wavelengths of blue light of cadmium at 15°C and 76 cms of mercury. The values are correct up to 1 part in  $10^9$ . After this the international standard metre is redefined to consist of 1,650,763.73 wavelengths of orange-red light of Krypton atom.

### Example 16

In a Michelson interferometer it is found that 200 fringes cross the field of view when the movable mirror is displaced through 0.049 mm. Calculate the wavelength of light used.

### Solution

$$N = 200, x = 0.049 \times 10^{-3} \text{ m}$$

Using

$$\lambda = \frac{2x}{N} = \frac{2 \times 0.049 \times 10^{-3}}{200}$$

$$= 4900 \times 10^{-10} \text{ m}$$

### Example 17

In a Michelson's interferometer the distance travelled by the mirror for two successive positions of maximum distances was 0.289mm. Calculate the difference between the wavelengths.  $\lambda = 5890 \text{ \AA}$ .

### Solution

We have

$$\lambda_1 - \lambda_2 = \frac{\lambda^2}{2x}$$

$$= \frac{(5890 \times 10^{-10})^2}{2 \times 0.289 \times 10^{-3}}$$

$$= 6 \times 10^{-10} \text{ m.}$$

### Example 18

In Michelson's interferometer experiment, calculate the various values of  $\theta$  corresponding to bright rings for  $d = 5 \times 10^{-3} \text{ cm}$ . Show that if  $d$  is decreased to  $4.997 \times 10^{-3} \text{ cm}$  the fringe corresponding to  $n = 200$  disappears. What will be the corresponding value of  $\theta$ . ( $\lambda = 5 \times 10^{-5} \text{ cm}$  given)

### Solution

The condition for bright ring is

$$2d \cos \theta = \left( n + \frac{1}{2} \right) \lambda$$

or

$$\cos \theta = \frac{\left( n + \frac{1}{2} \right) \lambda}{2d} = \frac{(2n+1)\lambda}{4d}$$

$$\cos \theta = \frac{(2n+1)5 \times 10^{-7}}{4 \times 5 \times 10^{-5}}$$

$$\cos \theta = \frac{2n+1}{400}$$

For the central bright,  $\cos \theta = 1 \therefore \theta = 0$

i.e.

$$\frac{2n+1}{400} = 1 \quad \text{or} \quad 2n+1 = 400$$

or

$$n = \frac{399}{2}$$

$$\theta = \cos^{-1} \frac{(2n+1)}{400}$$

Put  $2n + 1 = 400, 399, 398, \dots$  etc. we get bright fringes at these angles.

$\theta = 0^\circ, 4.052^\circ, 5.731^\circ, \dots$  corresponding to  $2n + 1 = 400, 399, 398, \dots$

(ii) Now  $d = 4.997 \times 10^{-5}$  m. The condition for dark is  $2d \cos \theta = n\lambda$

Then

$$\cos \theta = \frac{n\lambda}{2d} = \frac{n \times 5 \times 10^{-7}}{2 \times 4.997 \times 10^{-5}}$$

$$\cos \theta = \frac{n}{199.88} = \frac{n}{200}$$

For the central fringe  $\theta = 0$  and  $\cos \theta = 1$

$$\therefore 1 = \frac{n}{200}$$

or  $n = 200$

It shows that for  $n = 200$ , we obtain the central fringe and the condition for darkness is satisfied. Hence the 200<sup>th</sup> fringe disappears.

To find the angles  $\theta = \cos^{-1} \left( \frac{n}{200} \right)$

The angles correspond to  $n = 200, 199, 198, 197, \dots$  are  $0^\circ, 5.73^\circ, 8.11^\circ, 9.936^\circ, \dots$

## UNIVERSITY MODEL QUESTIONS

### Section A

(Answer questions in two or three sentences)

#### Short answer type questions

- What is the interference of light by a plane film?
- Why a thick film cannot produce interference when illuminated with white light?
- What are the conditions for brightness and darkness of normal incidence of light on a plane film producing interference?
- Write down the conditions of brightness and darkness of oblique incidence of light on a plane film producing interference.

