

NEUTRINO-ELECTRON ELASTIC SCATTERING RECO VS TRUE VARIABLES REVISITED AND FURTHER CUTS

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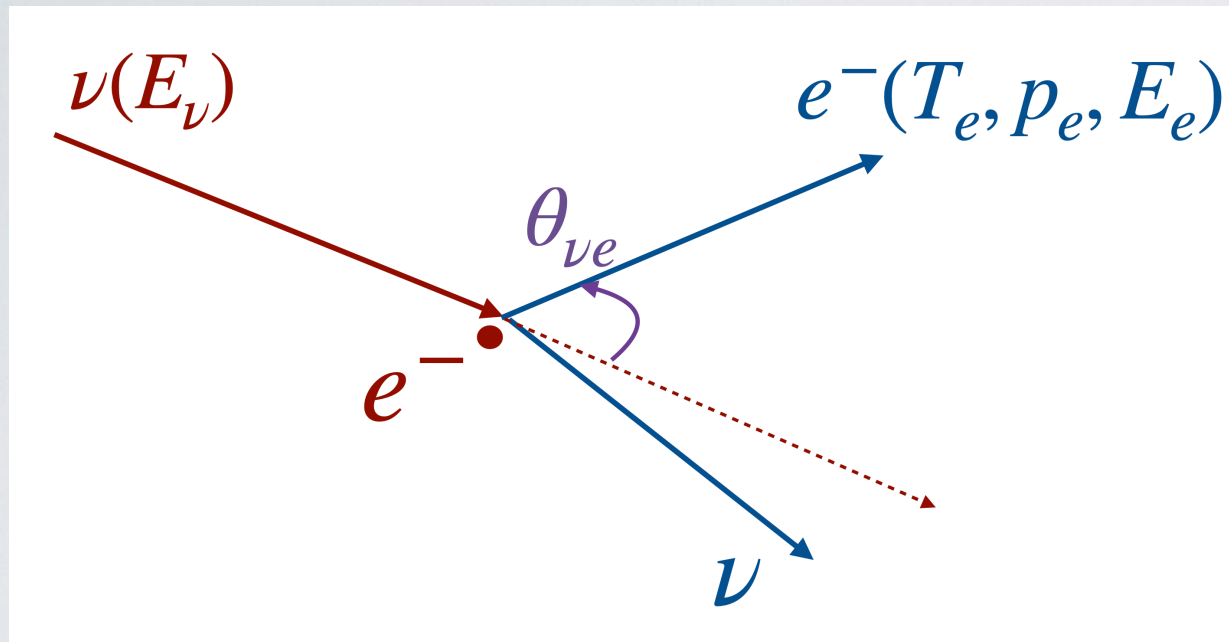
21 October 2022



INTRODUCTION

- Main update since last meeting: found bug in MC energy, now things look a lot more sensible
- Showing how the selection would benefit from a better reconstruction
- Estimated resolution needs for a good energy reconstruction

INTRODUCTION



$$E_\nu = \frac{m_e T_e}{p_e \cos \theta_{\nu e} - T_e}$$

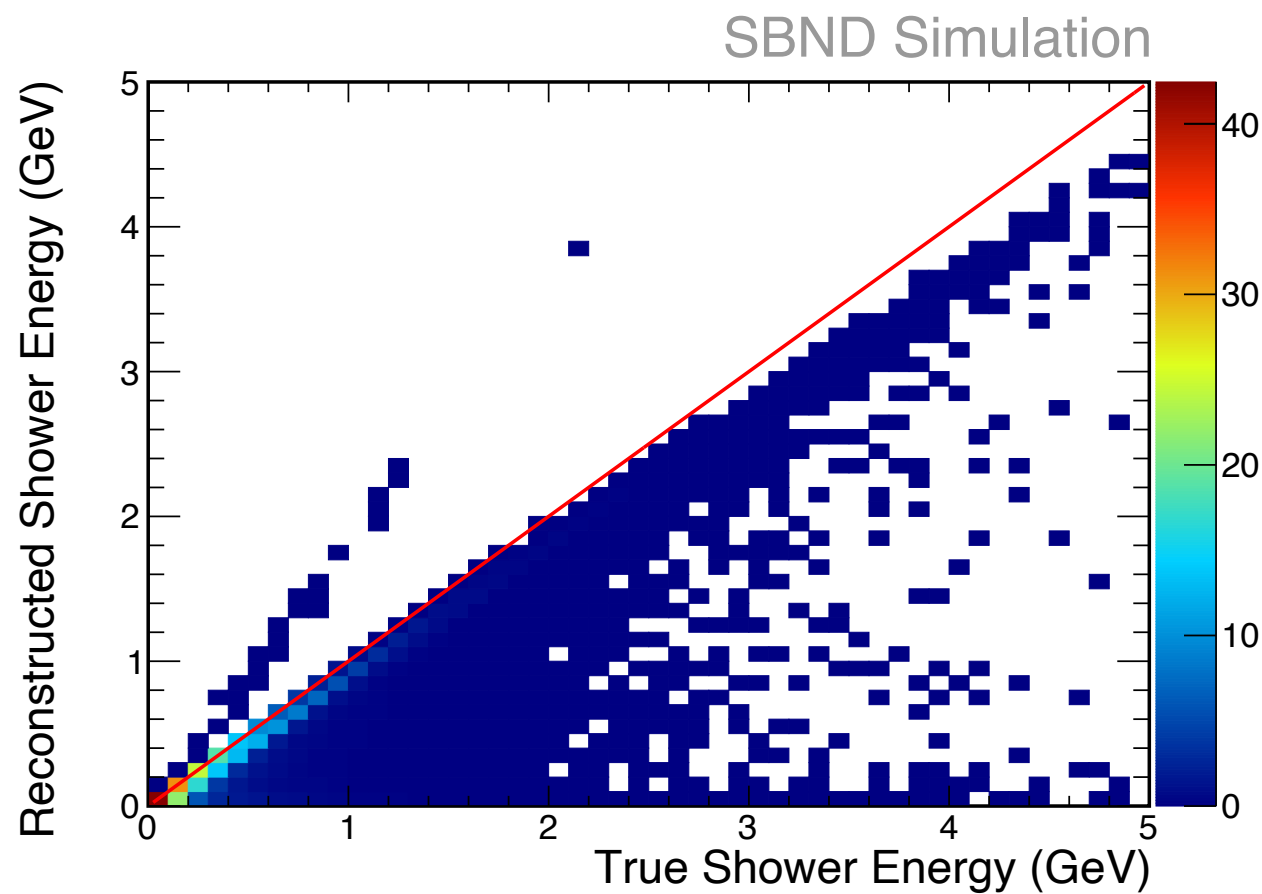
Annotations for the equation:

- $m_e T_e$ is annotated with $\mathcal{O}(10^{-4})$ and $\mathcal{O}(1)$.
- $p_e \cos \theta_{\nu e}$ is annotated with $\mathcal{O}(1)$ and ~ 1 .
- T_e is annotated with ~ 1 .

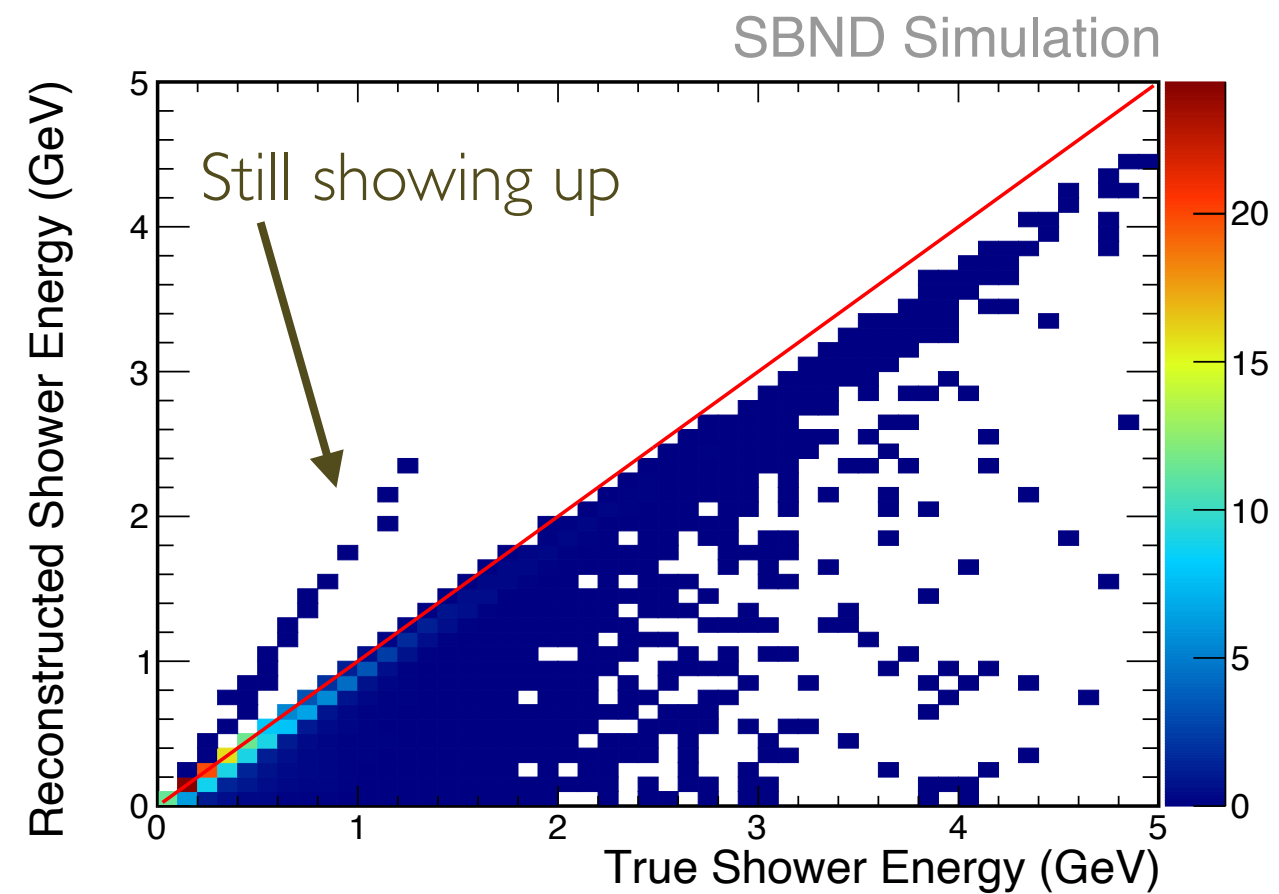
Caveat:

Measuring the incoming neutrino energy requires an **excellent** precision measuring both the energy and direction of the electron shower

SHOWER ENERGY (SIGNAL)



