

---

# **NECI Documentation**

***Release 0.1***

**Alavi Group**

August 01, 2008

# CONTENTS

<b>1</b>	<b>Introduction</b>	<b>1</b>
<b>2</b>	<b>Installation</b>	<b>2</b>
2.1	Download . . . . .	2
2.2	Compilation . . . . .	3
2.3	testcode . . . . .	5
<b>3</b>	<b>Run</b>	<b>6</b>
3.1	NECI . . . . .	6
3.2	CPMD-NECI . . . . .	6
<b>4</b>	<b>Theoretical review</b>	<b>7</b>
4.1	Introduction . . . . .	7
4.2	Graph evaluation . . . . .	8
<b>5</b>	<b>Input options</b>	<b>10</b>
5.1	Overview . . . . .	10
5.2	Non-block level options . . . . .	12
5.3	System . . . . .	12
5.4	PreCalc . . . . .	15
5.5	Calc . . . . .	17
5.6	Integrals . . . . .	28
5.7	Logging . . . . .	32
5.8	Examples . . . . .	34
<b>6</b>	<b>Output files</b>	<b>35</b>
6.1	MCPATHS and MCSUMMARY . . . . .	35

# Introduction

NECI is a rapidly developing code based on a post Hartree–Fock electronic structure method.

It calculates electron correlation via path-resummations in Slater Determinant space [[SumPaper](#)], [[StarPaper](#)], [[ThomPhDThesis](#)].

As a standalone package, NECI can perform calculations on electrons confined to a box, the uniform electron gas and the hubbard model.

NECI can also read in wavefunctions, or a set of integrals based on them, of molecular systems produced by another program (e.g. [[DALTON](#)] or [[MolPro](#)]) and run calculations using them as the basis for forming the necessary Slater Determinants.

Finally, NECI can also be compiled as a library for integration into existing codes. Currently this has been performed for the [[CPMD](#)] or [[VASP](#)] plane-wave packages, allowing calculations to be performed on periodic systems.

# Installation

Requirements:

- **subversion** The Alavi group currently only distributes code via the wmmm subversion repository (account required), which is kindly hosted by the Unilever Centre.
- LAPACK
- BLAS.
- FFTW 3.x.

## 2.1 Download

### 2.1.1 NECI

The NECI source code can be downloaded from:

```
svn checkout https://wmmm.ch.cam.ac.uk/svn2/groups/alavi/NECI/trunk NECI
```

Various branches also exist, but they may or may not be stable or under active development.

### 2.1.2 CPMD

Our modified version of CPMD (containing the necessary routines for integration with NECI) can be downloaded from:

```
svn checkout https://wmmm.ch.cam.ac.uk/svn2/groups/alavi/CPMD/branches/QMC/trunk CPMD
```

Again, there exist various branches and they may or may not be stable or under active development. It is suggested that you place NECI as a subdirectory of the CPMD source, as that is how the CPMD configuration files in the repository are set up.

### 2.1.3 VASP

Only users on the access list can download VASP from our repository:

```
svn checkout https://wmmm.ch.cam.ac.uk/svn2/groups/alavi/VASP/vasp.4.lib vasp.4.lib  
svn checkout https://wmmm.ch.cam.ac.uk/svn2/groups/alavi/VASP/vasp.5 vasp.5
```

## 2.2 Compilation

### 2.2.1 CPMD and NECI

For historical reasons, CPMD and NECI are much more closely interwoven than NECI is with VASP. In addition, the tools used to compile CPMD and NECI are very similar (but then they were written by the same people!).

The repository version of CPMD depends upon NECI, thus NECI must be compiled first. Various scripts take care of this for us.

CPMD and NECI have CONFIGURE subdirectories. Each file in a CONFIGURE subdirector contains the necessary information to compile the code using a certain compiler. Most of the time only the paths and flags for the FFTW, LAPACK and BLAS libraries, given in the LFLAGS variable, will need to be adjusted (if that). It is necessary to use the same compiler to compile CPMD and NECI. We have compiled and tested the codebases with the gfortran (4.2 and later), Portland and Intel compilers in 32 and 64 bit. The mkconfig.sh scripts in the CPMD and NECI directories produce the relevant makefiles, but this is better done via helper scripts.

Please note that not all the CPMD configure scripts will work: many of them were supplied with CPMD and have never been used with our modified version. Please only use platforms that exist in the NECI/CONFIGURE directory as well as the CPMD/CONFIGURE directory. It is easy to make your own configuration files using existing ones as a template.

Note that some of the platforms obtain LAPACK and BLAS as part of atlas, ACML or MKL, so this will need to be changed if different source libraries are used.

Two versions of CPMD (and the corresponding NECI library) exist, gcpmd.x (for gamma-point calculations) and kcpmd.x (for k-point calculations), to take advantage of some substantial memory savings when running gamma-point calculations, as all wavefunctions are then real. This is controlled via a C pre-processing statement. As neci.x is only for molecular calculations, it only exists in one form.

runmake.sh and compile are scripts for CPMD and NECI respectively which control the process of creating makefiles and compiling the codebases. They both refer to a “platform”, which is just a name of one of the configuration files in the CONFIGURE subdirectories.

compile by default generates new makefile (for both gamma-point and k-point compilations) and does a clean build of neci.x:

```
[NECI]$ ./compile -h
usage: ./compile [options] [platform]
Generate new makefiles and do a clean build of neci.x using platform as the configuration.
```

If platform is not specified, then the platform given in .compileconf is used.  
If .compileconf also doesn't exist, then the default (PC-PGI64) is used.

Options:

- d Compile with the compiler debug options on.
- f Fast: don't do a make clean before compiling.
- m Only make new makefiles.

In contrast, runmake.sh has a different default behaviour, in that it doesn't produce new makefiles by default, and does not do clean builds:

```
[CPMD]$ ./runmake.sh -h
Usage: ./runmake.sh [options]
Compile NECI (neci.x) and CPMD/NECI for gamma point (gcpmd.x) and k-point (kcpmd.x) calculations.
```

Options:

- c Recompile only CPMD routines.
- d Generate new makefiles for debugging. Recompile (at least) CPMD/qmc routines and all of NECI.
- g Compile only Gamma point code.
- k Compile only K-point code.
- m Generate new makefiles for NECI and CPMD using the platform given in .compileconf.

```
Defaults to PC-PGI64 if .compileconf does not exist.
-n Recompile only NECI routines.
-p [platform] Generate new makefiles for NECI and CPMD using the platform configuration.
-h Print this message.
```

Note that `runmake.sh` produces new makefiles for CPMD **and** for NECI, and compiles `neci.x`, the NECI libraries needed for CPMD, and `gcpmd.x` and `kcpmd.x`. To aid compilation, the `dest` subdirectories in the CPMD and NECI source directories contains the compiled objects for the gamma-point code and the `kdest` subdirectories contain the compiled objects for the k-point code.

Both the NECI and CPMD scripts default to compiling the codebases using the Portland 64-bit compiler, if a platform is not specified either via the command line or given in `.compileconf`, which is a text file which contains the name of the desired platform. Note that `runmake.sh` will use the same platform for both the CPMD and NECI makefiles.

The CPMD and NECI source directories also contain controlling Makefiles to further help the make process (and generally just act as wrappers for the `runmake.sh` and `compile` scripts). Run

```
make help
```

in each directory to see the various targets available.

To quickest way to compile both CPMD and NECI is to run:

```
[CPMD]$ make all
```

from within the CPMD source directory.

### File structure

NECI files:

**NECI/neci.x** Standalone neci-executable (links to NECI/dest/neci.x).

**NECI/dest/neci-cpmd.a** NECI library for CPMD gamma-point code.

**NECI/kdest/neci-vasp.a** NECI library for CPMD k-point code.

**NECI/dest/neci-cpmd.a** NECI library for VASP gamma-point code.

**NECI/kdest/neci-vasp.a** NECI library for VASP k-point code.

CPMD files:

**CPMD/gcpmd.x** Gamma-point executable of the CPMD-NECI code (links to CPMD/dest/cpmd.x). Must not be used for k-point calculations!

**CPMD/kcpmd.x** k-point executable of the CPMD-NECI code (links to CPMD/dest/cpmd.x). Must not be used for gamma-point calculations!

### 2.2.2 VASP

James has managed it. It's not completely pleasant. More to follow once it's been made easier!

## 2.3 testcode

testcode is a set of scripts written by James Spencer that is used to check that our programs produce the same results as they did before. It is useful both for development work, to ensure that regression issues are avoided, and testing successful compilations.

Every night the latest version of the codebase is checked out of the subversion repository and tested against a variety of compilers, giving confidence in the continued stability of the codebase.

testcode and the set of test jobs (both for NECI and CPMD-NECI), can be checked out of the subversion repository:

```
svn checkout https://wmm.ch.cam.ac.uk/svn2/groups/alavi/testcode testcode
```

Please see the testcode documentation for more details.

# Run

## 3.1 NECI

**Note:** How to obtain FCIDUMP/density-fitting input files?

```
neci.x input_file
```

If no file is given, then it takes input options from STDIN. This is rarely useful, however.

NECI prints output to SDTOUT, so output needs to be captured in some way:

```
neci.x input_file > output_file  
neci.x nput_file | tee output_file
```

## 3.2 CPMD-NECI

The converged Kohn–Sham orbitals obtained from a **OPTIMIZE WAVEFUNCTION** CPMD calculation can be used as input for a NECI calculation.

In contrast to the molecular case, NECI calculations based upon CPMD-generated wavefunctions are called from within CPMD itself. This allows us to take advantage of many routines that CPMD already possesses (FFT routines, initialisation, reading in the wavefunctions etc.).

To run, specify **QMC** in the **&CPMD** section of the CPMD input file. **RESTART WAVEFUNCTIONS OCCUPATION DENSITY COORDINATES LATEST** must also be specified. Running CPMD (assuming it has been correctly compiled with the appropriate NECI library) then calls NECI to read the NECI input file and perform the desired calculation.

For gamma-point calculations:

```
gcpmd.x input_file > output_file
```

For k-point calculations:

```
kcpm.d.x input_file > output_file
```

There are many other appropriate options that can be specified in the CPMD input file rather than the NECI input file. Please see the CPMD manula and the local CPMD documentation detailing additions the Alavi group has made.



# Theoretical review

## 4.1 Introduction

The energy of a system can be evaluated using a standard statistical mechanics result:

$$E = \frac{\text{Tr}[He^{-\beta H}]}{\text{Tr}[e^{-\beta H}]} \quad (4.1)$$

We choose to work in a Slater Determinant space, which is, by construction, anti-symmetric. In this space the energy expression becomes:

$$\begin{aligned} E &= \frac{\sum_{\mathbf{i}} \langle D_{\mathbf{i}} | He^{-\beta H} | D_{\mathbf{i}} \rangle}{\sum_{\text{veci}} \langle D_{\mathbf{i}} | e^{-\beta H} | D_{\mathbf{i}} \rangle} \\ &= \frac{\sum_{\mathbf{i}} w_{\mathbf{i}} \tilde{E}_{\mathbf{i}}}{\sum_{\mathbf{i}} w_{\mathbf{i}}} \end{aligned} \quad (4.2)$$

A given term in the numerator is simply the differential of the corresponding term in the denominator. There is a cleaner and more efficient way of evaluating the numerator than differentiation, but we will first turn our attention to the denominator.

