# Entanglement Entropy and Holography

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**Abstract:** The AdS/CFT correspondence, also called *holography*, is a physical duality between quantum gravity theories in anti-de Sitter (AdS) spacetimes and certain quantum field theories (QFTs) with conformal symmetry defined in the boundary of such space. The so-called holographic dictionary describes how quantities from each of these theories can be translated into quantities of the other. An important magnitude of the holographic dictionary is the entanglement entropy (EE) of boundary regions. This measures the degree of quantum entanglement between such regions and their complements. In this work, we study various aspects of EE in the holographic context. After a quick review of AdS/CFT and of general aspects of EE in QFT, we introduce the Ryu-Takayanagi formula, which computes the holographic EE of boundary regions in the semiclassical limit of the gravity side of the duality. We perform explicit calculations and general checks of the formula, review its generalizations to account for stringy and quantum corrections, and comment on its relation with black hole thermodynamics and the emergence of gravitational dynamics.

## I. INTRODUCTION

The search for a unification of quantum theories and gravity is a long-standing challenge in theoretical physics. The AdS/CFT correspondence —appeared in the late 1990s in the context of string theory [1]— provides such unification in the particular case of spacetimes with a timelike boundary. AdS/CFT is also a realization of the holographic principle, which establishes that quantum gravity theories must admit an equivalent description in terms of theories which live at the boundary of the corresponding spacetimes. In addition to its fundamental implications in the search for an ultimate theory of all the interactions, AdS/CFT provides an useful framework for proving aspects of quantum field theories (QFTs) in regimes in which the usual tools cannot be applied —*e.g.*, when the fields are strongly coupled— using the tools of classical general relativity. Complementarily, AdS/CFT allows us to study the quantum behavior of gravity by relating it to well-defined and well-understood QFTs. In this work we study an example of the first class of applications, namely, how the quantum entanglement of region algebras in holographic QFTs is encoded in a completely classical quantity from the gravity side.

In section IIsection*.2 we introduce the holographic principle and the AdS/CFT correspondence. The notion of entanglement entropy (EE) and some of its properties —including a generalization of the first-law of thermodynamics— in the context of conformal field theories (CFTs) is presented in section IIIsection*.5. In section IVsection*.7 we introduce the Ryu-Takayanagi formula, which allows to compute the holographic EE of boundary regions from the area of extremal surfaces in anti-de Sitter space. We evaluate the EE in the case of a circular region in CFT$_3$ explicitly and verify that the prescription fulfills the fundamental property of *strong subadditivity*. In section IVsection*.7 we provide an overview of how the Ryu-Takayangi formula gets corrected by stringy and quantum effects.

## II. HOLOGRAPHY AND ADS/CFT

### A. The Holographic Principle

Given a finite space region, one can imagine a process in which matter is continuously added into it. This will make the entropy increase. However, there is a limit to the amount of matter that can be introduced in the region, corresponding to the moment in which a black hole is formed. The entropy of a black hole only depends on its surface area and it is given, in Planck units, by [2, 3]

$$S_{\mathrm{BH}} = \frac{A_{\mathrm{H}}}{4G} \, , \qquad (1)$$

where $A_{\mathrm{H}}$ is the area of the event horizon of the black hole and $G$ is Newton's gravitational constant. As a consequence, the maximum entropy that a region can contain is given by its area divided by $4G$.

This implies that the degrees of freedom inside some region grow with the area of the boundary and not with the volume of the region, as one might have expected. This leads to the *holographic principle*, which states that in a quantum gravity theory all physics phenomena within some volume must be describable in terms of a theory defined on the boundary of the region [4].

### B. AdS/CFT Correspondence

The *AdS/CFT correspondence*, sometimes simply called *holography* or *gauge/gravity correspondence* [1], is

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an explicit realization of the holographic principle. It establishes the complete physical equivalence between quantum gravity theories living in anti-de Sitter spacetimes and certain types of CFTs living in the boundary of such spacetimes. If the gravitational theory is defined in $(d+1)$ spacetime dimensions, the dual CFT is defined in $d$ spacetime dimensions and, in a precise sense, the gravity theory will be a "hologram" of the CFT. AdS/CFT allows us to study aspects of each of these theories through the other. The so-called *holographic dictionary* maps quantities (observables) between the gravity theories and their dual CFTs [5]. For example, an empty AdS spacetime with no matter is dual to the vacuum state of the CFT, and an AdS spacetime with a black hole inside corresponds to a thermal state in the CFT.