5. What are Newton's rings? Give two of its uses.
6. Explain why the centre of Newton's rings is dark for reflected light.
7. Why an extended source of light is essential to observe colours in thin films?
8. Write down the cosine law and explain the symbols used.
9. What is reflectivity? What is its relation to amplitude reflection coefficient?
10. What is a non-reflecting film? How it can be achieved?
11. What are the conditions to be satisfied for a non-reflecting material?
12. What is an air wedge?
13. What are the conditions for constructive and destructive interferences formed in a film with non-parallel reflecting surfaces?
14. What are the uses of interference on wedge shaped film?
15. An extended source of light falls on a wedge shaped film. Discuss about the fringes so formed.
16. What is an interferometer?
17. What are the uses of Michelson's interferometer?
18. What is the function of a compensating plate in Michelson's interferometer?
19. What are the conditions to be satisfied to obtain constructive and destructive interferences in Michelson's interferometer?
20. What is the principle behind the standardisation of metre using Michelson's interferometer?
21. Why an extended source of light is essential to observe colours in thin films?
22. What do you understand by
  - (i) fringes of equal thickness      (ii) fringes of equal inclination
23. State conditions for obtaining
  - (i) circular fringes      (ii) straight lines fringes
  - (iii) white light fringes in Michelson's interferometer
24. What are the adjustments of the Michelson's interferometer for doing experiments with it?
25. Define standard metre in terms of wavelength of cadmium.
26. What are white light fringes?

#### Section B

*(Answer questions in a paragraph of about half a page to one page)*

#### Paragraph / Problem type questions

1. Derive the condition for brightness for normal incidence of light on a plane film.
2. Derive the cosine law.

3. Derive an expression for fringe width in wedge shaped film.
4. Explain the phenomenon of colours of thin films.
5. How will you determine the refractive index of a liquid by Newton's rings?
6. How will you test the optical flatness of a glass plate?
7. Prove that an excessive thin film appears black in the reflected light.
8. White light falls perpendicularly upon a film of soapy water whose thickness is  $5 \times 10^{-7}$  m and index of refraction 1.33. Which wavelength in the visible region will be reflected most strongly. [5320Å will be reflected]
9. Light of wavelength 5880Å is incident on a thin film of glass of  $\mu = 1.5$  such that the angle of refraction in the plate is  $60^\circ$ . Calculate the smallest thickness of the plate which will make it dark by reflection. [ $3.72 \times 10^{-7}$  m]
10. When light is incident normally on a glass plate of thickness  $0.5 \times 10^{-6}$  m and index of refraction 1.5. Which wavelengths in the visible region (400nm to 700nm) are strongly reflected by the plate [429nm, 600nm]
11. Using light of  $\lambda = 5.9 \times 10^{-7}$  m, it is found that in a thin film of air 7.4 fringes occur between two points. Deduce the difference of the film thickness between these points. [ $2.183 \times 10^{-6}$  m]
12. There is a wedge shaped air film between two glass plates. The plates are in contact along one edge and are separated by a wire of diameter 0.05 mm placed at a distance of 15 cm from the edge. Find the fringe width when light of wavelength 6000Å is incident normally. [ $9 \times 10^{-4}$  m]
13. A parallel beam of white light is allowed to fall on the slit of the collimator of a telescope and a thin film of air of thickness enclosed between two glass plates is placed in front of the slit with the surface of the air film perpendicular to the path of light. On seeing through the telescope 250 fringes are observed at the spectral region between wavelengths 4000Å and 6500Å. Calculate the thickness of the air film [ $0.013 \times 10^{-2}$  m]
14. The angle of an air wedge is  $24'$ . Light having wavelengths 4732Å and 5460Å is incident on the wedge normally. Find the distance from the edges in contact at which the maxima due to each wavelength coincide for the first time in the reflected light. [ $2.54 \times 10^{-4}$  cm]
15. A vertical soap film of length 10cm is viewed by reflected monochromatic light of wavelength 5893Å. Just before the film breaks there are 12 black and 11 bright fringes. Find the angle of wedge formed and thickness of film at the base just before it breaks. [ $2.44 \times 10^{-4}$  cm]
16. Newton's rings are formed by reflected light of wavelength 6250Å with a liquid between the plane and curved surfaces. If the diameter of the 10<sup>th</sup> bright ring is 5mm and