We can expand each term into a closed path of  $P$  steps through the discrete Slater Determinant space:

$$\begin{aligned} w_{\mathbf{i}, \mathbf{i}_1, \mathbf{i}_2, \dots, \mathbf{i}_P, \mathbf{i}}^P &= \sum_{\mathbf{i}_1} \sum_{\mathbf{i}_1} \cdots \sum_{\mathbf{i}_P} \langle D_{\mathbf{i}_1} | e^{\beta H/P} | D_{\mathbf{i}_2} \rangle \langle D_{\mathbf{i}_2} | e^{\beta H/P} | D_{\mathbf{i}_3} \rangle \cdots \langle D_{\mathbf{i}_P} | e^{\beta H/P} | D_{\mathbf{i}_1} \rangle \\ &= \sum_{\mathbf{i}_1} \sum_{\mathbf{i}_1} \cdots \sum_{\mathbf{i}_P} \rho_{\mathbf{i}_1 \mathbf{i}_2} \rho_{\mathbf{i}_2 \mathbf{i}_3} \cdots \rho_{\mathbf{i}_P \mathbf{i}_1}, \end{aligned} \quad (4.3)$$

where the  $\rho$  matrix consists of elements  $\rho_{\mathbf{ij}} = \langle D_{\mathbf{i}} | e^{\beta H/P} | D_{\mathbf{j}} \rangle$ .

Each path does not necessarily visit  $P - 1$  determinants: “hopping” terms are allowed. Due to the  $\rho$  matrix being diagonally dominant, paths containing small numbers of unique determinants will tend to have a much greater contribution to the overall energy.

The size of the Slater determinant space grows factorially with the number of electrons and virtual orbitals, making it impossible to sum together all the paths. Furthermore, the sign of a path is an incredibly poorly behaved quantity. It is possible to perform an analytical resummation of the paths into objects we term graphs, where each graph contains paths which only visit the vertices contained within the graph.

**Note:** To come: pictures of paths  $\rightarrow$  graph.

The expression for the energy now becomes a sum over Slater determinants and a sum graphs which originate from each Slater determinant:

$$E = \frac{\sum_{\mathbf{i}} \sum_G w_{\mathbf{i}}[G] \tilde{E}_{\mathbf{i}}[G]}{\sum_{\mathbf{i}} \sum_G w_{\mathbf{i}}[G]} \quad (4.4)$$

Furthermore, the graphs have a much better sign behaviour: there are many graphs with a definite-positive weight, at least for graphs with less than 5 vertices, which makes a Monte Carlo approach feasible.

The resummation of paths into graphs still leaves a sum that is far too large to be completely evaluated. There are various approximations we can apply.

1. Use a single reference reference approach, i.e. approximate the energy with:

$$E = \frac{\sum_G w_0[G] \tilde{E}_0[G]}{\sum_G w_0[G]} \quad (4.5)$$

where 0 refers to the reference (i.e. Hartree–Fock) determinant.

This sum, in general, still contains too many terms (and grows too rapidly with system size) to be of much use.

2. Truncate the sum at a certain graph size (e.g. restrict it to two or three vertices). This approach is referred to as a **VERTEX SUM** approach.
3. Find a large graph that is a good approximation to the ground state and is easy to evaluate. Our current model is the single and double excitation star, which contains all single and double excitations connected to the reference determinant but ignores any connections between the excited determinants. In other words, it couples all the single and double excitations together, but only through the reference determinant. This method is referred to as a **VERTEX STAR** approach. The star contains all graphs in the sum truncated at the two vertex level and much more but at no additional costly integrals to evaluate. This makes it a very attractive approach.

**Note:** To come: the propagation operator.

## 4.2 Graph evaluation

There are two approaches to evaluating the weight and energy contribution of a given graph: either by diagonalising the  $\rho$  matrix of the graph or by diagonalising the Hamiltonian matrix of the graph.

### 4.2.1 RHODIAG

Diagonalisation of the  $\rho$  matrix is referred to as **RHODIAG** in the input documentation.

The  $\rho$  matrix of the graph is the evaluation of the high-temperature thermal density operator on the space of Slater determinants spanned by the graph, or more formally:

$$\rho[G] = \sum_{\mathbf{ij} \in G} |D_{\mathbf{i}}\rangle \rho_{\mathbf{ij}} \langle D_{\mathbf{j}}| \quad (4.6)$$

We can obtain the eigenvectors and -values,  $\{v_k\}$  and  $\{\lambda_k\}$  of  $\rho[G]$  via matrix diagonalisation, and can then use them to evaluate the weight of the graph:

$$\begin{aligned} w_{\mathbf{i}}[G] &= \langle D_{\mathbf{i}} | e^{-\beta H[G]} | D_{\mathbf{i}} \rangle \\ &= \sum_{kl} \langle D_{\mathbf{i}} | v_k \rangle \langle v_k | e^{-\beta H[G]} | v_l \rangle \langle v_l | D_{\mathbf{i}} \rangle \\ &= \sum_{kl} \langle D_{\mathbf{i}} | v_k \rangle \langle v_k | \lambda_l^P | v_l \rangle \langle v_l | D_{\mathbf{i}} \rangle \\ &= \sum_{kl} \langle D_{\mathbf{i}} | v_k \rangle \lambda_l^P \delta_{kl} \langle v_l | D_{\mathbf{i}} \rangle \\ &= \sum_k \lambda_k^P \langle D_{\mathbf{i}} | v_k \rangle \langle v_k | D_{\mathbf{i}} \rangle \end{aligned} \quad (4.7)$$

where we have applied the identity operator,  $\sum_k |v_k\rangle \langle v_k|$  twice and used:

$$e^{-\beta H[G]} |v_l\rangle = \rho[G]^{P-1} \lambda_l |v_l\rangle. \quad (4.8)$$

In a similar fashion, the energy contribution,  $w_i \tilde{E}_i$  can be evaluated:

$$\begin{aligned}
 w_i \tilde{E}_i &= \langle D_i | H e^{-\beta H[G]} | D_i \rangle \\
 &= \sum_{j \in G} \langle D_i | H | D_j \rangle \langle D_j | e^{-\beta H[G]} | D_i \rangle \\
 &= \sum_{j \in G} \sum_{kl} \langle D_i | H | D_j \rangle \langle D_j | v_k \rangle \langle v_k | e^{-\beta H[G]} | v_l \rangle \langle v_l | D_i \rangle \\
 &= \sum_{j \in G} \sum_{kl} \langle D_i | H | D_j \rangle \langle D_j | v_k \rangle \lambda_l^P \delta_{kl} \langle v_l | D_i \rangle \\
 &= \sum_{j \in G} \sum_k \langle D_i | H | D_j \rangle \lambda_k^P \langle D_j | v_k \rangle \langle v_k | D_i \rangle
 \end{aligned} \tag{4.9}$$

The  $\rho$  matrix elements can be evaluated using a Taylor expansion with or without a Trotter approximation to improve the accuracy of the expansion.

### 4.2.2 HDIAG

Alternatively, we can use a slightly simpler approach which avoids having to evaluate  $\rho$  matrix by dealing with the Hamiltonian matrix directly. This method is referred to as **HDIAG** in the input documentation. The two approaches give essentially the same result. In an analogous fashion to the application of the *rho* matrix in the space of the graph, we consider the Hamiltonian to be a propagator acting in the space of the graph:

$$H[G] = \sum_{ij \in G} |D_i\rangle H_{ij} \langle D_j| \tag{4.10}$$

We can evaluate use this to evaluate the weight of the graph:

$$\begin{aligned}
 w_i[G] &= \langle D_i | e^{-\beta H[G]} | D_i \rangle \\
 &= \sum_{kl} \langle D_i | v_k \rangle \langle v_k | 1 - \beta H[G] + \frac{\beta^2 H[G]^2}{2!} - \frac{\beta^3 H[G]^3}{3!} + \dots | v_l \rangle \langle v_l | D_i \rangle \\
 &= \sum_{kl} \langle D_i | v_k \rangle (1 - \beta \lambda_l + \frac{\beta^2 \lambda_l^2}{2!} - \frac{\beta^3 \lambda_l^3}{3!} + \dots) \delta_{kl} \langle v_l | D_i \rangle \\
 &= \sum_k e^{-\beta \lambda_k} \langle D_i | v_k \rangle \langle v_k | D_i \rangle
 \end{aligned} \tag{4.11}$$

where now  $\{v_k\}$  and  $\{\lambda_k\}$  are eigenvectors and -values of the Hamiltonian matrix in the space of the graph.

Similarly, we can obtain the energy contribution of the graph:

$$\begin{aligned}
 w_i \tilde{E}_i &= \langle D_i | H e^{-\beta H[G]} | D_i \rangle \\
 &= \sum_{j \in G} \langle D_i | H | D_j \rangle \langle D_j | e^{-\beta H[G]} | D_i \rangle \\
 &= \sum_{j \in G} \sum_{kl} \langle D_i | H | D_j \rangle \langle D_j | v_k \rangle \langle v_k | 1 - \beta H[G] + \frac{\beta^2 H[G]^2}{2!} - \frac{\beta^3 H[G]^3}{3!} + \dots | v_l \rangle \langle v_l | D_i \rangle \\
 &= \sum_{j \in G} \sum_{kl} \langle D_i | H | D_j \rangle \langle D_j | v_k \rangle e^{-\beta \lambda_l} \delta_{kl} \langle v_l | D_i \rangle \\
 &= \sum_{j \in G} \sum_k \langle D_i | H | D_j \rangle e^{-\beta \lambda_k} \langle D_j | v_k \rangle \langle v_k | D_i \rangle
 \end{aligned} \tag{4.12}$$

# Input options

## 5.1 Overview

The NECI input file is keyword driven and requires a minimal amount of information.

The NECI input file is divided into various sections, or input blocks: system, precalc, calc, integral and logging. Of these, only the system and calc blocks are compulsory: all others are optional. Inside each input block, it is possible to set a variety of options. There are also three types of keywords that exist outside of an input block.

The order of the input blocks is not important (but certain orders are more logical than others), and nor is the order within a block, unless an option is only valid when a logical statement is true, in which case the relevant keyword for the logical statement must precede its related keywords.

General points to note:

- The input file is not case sensitive. In the input documentation, the keywords are given in capitals and **emphasised** for clarity and options or data required are in square brackets.
- Parameters which follow a keyword ought to be on the same line as the keyword, but this isn't a strict requirement.
- A new line is required for each keyword, unless the keyword is an option of another keyword, in which case it ought to be on the same line.
- Blank lines are ignored.
- Comments are enclosed in parentheses.
- Data items are terminated by space or comma.
- Only the variables relevant to the desired run are required.
- Unknown keywords return an error message and stop the run.
- Sensible defaults are set, reducing the amount of information required from the input file. There exist different sets of default options, allowing a large set of variables to be set with one command.

The overall structure, with a reasonably logical layout, is:

**TITLE**

**DEFAULTS**

**SYSTEM** [system type]

[System options]

**ENDSYS**

**PRECALC**

[PreCalc options]

**ENDPRECALC****CALC**

[Calc options]

**ENDCALC****INTEGRAL**

[Integral options]

**ENDINT****LOGGING**

[Integral options]

**ENDLOG****END**

**Warning:** This is a work in progress. Many places (especially, but not exclusively, where noted) need to be expanded and/or improved.

In addition, the following keywords are valid options, but are *not* documented:

- CALCREALPROD
- CALCRHOPROD
- DELTAH
- DERIV
- DETPOPS
- DIAGSHIFT
- EQUILSTEPS
- EXCHANGE-ATTENUATE
- HAPP
- LINROOTCHANGE
- MAXVERTICES
- MODMPTHEORY
- RESUMFCIMC
- RHOAPP
- RHOELEMS
- SAVEPREVARLOGGING
- SHIFTDAMP
- STARPROD
- STEPSSHIFT
- SUMPRODII

In contrast, the following options are documented, but are *not* valid input options:

- BANDGAP
- EXCHANGE-DAMPING
- STOCHASTICTIME
- MPMODTHEORY
- SAVEPRECALCLOGGING

## 5.2 Non-block level options

The following options exist outside of any input block:

**TITLE** Takes the rest of the line as the title and prints it to output. Useful for labelling the output.

**DEFAULTS [ DEFAULT FEB08 ]** Default: **DEFAULT**. NECI has a default set of defaults (the **DEFAULT** set), which are sensible, safe defaults. The **FEB08** set of defaults reflect further work, and change the defaults as follows:

- Fock-Partition-Lowdiag is set in the integral block.
- RhoEpsilon=  $10^{-8}$  in the calc block.
- MCPATHS is set to be on in the logging block.