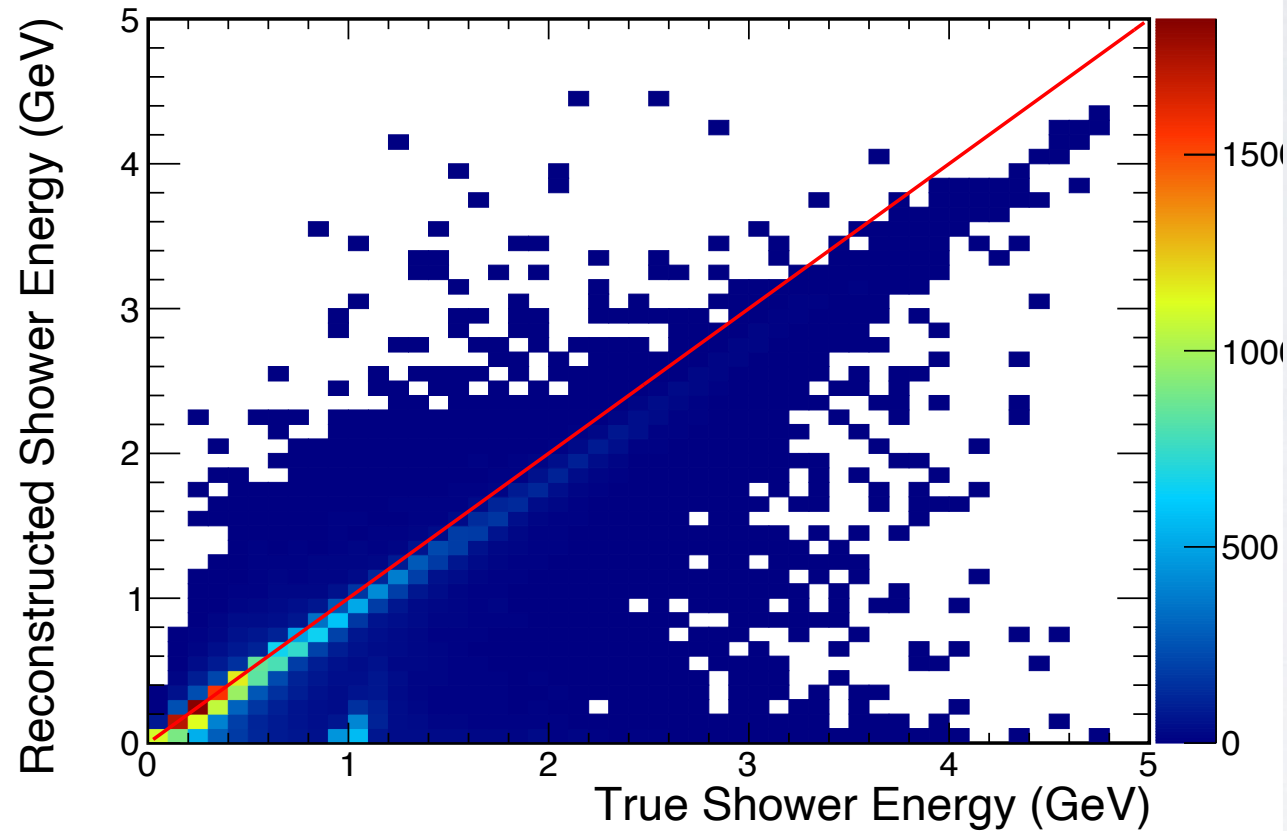
No cuts



One shower, no tracks

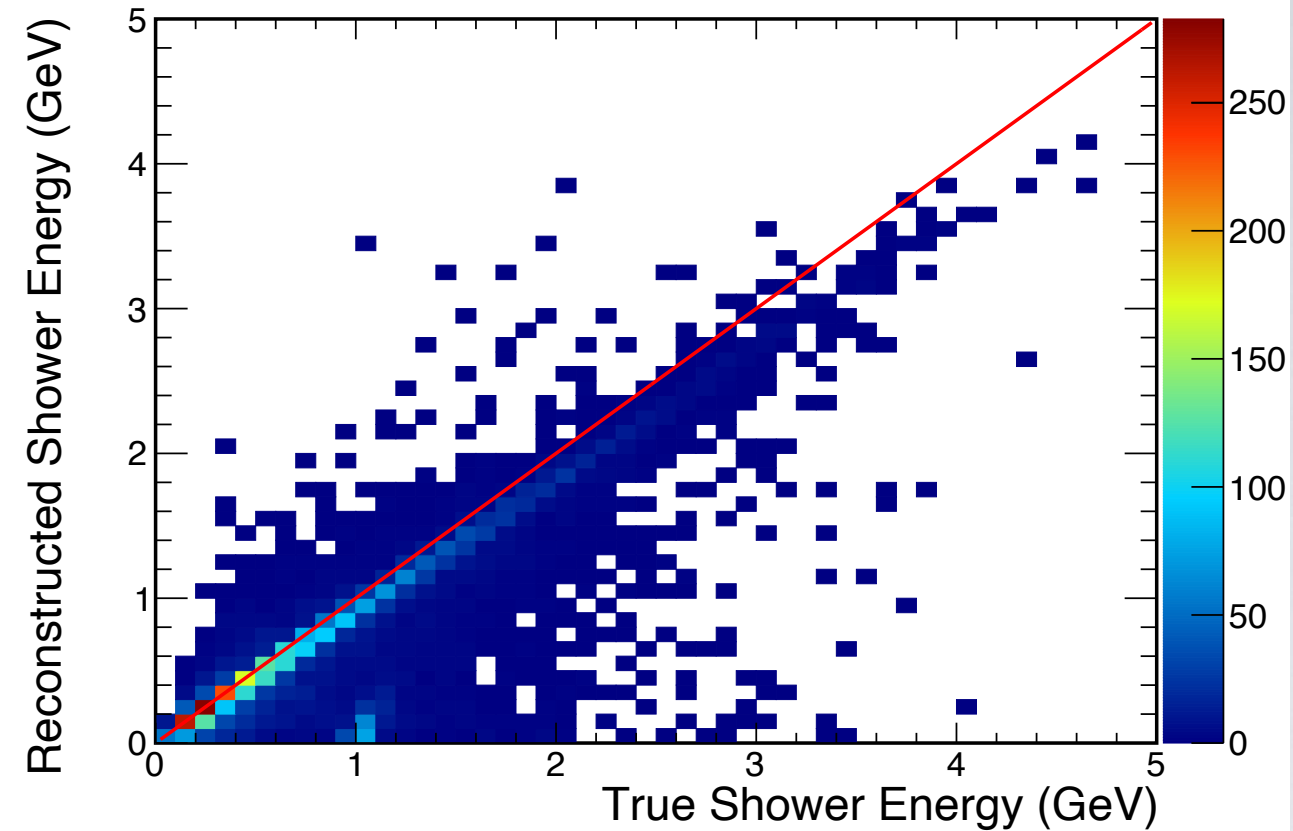
SHOWER ENERGY (BACKGROUND)

SBND Simulation



No cuts

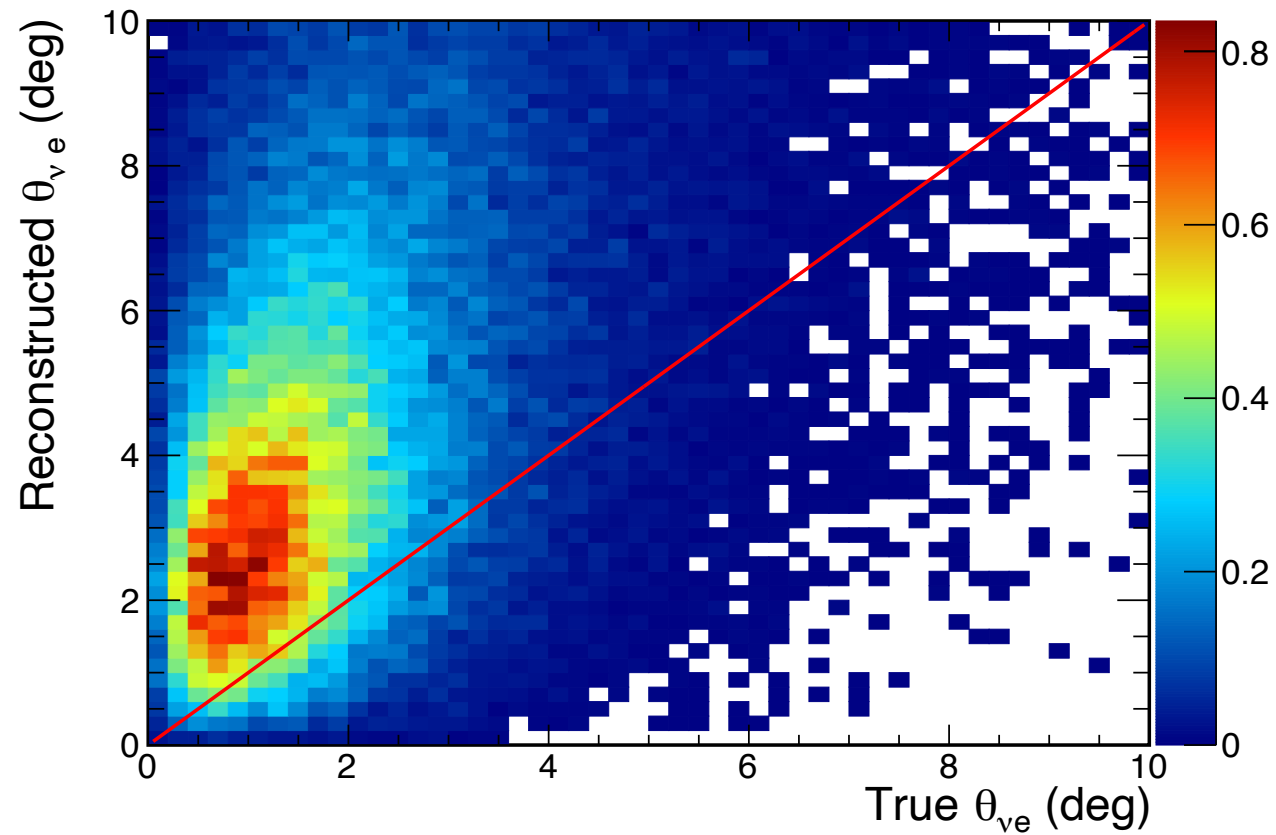
SBND Simulation



One shower, no tracks

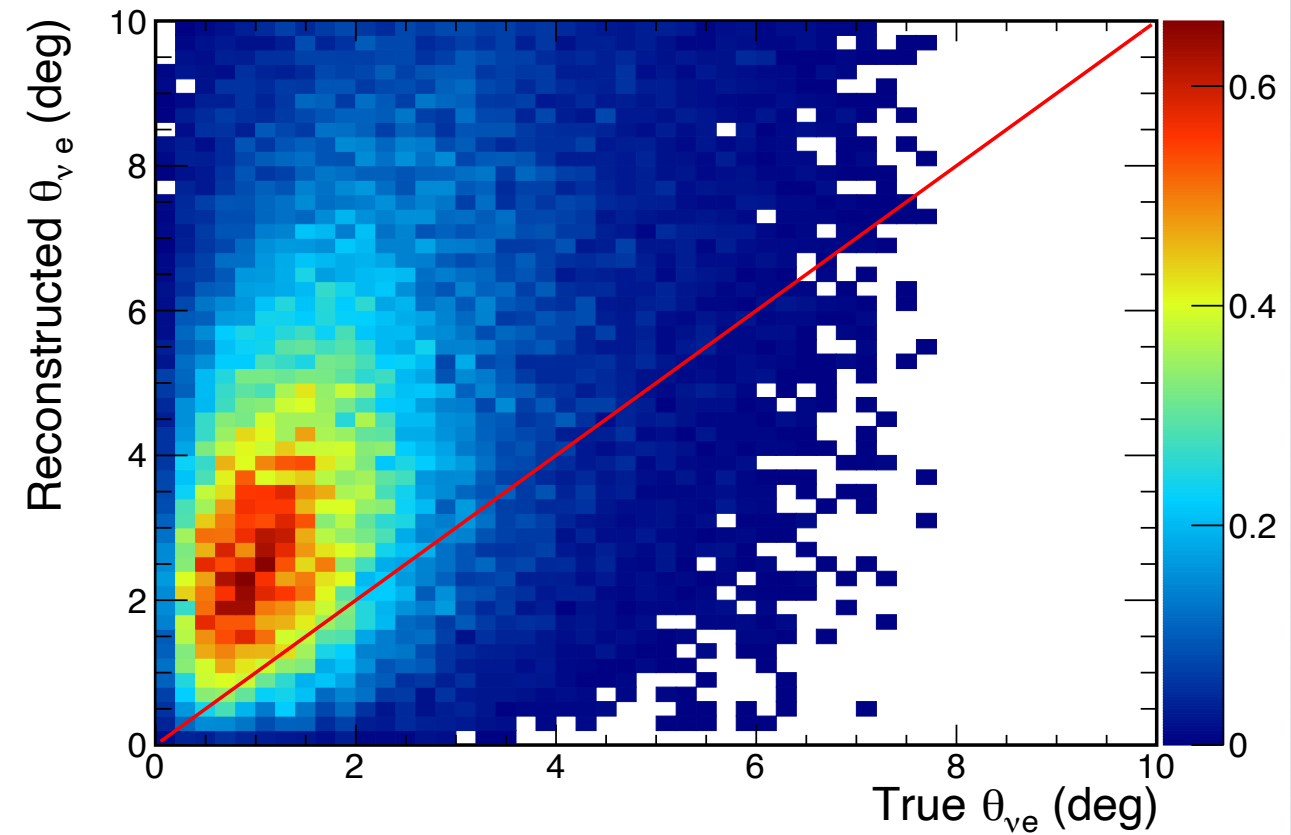
SHOWER DIRECTION (SIGNAL, ZOOM)