- radius of curvature of the curved surface is 1.2m. Calculate the refractive index of the liquid. [ $\mu = 1.4$ ]
17. In the Newton's rings arrangement, the radius of the curvature of the curved side of the planoconvex lens is 1m. For  $\lambda = 6 \times 10^{-7}$  m, what will be the radii of the 10<sup>th</sup> and 20<sup>th</sup> bright rings.
- $$\left[ r_{10} = 2.387 \times 10^{-3} \text{ m} \right]$$
- $$\left[ r_{20} = 3.421 \times 10^{-3} \text{ m} \right]$$
18. The diameter of the n<sup>th</sup> ring changes from 1.2cm to 1.04cm when the air space between the lens and the plate is replaced by a liquid. Find the refractive index of liquid. [ $\mu = 1.13$ ]
19. Newton's rings are formed with red light of  $\lambda = 670$  nm. The radius of the 20<sup>th</sup> ring is found to be  $1.1 \times 10^{-2}$  m. Find the radius of curvature of the 30<sup>th</sup> ring.
- $$\left[ r_{30} = 1.347 \times 10^{-2} \text{ m} \right]$$
20. In the Michelson's interferometer arrangement, if one of the mirrors is moved by a distance of 0.08mm, 250 fringes cross the field of view. Calculate the wavelength. [6000 Å]
21. The Michelson's interferometer experiment is performed with a source which consists of two wavelengths 4882Å and 4886Å. Through what distance does the mirror have to be moved between two positions of the disappearance of the fringes. [0.298mm]
22. In Michelson's interferometer calculate the various values of  $\theta$  corresponding to dark rings. For  $d = 0.3$  mm and  $\lambda = 6 \times 10^{-5}$  cm. [0°, 2.56°, 3.62°, 4.44°, 5.13°...]

### Section C

(Answer questions in about one or two pages)

#### Long answer type questions (Essays)

- Derive an expression for the conditions of brightness and darkness produced under oblique incidence of light on a plane film producing interference due to reflected light.
- Explain the formation of Newton's rings. How can these be used to determine the wavelength of monochromatic light.
- What are Newton's rings? Derive an expression for the radii of rings.
- Describe Michelson's interferometer. How will you determine the wavelength of monochromatic light with the help of Michelson's interferometer?
- Describe the construction with diagram and outline the theory of Michelson's interferometer. Discuss the nature of interference pattern produced.
- How can Michelson's interferometer is used for
  - Determining the difference of wavelengths of two nearby spectral lines.
  - Calibrating standard metre.
- Explain the theory of interference formed in a wedge shaped film. Derive an expression for fringe width. How will you determine the wavelength of given light using air wedge?

#### Hint to problems

1 to 7 see book work

8. The condition for brightness is

$$2\mu t \cos r = (2n+1) \frac{\lambda}{2}$$

$$\therefore \lambda = \frac{4\mu t \cos r}{(2n+1)}, r=0 \quad \therefore \cos r = 1$$

Put  $n = 0, n = 1, 2, 3$  in the above we get

$$\lambda_0 = 2660 \text{ Å}$$

$$\lambda_1 = 8866 \text{ Å}$$

$$\lambda_2 = 5320 \text{ Å}$$

$$\lambda_3 = 3800 \text{ Å}$$

out of these only  $\lambda_2$  lies in the visible spectrum. (See also example 2)

9. For the film to appear dark by reflected light, we have

$$2\mu t \cos r = n\lambda$$

For smallest thickness  $n = 1$

$$\therefore t = \frac{\lambda}{2\mu \cos r} = \frac{5880 \times 10^{-10}}{2 \times 1.5 \times \cos 60^\circ} = 3.72 \times 10^{-7} \text{ m}$$

10. The condition for brightness is

$$2\mu t \cos r = (2n+1) \frac{\lambda}{2} \quad n = 0, 1, 2, \dots$$

$$\cos r = 1 \quad \therefore r = 0$$

$$\therefore 2\mu t = (2n+1) \frac{\lambda}{2}$$

When  $\lambda = 400$  nm  $n = 3.25$

When  $\lambda = 700$  nm  $n = 1.66$

Hence  $n$  can take values 2 and 3.