This can be specified anywhere in the input file outside of an input block. All other options in the input file override the defaults.

**END** End of input file. Not required, unless there is text after the input (e.g. comments or notes) which is not commented out or if the input file is given via STDIN.

## 5.3 System

**SYSTEM [system type]** Starts system block. The system type must be provided and specifies the basis upon which NECI performs a calculation. ORDER is only valid for some system types—see below.

[System options—see below.]

**ENDSYS** End the system input block.

The available system types fall into three categories:

- Read in data produced by a molecular computational chemistry package:

**READ** Read in an FCIDUMP file containing the molecular integrals.

**GENERIC** Synonym for **READ**.

**BINREAD** Read in an unformatted FCIDUMP file containing the molecular integrals.

**STARSTOREREAD** Only the integrals required for a double-excitation star calculation are read in from an FCIDUMP. The one-electron integrals, which we call TMat elements, are stored as integrals involving spatial orbitals, meaning that UHF is no longer available. In addition, only non-zero one-electron integrals are stored. The memory required to store the coulomb integrals is massively reduced, from  $\frac{M^4}{8}$  to just  $\frac{N^2 M^2}{2}$ , where  $M$  and  $N$  are the total number of orbitals and the number of occupied orbitals respectively. We only store the  $\langle ij|ab \rangle$  integrals in the UMAT array, where  $i$  and  $j$  are occupied, as well as the  $\langle ii|jj \rangle$  and  $\langle ij|ij \rangle$  integrals over all states in the UMAT2D array. Can only be used for the 2-vertex sum and the 2-vertex star calculations.

**STARBINREAD** Like **STARSTOREREAD**, but FCIDUMP is a binary file.

**DFREAD** Read in a set of density fitted coefficients and coulomb integrals from files SAV\_DFASOL and SAV-Ta-INT (generated by [CamCasp]). One-electron integrals are read in from HONEEL, which also contains  $\langle ij|ij \rangle$  and  $\langle ij|ji \rangle$  integrals (generated by readintOCC.x—a local package).

- Use a model system:

**BOX** Run a calculation on electrons confined to a box. See [TwoElBox] for more details.

**HUBBARD** Run a Hubbard model calculation.

**UEG** Run a uniform electron gas calculation.

- Periodic systems:

**CPMD** Perform a calculation based upon the Kohn–Sham wavefunctions produced by CPMD.  
Only available in a combined CPMD-NECI executable.

**VASP** Perform a calculation based upon the Hartree–Fock wavefunctions produced by VASP.  
Only available in a combined VASP-NECI executable.

**ORDER** If **ORDER** is specified directly after **READ**, **GENERIC**, **STARBINREAD**, **STARSTOREREAD**, **BINREAD**, or **DFREAD** then perform a quick HF calculation in the space of the orbitals. The orbitals are then reordered according to the HF energies, rather than using the orbital energies read in. If **CPMD** is followed by **ORDER**, then the CPMD orbitals are ordered, not according to their Kohn–Sham eigenvalues, but instead according to their one-electron energies (i.e. with no exchange or correlation). **ORDER** is not valid for any other system type.

### 5.3.1 General options

**BANDGAP** Perform calculations for systems containing NEL, NEL+1, and NEL-1 electrons and extract the band gap energy.

**COULOMB [FCOUL]** Multiply the strength of the coulomb interaction by FCOUL.

**COULOMB-DAMPING ENERGY [ $\mu$   $\beta$ ]** Damp the two-electron coulomb integrals,  $\langle ab||cd \rangle$  with factor  $f(E_a)f(E_b)f(E_c)f(E_d)$  where  $f(E_a) = \text{erfc}(\beta * (E_a - \mu))$ . A  $\beta$  of 1 gives a damping range of 2; a  $\beta$  of 40 gives a damping range of 0.05.

**COULOMB-DAMPING ORBITAL [ORB  $\beta$ ]** Damp the coulomb integrals as above, with MU set to be halfway between the energies of ORB and ORB+1.

**Note:** **COULOMB-DAMPING** is now disabled [26/7/06].

**CSF [STOT]** Default off. Default STOT=0.

If specified, work in CSFs rather than determinants. CSFs might not function properly for some Monte Carlo, but should work for vertex sums and diagonalization. STOT is twice the magnitude of spin to restrict the resultant space.

**ELECTRONS [NEL]** Specify the number of electrons. Required for all system types apart from CPMD- or VASP-based calculations.

**ENERGY-CUTOFF EMax** Default off.

Reject basis functions with an (unscaled) energy larger than EMax.

**EXCHANGE [ON | OFF]** Default ON.

Specify whether to include Exchange in the Slater-Condon rules. If off, we are effectively reduced to a using Hartree multi-electron wavefunctions rather than Slater determinants.

**NEL [NEL]** Synonym for **ELECTRONS**.

**NOSYMMETRY** Ignore all spatial symmetry information. This does not apply to periodic calculations.

**SPIN-RESTRICT [LMS]** Default off. Default LMS=0. Turns spin restriction on, limiting the working space to the z-component of spin being LMS\*2.

**STORE-AS-EXCITATIONS** Store determinants as a 4-integer list of orbitals excited from, and orbitals excited to, in comparison to the reference determinant, rather than as an n-electron list of the occupied orbitals in the determinant. This means that the scaling is reduced to  $N^2 M^2$  rather than  $N^3 M^2$ , as we run through the list for each excitation. Currently only working for the 2-vertex star Fock-Partition-Lowdiag calculations.

**SYM** [ $l_x, l_y, l_z$  **iSym**] Default off.

If specified, limit the working Slater determinant space to the set of determinants with the specified symmetry quantum numbers. The symmetry of a given orbital is specified in one of two ways:

**model system calculations:** 3 quantum numbers,  $l_x, l_y, l_z$ .

**molecular or periodic calculations:** Symmetry label, **iSym**, which corresponds to an irreducible representation of the symmetry group.

The symmetry label(s) of each orbital is included in the output, from which the symmetry of the desired set of Slater determinants can be evaluated (albeit in a somewhat laborious manner). All four numbers are required, but only the relevant one(s) are used.

For Abelian symmetry groups, each symmetry is printed out in terms of a propagating vector. Internally an integer label is still used, according to the formula:

$$i_{\text{SYM}} = \sum_{i=1}^3 p_i * 2^{15^{i-1}} \quad (5.1)$$

where  $p_i$  are the components of the propagating vector.

**USEBRILLOUINTHEOREM** Apply Brillouin's theorem: the net effect of single-excitations of the Hartree–Fock determinant coupled to the Hartree–Fock determinant is zero, so explicitly exclude such single excitations.

### 5.3.2 Model system options

The following apply to electron in a box, Hubbard model and uniform electron gas calculations, unless otherwise noted.

**BOXSIZE** [A [BOA COA] ] Required for **UEG** and **BOX** calculations. BOA and COA optional. Default BOA=COA=1.

Set lattice constants a, b and c respectively, where b and c are defined as a ratio of a.

**CELL** [NMAXX NMAXY NMAXZ] Maximum basis functions for each dimension. For **HUBBARD** and **UEG**, functions range from -NMAXi to NMAXi, but for **BOX**, they range from 1 to NMAXi, where i=X,Y,Z.

### 5.3.3 Box options

**ALPHA** [ $\alpha$ ] Sets TALPHA=.true. and defines  $\alpha$ .

Integrate out the Coulomb singularity by performing part in real space and part in Fourier space, with the division according to the screening parameter  $\alpha$ . See [TwoElBox].

**MESH** [NMSH] Default NMSH=32.

Number of mesh points used for calculating integrals.

### 5.3.4 Hubbard options

**B** [BHUB] Default=0.

Sets B (hopping or kinetic energy) parameter for the Hubbard model.

**U** [UHUB] Default=0.

Sets U (on-site repulsion) parameter for the Hubbard model.

**REAL** Set Hubbard model to be in real space.

**APERIODIC** Hubbard model is set to be not periodic.

**TILT** [ITILTX ITILTY] Default off.

The Hubbard model is tilted and the unit vectors are (x,y)=(ITILTX,ITILTY) and (-y,x). Require  $x \geq y$ .



### 5.3.5 UEG options

**EXCHANGE-CUTOFF** [ $R_c$ ] Use the method detailed in [AttenEx] for calculating the exchange integrals.

Sets cutoff distance  $R_c$  for the exchange electron-electron potential. If  $R_c$  is not explicitly set, it will be set to be equivalent to a sphere of the same volume as the cell,  $R_c = (\frac{\Omega}{4\pi/3})^{1/3}$ .

**EXCHANGE-DAMPING** [ $R_c$ ] Sets cutoff parameter  $R_c$  for attenuated potential  $V(r) = \frac{\text{erfc}(r/R_c)}{r}$ . If  $R_c$  is not explicitly set, it will be set to be equivalent to a sphere of the same volume as the cell,  $R_c = (\frac{\Omega}{4\pi/3})^{1/3}$ .

## 5.4 PreCalc

**Note:** George, please improve! My interpretations also need to be checked...

**PRECALC** Start pre-calculation block. This chooses which weighting parameters to use in Monte Carlo calculation, in order to give minimum variance. This is an optional input block, and is not required if the default parameters are to be used, or are specified explicitly in the **CALC** input block. Currently, only the **IMPORTANCE** parameter, the **C EXCITWEIGHTING** parameter, and the a & b optimal parameters are searched for simultaneously, optimised and parsed through the the main program.

[PreCalc options—see below.]

**ENDPRECALC** End the precalc input block.

### 5.4.1 PreCalc Options

**Note:** George: What on earth are the following:

- A,B etc parameters.
- XXX.
- U matrix.

**VERTEX** [**HDIAG RHODIAG**] [**SUM MC**] Similar to the methods section in the **CALC** block, specify the method to use at the next vertex level (the first entry is for the second vertex level) in searching for the best parameters, using the hamiltonian matrix diagonalisation (**HDIAG**) or the  $\rho$  matrix diagonalisation (**RHODIAG**). After this is specified, on the same line, specify whether to calculate the expected variance using the full sum at this vertex level (**SUM**), or using a Monte Carlo sum (**MC**).

Currently, only the **HDIAG** routine works when performing a MC expected variance, though the diagonalisation of the *rho* matrix now works with the full sum.

For example:

```
**VERTEX** **HDIAG** **SUM**
XXX
XXX
**VERTEX** **HDIAG** **MC**
XXX
```

where XXX are the vertex options (see below).

If no further options are specified for a given vertex level, the optimum values of all the **EXCITWEIGHTING** variables will be found, but not used in the main program.

**GRIDVAR** [**A\_ExcitFromStart**] [**A\_ExcitFromEnd**] [**A\_ExcitFromStep**] [**B\_ExcitToStart**] [**B\_ExcitToEnd**] [**B\_ExcitToStep**]

Produce a 3D map of the variance landscape, but do not explicitly calculate the minimum. Values for the A parameter start, end, and step must be specified, followed by the same for the B parameter.

**LINEVAR** [**G\_VMC\_PIS**start] [**G\_VMC\_PIE**nd] [**G\_VMC\_PIS**tep] Same as **GRIDVAR**, but produce a 1D line for one variable - currently only working for the **IMPORTANCE** parameter and the U-matrix element parameter. Shows change in expected variance over the designated range of values.

**MEMORISE** All the graphs, their excitation generators, weights, energies, and unbiased probabilities should be stored in the memory. Speeds up the calculation of the variance by around twofold. However, this cannot be used for large systems. ~31000 4v graphs, or 22000 3v graphs was the maximum for the nitrogen dimer with the VQZ basis .

Currently only available for **MC** precalculations. If this is not set, then only the first node excitation generators are stored—there should be enough memory for this

**PREGRAPHEPSILON** [**PREWEIGHTE**PS] Default  $10^{-8}$ .

Gives the threshold above which the weight of a graph must be if they are to be included in the full precalc variance calculation.