SBND Simulation



No cuts

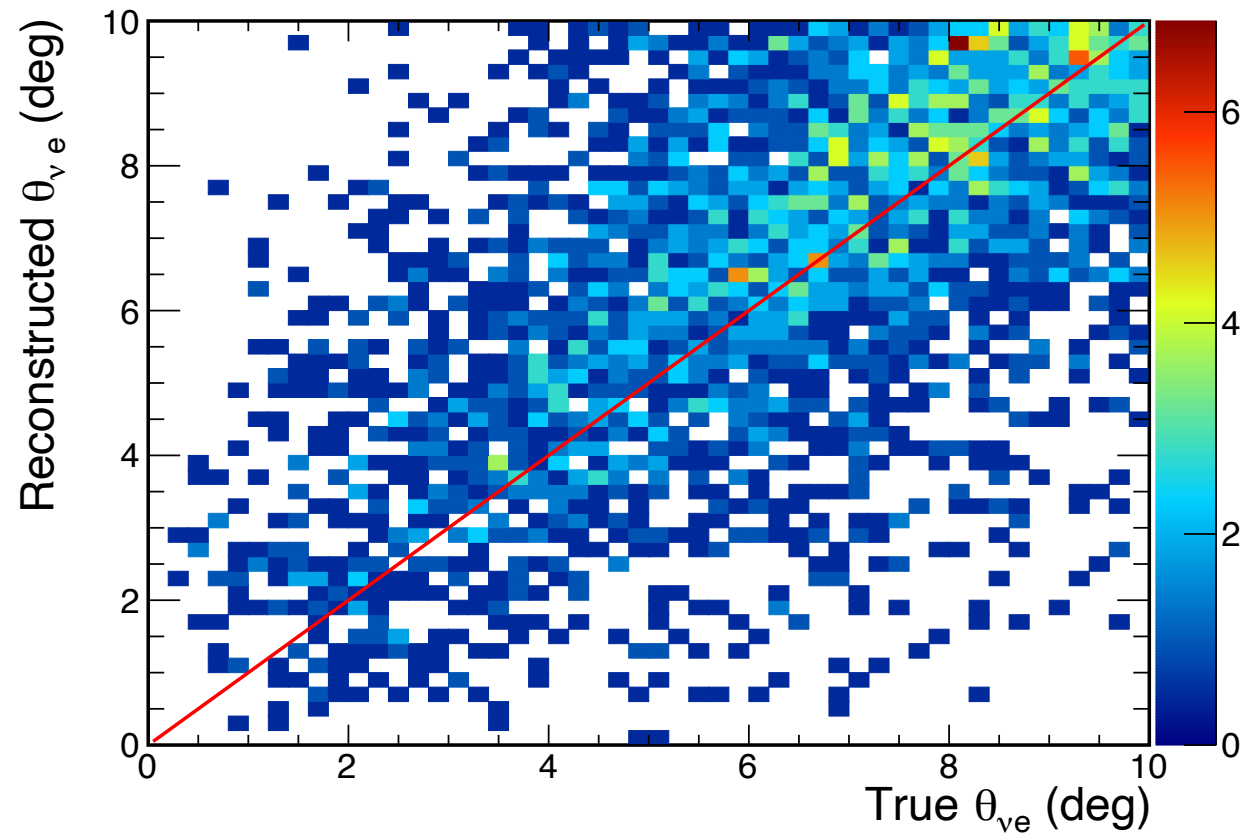
SBND Simulation



One shower, no tracks

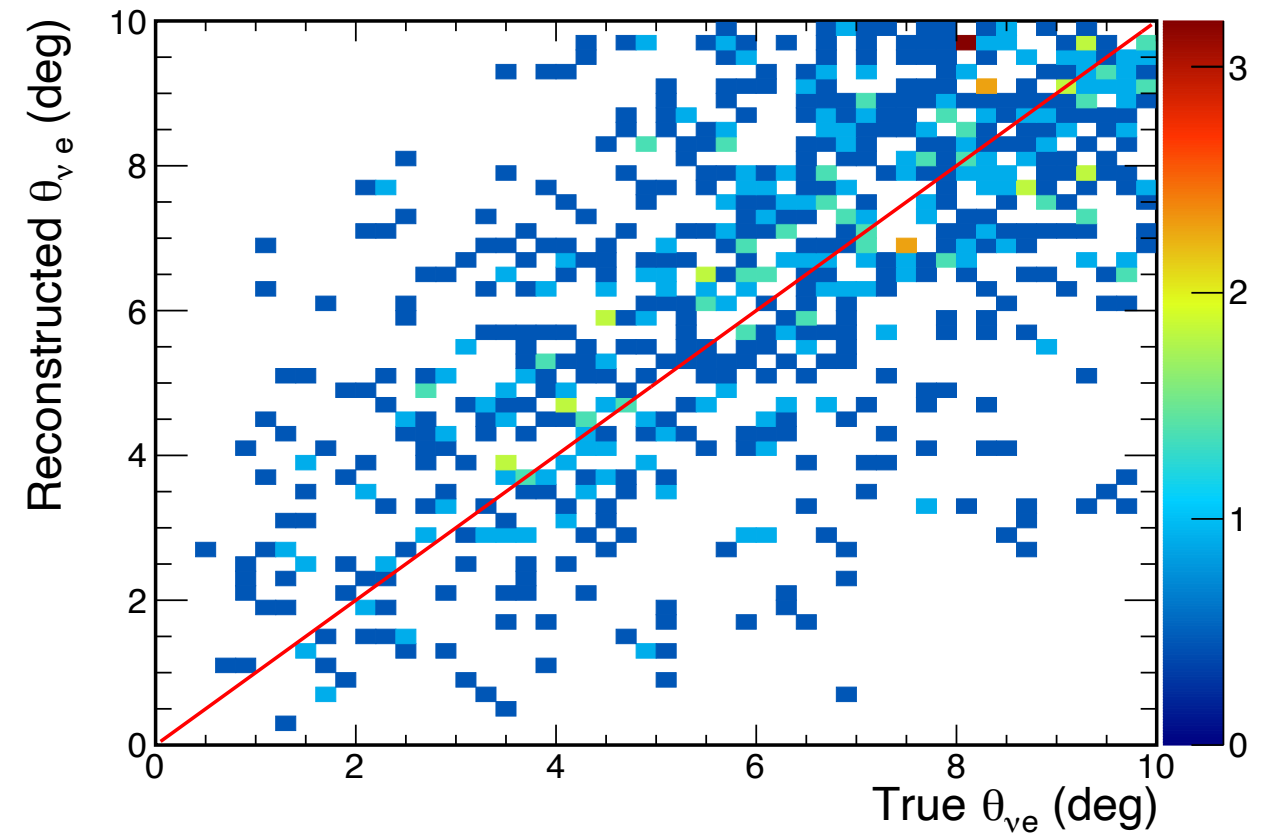
SHOWER DIRECTION (BACKGROUND)

SBND Simulation



No cuts

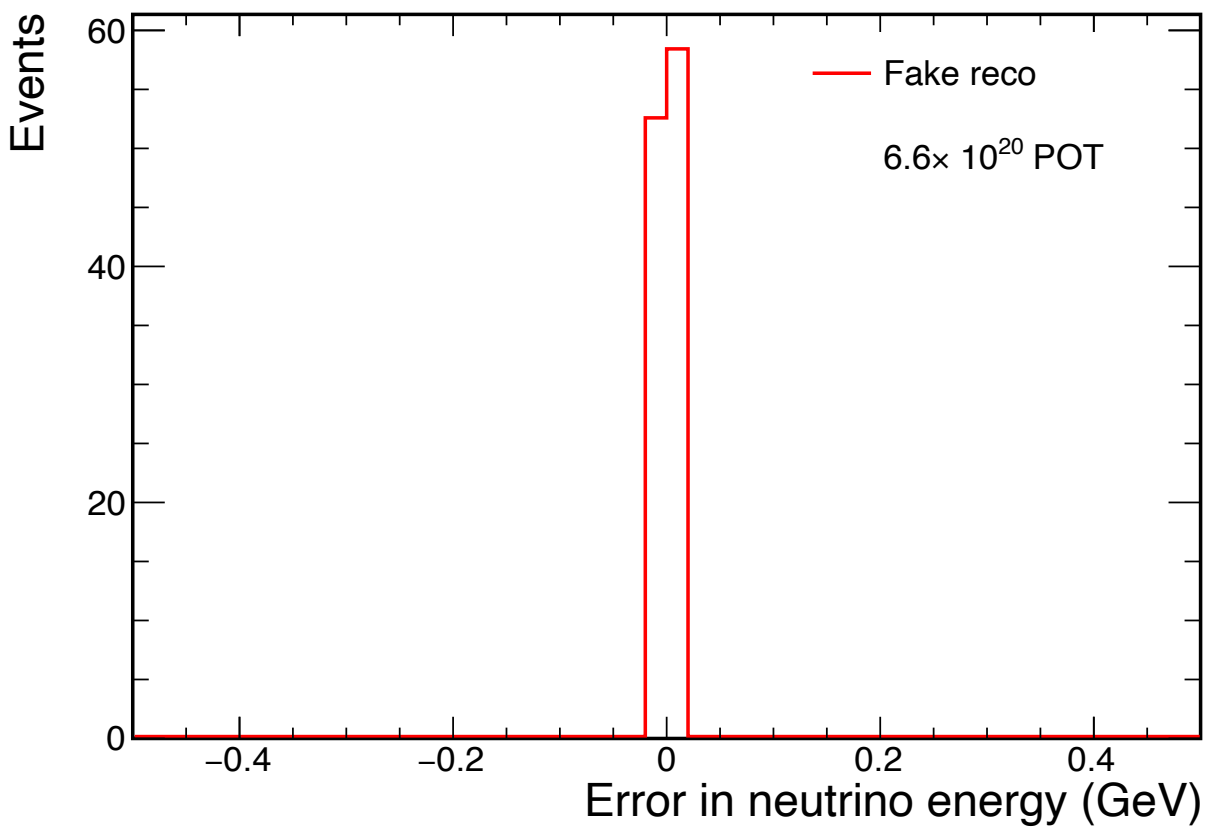
SBND Simulation



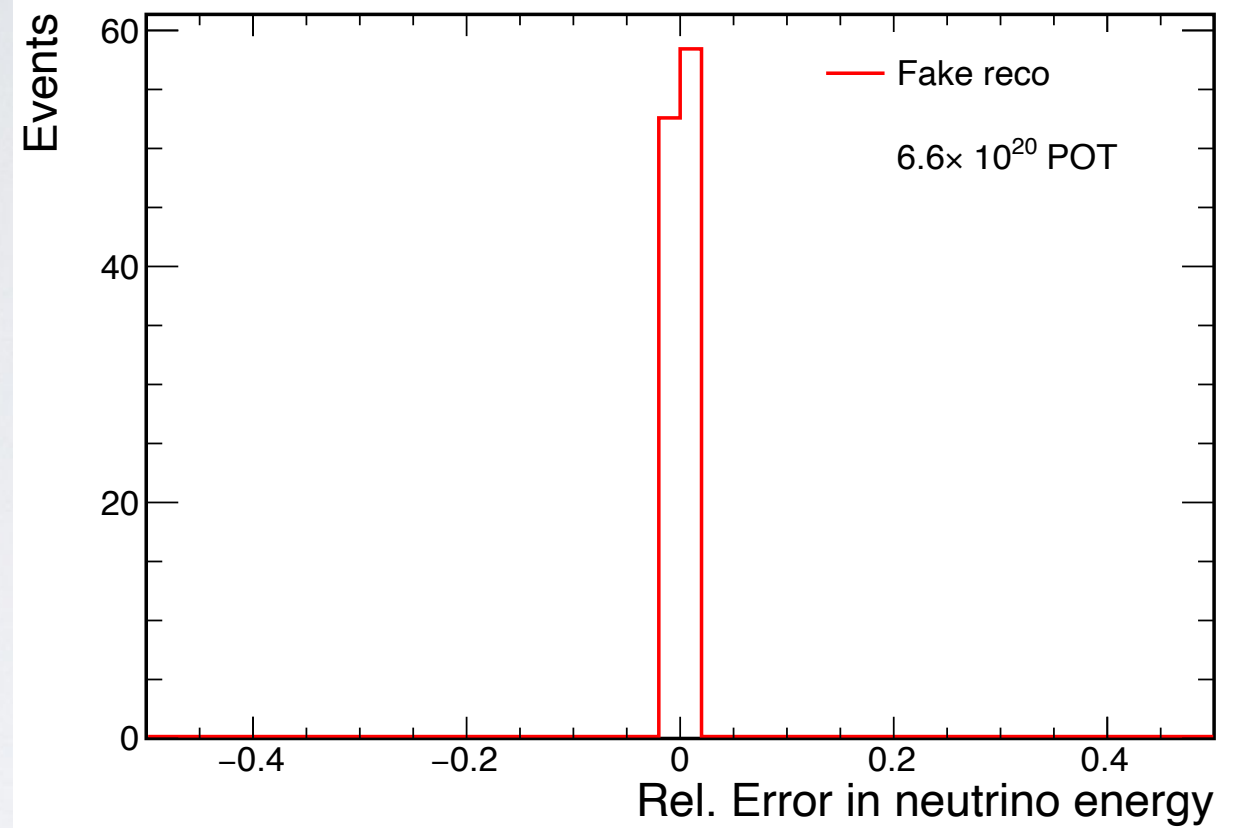
One shower, no tracks

HOW GOOD DO WE *NEED*
THE RECONSTRUCTION TO
BE?