An anti-de Sitter spacetime is a maximally symmetric spacetime with negative curvature, which solves Einstein's field equations with a negative cosmological constant. The metric of an AdS spacetime of $(d+1)$ dimensions in Poincaré coordinates is

$$\mathrm{d}s^2_{\mathrm{AdS}_{(d+1)}} = \frac{L^2}{z^2}\left(-\mathrm{d}t^2 + \mathrm{d}z^2 + \sum_{i=1}^{d-1}\mathrm{d}x_i^2\right)\,, \qquad (2)$$

with the time and spatial dimensions $t, x_i \in (-\infty, +\infty)$, an extra dimension $z \in (0, +\infty)$, sometimes called *holographic coordinate*, and where $L$ is the *AdS radius*. Fixing the coordinate $z$, the metric reduces to the one of $d$-dimensional Minkowski spacetime "weighted" by the constant factor $1/z^2$. Hence, AdS can be foliate by Minkowski spacetimes living at different values of $z$.

$\mathrm{AdS}_{d+1}$ can be represented as a cylinder where each slice corresponds to a constant time and where $z$ grows radially towards the center [6]. Each slice has a $d$-dimensional boundary $\partial\mathrm{AdS}_{(d+1)}$ where the $\mathrm{CFT}_d$ lives (Fig. 1AdS$_3$ spacetime. In the conformal boundary lives the $\mathrm{CFT}_2$figure.1).
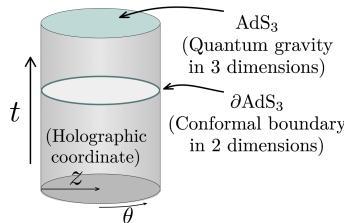


FIG. 1: AdS$_3$ spacetime. In the conformal boundary lives the CFT$_2$.

CFTs, on the other hand, are QFTs that are invariant under conformal transformations. These are angle-preserving transformations which leave the metric invariant up to an overall factor [7]. In particular, the Poincaré group is a subgroup of the conformal group, but there are additional transformations corresponding to dilatations and special conformal transformations. The number of generators of a $d$-dimensional CFT coincides with the number of isometries of a $(d+1)$-dimensional AdS spacetime. This is a hint of the holographic duality.

The first instance of the AdS/CFT correspondance ever described was the duality between $d = 4$, $\mathcal{N} = 4$ Super Yang-Mills theory and type-IIB string theory on AdS$_5\times$S$_5$ [1], but many other examples are known by now. Many general rules of the duality can be exploited without specifying the full field content of the theories and here we will make use of this fact.

The gauge/gravity correspondence is valid independently of the intensity of the gravitational coupling. Interestingly, a strongly coupled CFT with a large number of colors is dual to a classical gravitational theory. As it will be seen shortly, in this situation, it is possible to explain classical gravitational fenomena by highly quantum features, and vice versa, using the holographic dictionary.

## III. ENTANGLEMENT ENTROPY IN CFT

Given some quantum system composed of two subsystems $A$ and $B$ in a pure state $|\Psi\rangle$, there are two possibilities. If one can write $|\Psi\rangle$ as a single tensor product of pure states, $|\Psi\rangle = |\Phi\rangle_A \otimes |\tilde{\Phi}\rangle_B$, one say that the state is separable. If this is not possible, $|\Psi\rangle \neq |\Phi\rangle_A \otimes |\tilde{\Phi}\rangle_B$, one say that the state is entangled. In the latter case, one cannot describe neither of them independently without losing information (in other words, if one trace over one of the subsystems, the reduced state of the other will not be pure). The two form an inseparable entity.

The *entanglement entropy* (EE) is a measure of the degree of quantum entanglement between two subsystems [8]. It is defined by the von Neumann entropy of the reduced density matrix $\rho_A$ of one of the subsystems as

$$S(A) \equiv -\,\mathrm{tr}_A(\rho_A \log \rho_A)\,, \qquad (3)$$

being $\rho_A = \mathrm{tr}_B |\Psi\rangle\langle\Psi|$. If $\lambda_i$ are the eigenvalues of $\rho_A$, then the entanglement entropy would take the simplified form $S = -\sum_i \lambda_i \log \lambda_i$.