If the graph weight is below the threshold, then the probability of obtaining the graph does not need to be calculated, and so the optimisation routine is faster (but less accurate) at higher thresholds.

**TOTALERROR** [**Desired Error from main calculation**] Calculates the required number of cycles so that the final error from the Monte Carlo calculation is equal to the error specified.

Only valid if the highest vertex level in the precalculation stage is the same as the highest vertex level in the main MC calculation, i.e. the highest MC vertex level is independently optimised.

The optimum vertex splitting is also calculated.

**TOLERANCE** [**TOLERANCE**] Default 0.1.

The fractional precision to which the optimum parameter is obtained, using the minimisation method.

**TRUECYCLES** [**No.of cycles**] Specify the total number of MC cycles that we want to use in the main calculation. The precalculation stage will then automatically split the cycles between the vertex levels in the main calculation, according to how the use statements indicate the parameters are to be split.

The cycles are then split so to best minimise the overall variance of the run, assuming that the variance of the whole run is simply the sum of the variances of the individual vertex results.

## 5.4.2 Vertex options

For the specified vertex level:

**CYCLES** [**nCYCLES**] Specify the number of graphs generated in the MC algorithm for each expected variance calculation in the parameter minimisation algorithm.

Applies only to a vertex level which evaluates the expected variance using the MC algorithm in the **VERTEX** line.

**NONE** Perform no calculations or optimisations.

**UEPSILON** [**UEPSILON**] Default **UEPSILON** is 0.

Find the optimum **C EXCITWEIGHTING** coefficient and pass through to the main program, unless setting **C** to be zero changes the expected variance by an amount less than **UEPSILON**, in which case **C** is set to be zero.

The calculation of the U matrix elements can be time-consuming in a real MC simulation, yet can have a negligible effect on the final result. In these cases, setting the **C** coefficient to be zero makes the full MC simulation much faster.

With the the default value of **UEPSILON**, the optimum value of **C** will always be used in the main program.

**FINDC** Find the optimum **C EXCITWEIGHTING** parameter.

The **C EXCITWEIGHTING** parameter. will only be found if this flag is set, or if a **UEPSILON** is set.

By default, the optimisation algorithm will only seek to find the values which give the minimum expected variance by varying the **A** and **B EXCITWEIGHTING** parameters, (or the parameters in the weighting scheme specified).

**FINDD** Find the optimum D **EXCITWEIGHTING** parameter for this vertex level (g\_VMC\_ExcitToWeight2).

**USED** Pass the optimum D **EXCITWEIGHTING** parameter found at this vertex level through to the main calculation.

**FINDIMPORT** Run the optimisation algorithm for the IMPORTANCE parameter.

The optimised value will be found printed out, but will not be passed through to the main calculation.

Can only be set for vertex levels of three or higher for obvious reasons.

**Note:** And the obvious reasons are?

**USEIMPORT** Find the optimal IMPORTANCE parameter, and use in the main calculation.

As for **FINDIMPORT**, can only be set for vertex levels of three or higher.

**USE [MC\_VERTEX\_LEVEL\_1] [MC\_VERTEX\_LEVEL\_2] ...** Use in the main calculation the parameters calculated from the specified precalc vertex levels, rather than any other, are to be passed through and used in the main program when performing a MC at one of the vertex levels specified. Any given vertex level can only be specified once in all the **USE** statements.

Vertex levels in the main calculation which are not specified in one of the precalc **USE** statements, will use the parameters which are given in the **CALC** block of the input file.

A **USE** statement on its own will only calculate the A and B **EXCITWEIGHTING** parameters (or any of the other weighting scheme parameters specified) and use them for all MC vertex levels in the main calculation, unless **FINDC** or **UEPSILON** is specified, in which case for the C parameter to also be used.

**Note:** This is somewhat confusing. How does it fit in with USEIMPORT etc.?

## 5.5 Calc

**CALC** Start calculation block. This chooses what calculation to do.

[Calculation options—see below.]

**ENDCALC** End the calculation input block.

### 5.5.1 General options

**ALLPATHS** Choose all determinants (i.e. set NPATHS = -1).

**BETA [BETA]** Set  $\beta$ .

**BETAOVERP [BETAP]** Default= 1.d-4.

Set  $\beta/P$ .

**DELTABETA [DBETA]** Set  $\delta\beta$ . If given a negative value, calculate it exactly.

**Note:** What is this used for?

**DETINV [DETINV]** Specify the root determinant for which the complete vertex series is worked out, using the determinant index obtained from a previous calculation. If **DETINV** is negative, the NPATHS calculations are started at this determinant.

**EXCITE [ICILEVEL]** Default 0.

Excitation level at which to truncate determinant list. If ICILEVEL=0 then all determinants are enumerated.

**EXCITATIONS [OLD NEW]** For generation of up to double excitations use the old (completely reliable), or new (faster, but does not work for more than 2-vertex level SUMS) routine

**Note:** You can now use the **NEW** routines for all methods, right? What is the difference between **NEW** and **OLD**? (If it doesn't say, how else can a user make an informed decision as to which to use?)

**EXCITATIONS [SINGLES DOUBLES]** Default is to use all excitations.

Restricts determinants which are allowed to be connected to the reference determinant to be either single or double excitations of the reference determinant.

Applies only to the **VERTEX [SUM STAR] NEW** methods.

**HAMILTONIAN [STAR]** Store the Hamiltonian. This is defaulted to ON if **ENERGY** is set, but can be used without **ENERGY**.

**STAR** Only the connections between the root determinant and its excitations should be included in the Hamiltonian and not off-diagonal elements between excited determinants.

**MAXVERTICES [MAXVERTICES]** Give the vertex level of the calculation. Cannot be used in conjunction with a **METHODS** block.

**METHOD [Method option(s)]** Specify the method for a graph theory calculation. See Method options for the available methods.

Can only be specified once if used outside of the methods block, in which case the given method is applied to all vertex levels.

**METHODS** Begin a methods block. This allows a different method for each vertex level. Each vertex level can contain **EXCITATIONS**, **VERTICES**, **CYCLES** and **CALCVAR** keywords. Each **METHOD** line and the options that follow it detail the calculation type for the next vertex level, with the first **METHOD** line used for the the second-vertex level, unless over-ridden with the **VERTICES** option.

The block terminates with **ENDMETHODS**.

For example:

```
METHODS
  METHOD VERTEX SUM NEW
  EXCITATIONS DOUBLES
  METHOD VERTEX STAR POLY
  EXCITATIONS SINGLES
  VERTICES 2
ENDMETHODS
```

sets the first method, at the two-vertex level, to be a complete 2-vertex sum of only doubles, and the second method, overridden to be also at the two-vertex level, to be a vertex star of singles.

Similarly:

```
METHODS
  METHOD VERTEX SUM NEW
  METHOD VERTEX SUM MC
  [Monte Carlo options]
ENDMETHODS
```

performs a full sum at the two-vertex level and a Monte Carlo calculation at the three-vertex level.

**ENDMETHODS** Terminate a methods block.

**PATHS [option]** Select the number of determinants taken to be the root of the graph. Usually set to 1. Valid options:

**NPATHS** Choose the first NPATHS determinants and calculate RHOP11 etc.

**ALL** Choose all determinants (same as ALLPATHS).

**ACTIVE** Choose only the active space of determinants: the degenerate set containing the highest energy electron.

**ACTIVE ORBITALS nDown nUp** Set the active space to be nDown and nUp orbitals respectively from the Fermi level

**ACTIVE SETS nDown nUp** Set the active space to be nDown and nUp degenerate sets respectively from the Fermi level

**RHOEPSILON [RHOEPSILON]** Set the minimum significant value of an element in the *rho* matrix as a fraction of the maximum value in the *rho* matrix. Matrix elements below this threshold are set to be 0.

**STARCONVERGE [STARCONV]** Default 1.d-3.

Set the convergence criteria for whether a roots to the star graph is significant.

**TROTTER** Default.

Perform a Trotter decomposition to evaluate the *rho* matrix elements.

**TIMESTEPS [I\_P]** Set P, the timesteps into which  $e^{-\beta H}$  is split. Automatically sets  $\beta/P = 0$  (as required) but returns an error message if **BETAOVERP** is also used.

**WORKOUT [NDETNWORK]** Sets the number of determinants which are worked out exactly.

**Note:** What is this used for?

**VERTICES** Only available inside a methods block.

By default, each method takes a number of vertices corresponding to its index within the methods block, the first methods corresponding to the 2-vertex level, the second to the 3-vertex level, and so on. **VERTICES** overrides this, and allows the vertex level of each method to be explicitly specified, enabling, for example, the 2-vertex level to be split up and the contributions from single and double excitations of the reference determinant to be handled separately.

## 5.5.2 Method options

**VERTEX SUM [OLD NEW HDIAG] [SUB2VSTAR] [LOGWEIGHT]** Calculate the vertex sum approximation.

**OLD** Diagonalise the  $\rho$  matrix using the original method.

**NEW** Diagonalise the  $\rho$  matrix using a more modern, more efficient method. Recommended.

**HDIAG** Diagonalise the Hamiltonian matrix instead of the *rho* matrix in order to calculate the weight and energy contribution of each graph.

**SUB2VSTAR** Remove paths which were present in the 2-vertex star for each graph. If this is specified for ANY vertex level, it applies to all **SUM** and MC vertex levels.

**LOGWEIGHT** Form Q as a multiplication of factors from graphs. This results in the quantity  $\log w$  being used instead of  $w$ , which also translates to the energy expression only involving  $\bar{E}$  not weights. Hopefully this is size-consistent.

**Warning:** **SUB2VSTAR** and **LOGWEIGHT** are experimental options.

**VERTEX [MC MCMETROPOLIS MCDIRECT MCMP] [HDIAG]** Perform a Monte Carlo calculation.

**MCDIRECT** Perform direct stochastic sampling for the graph theory vertex sum method, dividing each freshly generated graph by its normalized generation probability.

If **MULTIMCWEIGHT** is specified then the sampling generates graphs from all weighted levels using the weighting - a single MC calculation is performed.

If **MULTIMCWEIGHT** is not specified (default), a separate MC calculation is performed at each vertex level. Combined statistics are printed.

**Warning:** **MULTIMCWEIGHT** is not documented. Use with great caution.

**MCMP** Perform direct stochastic sampling, as in **MCDIRECT**, but for the Moller–Plesset method.

**MC or MCMETROPOLIS** Perform Metropolis Monte Carlo.

This may be performed in a number of ways. The way is chosen by the location of the **VERTEX MC** command.

**Warning:** The following options appear in INPUT\_DOC but, however, are incredibly poorly documented. In particular:

- No detail on the arguments the options take (e.g. **BIAS**).
- Some options documented don't exist (e.g. **SINGLE**, **BIAS**, **MULTI**, **STOCHASTIC-TIME**).
- Sufficient tests are not present in the test suite.

Do not use.

The “options” are:

```

**STOCHASTICTIME**
    may also be specified to perform stochastic
    time simulations with a given **BIAS**

**SINGLE**
    MC is performed at a single vertex level using a composite
    1-vertex graph containing a full sum previously performed.

**BIAS**
    is used to choose whether a step selects a composite
    (all lower levels) or a normal (this level) graph. Stochastic
    time MC is performed. This can only be specified in the
    **METHODS** section, and only at the last vertex level.
    Uses **EXCITWEIGHTING** for excitation generation weighting
    and **IMPORTANCE** for graph generation weighting

**MULTI**
    MC is performed at a multiple vertex levels, but still
    using a composite 1-vertex graph containing a full sum
    previously performed. MULTI should be specified in all the
    (contiguous) vertex levels to be included (not composited)
    in the MC. **BIAS** is used to choose whether a step
    selects a composite (all lower levels) or a normal (the
    **MULTI** levels) graph. **MULTIMCWEIGHT** is specified
    for each **MULTI** level, and gives a relative weighting
    of selecting the vertex level graphs once a non-composite
    graph is chosen. Stochastic time MC is performed.
    This can only be specified in the **METHODS** section.
    Once **MULTI** has been specified, it must be specified
    on all subsequent vertex levels in a **METHODS** section.
    Uses **EXCITWEIGHTING** for excitation generation weighting
    and **IMPORTANCE** for graph generation weighting

**FULL**
    Does MC at all levels using BIAS to bias the levels,
    **EXCITWEIGHTING** for excitation generation, and
    **IMPORTANCE** to for graph generation weighting. This is
    only available *WITHOUT* a **METHODS** section. If **HDIAG**
    is specified, the H-diagonalizing routine is used, otherwise,
    the rho-diagonalizer is used. **HDIAG** is automatically
    specified for **MCMP**.