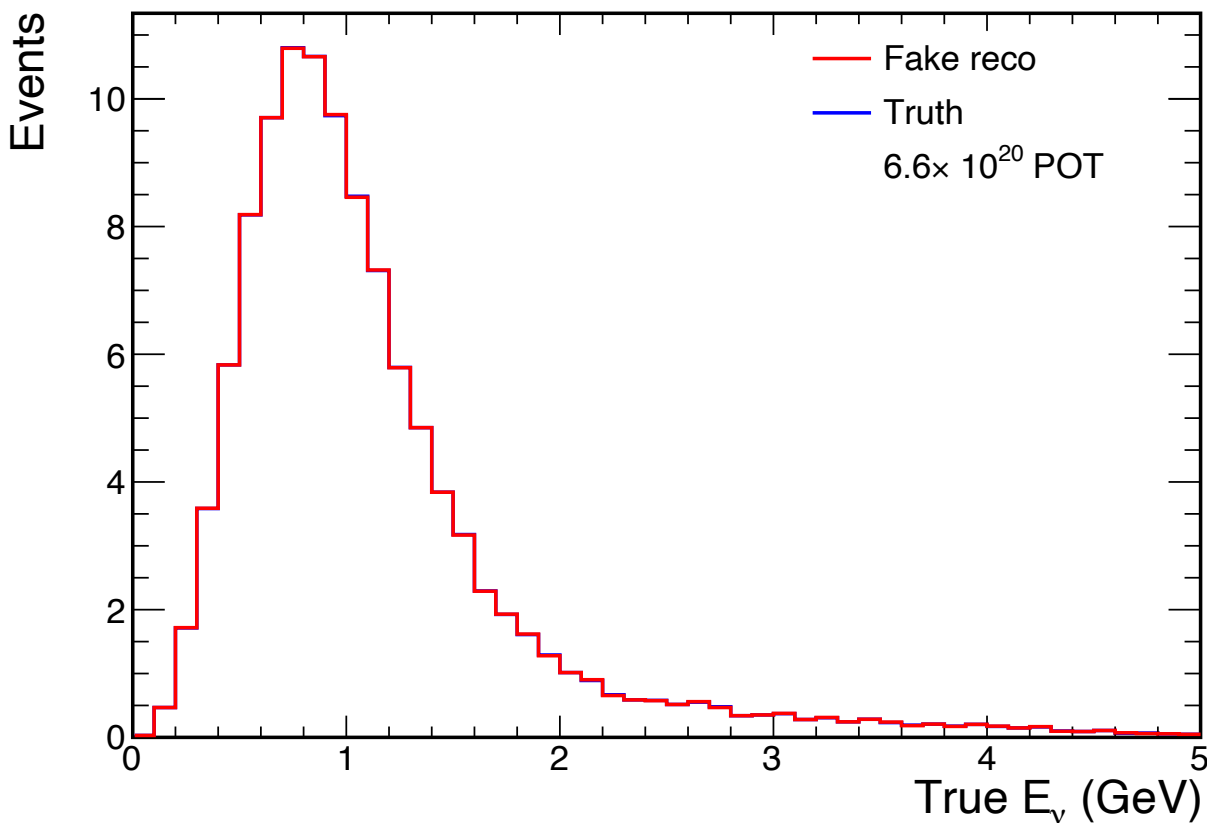
SBND Simulation



SBND Simulation



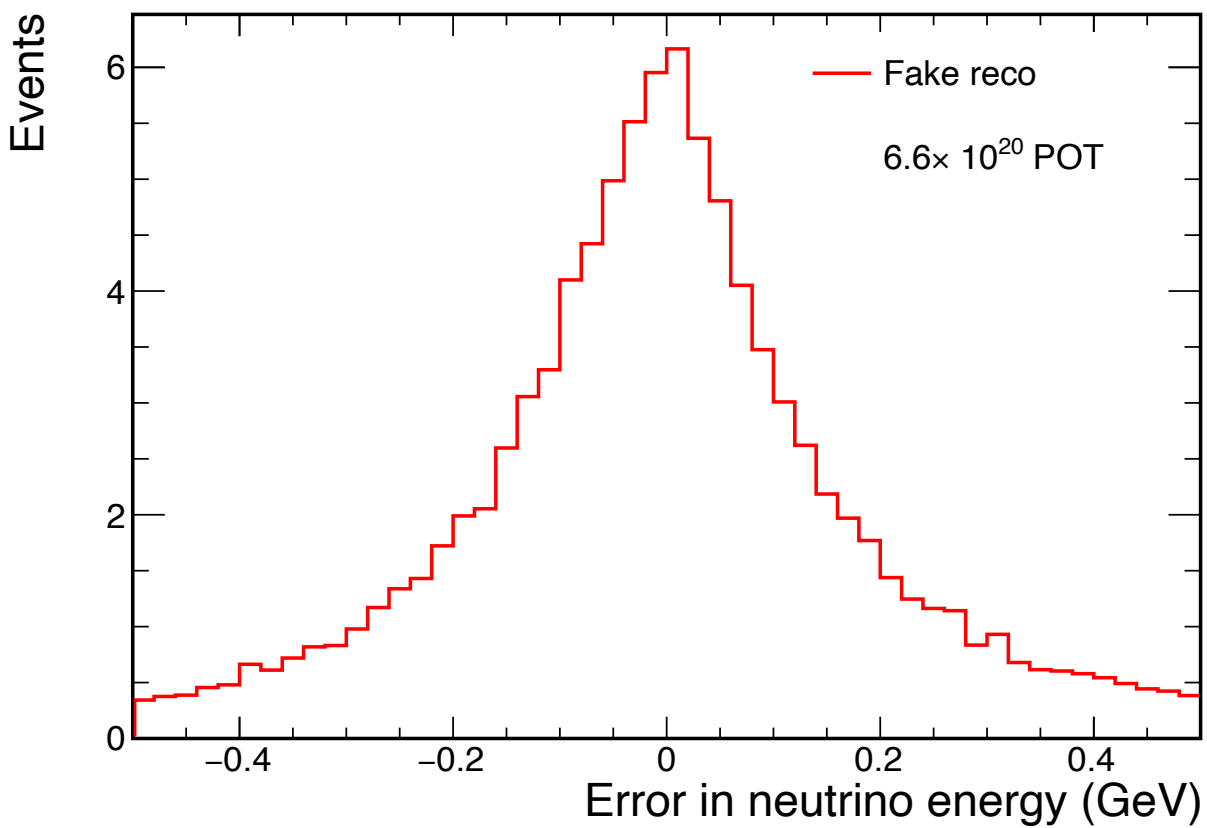
SBND Simulation



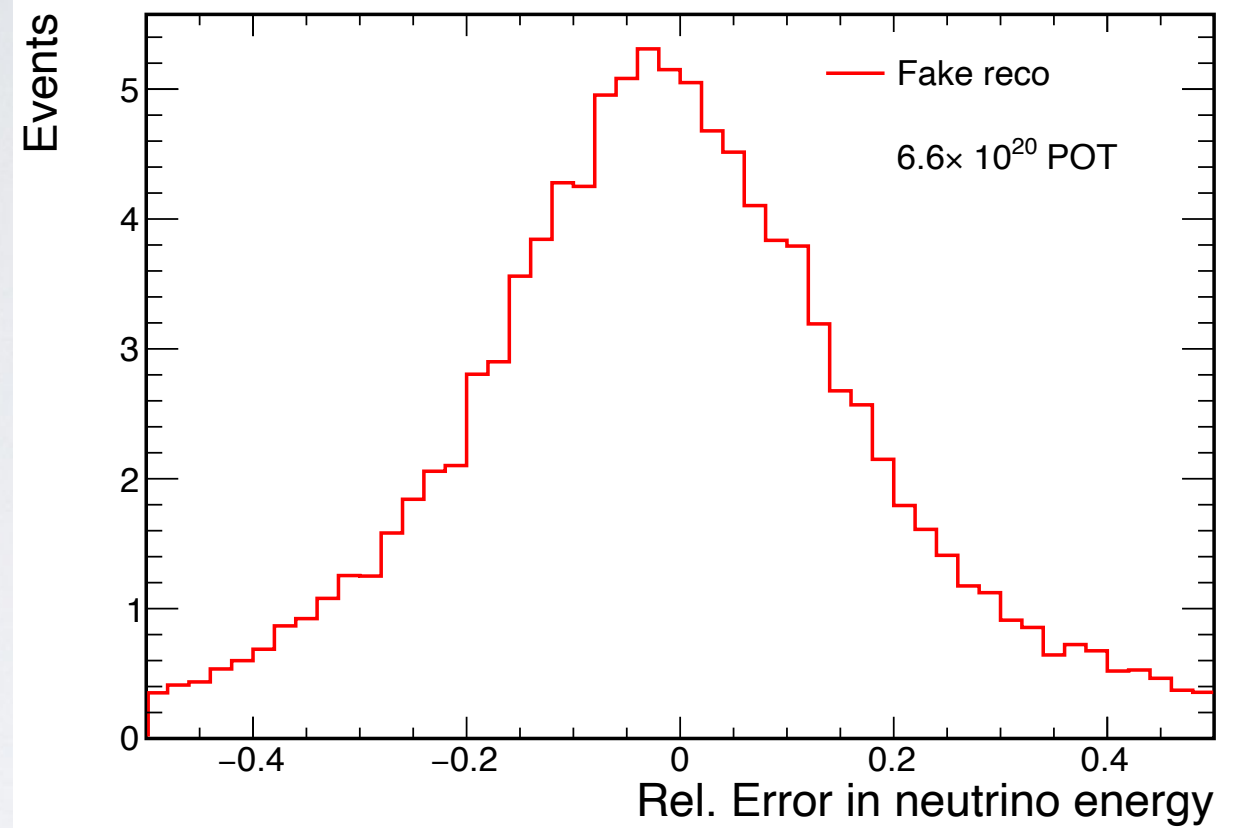
Using all the true information in

$$E_{\nu} = \frac{m_e T_e}{p_e \cos \theta_{\nu e} - T_e}$$

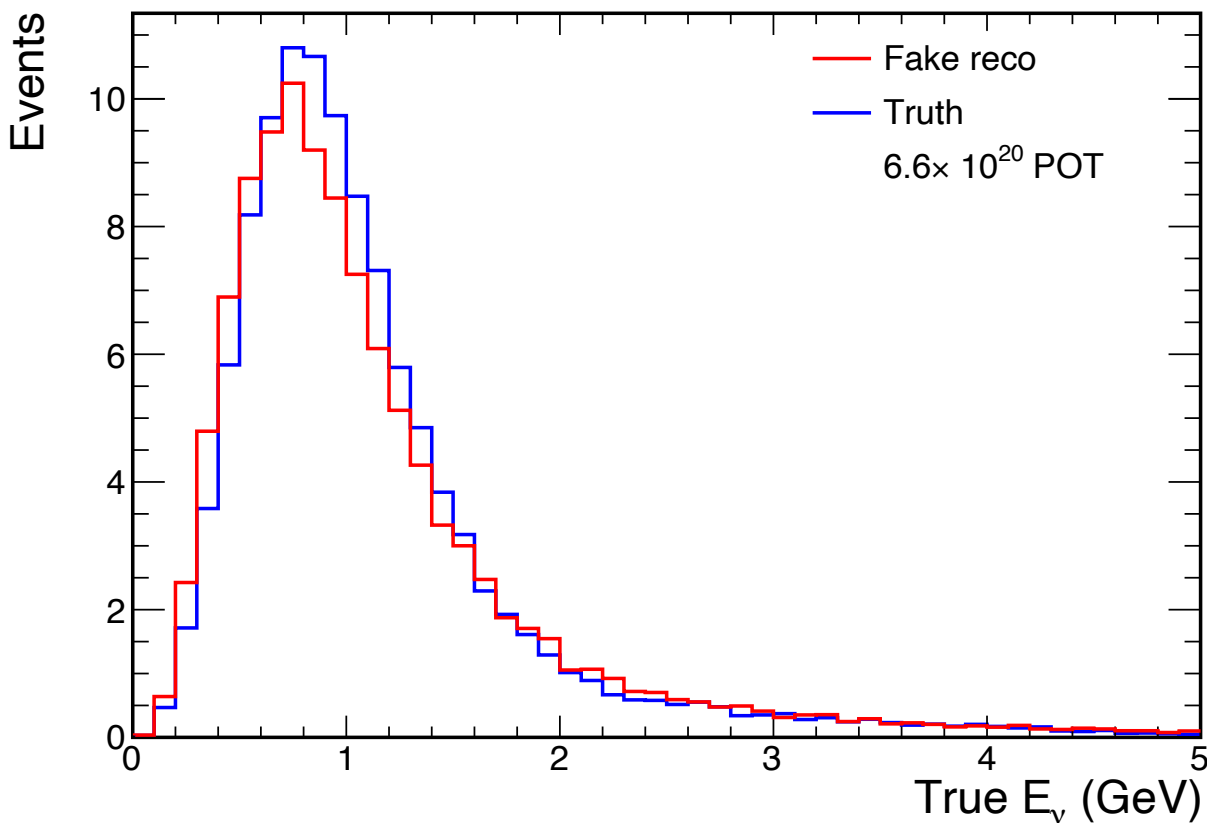
SBND Simulation



SBND Simulation