When  $n = 2$   $\lambda_1 = 429$  nm, when  $n = 3$ ,  $\lambda_2 = 600$  nm

11. We have  $2\mu t = (2n+1) \frac{\lambda}{2}$  ....(1)

assumed normal incidence

If  $\delta t$  is the thickness for the shift of 7.4 fringes, then

$$2\mu(t + \delta t) = [2(n + 7.4) + 1] \frac{\lambda}{2} \quad \dots(2)$$

$$\text{eqn 2 - eqn 1} \Rightarrow 2\mu\delta t = 7.4\lambda$$

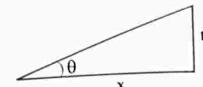
$$\therefore \delta t = \frac{7.4\lambda}{2\mu} = 2.183 \times 10^{-6} \text{ m}$$

12. If  $t$  be the diameter of the wire, then

$$\theta = \frac{t}{x} \quad \left( \text{Angle} = \frac{\text{Arc}}{\text{Radius}} \right)$$

$$\beta = \frac{\lambda}{2\theta} = \frac{\lambda}{2t}x$$

$$\beta = \frac{6000 \times 10^{-10} \times 15 \times 10^{-2}}{2 \times 0.05 \times 10^{-3}} = 9 \times 10^{-4} \text{ m}$$



13. The condition for constructive interference in a thin film where the interference is due to transmitted rays is

$$2\mu t \cos r = n\lambda$$

$$\text{or } 2t = n\lambda \quad (\because \mu = 1, \cos r = 1)$$

$$2t = n_1 \lambda_1$$

$$2t = n_2 \lambda_2$$

$$\text{or } n_1 = \frac{2t}{\lambda_1} \text{ and } n_2 = \frac{2t}{\lambda_2}$$

$$\therefore n_1 - n_2 = 2t \left( \frac{1}{\lambda_1} - \frac{1}{\lambda_2} \right)$$

$n_1 - n_2 = 250$ ,  $\lambda_1$  and  $\lambda_2$  are given find  $t$ .

14. For maximum intensity we have

$$2\mu t \cos r = \left( n + \frac{1}{2} \right) \lambda \quad (\because \mu = 1, r = 0)$$

$$\text{or } 2t = \left( n + \frac{1}{2} \right) \lambda_1 \quad \dots(1)$$

$$2t = \left( n + \frac{3}{2} \right) \lambda_2 \quad \dots(2)$$

$$\therefore n = 7$$

$$t = x\theta$$

using

Putting this in eqn (1) or (2), solve it for  $x$ .

$$x = 2.54 \times 10^{-2} \text{ m}$$

$$15. \beta = \frac{\text{length}}{\text{No. of fringes}} \quad \beta = \frac{10}{11} \text{ cm}$$

$$\text{Using } \beta = \frac{\lambda}{2\mu\theta} \text{ find } \theta, \quad \theta = 2.44 \times 10^{-5} \text{ radian}$$

When the film breaks the thickness at the top is zero

$$\therefore \text{Thickness at the bottom } t = x_0 \theta$$

$$= 10 \times 2.44 \times 10^{-5} = 2.44 \times 10^{-4} \text{ cm}$$

16. For the  $n^{\text{th}}$  bright ring in air

$$D_n^2 = 4R \left( n - \frac{1}{2} \right) \lambda_{\text{air}}$$

$$\text{In liquid} \quad D_n'^2 = 4R \left( n - \frac{1}{2} \right) \lambda_{\text{liquid}}$$

$$\frac{\lambda_{\text{air}}}{\lambda_{\text{liquid}}} = \mu \quad \therefore \quad D_n'^2 = 4R \left( n - \frac{1}{2} \right) \frac{\lambda_{\text{air}}}{\mu}$$

$$\text{or} \quad \mu = \frac{4R \left( n - \frac{1}{2} \right) \lambda_{\text{air}}}{D_n'^2}$$

$$17. \quad r_n = \sqrt{\left( n - \frac{1}{2} \right) \lambda R}$$

$$18. \quad D_n^2 = 4nR\lambda \quad \text{or} \quad D^2 \propto \lambda$$

$$\frac{D_n^2}{D_{\text{liquid}}^2} = \frac{\lambda_{\text{air}}}{\lambda_{\text{liquid}}} = \mu$$

$$19. \quad r_n \propto \sqrt{n}$$

$$\therefore \frac{r_{30}}{r_{20}} = \sqrt{\frac{30}{20}}$$

$$20. \quad N = 250, \quad x = 0.08 \times 10^{-3} \text{ m}$$

$$\lambda = \frac{2x}{N} = \frac{2 \times 0.08 \times 10^{-3}}{250} = 6400 \times 10^{-10} \text{ m}$$