```

**VERTEX SUM READ** Read in from pre-existing MCPATHS file for that vertex level. Only really useful in a **METHODS** section.

**VERTEX STAR [ADDSINGLES COUNTXCITS] [star method] [OLD NEW [H0] ]** Construct a single and double excitation star from all determinants connected to the root (ignoring connections between those dets). See [[StarPaper](#)] for more details.

**ADDSINGLES** Extend the star graph approach.

Add the single excitations which are en-route to each double excitation to that double excitation as spokes, and prediagonalize the mini-star centred on each double excitation. For example, if the double excitation is (ij->ab), then singles (i->a),(i->b),(j->a) and (j->b) are created in a star with (ij->ab), the result diagonalized, and the eigenvalues and vectors used to create a new spoke of the main star graph. Only works with **NEW**.

**COUNTEXCITS** Run through all the symmetry allowed excitations first and count the connected determinants on the star. Enables the memory requirements to be reduced as only connected determinants need to be stored. However, the time taken is increased, as it is necessary to run through all determinants in the star twice. Especially useful for large systems with memory restraints, when density fitting has necessarily turned off symmetry. Also useful if a **RHOEPSILON** has been set to a large value so that many of the symmetry allowed excitations will be counted as disconnected.

**Note:** Useful for periodic calculations? Does it need just the symmetry info or the transition matrix elements as well?

**OLD** Use a pre-generated list of determinants using the excitation routine version specified in **EXCITATIONS OLD** or **EXCITATIONS NEW**.

**NEW** Generate determinants on the fly without storing them, using the **NEW** excitation routine. Much more memory efficient.

**NEW H0** Use the zeroth order N-particle Hamiltonian (shifted such that  $H_{ii}^0 = H_{ii}$ ) rather than the fully interacting Hamiltonian to generate the roots of the polynomial.

**Note:** And you'd want to use **NEW H0** why exactly?

The available star methods are:

**DIAG** Perform a complete diagonalization on the resultant matrix. This can be very slow. However, by specifying **LANCZOS** in the **CALC** block, you can do a Lanczos diagonalisation, which scales much better. **EIGENVALUES** can also be specify to only evaluate the first few eigenvalues.

**POLY** Use the special properties of the matrix to find the roots of the polynomial and uses them to calculate the relevant values. This is order  $N_{\text{graph}}^2$ .

**Note:**  $N_{\text{graph}} = n_{\text{Dets}}$

**POLYMAX** Similar to **POLY** but only finds the highest root of the polynomial, so is order  $N_{\text{graph}}$ . It can be used when  $P$  is very large (i.e.  $\beta$  is very large, e.g. 40).

**POLYCONVERGE** Similar to **POLY** but adds  $i$  out of  $N$   $\lambda_i$  roots, such that  $(N-i)\lambda_i^P < 10^{-3}$ , i.e. we evaluate enough roots such that a very conservative error estimate of the contribution of the remaining roots is negligible.

**POLYCONVERGE2** Similar to **POLYCONVERGE** but requires  $w(1..i)(N-i)\lambda_i^P < 10^{-3}$ , where  $w(1..i)$  is the cumulative sum of  $\lambda_i^P$ , which should be a better estimate of the convergence.

The following are experimental star methods:

**MCSTAR** Use a basic implementation of the spawning algorithm in order to sample the star graph stochastically. The sampling uses elements of the Hamiltonian matrix rather than the *rho* matrix, so there will be some differences in the converged energy compared to a **VERTEX STAR NEW** calculation.

Many of the **FCIMC** options are also available with MCStar, and there are also some extra one.

**NODAL** Prediagonalise a completely connected set of virtuals for each set of occupied (i,j) spin-orbitals. The diagonalised excitations are then solved as a star graph. Must be used with **NEW**.

**STARSTARS** Use an approximation that the change of eigenvalues and the first element of the eigenvectors of the star graph is linear with respect to multiplying the diagonal elements by a constant. Once this scaling is found, all stars of stars are prediagonalised, and reattached to the original graph. This results in  $N^2$  scaling, where  $N$  is the number of excitations.

**TRIPLES** Prediagonalise an excited star of triple excitations from each double excitation, reattach the eigenvectors, and solves the complete star. Currently only available with '**NEW**', '**COUNTEXCITS**' and '**DIAG**'.

## Experimental methods

**VERTEX FCIMC [MCDIFFUSION]** Perform Monte Carlo calculations over pure determinant space, which is sampled using a series of ‘particles’ (or ‘walkers’).

The walkers are not necessarily unique and must be sorted at every iteration. Each walker has its own excitation generator.

**MCDIFFUSION** is a completely particle-conserving diffusion algorithm and is much more experimental.

**FCIMC** and **MCDETS** calculations share many of the same options (see Walker Monte Carlo options, below).

**VERTEX GRAPHMORPH [HDIAG]** Set up an initial graph and systematically improve it, by applying the *rho* matrix of the graph and its excitations as a propagator on the largest eigenvector of the graph. From this, an improved graph is stochastically selected, and the process is repeated, lowering the energy. If **HDIAG** is specified, it is the hamiltonian matrix elements which determine the coupling between determinants, and it is the hamiltonian matrix which is diagonalised in each iteration in order to calculate the energy.

**Note:** **GRAPHMORPH** has not been tested with complex wavefunctions. It will almost certainly not work for them.

**VERTEX MCDETS** Perform Monte Carlo calculations over pure determinant space, which is sampled using a series of ‘particles’ (or ‘walkers’).

**MCDETS** is similar to **FCIMC** but maintains at most one ‘particle’ at each determinant, which may then contain subparticles (which correspond to the individual ‘walkers’ in **FCIMC**), in a binary tree. This makes some efficiency savings where the same information about a determinant is not duplicated.

**FCIMC** and **MCDETS** calculations share many of the same options (see Walker Monte Carlo options, below).

**VERTEX RETURNPATHMC** Use a spawning algorithm which is constrained in three ways:

- #. a particle can only be spawned where it will increase its excitation level with respect to the reference determinant or back to where it was spawned from.
- #. they will spawn back to where their parents were spawned from with probability **PRet**, which is specified using **RETURNBIAS**.
- # length of spawning chain must be less than the maximum length given by **MAXCHAIN-LENGTH**.

**Note:** How can a particle be restricted to spawning to spawning at most back to where it was spawned from *and* have a probability of spawning back to where its parent was spawned from? Documentation *must* be clearer.

This attempts to circumvent any sign problem in the double excitations and the HF, and hopefully this will result in a more stable MC algorithm. It remains to be seen if this approach is useful. Should revert to the star graph in the limit of the return bias tending to 1 or the length of the spawn chain tending to 1.

**Note:** **FCIMC**, **GRAPHMORPH**, **MCDETS** and **RETURNPATHMC** have not been tested with complex wavefunctions. It will almost certainly not work for them.

All four are experimental options under development.

## 5.5.3 Walker Monte Carlo options

The following options are applicable for both the **FCIMC** and **MCDETS** methods:

**Note:** I have made some guesses on the following option names. Clearly some keys are broken on George’s keyboard. Specifically:

```
StepsSft --> STEPSSHIFT
SftDamp  --> SHIFTDAMP
DiagSft  --> DIAGSHIFT
```



I also had to guess about **BINCANCEL**. It seems to be a **FCIMC** option, but was placed with **MCSTAR** (and was with all the **VERTEX STAR** methods).

This section needs to be extended substantially.

**DIAGSHIFT** [**DiagSft**] Set the initial value of the diagonal shift.

**INITWALKERS** [**nWalkers**] Default 3000.

Set the initial population of walkers.

**NMCYC** [**NMCYC**] Set the total number of timesteps to take.

**SHIFTDAMP** [**SftDamp**] Damping factor (OF WHAT?!). <1 means more damping.

**STEPSSHIFT** [**StepsSft**] Default 100.

Set the number of steps taken before the diagonal shift is updated.

**TAU** [**TIMESTEP**] Default 0.0.

Set the number of evolution timesteps before the EShift is updated. Must be set such that  $\max(H_{jj}, |H_{ij}|) * \tau < 1$ .

**Note:** Brilliantly, there's a comment in the code which states (contradictorily):

```
For FCIMC, this can be considered the timestep of the simulation.
If not set, it will default to the value which will allow the
fastest destruction of walkers.
```

The following options are only available in **FCIMC** calculations:

**READPOPS** Read the initial walker configuration from the file POPSFILE. **DIAGSHIFT** and **INITWALKERS** given in the input will be overwritten with the values read in from POPSFILE.

**SCALEWALKERS** [**fScaleWalkers**] Scale the number of walkers by fScaleWalkers, after having read in data from POPSFILE.

**STARTMP1** Set the initial configuration of walkers to be proportional to the MP1 wavefunction.

**GROWMAXFACTOR** [**GrowMaxFactor**] Default 9000.

Set the factor by which the initial number of particles are allowed to grow before they are culled.

**CULFACTOR** [**CullFactor**] Default 5.

Set the factor by which the total number of particles is reduced once it reaches the GrowMaxFactor limit

**NOBIRTH** Force the off-diagonal *rho* (?) matrix elements to become zero, and hence obtain an exponential decay of the initial populations on the determinants, at a rate which can be exactly calculated and compared against.

**MCDIFFUSE** [**Lambda**] Default 0.0.

Set the amount of diffusion compared to spawning in the **FCIMC** algorithm.

**FLIPTAU** [**FlipTauCyc**] Default: off.

Cause time to be reversed every FlipTauCyc cycles in the **FCIMC** algorithm. This might help with under-sampling problems.

**NON-PARTCONSDIFF** Use a separate partitioning of the diffusion matrices, in which the antidiffusion matrix (+ve connections) create a net increase of two particles.

**FULLUNBIASDIFF** Fully unbiased for the diffusion process by summing over all connections.

**NODALCUTOFF** [**NodalCutoff**] Constrain a determinant to be of the same sign as the MP1 wavefunction at that determinant, if the normalised component of the MP1 wavefunction is greater than the NodalCutoff value.

**NOANNIHIL** Remove the annihilation of particles on the same determinant step.

The following option are only available in **MCSTAR** calculations:

**BINCANCEL** This is a separate method to cancel down to find the residual walkers from a list, involving binning the walkers into their determinants. This has to refer to the whole space, and so is much slower. See also the **WAVEVECTORPRINT** and **POPSFILE** options in the **LOGGING** block.

### 5.5.4 Return Path Monte Carlo options

**MAXCHAINLENGTH [CLMAX]** Set the maximum allowed chain length before a particle is forced to come back to its origin.

**RETURNBIAS [PRet]** Set the bias at any point to spawn at the parent determinant.

### 5.5.5 Perturbation theory options

**MPTHEORY** In addition to doing a graph theory calculation, calculate the Moller–Plesset energy to the same order as the maximum vertex level from the reference determinant (e.g. with 2-vertex sum the MP2 energy is obtained, with 3-vertex the MP3 energy etc. Note that the MP2 energy can be obtained in conjunction with a **VERTEX STAR** calculation.

**Note:** INPUT\_DOC has this to say:

```
Instead of a normal path-integral expansion **MC**, do a
Moller--Plesset. Requires **HDIAG**, and **BIAS**=0.D0. Can be
used without a **METHODS** section. If a **METHODS** section is
needed to specify different numbers of cycles at each level, then
**MCDIRECTSUM** must also be set.
```

I am sure this is out of date...