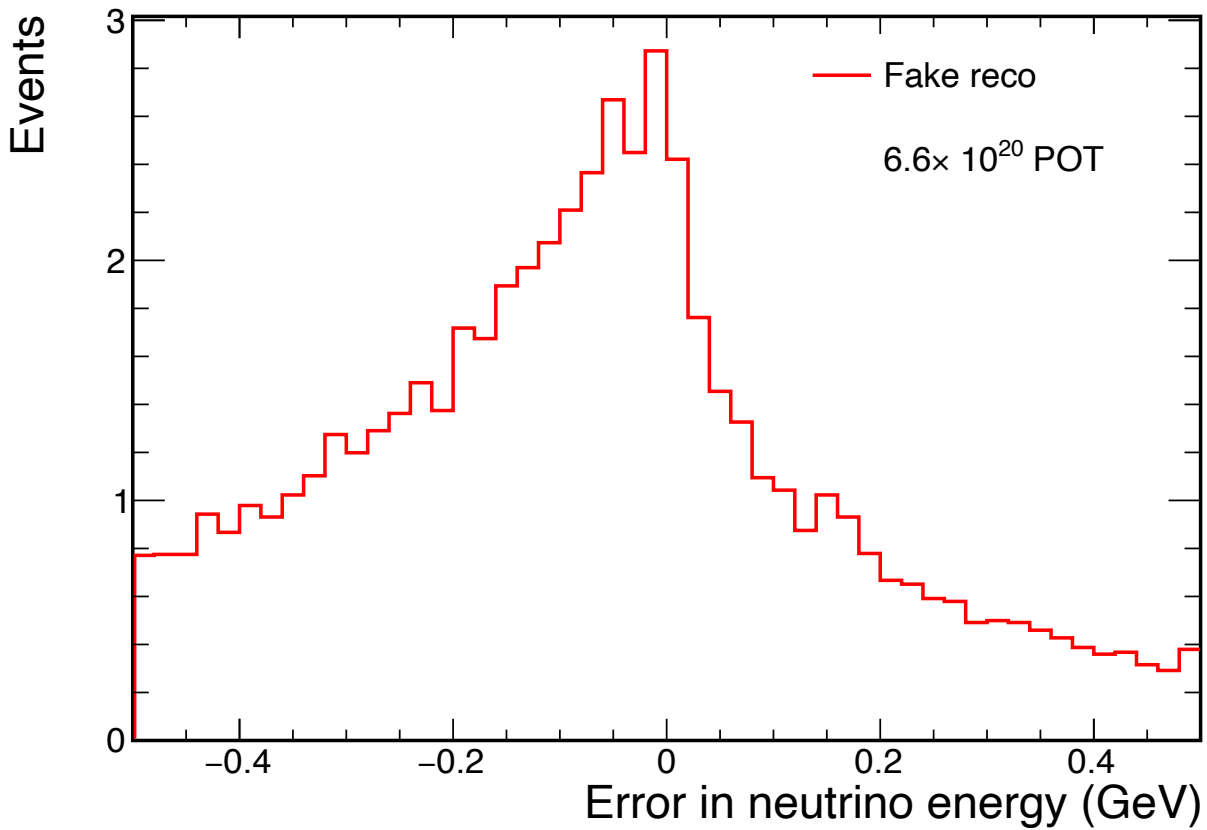
SBND Simulation



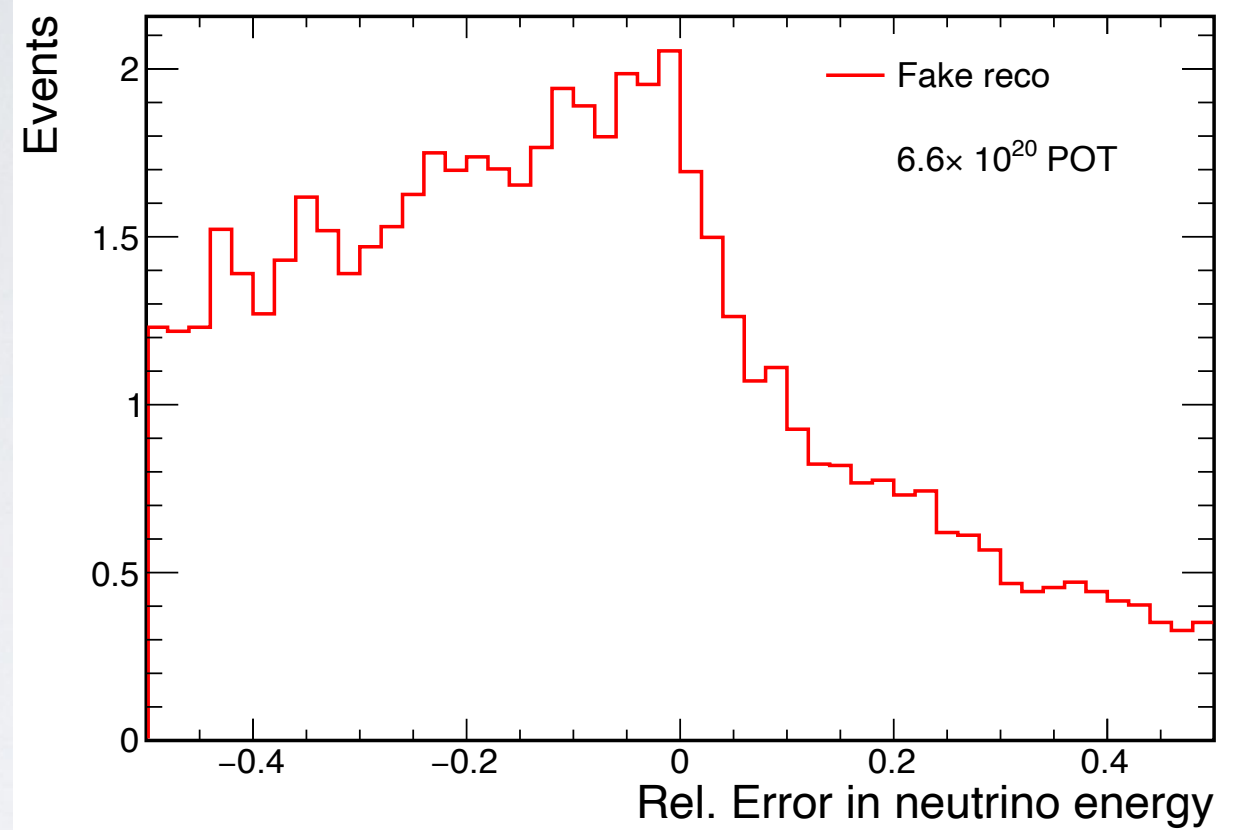
Smearing energy reconstruction by a gaussian 10% (unbiased)

We can absorb this with unfolding (reco-to-true matrix)

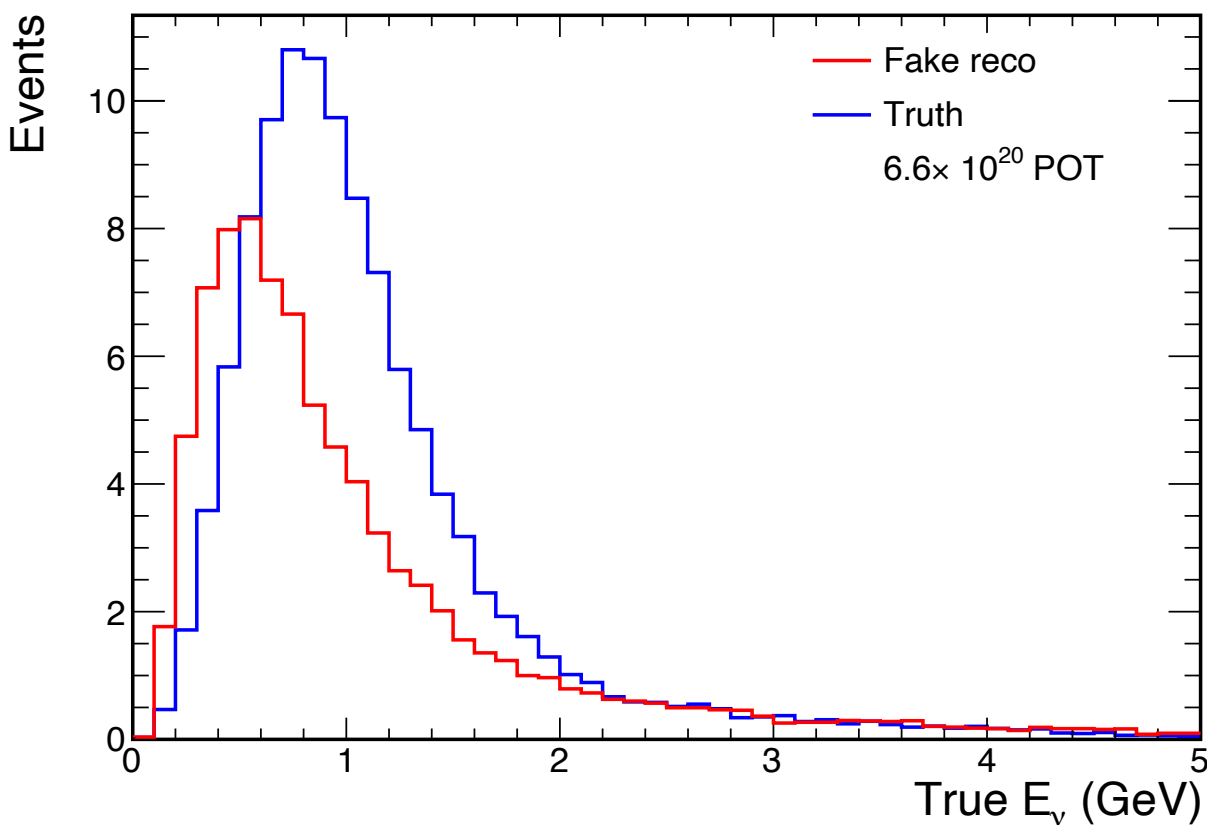
SBND Simulation



SBND Simulation



SBND Simulation



Smearing the angle by 1° (unbiased)