The von Neumann entropy is always positive, and it is null for a pure state. Hence, the entanglement entropy of separable states vanishes, as it should.

The natural subsystems in QFT are spacetime regions. For any QFT, given a global state and some region $A$, there is a regulated sense in which it induces a density matrix $\rho_A$ from which one can compute its EE with respect to its causal complement. This EE is intrinsically divergent, since the region is separated from its vicinity by a zero-dimensional boundary. Nonetheless, one can regulate it and obtain physically meaningful results.

The general expression of the entanglement entropy for an arbitrary region in a $d$-dimensional CFT is —see *e.g.,* [8],

$$S_A^{\mathrm{CFT}_d} = c_{d-2}\left(\frac{H}{\delta}\right)^{d-2} + c_{d-1}\left(\frac{H}{\delta}\right)^{d-4} + \dots \qquad (4)$$

$$+ \begin{cases} c_1\dfrac{H}{\delta} + (-1)^{\frac{d-1}{2}} s_{\mathrm{univ}} & \text{for odd } d\,, \\[2ex] c_2\dfrac{H}{\delta} + (-1)^{\frac{d-2}{2}} s_{\mathrm{univ}} \log\left(\dfrac{H}{\delta}\right) + c_0 & \text{for even } d\,, \end{cases}$$

where $H$ is the characteristic length of region $A$, $\delta$ is an ultraviolet cut-off, $c_i$ are coefficients that are non-universal (not well-defined in the continuum, i.e., dependent on the definition of $\delta$), and $s_{\text{univ}}$ are universal coefficients that contain well-defined ("universal") information about the corresponding QFT.

### A.  The first law of entanglement entropy

The EE satisfies an interesting relation known as the *first law of entanglement entropy*. Here we derive it and show that it incorporates the usual first-law of thermodynamics as a particular case.

Consider a family of states $|\Psi(\lambda)\rangle$ defined by the parameter $\lambda$ of a general quantum system with a subsystem $A$. The first order variation of the EE is [9]

$$\frac{\mathrm{d}}{\mathrm{d}\lambda} S_A = -\operatorname{tr}\left(\log \rho_A \frac{\mathrm{d}}{\mathrm{d}\lambda}\rho_A\right) - \operatorname{tr}\left(\frac{\mathrm{d}}{\mathrm{d}\lambda}\rho_A\right) \ , \quad (5)$$

where the last term vanishes since the trace of the density matrix is one for any state.

Now, given any state $\rho$, its *modular hamiltonian* is defined as $H = -\log\rho$. Using this, the last equation can be rewritten as

$$\frac{\mathrm{d}}{\mathrm{d}\lambda} S_A = \frac{\mathrm{d}}{\mathrm{d}\lambda}\langle H_A \rangle \ , \quad (6)$$

where $\langle H_A \rangle \equiv \operatorname{tr}(H_A \rho_A)$ is the expectation value of the modular Hamiltonian of $\rho_A$ in the same state. This is the first law of EE. It is called like this because it represents a quantum generalization of the first law of thermodynamics. One can see this by considering the specific case where the unperturbed density matrix is in a thermal state. In that case, the modular Hamiltonian is related to the usual Hamiltonian of the system $H$ by $\rho_A = e^{-H/T}$, where $T$ is the temperature, and the EE becomes a thermal entropy. Applying Eq. 6equation.3.6 and the definition of the modular hamiltonian, it is immediate to obtain that, in this case,

$$\frac{\mathrm{d}\langle H \rangle}{\mathrm{d}\lambda} = T\frac{\mathrm{d}S_A}{\mathrm{d}\lambda} \ . \quad (7)$$

In other words, $\mathrm{d}E = T\,\mathrm{d}S$. Therefore, we see that the usual first law of thermodynamics is a consequence of the more fundamental first law of EE.

## IV.  HOLOGRAPHIC ENTANGLEMENT ENTROPY

### A.  Ryu-Takayanagi formula

For general CFTs, it is very difficult to compute the EE of a region. On the other hand, it turns out to be rather easy to do it for holographic CFTs. Remarkably,

in the holographic context, an essentially quantum quantity such as EE can be obtained from areas of extremal surfaces on AdS spacetime.