**EPSTEIN-NESBET** Apply Epstein–Nesbet perturbation theory, rather than Moller–Plesset. Only works for **VERTEX SUM NEW** and **VERTEX SUM HDIAG** and only at the 2-vertex level.

**LADDER** Use ladder diagram perturbation theory, rather than Moller–Plesset. The energy denominator is  $E_0 - E_I + |H_{0I}|^2$ . Only works for **VERTEX SUM NEW** and **VERTEX SUM HDIAG** and only at the 2-vertex level.

**MPMODTHEORY** Perform a hybrid of Epstein–Nesbet and Moller–Plesset theory, which includes only the  $\langle ij||ijket + \langle ab||abket$  terms in the denominator. Only works for **VERTEX SUM NEW** and **VERTEX SUM HDIAG** and only at the 2-vertex level.

### 5.5.6 Diagonalisation options

Options for performing a full diagonalisation in the space of the full basis of spin orbitals.

**Warning:** This quickly becomes prohibitively expensive as system size increases.

**ACCURACY [B2L]** Desired level of accuracy for Lanczos routine.

**BLOCK [ON OFF]** Default off.

Determines whether the Hamiltonian is calculated for each block or not. This only works for **COMPLETE**.

**BLOCKS [NBLK]** Set number of blocks used in Lanczos diagonalisation.

**COMPLETE** Perform a full diagonalisation working out all eigenvectors and eigenvalues. if **HAMILTONIAN** is **OFF**, discard the eigenvectors and eigenvalues after having used them for calculation. Relevant options are **HAMILTONIAN** and **BLOCK**.

**Note:** When would it be advantageous to save the eigenvalues and -vectors are a diagonalisation?

**EIGENVALUES [NEVAL]** Required number of eigenvalues.

**ENERGY** Calculate the energy by diagonalising the Hamiltonian matrix. Requires one of **COMPLETE**, **LANCZOS**, or **READ** to be set.

Exact  $E(\text{Beta})$  is printed out as:

$$E(\text{Beta}) = \frac{\sum_{\alpha} E_{\alpha} e^{-\beta E_{\alpha}}}{\sum_{\alpha} e^{-\beta E_{\alpha}}} \quad (5.2)$$

The result will, of course, change depending upon the symmetry subspace chosen for diagonalization for finite temperatures.

The diagonalization procedure creates a list of determinants, which is printed out to the DETS file.

The weight,  $w_i$  and weighted energy,  $w_i \tilde{E}_i$  are also calculated for all NPATH determinants.

**Note:** **ENERGY** was documented twice in the INPUT\_DOC file. This is not particularly helpful...

I have (hopefully) combined them correctly.

**KRYLOV [NKRY]** Set number of Krylov vectors.

**LANCZOS** Perform a Lanczos block diagonalisation on the Hamiltonian matrix.

Relevant parameters are **BLOCKS**, **KRYLOV**, **ACCURACY**, **STEPS** and **EIGENVALUES**.

**READ** Read in eigenvectors and eigenvalues of the Hamiltonian matrix from a previous calculation.

**STEPS [NCYCLE]** Set the number of steps used in the Lanczos diagonalisation.

## 5.5.7 Graph morphing options

A new approach developed by George Booth. Take an initial starting graph and over many iterations allow the determinants contained within the graph to change, so that the resultant graph is a better approximation to the true ground state.

**GRAPHBIAS [GraphBias]** If at each iteration the graph is being completely renewed, then this bias specifies the probability that an excitation of the previous graph is selected to try and be attached, rather than one of the determinants in the previous graph.

**GRAPHSIZE [NDets]** Specify the number of determinants in the graph to morph.

**GROWGRAPHSEXPO [GrowGraphsExpo]** Default is 2.D0.

The exponent to which the components of the excitations vector and the eigenvector are raised in order to turn them into probabilities. The higher the value, the more that larger weighted determinants will be favoured, though this might result in the graph growing algorithm getting stuck in a region of the space.

**GROWINITGRAPH** Grow the initial graph non-stochastically from the excitations of consecutive determinants.

**INITSTAR** Set up the completely connected two-vertex star graph, and use as the starting point for the morphing.

Automatically changes the NDets parameter to reflect the number of double excitations in the system.

**ITERATIONS [Iters]** The number of graph morphing iterations to perform.

**MAXEXCIT [iMaxExcitLevel]** Limit the size of the excitation space by only allowing excitations out to iMax-ExcitLevel away from HF reference determinant.

**MCEXCITSPACE [NoMCExcits]** Stochastically sample the space of excitations from each determinant in the graph with NoMCExcits determinants chosen per determinant.

**MOVEDETS [NoMoveDets]** Grow the graphs using an alternative Monte Carlo, where a number of determinants are deleted from the previous graph and reattached elsewhere in the graph in a stochastic manner, according to the probabilities given by the application of the *rho* propagator to the eigenvector of the previous graph.

**NOSAMEEXCIT** Ignore the connections between determinants which are of the same excitation level in comparison to the reference determinant. Currently only available in conjunction with **INITSTAR**, so the starting graph is simply the doubles star graph (with no cross connections).

**ONEEXCITCONN** Grow the graph by attaching only determinants which differ by one excitation level to the connecting vertex in the previous graph. Currently not implemented with MoveDets.

**SINGLESEXCITSPACE** Restrict the space into which the current graph is allowed to morph to just single excitations of the determinants in the current graph. This should reduce the scaling of the algorithm.

### 5.5.8 Monte Carlo options

Options for performing a Monte Carlo calculation on a vertex sum (as specified in the **METHODS** section).

The Monte Carlo routines have only ever been tested for molecular and model systems and probably are not currently functional for **CPMD** or **VASP** based calculations.

See the reports by Ramin Ghorashi ([[RGPIII](#)]) and George Booth ([[GHBCPGS](#)]).

**CALCVAR** Only available for performing full vertex sums using the **HDIAG** formulation to evaluate the thermal density matrix elements.

Calculate a theoretical approximation to the expected variance if a non-stochastic MC run were to be performed, with the parameters given, at the chosen vertex level. Currently the expected variance is sent to **STOUT** as a full variance for the total energy ratio. Causes the calculation to take longer since the generation probabilities of the graphs must all be calculated. The sum over graphs of the generation probabilities is also printed out for each vertex level. This should equal 1, since we are working with normalised probabilities.

**POSITION [IOBS JOBS KOBS]** Sets the position of the reference particle.

**CIMC** Perform a configuration interaction space Monte Carlo.

**BETAEQ [BETAEQ]** Default is set to be  $\beta$ , as set above.

Set  $\beta$  to have a different value for the equilibration steps.

**Note:** What are the equilibration steps?

**BIAS [G\_VMC\_FAC]** Default 16.

Vertex level bias for **FULL MC**. Positive values bias toward larger graphs, negative values towards smaller graphs.

For **SINGLE** and **MULTI** level MC (using a composite 1-vertex graph containing a full sum previously performed), this is the probability of generating a graph which is not the composite graph. The default is invalid, and this must be set manually. Stochastic time MC is used. If **BIAS** is negative, then  $|\text{BIAS}|$  is used, but stochastic-time MC is not performed.

**Note:** **BIAS** seems to do two very different things if it is set to a negative value. Please clarify.

**DETSYM [MDK(I), I=1,4]** The symmetry of the **CIMC** determinant.

**Note:** Specify the symmetry how?

**Note:** If any if the **CIMC** options are set without **CIMC** being specified, the code will return an error and exit.\*\*

**EQSTEPS [IEQSTEPS]** The number of equilibration sets for the CI space Monte Carlo routine.

**GRAPHEPSILON [GRAPHEPSILON]** Default 0.0.

The minimum significant value of the weight of a graph.

Ignore the contributions to the weight and  $\tilde{E}$  of all graphs with a weight that is smaller in magnitude than **GRAPHEPSILON**.

**IMPORTANCE [G\_VMC\_PI]** Ddefault 0.95.

Set the generation probability for the MC routine. This is the probability that new determinants are excitations of the pivot, i.

**MCDIRECTSUM** Perform Monte Carlo on graphs summing in energies weighted with the weight/generation probability of the graph.

**PGENEPSILON [PGENEPSILON]** Default 0.0.

Set the minimum significant value of the generation probability of a graph.

Because for larger graphs, the calculation of the generation probability is subject to numerical truncation errors, generation probabilities which are lower than a certain value are unreliable, and can cause the Monte Carlo algorithm to get stuck: if a graph had a very small generation probability, it would be difficult for a Monte Carlo run to accept a move to a different graph. If the magnitude of the generation probability of a graph is smaller than PGENEPSILON, then a new graph is generated.

Setting this too high could cause problems in the graph generation phase, so NECI will exit with an error if it generates 10000 successive graphs each with generation probabilities below PGENEPSILON.

**SEED [G\_VMC\_SEED]** Default -7.

Set the random seed required for the Monte Carlo calculations.

**STEPS [IMCSTEPS]** Set the number of steps for the CI space Monte Carlo routine.

**VVDISALLOW** Disallow V-vertex to V'-vertex transitions in stochastic time Monte Carlo: i.e. allow only transitions to graphs of the same size.

## Weighting schemes

By default the vertex sum Monte Carlo algorithm selects excitations with no bias. The variance of a Monte Carlo calculation can be reduced by preferentially selecting for certain types of excitation.

**EXCITWEIGHTING [g\_VMC\_ExcitFromWeight g\_VMC\_ExcitToWeight G\_VMC\_EXCITWEIGHT] [g\_VMC\_ExcitToW**

Default 0.d0 (unweighted) for all values.

A weighting factor for the generation of random excitations in the vertex sum Monte Carlo. A parameter set to zero has a corresponding weighting factor of 1.

For generating an excitation from occupied spin orbitals i and j to unoccupied spin orbitals k and l:

- the probability of choosing pair (ij) is proportional to

$$e^{(E_i + E_j)g\_VMC\_ExcitFromWeight} \quad (5.3)$$

- the probability of choosing pair (kl) is proportional to

$$e^{-(E_k + E_l)g\_VMC\_ExcitToWeight} e^{(\langle ij|U|kl \rangle * G\_VMC\_EXCITWEIGHT | E_i + E_j - E_k - E_l | g\_VMC\_ExcitToWeight2} \quad (5.4)$$

**POLYEXCITWEIGHT [g\_VMC\_ExcitFromWeight g\_VMC\_PolyExcitToWeight1 g\_VMC\_PolyExcitToWeight2 G\_VMC\_P**

Default 0.0 for all values (i.e. unweighted: all weighting factors are set to 1).

Weighting system for the choice of virtual orbitals in the excitations.

The probability of choosing the pair of spin orbitals, kl, to excite to is set to be constant for  $E_k + E_l$  is less than  $g\_VMC\_PolyExcitToWeight1$ . For higher energy virtual orbitals, the weighting applied is a decaying polynomial which goes as:

$$(E_k + E_l - g\_VMC\_PolyExcitToWeight1 + 1)^{-g\_VMC\_PolyExcitToWeight2} \quad (5.5)$$

$g\_VMC\_PolyExcitToWeight1$  is constrained to be not more than the energy of the highest virtual orbital.

**POLYEXCITBOTH** [**g\_VMC\_PolyExcitFromWeight1** **g\_VMC\_PolyExcitFromWeight2** **g\_VMC\_PolyExcitToWeight1** **g\_VMC\_PolyExcitToWeight2** **G\_VMC\_EXCITWEIGHT**]

Identical to **POLYEXCITWEIGHT**, except that the polynomial weighting function applies also to the occupied orbitals. This means that there is another variable, since now the ‘ExcitFrom’ calculation also needs a value for sigma, and for the exponent. The sigma variables are now both under similar constraints as specified above, which means that they cannot be larger or smaller than the highest and lowest energy orbital respectively. This prevents the PRECALC block from getting stuck, or from finding local variance minima.