Improving the direction reconstruction is the priority

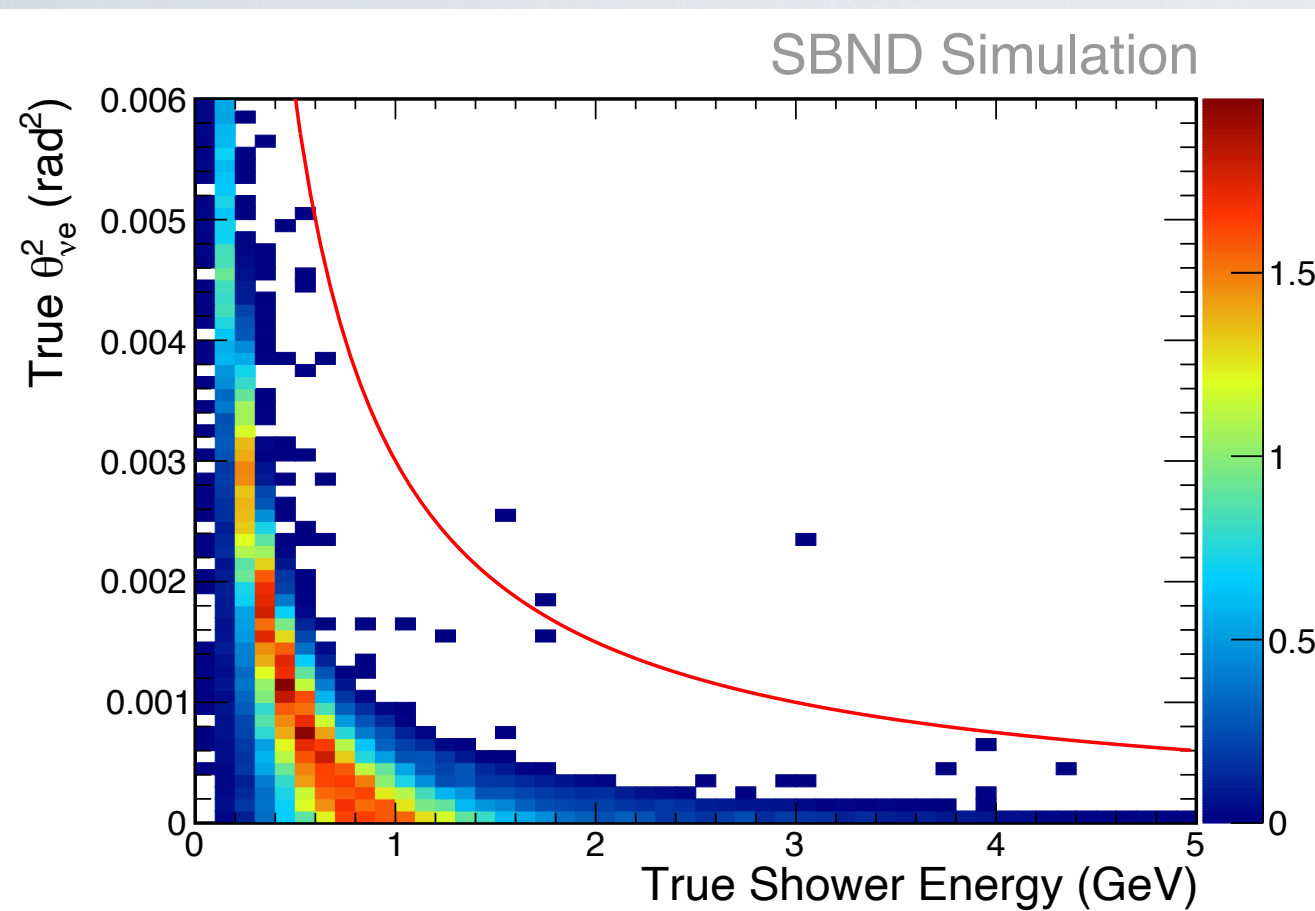
ABOUT THE KINEMATIC CUT

In the relativistic limit:

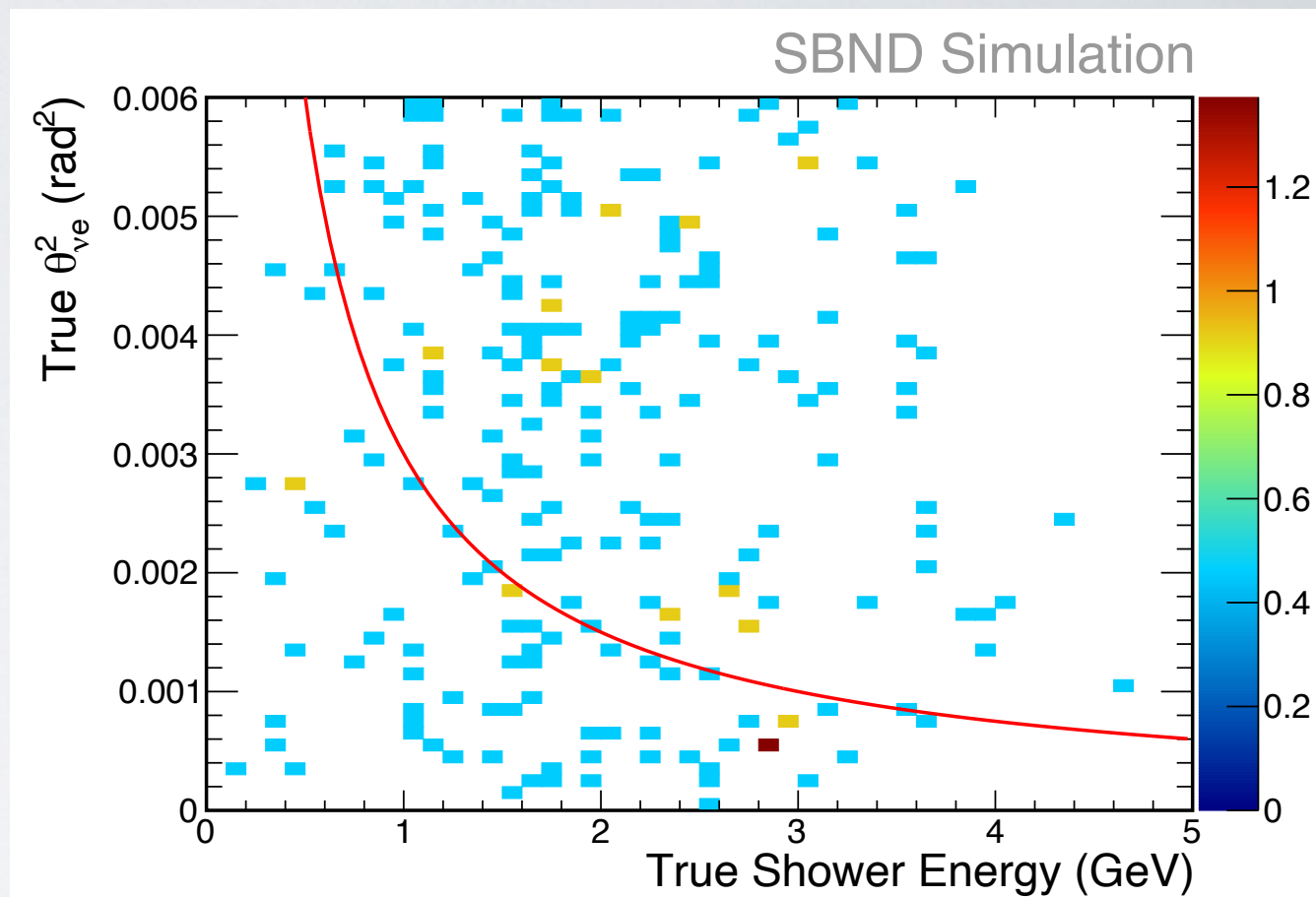
$$1 - \cos \theta = \frac{m_e (1 - y)}{E_e}, \text{ where } y = \frac{T_e}{E_\nu}$$

$$E_e \theta^2 \simeq 2m_e (1 - y) \leq 2m_e = 0.001 \text{ GeV}$$

SELECTION (TRUE VARIABLES)