Given a gravity theory defined on $(d+1)$-dimensional AdS spacetime, the dual CFT will live at its conformal boundary. This is a $d$-dimensional Minkowski space, which lies, in Poincaré coordinates, at $z = \delta \ll 1$. The EE for a region $A$ in the holographic CFT can be computed using the so-called Ryu-Takayanagi formula [10]:

$$S_A = \frac{\operatorname{Area}(\gamma_A^{\min})}{4G} \ , \quad (8)$$

where $\gamma_A^{\min}$ is the surface of minimal area defined on AdS spacetime connected to the $(d-1)$-dimensional boundary $\partial A$ of the region $A$, and $G$ is the $(d+1)$-dimensional Newton constant (see Fig. 2Region $A$ (dark blue) and its boundary $\partial A$ inside a $z = \delta$ AdS slide (grey) and its respective $\gamma_A$ (light blue) inside the AdS spacetimefigure.2).

The area of $\gamma_A^{\min}$ is obtained by

$$\operatorname{Area}(\gamma_A^{\min}) = \int_{\gamma_A} \sqrt{h} \ \mathrm{d}^d y \ , \quad (9)$$

where $y$ are the $d$ coordinates that represent the surface $\gamma_A^{\min}$ and $h$ is the determinant of the metric $h_{ij} = (\partial x^\mu/\partial y^i)(\partial x^\nu/\partial y^j)\,g_{\mu\nu}$ induced on the surface by the surrounding spacetime.

The Ryu-Takayanagi formula is valid for generic systems, and provides a hint on how the geometry of spacetime can emerge from mere quantum information.

As one can verify, the Ryu-Takayanagi formula for a $(d+1)$-dimensional anti-de Sitter reproduces the expected general behaviour of the EE (Eq. 4equation.3.4) for a $d$-dimensional conformal field theory [11, 12]. Let us see this explicitly with an example.
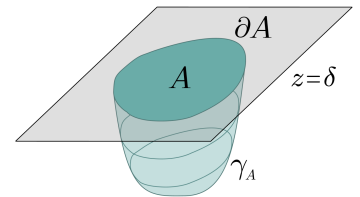


FIG. 2: Region $A$ (dark blue) and its boundary $\partial A$ inside a $z = \delta$ AdS slide (grey) and its respective $\gamma_A$ (light blue) inside the AdS spacetime.

### B.  Entanglement Entropy for a Disk in CFT$_3$

In this section, we compute the EE for a circular region in a holographic CFT$_3$ dual to Einstein gravity. We will use this calculation to verify the general expression of the EE for a QFT, identify the universal coefficient for this kind of theory, and illustrate how AdS provides us with a geometric ultraviolet regulator.

Let $A$ be a disk-shaped region of radius $R$ defined in the conformal boundary of pure AdS$_4$. This 2-dimensional region is defined in polar coordinates as

$A = \{(r,\theta,z,t) \mid t = 0, z = \delta, r \leq R\}$. We parametrize the bulk surface of minimal area $\gamma_A^{\min}$ as:

$$\gamma_A^{\min} = \{(r,\theta,z,t) \mid t = 0, z = f(r,\theta)\} \ ,$$

where $f(r,\theta)$ is certain function we need to identify. There is no property on the AdS spacetime theory that could prevent the symmetry of $\partial A$ on the coordinate $\theta$ from being transfered to $\gamma_A^{\min}$. Hence, $z = f(r)$.

The metric corresponding to AdS$_4$ spacetime reads

$$\mathrm{ds}^2_{\mathrm{AdS}_4} = g_{\mu\nu}\mathrm{d}x^\mu\mathrm{d}x^\nu = \frac{L^2}{z^2}[-\mathrm{d}t^2 + \mathrm{d}z^2 + \mathrm{d}r^2 + r^2\mathrm{d}\theta^2] \ .$$