**Note:** What is sigma?

**CHEMPOTWEIGHTING** [**g\_VMC\_PolyExcitFromWeight2** **g\_VMC\_PolyExcitToWeight2** **G\_VMC\_EXCITWEIGHT**]

Weighting is of the same form as **POLYEXCITBOTH**, but sigma is now constrained to be at the chemical potential of the molecule. Has only two parameters with which to minimise the expected variance.

**CHEMPOT-TWOFROM** [**g\_VMC\_ExcitWeights(1)** **g\_VMC\_ExcitWeights(2)** **g\_VMC\_ExcitWeights(3)** **G\_VMC\_EXCITWEIGHT**]

When choosing the electron to excite, use a increasing polynomial up to the chemical potential and a decaying polynomial for spin orbitals above the chemical potential, in order to encourage mixing of the configurations around the HF state. Contains three adjustable parameters and testing needs to be done to see if this is beneficial. Expected to make more of a difference as the vertex level increases.

**Note:** What is the actual weighting form of **CHEMPOT-TWOFROM**?

**UFORM-POWER** New power form for the U-matrix element weighting using the appropriate **EXCITWEIGHT** element, which is believed to be better. This uses the form  $W = U^{\text{EXCITWEIGHT}}$ , rather than the exponential form.

## 5.5.9 Experimental options

**Note:** More documentation on these options needed.

**EXCITATIONS FORCEROOT** Force all excitations in **VERTEX [SUM STAR] NEW** calculations to come from the root.

**EXCITATIONS FORCETREE** Disallow any excitations in a **VERTEX SUM NEW** which are connected to another in the graph, forcing a tree to be produced. Not all trees are produced however.

**FULLDIAGTRIPS** An option when creating a star of triples, to do a full diagonalisation of the triples stars, without any prediagonalisation. Very very slow...

**LINEPOINTSSTAR [LinePoints]** Set the number of excited stars whose eigenvalues are evaluated when using **StarStars**, in order to determine linear scaling.

**NOTRIPLES** Disallow triple-excitations of the root determinant as the 3rd vertex in **HDIAG** calculations at the third vertex level and higher.

## 5.6 Integrals

**Note:** **INSPECT** and **ORDER** options currently make no sense.

**INTEGRAL** Starts the integral block.

[Integral options—see below.]

**ENDINT** End of integral block.

### 5.6.1 General options

**FREEZE [NFROZEN NTFROZEN]** Set the number of frozen core states and frozen excited states respectively. Both must be a multiple of two - an error is returned if this is not the case. The Slater determinant space of a calculation does not include determinants which contain excitations from frozen core states or excitations to frozen excited states.

**INSPECT [SPECDET(I), I=1,NEL-NFROZEN]** Investigate the specified determinant.

**ORDER [ORBORDER(I), I=1,8]** Set the preliminary ordering of basis functions for an initial guess at the reference determinant. There are two ways of specifying open orbitals:

1. If `orborder2(I,1)` is integral, then if it's odd, we have a single.
2. Non-integral. The integral part is the number of closed orbitals, and the fractional\*1000 is the number of open orbitals. e.g. 6.002 would mean 6 closed and 2 open which would have `orborder(I,1)=6`, `orborder(I,2)=4` but say 5.002 would be meaningless as the integral part must be a multiple of 2

### 5.6.2 Density fitting options

**DFMETHOD [method]** control the Density fitting method. Possible methods are:

**DFOVERLAP**  $(ij|lab) = (ij|P)(P|lab)$

**DFOVERLAP2NDORD**  $(ij|lab) = (ij|P)(P|lab) + (ij|P)(P|lab) - (ij|P)(P|Q)(Q|lab)$

**DFOVERLAP2**  $(ij|lab) = (ij|P)(P|Q)(Q|lab)$

**DFCOULOMB**  $(ij|lab) = (ij|P)[(P|Q)^{-1}](Q|lab)$

where the sums over P and Q are implied.

All methods are precontracted to run in order(nBasis) except DFOVERLAP2NDORD.

**DMATEPSILON DMatEpsilon (default 0)** The threshold for density matrix elements, below which small density matrix elements are ignored, and consequently speeds up calculations.

### 5.6.3 Hartree–Fock options

The Hartree–Fock options have only been tested for molecular and model systems. They allow the Hartree–Fock orbitals (in the space of the original basis) to be used in a graph calculation instead of the original basis.

**Note:** James has never used these options. Please can those who have document them in more detail.

**HF** Use a Hartree–Fock basis.

**CALCULATE** Calculate the Hartree–Fock basis rather than reading it in. By default, the Hartree–Fock calculation is performed before any freezing of orbitals, i.e. in the full original basis.

**HFMETHOD [HFMETHOD]** Default: **SINGLES**.

Specify the method for the Hartree–Fock routine. Options are:

**STANDARD** Use normal Hartree–Fock process.

**DESCENT [SINGLES, OTHER]** Use singles or other gradient descent.

**MODIFIED** Modify virtuals. Experimental.

**MAXITERATIONS [NHFIT]** Set the maximum number of Hartree–Fock iterations.

**MIX [HFMIX]** Set the mixing parameter for each Hartree–Fock iteration.

**POSTFREEZEHF** Do Hartree–Fock after freezing instead of before (still needs **HF** and **CALCULATE**). The Hartree–Fock calculation is performed only in the space of the unfrozen orbitals.

**RAND [HFRAND]** Default 0.01.

Set the maximum magnitude of the random numbers added to the starting density matrix. Use to perturb away from an initially converged Hartree–Fock solution.

**READ [MATRIX BASIS]** Read in U matrix and/or Hartree–Fock basis in terms of the original basis.

**RHF** Use restricted Hartree-Fock theory.

**THRESHOLD [ ENERGY [HFEDELTA] ORBITAL [HFCDELTA] ]** Set the convergence threshold for the energy and/or the orbitals.

**UHF** Use unrestricted Hartree-Fock theory.

## 5.6.4 Partitioning options

If the weight and energy contribution from a graph are evaluated from diagonalising the  $\rho$  matrices, then various schemes are available to deal with the  $e^{-\beta\hat{H}/P}$  operator.

**Note:** More detail on these needed.

**FOCK-PARTITION** For calculation of  $\rho$  operator with the Trotter approximation, partition the Hamiltonian according to the N-electron Fock operator and Coulomb perturbation.

**FOCK-PARTITION-LOWDIAG** For calculation of  $\rho$  operator with Trotter approximation, partition the Hamiltonian according to the N-electron Fock operator and coulomb perturbation. Take just the first order approximation (i.e. ignore the  $\beta/P$  term) for the diagonal terms of the  $\rho$  matrix.

**FOCK-PARTITION-DCCORRECT-LOWDIAG** For calculation of  $\rho$  operator with Trotter approximation, partition the Hamiltonian according to the N-electron Fock operator and Coulomb perturbation. Remove the Coulomb double counting in the Fock operator. Take just the first order approximation (i.e. ignore the  $\beta/P$  term) for the diagonal terms of the  $\rho$  matrix.

**DIAG-PARTITION** Default partitioning scheme.

For calculation of  $\rho$  operator with Trotter approximation, partition the Hamiltonian as the diagonal and non-diagonal matrix elements between the determinants.

**RHO-1STORDER** Calculate rho elements to only 1st order Taylor expansion (without applying a Trotter approximation).

## 5.6.5 VASP and CPMD options

There are too many 2-electron integrals to store for periodic systems (**CPMD** or **VASP** based calculations). Instead, as many integrals as possible are cached. Each four-index integral is reduced to two indices, A and B. Each A index has so many slots associated with it in which the integral involving A and B can be stored. The cache stores as many integrals as possible. If the cache is full and a new integral is calculated, then an element in the cache is over-written.

The efficiency of a calculation is heavily dependent on the size of the integral cache.

**UMATCACHE [SLOTS] [nSlots]** Default nSlots=1024.

Set the number of slots for each A index.

The total amount of memory used by the cache will be in the order of  $NSLOTS * NSTATES * (NSTATES - 1) / 2$  words.

If nSlots=0, then disable caching of integrals calculated on the fly, but retain precomputation of 2-index 2-electron integrals ( $\langle ij|ij \rangle$  and  $\langle ij|ji \rangle$ ).

If nSlots=-1, no 2-electron integrals are stored.

Disabling the cache is very expensive.

The keyword **SLOTS** is optional and is present to contrast with the **MB** keyword.



**UMATCACHE MB [MB]** Number of megabytes to allocate to the UMAT cache. The number of slots is then set accordingly.

**NOUMATCACHE** Disable all UMAT caching (identical to **UMATCACHE -1**).

### 5.6.6 Experimental options

**Note:** Please document in more detail!

**NRCONV [NRCONV]** Default  $10^{-13}$ .

This sets the convergence criteria for the Newton-Raphson algorithm in findroot. This takes place after initial bounds for the root are calculated using regular falsi (see above). Values smaller than  $10^{-15}$  tend to create a fault since the Newton-Raphson algorithm cannot converge given the number of iterations allowed.

**NRSTEPSMAX [NRSTEPSMAX]** This sets the maximum number of Newton Raphson steps allowed to try and converge upon a root to the accuracy given in **NRCONV**. This is only applicable for the star graph, when trying to find the roots of the polynomial using **POLY NEW**, **POLY OLD** or **POLYCONVERGE**. Default value is 50.

**RFCONV [RFCONV]** Default  $10^{-8}$ .

Set the convergence criteria for the Regular falsi algorithm in findroot. Only used with a star calculation which involves calculating the roots of a polynomial to find the eigenvalues. A Newton-Raphson convergence takes place after.

**INCLUDEQUADRHO** This changes the rho matrix for stars so that it includes the square of the eigenvalues - rho -> rho + rho<sup>2</sup>/2. This is in an attempt to improve size consistency for the star graph. No change for large beta, and only very small changes for smaller betas.

**EXPRHO** The rho matrix is exponentiated, 1 is subtracted, and this is used as the matrix to be diagonalised. This is the full expansion for which **INCLUDEQUADRHO** is a truncation. Again, this is used to achieve size consistency with the star, although seems to have little effect, and no effect at high beta.

**DISCONNECTNODES** If using a nodal approximation, the connections between determinants in the same nodes are ignored - should then be equivalent to the original star calculation.

**CALCEXCITSTAR** Used with **STARSTARS**, it explicitly calculates each excited star and diagonalises them separately. This removes the approximation of cancelling fictitious excitations if the original star is used as a template for higher excitations. Scaling is bad, and all matrix elements have to be calculated exactly.

**STARNODOUBS** Only to be used with **CALCEXCITSTAR** when explicitly calculating excited stars, it forbids the excited stars to have excitations which are double excitations of the Hartree-Fock determinant.

**STARQUADEXCITS** Only to be used with **CALCEXCITSTAR**, when calculating the excited stars, it only allow the excited stars to have excitations which are quadruple excitations of the Hartree-Fock determinant.

**QUADVECMAX** Used with **STARSTARS**, it uses only the largest first element of the eigenvectors as the connection to each excited star. This means that for each excited star, only one connection is made back to the original star, meaning that the scaling is reduced. This seems to be a good approximation.

**QUADVALMAX** Same as **QUADVECMAX**, only the largest eigenvalue for each excited star is used. Seems to be little difference in results.

**DIAGSTARSTARS** Used with **STARSTARS**, it performs a full diagonalisation on each excited star, using the original star as a template, i.e. same excitations, and same offdiagonal elements. All that occurs is that the diagonal elements are multiplied by rho<sub>jj</sub>. Large Scaling.

**EXCITSTARSROOTCHANGE** Used with **DIAGSTARSTARS** only at the moment, when this is set, only the root element of the excited star matrices changes when constructing excited stars with roots given by rho<sub>jj</sub>. The remainder of the excited star matrix is identical to the original star matrix.