21. We have  $\lambda_1 - \lambda_2 = \frac{\lambda_1 \lambda_2}{2x}$

$$2x = \frac{4886 \times 10^{-10} \times 4882 \times 10^{-10}}{\lambda_1 - \lambda_2}$$

$$2x = \frac{4886 \times 4882 \times 10^{-20}}{4 \times 10^{-10}}$$

$$= 5963 \times 10^{-5} \text{ m}$$

Distance moved  $x = 0.298 \text{ mm}$

22. The condition for dark ring is given by

$$2d \cos \theta = n\lambda$$

$$\text{or } \cos \theta = \frac{n\lambda}{2d} = \frac{n \times 6 \times 10^{-7}}{2 \times 0.3 \times 10^{-3}} = \frac{n}{1000}$$

$$\text{i.e. } \cos \theta = \left( \frac{n}{1000} \right)$$

For the central dark  $\cos \theta = 1$  ( $\therefore \theta = 0^\circ$ )

$$1 = \frac{n}{1000} \Rightarrow n = 1000$$

Then  $n = 999, 998, 997, 996, \dots$  etc. give the other angles.

$$\theta = \cos^{-1} \left( \frac{n}{1000} \right)$$

For  $n = 1000, \theta = 0^\circ$

$n = 999, \theta = 2.56^\circ$

$n = 998, \theta = 3.62^\circ$

$n = 997, \theta = 4.44^\circ$

$n = 996, \theta = 5.13^\circ$  etc.

### IMPORTANT FORMULAE

1. Normal incidence of light on a plane film:

$$(i) \text{ Condition for brightness: } 2\mu t = \left( n + \frac{1}{2} \right) \lambda, \quad n = 0, 1, 2, \dots$$

$$(ii) \text{ Condition for darkness: } 2\mu t = n\lambda, \quad n = 1, 2, 3, \dots$$

2. Oblique incidence of light on a plane film:

$$(i) \text{ The optical path difference between the two reflected rays: } \Delta = 2\mu t \cos r$$

$$(ii) \text{ Condition for brightness: } 2\mu t \cos r = \left( n + \frac{1}{2} \right) \lambda, \quad n = 0, 1, 2, 3, \dots$$

$$(iii) \text{ Condition for darkness: } 2\mu t \cos r = n\lambda, \quad n = 1, 2, 3, \dots$$

3. The cosine law:

The effective path difference between the wave reflected from the lower surface and the upper surface of a film:  $\Delta = 2\mu t \cos r$  known as cosine law.

$$4. \text{ Reflection coefficient: } R = \left( \frac{\mu_1 - \mu_2}{\mu_1 + \mu_2} \right)^2$$

$$5. \text{ Transmission coefficient: } T = \frac{4\mu_1^2}{(\mu_1 + \mu_2)^2}$$

Remember that  $R + T = 1$

6. Interference by a film with two non-parallel reflecting surfaces - wedge shaped.

$$(i) \text{ Condition for constructive interference: } 2\mu t = \left( n + \frac{1}{2} \right) \lambda, \quad n = 0, 1, 2, 3, \dots$$

$$(ii) \text{ Condition for destructive interference: } 2\mu t = n\lambda, \quad n = 1, 2, 3, \dots$$

$$(iii) \text{ Expression for fringe width: } \beta = \frac{\lambda}{2\beta\theta}$$

7. Newton's rings:

$$(i) \text{ The diameter of the } n\text{th dark ring: } D_n = 2\sqrt{n\lambda R}, \quad n = 1, 2, 3, \dots$$

(ii) Wavelength of the light used:

$$\lambda = \frac{D_{n+k}^2 - D_n^2}{4kR} \text{ in air}$$

$$\lambda = \mu \frac{D_{n+k}^2 + D_n^2}{4kR} \text{ in medium}$$

(iii) Radius of curvature of a convex lens:  $R = \frac{fd}{f-d}$  Boy's method

## 8. Michelson's interferometer:

(i) Condition for constructive interference:  $2d \sin \theta = \left( n + \frac{1}{2} \right) \lambda, n = 0, 1, 2, \dots$ (ii) Condition for destructive interference:  $2d \sin \theta = n\lambda, n = 1, 1, 3, \dots$ (iii) Difference between the wavelength:  $\lambda_1 - \lambda_2 = \frac{\lambda^2}{2x}, \lambda = \sqrt{\lambda_1 \lambda_2}$