The induced metric of the surface $\gamma_A^{\min}$ is in turn given by

$$\mathrm{ds}^2_{\gamma_A^{\min}} = h_{ij}\mathrm{d}x^i\mathrm{d}x^j = \frac{L^2}{f(r)^2}\left[(1 + \dot{f}(r)^2)\,\mathrm{d}r^2 + r^2\mathrm{d}\theta^2\right] \ ,$$

where $\dot{f}(r) \equiv \partial f/\partial r$. The determinant of the induced metric will be $h = L^4 r^2(1 + \dot{f}(r)^2)/f(r)^4$. The minimal value of the integral over the polar coordinates of the square root of the induced metric will correspond to the area of $\gamma_A^{\min}$. So, by the Ryu-Takayanagi formula, the entanglement entropy related to the region $A$ will be

$$S_A = \min\int_{\gamma_A} \frac{\sqrt{h}\,\mathrm{d}x^\rho}{4G} = \frac{\pi L^2}{2G}\min\int_r \mathrm{d}r\frac{r\sqrt{1 + \dot{f}(r)^2}}{f(r)^2} \ ,$$

where in the second line we performed the angular integration. The resulting integral in $r$ must be evaluated on the minimal surface. In order to find it, we extremize the functional using the Euler-Lagrange equations for the effective Lagrangian $\mathcal{L}[r, f(r), \dot{f}(r)]$. They read

$$\frac{\partial\mathcal{L}}{\partial f} - \frac{\mathrm{d}}{\mathrm{d}r}\left[\frac{\partial\mathcal{L}}{\partial\dot{f}}\right] = 0$$

$$\longrightarrow (1 + \dot{f}^2)\left(-2r - f\dot{f} - rf\ddot{f}\right) + rf\dot{f}^2\ddot{f} = 0 \ .$$

One can prove that $f(r) = \sqrt{R^2 - r^2}$ is a solution of the previous relation and corresponds to the function that minimizes the functional of the entanglement entropy and connects to the boundary region $A$. Hence, the surface of minimal area is found to be a half sphere.

To compute the integral of the EE expression, we should first determine its limits carefully. The inferior bound corresponds to the lower part of the half sphere inside the bulk, where $r_{\min} = 0$. The superior bound corresponds to the connexion of the surface with the conformal boundary, that is, $z = \delta = \sqrt{R^2 - r_{\max}^2}$. Observe that if we had not included $\delta$ and had integrated all the way to $z = 0$, the result would have diverged. This is precisely what we expect for the EE, which becomes divergent in the continuum. In this case, the geometric cutoff introduced in the holographic coordinate, $z = \delta$ plays the role of the ultraviolet regulator of the EE. The

final result for the EE of the disk is

$$S_A = \frac{\pi L^2}{2G}\int_0^{\sqrt{R^2-\delta^2}} \mathrm{d}r\frac{r}{f(r)^2}\sqrt{1 + \dot{f}(r)^2} =$$

$$= \frac{\pi L^2}{2G}\frac{R}{\delta} - \frac{\pi L^2}{2G} + \mathcal{O}(\delta) \ .$$

This coincides with the general expression expected for the EE of a 3-dimensional CFT. Repeating the calculation for various regions and in different dimensions, the corresponding structure found is always consistent [10, 11].

From the above expression we can identify the universal contribution to the EE for a 3-dimensional holographic theory dual to Einstein gravity. The result reads

$$s_{\mathrm{univ}} = \frac{\pi}{2}\frac{L^2}{G} \ . \tag{10}$$

In the case of a disk region, like the one we just considered, $s_{\mathrm{univ}}$ is related to another important quantity, namely, the Euclidean free energy on a three-dimensional sphere. One has $s_{\mathrm{univ}} = -\log Z_{\mathbb{S}^3}$ for general CFTs [13]. Hence, we find that for holographic theories dual to four-dimensional Einstein gravity, the sphere free energy is given in terms of the AdS radius and the Newton constant by (10equation.4.10).

### C.  Strong subadditivity

*Strong subadditivity* is a fundamental general property of EE. It relates the EEs of two regions $A$ and $B$ with the ones of their union $A \cup B$ and intersection $A \cap B$ (Fig. 3Representation of the surfaces $\gamma_A^{\min}$ and $\gamma_B^{\min}$, their union $\gamma_{A\cup B}$ and intersection $\gamma_{A\cap B}$, and the bulk region $r_A$figure.3):

$$S(A) + S(B) \geq S(A \cup B) + S(A \cap B) \ , \tag{11}$$

This inequality is fulfilled in any quantum mechanical theory, but it is remarkably difficult to prove in general. An important test of the validity of the Ryu-Takayanagi formula is whether it fulfills it. It turns out to be particularly easy to prove that it does [14].