**RMROOTEXCITSTARSROOTCHANGE** Another option for use with **DIAGSTARSTARS**, when this is set, the same occurs as for **EXCITSTARSROOTCHANGE**, apart from the fact that the root is removed as an excited determinant in each excited star.

## 5.7 Logging

**Note:** All the logging options should say to which file they print output. Please correct this!

**LOGGING** Start the logging input block. The logging options allow additional (potentially expensive, potentially verbose) information to be printed out during a calculation. By default, all logging options are turned off.

[Logging options—see below.]

**ENDLOG** End the logging input block.

### 5.7.1 General options

**FMCPR [LABEL, RHO, 1000, EXCITATION]** More than one of the options can be specified.

Log the following to the PATHS file:

**LABEL** Logs the determinants contained by each graph as each determinant is generated in the format:  $[(D_0), (D_1), \dots, (D_v),]$  where each determinant given as a comma-separated list of the indices of the occupied orbitals: e.g.  $D_0 = (1, 2, 9, 10,)$ .

If CSFs are being used, then the CSF is printed. There is no newline after this.

For **MULTI MC** or **SINGLE MC**, only the non-composite graphs are printed.

**EXCITATION** Log each graph in excitation format instead of full format above. The format is  $[A(i, j) \rightarrow (a, b), B(k, l) \rightarrow (c, d), \dots, C(m, n) \rightarrow (e, f)]$

**where** A, B, ..., C are determinants in the graph from which the excitation is made. i, j, ... are the orbitals within that determinant which are excited from, where  $(i < j, k < l, \dots)$ . a, b, ... are the orbitals they are excited to, where  $(a < b, c < d, \dots)$ .

This format does not in general provide a unique way of specifying multiply connected graphs, but the first possible determinant to which the next det in the graph is connected is chosen, so what is output should be unique. Single excitations are written as e.g.  $A(i, 0) \rightarrow (a, 0)$ .

**RHO** Log the  $\rho$  matrix for each graph in the form:  $(\rho_{11}, \rho_{12}, \dots, \rho_{1v}, | \rho_{21}, \rho_{22}, \dots, \rho_{2v}, | \dots | \rho_{v1}, \rho_{v2}, \dots, \rho_{vv}, |)$ , where the graph consists of  $v$  vertices. A newline is appended.

**XIJ** Log the  $x_{ij}$  matrix, which contains the generation probabilities of one determinant in the graph from all the others. For MC this is already generated, but for full sums this must be generated, so will be slower. Generation probabilities are set with the **EXCITWEIGHTING** option.

**The format is:**  $\{x_{11}, x_{12}, \dots, x_{1v}, | \dots | x_{v1}, x_{v2}, \dots, x_{vv}, | \}$

In general  $x_{ij} \neq x_{ji}$ . The  $x_{kk}$  element lists the number of possible excitations from  $k$  determinant. The matrix is followed by a newline.

**After all these possible options, the following are printed:** Weight [pGen] ETilde\*Weight Class [Accepted]

pGen is only printed for: Monte Carlo, or if doing a full sum and the XIJ logging option is set. Accepted is only printed for Monte Carlo calculations. A newline is placed at the end of this data. For Monte Carlo calculations, the values printed depend on the options. If **LABEL** is set, then all generated graphs and their values are printed, otherwise only values of accepted graphs are printed.

**CALCPATH [LABEL RHO]** Log CALCPATH\_R to PATHS. Either just label logging or also log the  $\rho$  matrix, with the same format as above.

**HAMILTONIAN** Log HAMIL, PSI and PSI\_LONG.

**HFBASIS** Log HFBASIS.

**HFLOGLEVEL [LEVEL]** Default 0.

If LEVEL is set to be positive, the density matrices, fock matrices and eigenvectors during a Hartree–Fock calculation are printed out to SDOUT.

**MCPATHS** Log MCPATHS data to the MCPATHS file for full vertex sum and MCSUMMARY file when using a METHODS section. Also log to the RHOPII file.

**PSI** Log PSI\_COMP.

**TIMING** [**iGlobalTimerLevel**] (**default 40**) Timing information is only recorded for routines with level less than or equal to iGlobalTimerLevel. Less than 10 means general high level subroutines. Greater than 50 means very low level. Routines without a level are always timed (most of them).

**XIJ** Synonym for **FMCPR XIJ**.

## 5.7.2 FCIMC options

**POPSFILE** Print out the determinants at the end of the MC run. A calculation can then be restarted at a later date by reading the determinants back in using **READPOPS** in the **CALC** section. at a later date.

## 5.7.3 GraphMorph options

**DISTRIBS** Write out the distribution of the excitations in each graph as it morphs over the iterations. The first column is the iteration number, and then subsequent columns denote the number of n-fold excitations in the graph.

## 5.7.4 PRECALC options

**PREVAR** Print the vertex level, Iteration number, parameter, and expected variance, for each parameter which was searched for in the **PRECALC** block, showing the convergence on the optimum value, to the **PRECALC** file.

**SAVEPRECALCLOGGING** Allows different logging levels to be used in the **PRECALC** block than for the main calculation.

All logging options specified before **SAVEPRECALCLOGGING** are only used in the the **PRECALC** part of the calculation. All logging options specified after **SAVEPRECALCLOGGING** are only used in the the main part of the calculation.

## 5.7.5 Monte Carlo options

**BLOCKING** Perform a blocking analysis on the MC run. An MCBLOCKS file will be produced, which lists  $\log(2)[\text{blocksize}]$ , the average of the blocks, the error in the blocks (where the blocks are the energy ratio), and the full error, treating the energy estimator as a correlated ratio of two quantities.

**VERTEX** [**EVERY n**] Log the vertex MC with  $\tilde{E}$  every n (real) cycles and/or log the vertex MC contribution every cycle. Setting  $\Delta = \tilde{E} - \tilde{E}_{\text{ref}}$ , where  $\tilde{E}_{\text{ref}}$  is usually the 1-vertex graph:

**EVERY** write a VMC file with the following info, with a new line each time the current graph changes:  
tot # virt steps, # steps in this graph, #verts, Class, Weight, Delta, <sign(W)>, <Delta sign(W)>, ~standard deviation <Delta sign>/<sign>, pgen

**n:** write a VERTEXMC file with the following info: 0, #graphs, <sign(W)>, stdev(sign(W)), <Delta>, <sign Delta>/<sign>, <Delta^2>, acc ratio, trees ratio, nontree+ ratio, non-tree- ratio, <Delta sign(W)>, E~ reference, #sequences, w reference

**Note:** George, what are most of these values?

**WAVEVECTORPRINT** [**nWavevectorPrint**] Relevant only for Monte Carlo star calculations.

Calculate the exact eigen-vectors and -values initially, and print out the running wavevector every nWavevectorPrint Monte Carlo steps. This slows the calculation down substantially.

## 5.8 Examples

To be added.

# Output files

## 6.1 MCPATHS and MCSUMMARY

The MCPATHS logging file (or MCSUMMARY if a **METHODS** block is used) has the following layout:

```
<Header Line>
<Vertex Sum Section for Det 1>
<Vertex Sum Section for Det 2>
...
<MC Summary information>
```

The Header Line is:

```
\ "Calculating XXX W_Is...\ "
```

where XXX is the number of determinants calculated, and the number of Vertex Sum sections.

Each Vertex Sum Section consists of:

```
(Determinant Calculated)
<method level 1 line>
<method level 2 line>
...
```

The Method level line is:

```
<level> <weight> <cumlweight> <timing> <GraphsSummed> [<PartGraphs>] <w E~> [<MP2 contrib> [<
```

**where:** **level** The vertex level as specified by the METHODS section.

**weight** The contribution of this level to  $s_i$  (see RHOP2 or RHOP2ex file section). This can be further analysed in CLASSPATHS{,2}.

**cumlweight** The sum of weights up to and including this level.

**timing** (double) The number of seconds calculating this level took.

**GraphsSummed** The total number of graphs summed together at this level.

**PartGraphs** Recursively, how many times FMCPR? is called. It is called once as each node is added to a graph.

$w\tilde{E}$  The contribution of this level to  $w\tilde{E}$ . This can be further analysed in CLASSPATHS{,2}

**MPn contrib** The contribution of graphs at this level to MPn theory.

There are is one method level line for each method level specified in the **METHODS** section, plus one for the 1-vertex graph.

The MC Summary information is split into 4 parts:

TotalStats:

GRAPHS (V) . . . WGHT- (V)

GenStats:

GEN-> \*

AccStats:

ACC-> \*

Sequences:

Sequences, Seq Len

**TotalStats:** The output is split into columns depending on the levels sampled in the Monte Carlo:

DataType Total 1-vertex 2-vertex 3-vertex ...

**The DataTypes are:** **GRAPHS(V)** The number of graphs sampled with this number of vertices.

**TREES(V)** The number of trees sampled with this number of vertices. Trees contain no cycles.

**NON-TR+(V)** The number of non-trees sampled with this number of vertices whose weight is positive.

**NON-TR-(V)** The number of non-trees sampled with this number of vertices whose weight is negative.

**WGHTT(V)** The total weight of trees with this number of vertices.

**WGHT+(V)** The total weight of positive non-trees with this number of vertices.

**WGHT-(V)** The total weight of negative non-trees with this number of vertices.

**GenStats:** Statistics on the number of graph-graph transitions generated.

**The columns correspond to the graph FROM which the transition was generated:** DataType 1-vertex 2-vertex 3-vertex ...

**The rows correspond to the graph TO which the transition was generated:** GEN-> 1 GEN-> 2 ...

**AccStats:** Format as GenStats, but has the number of transitions accepted.

**Sequences:** Records the number of sequences of consecutive graphs accepted with the same weight.

# BIBLIOGRAPHY

- [SumPaper] A combinatorial approach to the electron correlation problem, Alex J. W. Thom and Ali Alavi, J. Chem. Phys. 123, 204106, (2005).
- [StarPaper] Electron correlation from path resummations: the double-excitation star, Alex J. W. Thom, George H. Booth, and Ali Alavi, Phys. Chem. Chem. Phys., 10, 652-657 (2008).
- [ThomPhDThesis] Towards a quantum Monte Carlo approach based on path resummations, Alex J.W. Thom, PhD Thesis (2006).
- [DALTON] DALTON, a molecular electronic structure program, Release 2.0 (2005), see <http://daltonprogram.org/>
- [MolPro] MOLPRO is a package of ab initio programs written by H.-J. Werner, P. J. Knowles, R. Lindh, F. R. Manby, M. Schütz, P. Celani, T. Korona, G. Rauhut, R. D. Amos, A. Bernhardsson, A. Berning, D. L. Cooper, M. J. O. Deegan, A. J. Dobbyn, F. Eckert, C. Hampel, G. Hetzer, A. W. Lloyd, S. J. McNicholas, W. Meyer, M. E. Mura, A. Nicklaß, P. Palmieri, R. Pitzer, U. Schumann, H. Stoll, A. J. Stone, R. Tarroni, and T. Thorsteinsson, see <http://www.molpro.net>
- [CPMD] CPMD, <http://www.cpmc.org/>, Copyright IBM Corp 1990-2008, Copyright MPI für Festkörperforschung Stuttgart 1997-2001.
- [VASP] VASP
- [TwoElBox] Two interacting electrons in a box: An exact diagonalization study, Ali Alavi, JCP 113 7735 (2000).
- [AttenEx] Efficient calculation of the exact exchange energy in periodic systems using a truncated Coulomb potential, James Spencer and Ali Alavi, PRB, 77 193110 (2008).
- [CamCasp] Cambridge package for Calculation of Anisotropic Site Properties, Alston Misquitta and Anthony Stone. <http://www-stone.ch.cam.ac.uk/programs.html#Camcasp>
- [StarPaper] Electron correlation from path resummations: the double-excitation star, Alex J. W. Thom, George H. Booth, and Ali Alavi, Phys. Chem. Chem. Phys., 10, 652-657 (2008).
- [GHBCPGS] CPGS report, George Booth.
- [RGPtIII] Part III report, Ramin Ghorashi.