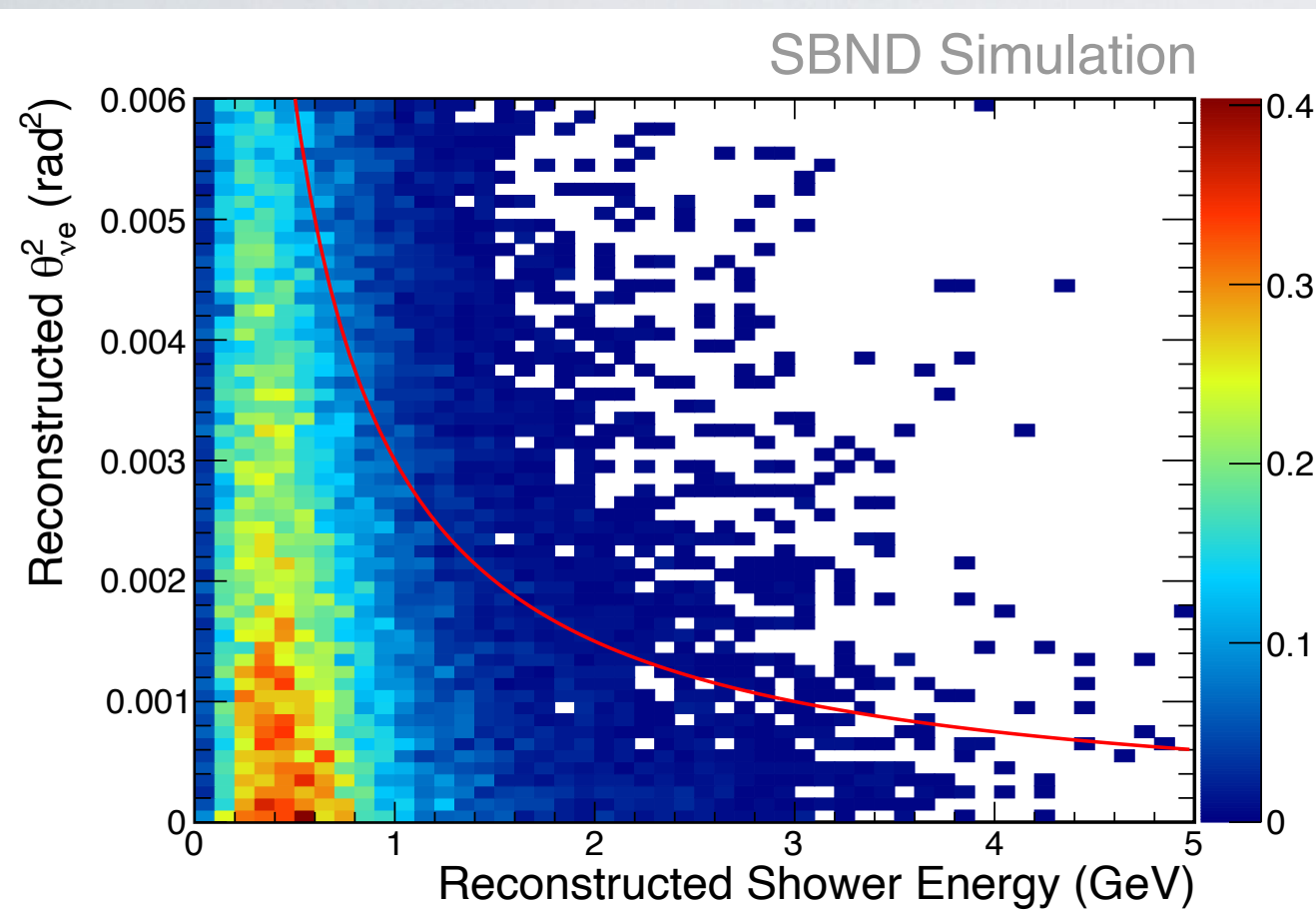
Signal



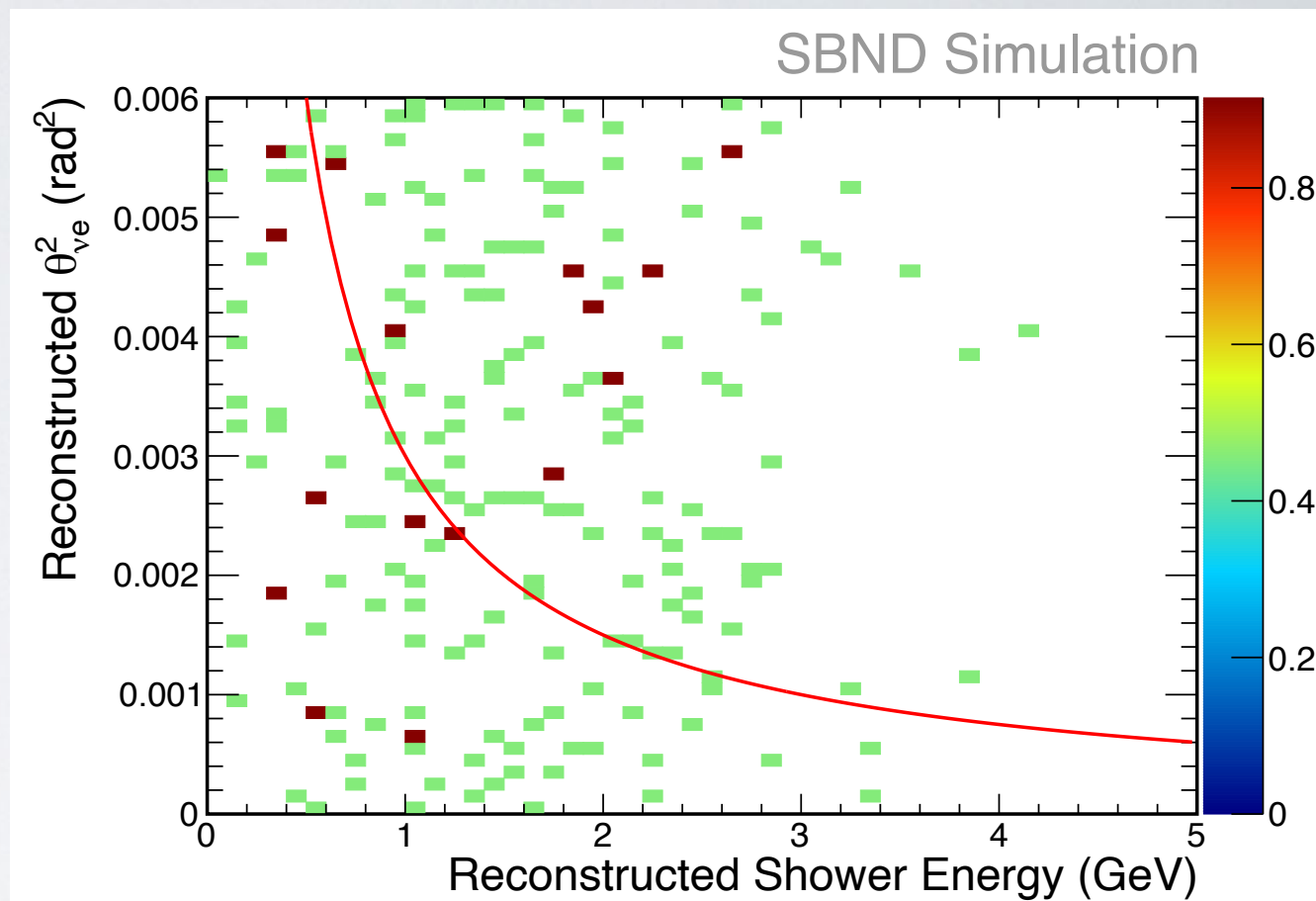
Background

Optimum is at $E\theta^2 < 0.0011 \text{ GeV} \cdot \text{rad}^2$, very close to the theoretical boundary

SELECTION (RECO VARIABLES)



Signal



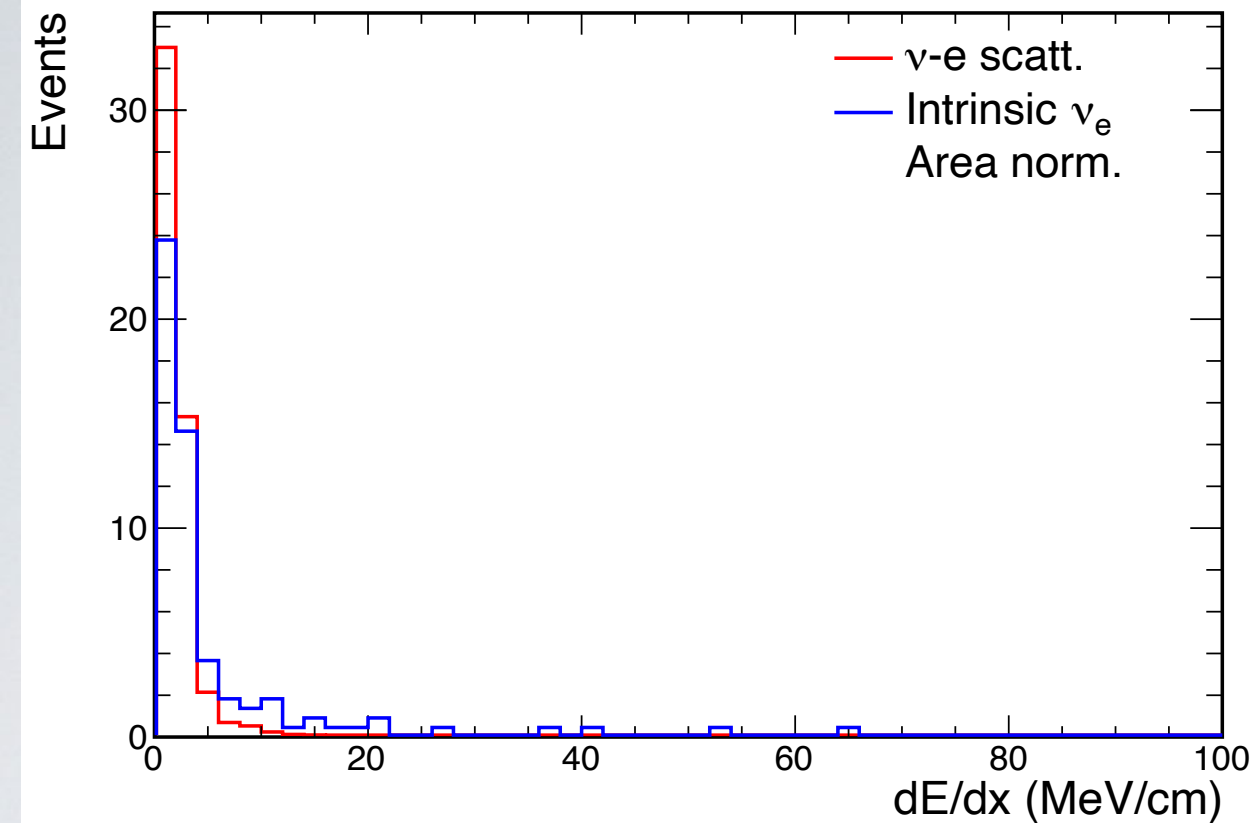
Background

Optimum is at $E\theta^2 < 0.0029 \text{ GeV} \cdot \text{rad}^2$

Checked linear cut and we gain nothing (~same efficiency and purity)

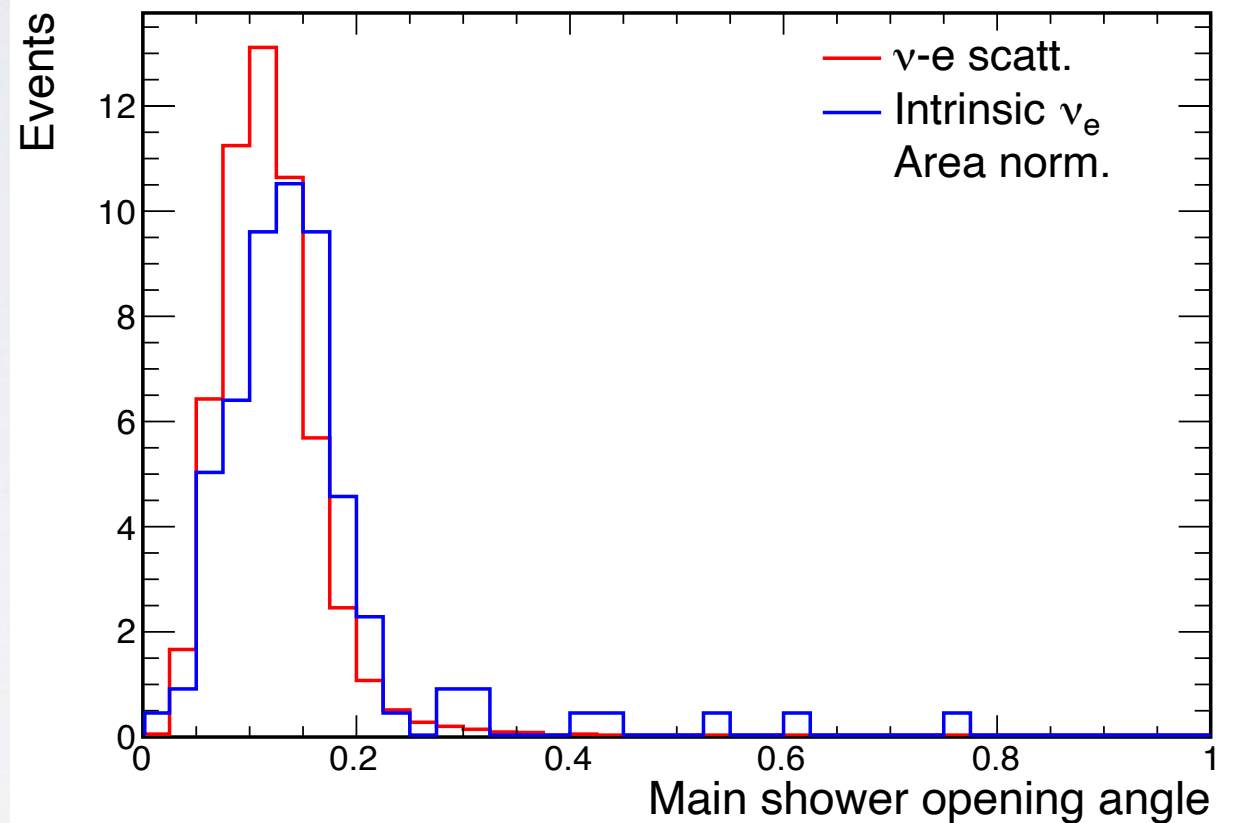
ADDITIONAL VARIABLES?

SBND Simulation



Cut < 30 MeV/cm

SBND Simulation

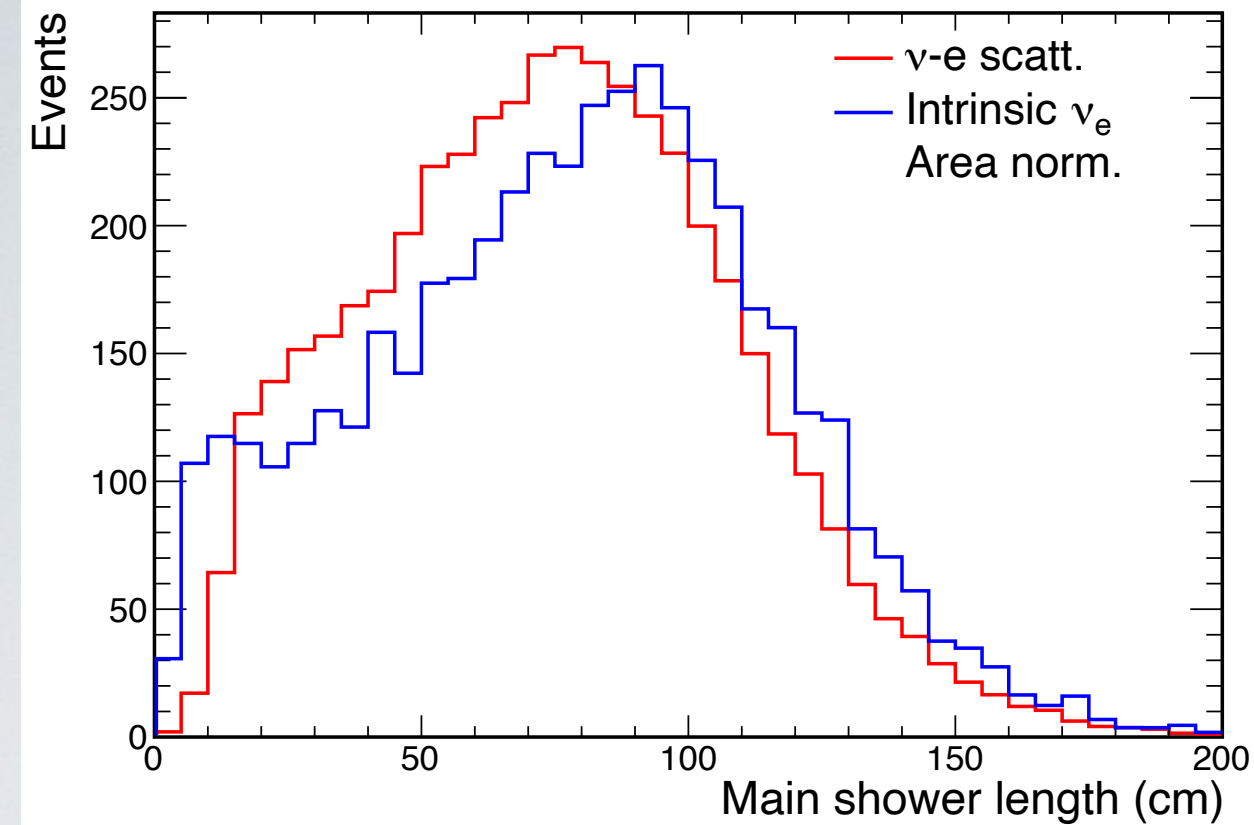


Cut < 0.5

Preliminary cuts (i.e., eyeballed), on top of previous cuts
Underlying idea: what's left is events where a hadronic shower is collinear