We define $r_A$ and $r_B$



FIG. 3: Representation of the surfaces $\gamma_A^{\min}$ and $\gamma_B^{\min}$, their union $\gamma_{A\cup B}$ and intersection $\gamma_{A\cap B}$, and the bulk region $r_A$.

to be the bulk regions inside $\gamma_A^{\min}$ and $\gamma_B^{\min}$, respectively. Their union and intersection will be $r_{A\cup B} \equiv r_A \cup r_B$ and $r_{A\cap B} \equiv r_A \cap r_B$. The boundaries of these regions can be decomposed as

$$\partial r_{A\cup B} = (A \cup B) \cup \gamma_{A\cup B} \ , \ \partial r_{A\cap B} = (A \cap B) \cup \gamma_{A\cap B} \ ,$$
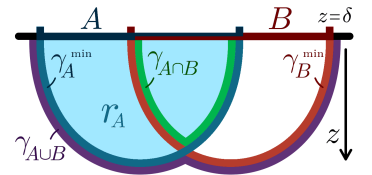
where $\gamma_{A\cup B}$ and $\gamma_{A\cap B}$ are the segments inside the bulk of the boundary surfaces of $r_{A\cup B}$ and $r_{A\cap B}$. These surfaces are connected to $\partial(A\cup B)$ and $\partial(A\cap B)$, respectively, but nothing says that they should be their corresponding minimal area surfaces $\gamma_{A\cup B}^{\min}$ and $\gamma_{A\cap B}^{\min}$. The areas of $\gamma_{A\cup B}$ and $\gamma_{A\cap B}$ correspond to upper bounds for the minimal possible area. Also, observe that the sum of the areas of $\gamma_{A\cup B}$ and $\gamma_{A\cap B}$ trivially equals the sum of the areas of $\gamma_A^{\min}$ and $\gamma_B^{\min}$. Hence,

$$\text{Area}(\gamma_A^{\min}) + \text{Area}(\gamma_B^{\min}) \geq \text{Area}(\gamma_{A\cup B}^{\min}) + \text{Area}(\gamma_{A\cap B}^{\min}) \ ,$$

Therefore, the Ryu-Takayanagi formula (Eq. 8equation.4.8) implements the strong subadditivity property in a simple geometric way which makes use of the minimization principle of AdS holographic surfaces.

### D.   Corrections to the Ryu-Takayanagi formula

The Ryu-Takayanagi formula (8equation.4.8) provides the EE for holographic theories dual to Einstein gravity in general dimensions. Nevertheless, the Einsteinian description will break down if one moves away from the strongly coupled and large-number of colors regime. From the gravity side, this is manifest in the appearance of both stringy and quantum corrections to classical Einstein gravity.

Stringy corrections appear as higher-curvature corrections to the Einstein-Hilbert action. At the level of the holographic EE, these produce corrections to the Ryu-Takayanagi area formula, in a way similar to how the Bekenstein-Hawking formula for the entropy of a black hole (1equation.2.1) is replaced by the Wald formula [15] in the presence of higher-curvature corrections. However, replacing the functional of the entanglement entropy by the Wald one does not work in general. The full formula takes the schematic form [16]

$$S_A = S_{\text{Wald}} + S_{\text{Anomaly}} \tag{12}$$

where $S_{\text{Wald}}$ reduces to the Ryu-Takayanagi one in the Einstein gravity case, but otherwise contains terms involving various components of the Riemann tensor, and $S_{\text{Anomaly}}$ simpliy vanishes for Einstein gravity, but includes terms involving extrinsic curvatures of the generalized minimal surface for more general theories.

One can also consider corrections to the Ryu-Takayanagi formula related to quantum mechanical effects in the bulk. This quantum corrections are essentially given by the entanglement entropy of the quantum fields living inside the bulk region bounded by the minimal area surface, $r_A$ —see Fig. 3Representation of the surfaces $\gamma_A^{\min}$ and $\gamma_B^{\min}$, their union $\gamma_{A\cup B}$ and intersection $\gamma_{A\cap B}$, and the bulk region $r_A$figure.3. The corrected formula reads [17]

$$S_A = \frac{\text{Area}(\gamma_A^{\min})}{4G} + S_{r_A} + \mathcal{O}(G) \ , \tag{13}$$

where the correction $S_{r_A}$ is order $\mathcal{O}(G^0)$ and we have also included a putative subleading correction.