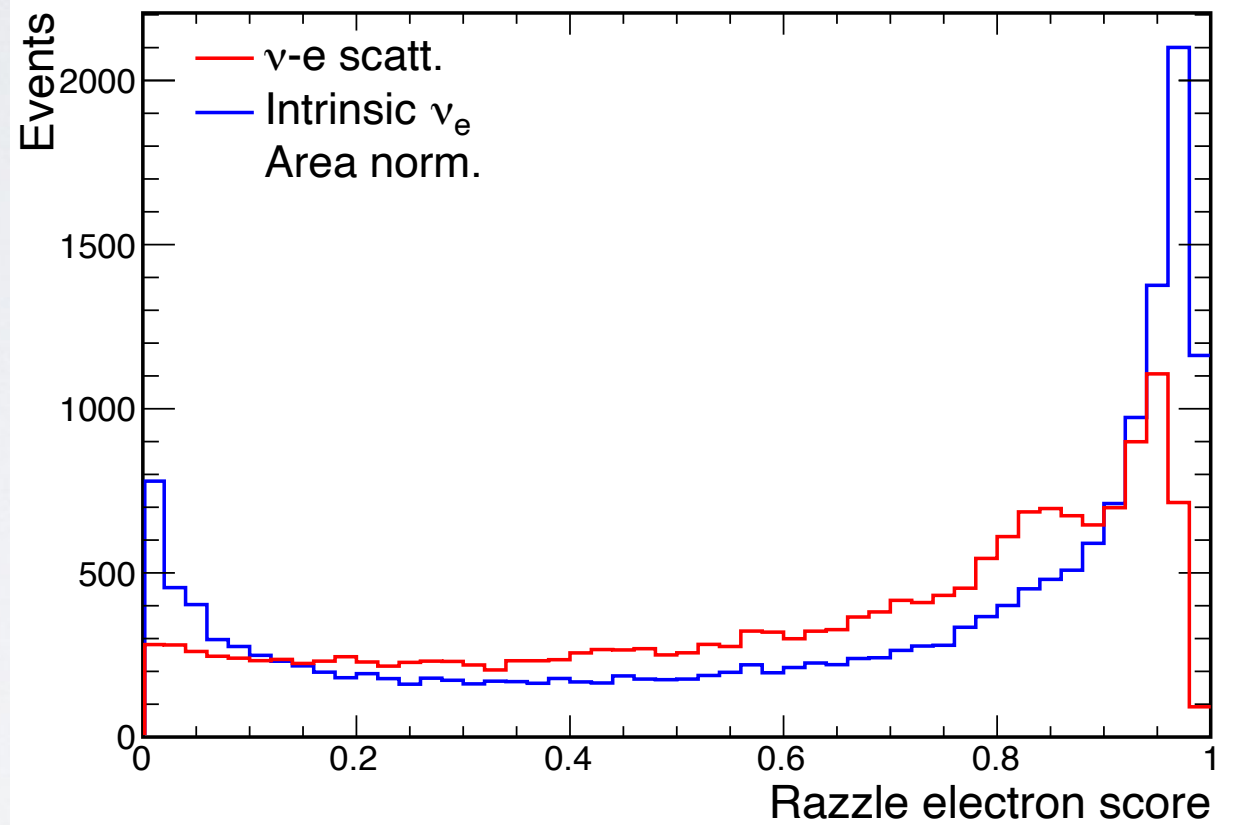
ADDITIONAL VARIABLES?

SBND Simulation



Cut > 10 cm

SBND Simulation



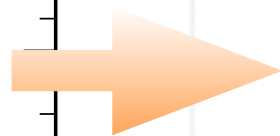
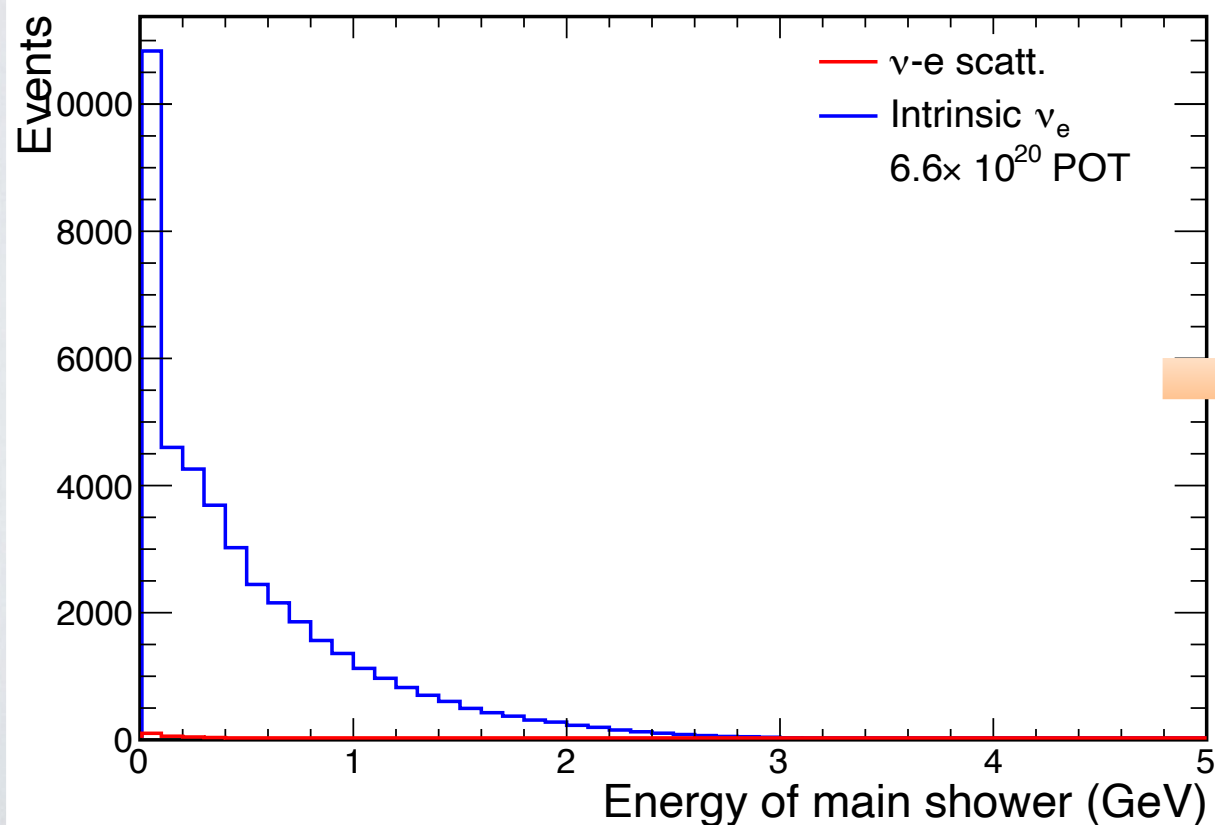
Cut > 0.2

Preliminary cuts (i.e., eyeballed), on top of previous cuts

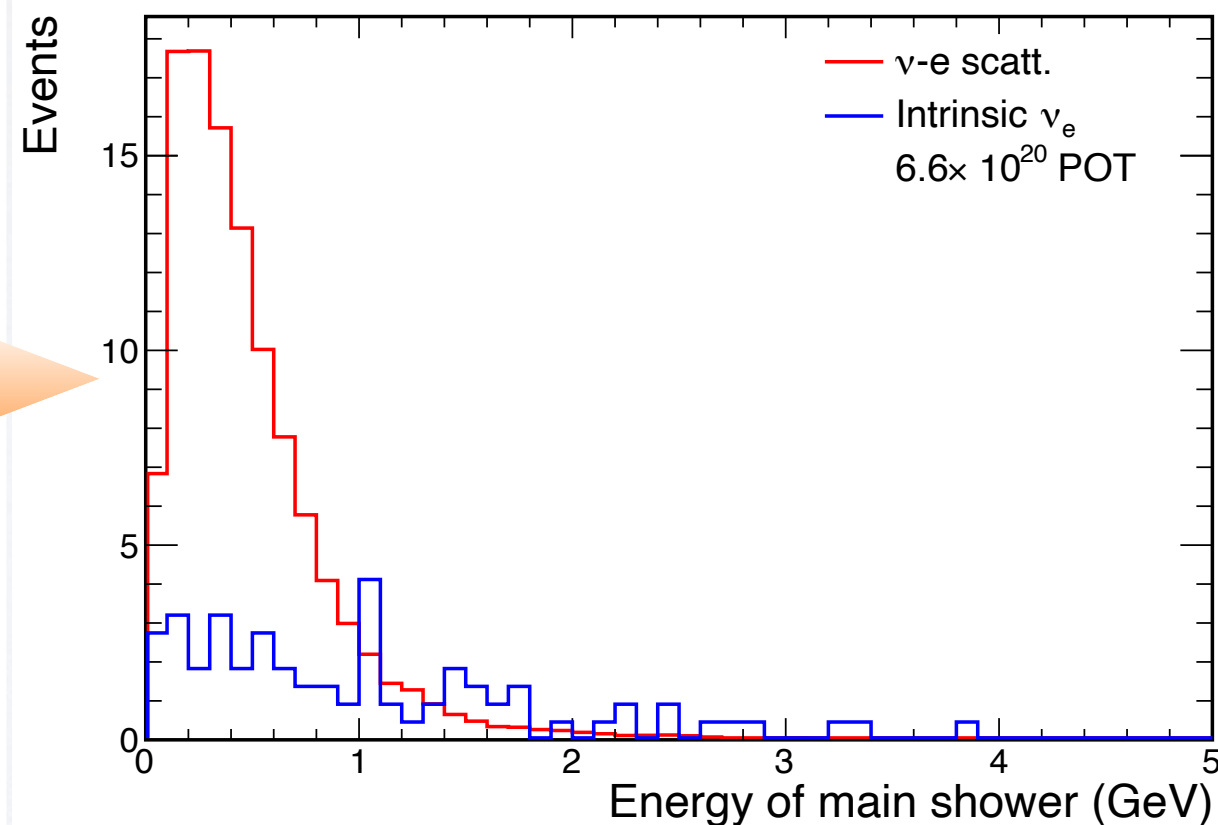
CUTFLOW

Cut	Efficiency	Purity
No cut	100%	0.8%
One shower	68.5%	1.4%
No tracks	59.0%	4.3%
$E\theta^2 < 0.003$	32.7%	68.9%
All remaining	30.4%	74.3%

SBND Simulation



SBND Simulation



CONCLUSIONS

- In order to reconstruct the energy of the incoming neutrino with good precision, the main goal lies in improving the reconstruction (particularly direction)
- Event selection would also greatly benefit from such improvement
 - Next step, test out using Marina's direction reconstruction algorithm
- There seems to be some room for improvement in the selection but most of the low-hanging fruit has already been picked