## V.   CONCLUSIONS

We have explored various aspects of the Ryu-Takayanagi formula (see the introduction for a summary), a remarkable entry in the holographic dictionary which allows to compute a genuinely quantum quantity like the EE of boundary regions, in terms of a fully classical quantity related to the geometry of spacetime, namely, the area of extremal surfaces in AdS. This connection between spacetime geometry and entanglement can be made more explicit. Remarkably, one can show that the first law of EE implies, for holographic CFTs, that perturbations of the bulk metric satisfy the linearized Einstein equations [18]. In a sense, the entanglement structure of the boundary theory controls the dynamics of the gravitational field, which therefore becomes an emergent phenomenon.

[1] J. M. Maldacena, International Journal of Theoretical Physics **38**, 1113 (1999), ISSN 00207748, arXiv:hep-th/9711200.

[2] J. D. Bekenstein, Physical Review D **7**, 2333 (1973), ISSN 0556-2821.

[3] S. W. Hawking, Communications in Mathematical Physics **43**, 199 (1975), ISSN 1432-0916.

[4] G. 't Hooft, *Dimensional Reduction in Quantum Gravity* (2009), arXiv:gr-qc/9310026.

[5] E. Witten, *Anti De Sitter Space And Holography* (1998), arXiv:hep-th/9802150.

[6] S. W. Hawking, G. F. R. Ellis, S. W. Hawking, and G. F. R. Ellis, *The large scale structure of space-time*, Cambridge monographs on mathematical physics (Cambridge Univ.

Press, Cambridge, 2008), 21st ed., ISBN 978-0-521-09906-6 978-0-521-20016-5.

[7] P. Di Francesco, P. Mathieu, and D. Sénéchal, *Conformal Field Theory*, Graduate Texts in Contemporary Physics (Springer, New York, NY, 1997), ISBN 978-1-4612-7475-9 978-1-4612-2256-9.

[8] T. Nishioka, Reviews of Modern Physics **90**, 035007 (2018), ISSN 0034-6861, 1539-0756, arXiv:1801.10352 [hep-th].

[9] M. Van Raamsdonk, in *New Frontiers in Fields and Strings* (2017), pp. 297–351, arXiv:1609.00026 [gr-qc, physics:hep-th, physics:quant-ph].

[10] S. Ryu and T. Takayanagi, Physical Review Letters **96**, 181602 (2008), ISSN 0031-9007, 1079-7114, arXiv:hep-th/0603001.

[11] S. Ryu and T. Takayanagi, Journal of High Energy Physics **2006**, 045 (2006), ISSN 1029-8479, arXiv:hep-th/0605073.

[12] T. Nishioka, S. Ryu, and T. Takayanagi, Journal of Physics A: Mathematical and Theoretical **42**, 504008 (2009), ISSN 1751-8113, 1751-8121, arXiv:0905.0932 [cond-mat, physics:gr-qc, physics:hep-th, physics:quant-ph].

[13] H. Casini, M. Huerta, and R. C. Myers, Journal of High Energy Physics **2011**, 36 (2011), ISSN 1029-8479, arXiv:1102.0440 [gr-qc, physics:hep-th].

[14] M. Headrick and T. Takayanagi, Physical Review D **76**, 106013 (2007), ISSN 1550-7998, 1550-2368, arXiv:0704.3719 [cond-mat, physics:hep-th, physics:quant-ph].

[15] V. Iyer and R. M. Wald, Physical Review D **50**, 846 (1994), publisher: American Physical Society.

[16] X. Dong, Journal of High Energy Physics **2014**, 44 (2014), ISSN 1029-8479.

[17] T. Faulkner, A. Lewkowycz, and J. Maldacena, Journal of High Energy Physics **2013**, 74 (2013), ISSN 1029-8479, arXiv:1307.2892 [hep-th].

[18] T. Faulkner, M. Guica, T. Hartman, R. C. Myers, and M. Van Raamsdonk, Journal of High Energy Physics **2014**, 51 (2014), ISSN 1029-8479, arXiv:1312.7856 [hep-th].