

Electromagnetism

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Contents

1	Introduction	3
1.1	Charges and currents	3
1.2	Fields and forces	4
1.3	Maxwell's equations	4
1.4	Units	5
2	Electrostatics	5
2.1	Gauss' Law	5
2.1.1	Spherical symmetry	6
2.1.2	Cylindrical symmetry	7
2.1.3	Planar symmetry	7
2.1.4	Surface charge and discontinuity	8
2.2	The electrostatic potential	8
2.2.1	Point charge	9
2.2.2	Electric dipole	9
2.2.3	Field lines and equipotentials	10
2.2.4	Dipole in an external field	10
2.2.5	Multipole expansion	10
2.3	Electrostatic energy	11
2.4	Conductors	12
3	Magnetostatics	14
3.1	Ampere's law	14
3.2	The magnetic vector potential	16
3.3	The Biot-Savart law	16
3.4	Magnetic dipoles	17
3.5	Magnetic forces	20
4	Electrodynamics	22
4.1	Faraday's law of induction	22
4.2	Ohm's law	26
4.3	Time-dependent electric fields	26
4.4	Electromagnetic waves	27
4.5	Electromagnetic energy	30

1 Introduction

1.1 Charges and currents

Electric charge is a physical property of elementary particles. It is:

- (i) A signed quantity, it can either be positive, negative, or zero.
- (ii) It is quantised to integer multiples of the elementary charge.
- (iii) It is a conserved quantity even if particles are created or destroyed.

By convention the electron has charge $-e$, the proton has charge $+e$ and the neutron has no charge. On macroscopic scales, the number of particles is so large that charge can be considered to have a continuous electric charge density $\rho(\mathbf{x}, t)$. The total charge in a volume V is then

$$Q = \int_V \rho dV.$$

The *electric current density* $\mathbf{J}(\mathbf{x}, t)$ is the flux of electric charge per unit area. The current following through a surface S is

$$I = \int_S \mathbf{J} \cdot d\mathbf{S}.$$

Consider a time-independent volume V with boundary S . Since charge is conserved, we have that

$$\begin{aligned} \frac{dQ}{dt} &= -I \\ \frac{d}{dt} \int_V \rho dV + \int_S \mathbf{J} \cdot d\mathbf{S} &= 0 \\ \int_V \left(\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{J} \right) dV &= 0 \end{aligned}$$

Since this is true for any V , we have that

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{J} = 0.$$

This *equation of charge conservation* has the typical form of a conservation law.

The discrete charge distribution of a single particle of charge q_i ; and position vector $\mathbf{x}_i(t)$, is

$$\begin{aligned} \rho &= q_i \delta(\mathbf{x} - \mathbf{x}_i(t)), \\ \mathbf{J} &= q_i \dot{\mathbf{x}}_i \delta(\mathbf{x} - \mathbf{x}_i(t)). \end{aligned}$$

For N particles, it is

$$\begin{aligned} \rho &= \sum_{i=1}^N q_i \delta(\mathbf{x} - \mathbf{x}_i(t)) \\ \mathbf{J} &= \sum_{i=1}^N q_i \dot{\mathbf{x}}_i \delta(\mathbf{x} - \mathbf{x}_i(t)). \end{aligned}$$

As an exercise we can see that these satisfy the equation of charge conservation.

1.2 Fields and forces

Electromagnetism is a *field theory*.

Charged particles don't interact directly, but rather by generating fields around them, which are then experienced by other charged particles. In general we have two time-dependent vector fields, the electric field $\mathbf{E}(\mathbf{x}, t)$, and the magnetic field $\mathbf{B}(\mathbf{x}, t)$.

The *Lorentz force* on a particle of charge q and velocity \mathbf{v} is

$$\mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B}).$$

1.3 Maxwell's equations

In this course we will explore some consequences of Maxwell's equations.

Definition. (Maxwell's equations)

$$\begin{aligned}\nabla \cdot \mathbf{E} &= \frac{\rho}{\varepsilon_0} \\ \nabla \cdot \mathbf{B} &= 0 \\ \nabla \times \mathbf{E} &= -\frac{\partial \mathbf{B}}{\partial t} \\ \nabla \times \mathbf{B} &= \mu_0 \left(\mathbf{J} + \varepsilon_0 \frac{\partial \mathbf{E}}{\partial t} \right).\end{aligned}$$

Remark. We have some properties about these equations.

- Coupled linear PDEs in space and time,
- Involve two positive constants:
 - (i) ε_0 (vacuum permittivity)
 - (ii) μ_0 (vacuum permeability)
- Charges (ρ) and currents (\mathbf{J}) are the sources of electromagnetic fields.
- Each equation is an equivalent integral form (see later) related via the divergence or Stokes' theorem.
- These are the *vacuum* equations that apply on microscopic scales or in a vacuum. A related macroscopic version applies in media (Part II Electrodynamics).
- The equations are consistent with each other and with charge conservation. We will show this now.
 - (i) Taking the divergence of the third equation, this agrees with the time derivative of the second equation.
 - (ii) For charge conservation, we have that

$$\begin{aligned}\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{J} &= \frac{\partial}{\partial t} (\varepsilon_0 \nabla \cdot \mathbf{E}) + \nabla \cdot \left(-\varepsilon_0 \frac{\partial \mathbf{E}}{\partial t} + \frac{1}{\mu_0} \nabla \times \mathbf{B} \right) \\ &= 0.\end{aligned}$$

1.4 Units

The SI unit of electric charge is the coulomb (C). The elementary charge is exactly

$$e = 1.602\,176\,634 \times 10^{-19} \text{ C}.$$

The SI unit of electric current is the ampere or amp (A) which is equal to 1 C s^{-1} .

The SI base units needed in electromagnetism and then the second, metre, kilogram, and ampere. From the Lorentz force law we see that the units of \mathbf{E} and \mathbf{B} must be

$$\text{kg m s}^{-3} \text{A}^{-1} \quad \text{and} \quad \text{kg s}^{-2} \text{A}^{-1}.$$

We sometimes refer to the units of \mathbf{B} as the *Telsa* (T).

From Maxwell's equations we can work out the units of ε_0 and μ_0 . The values of these constants can be calculated via experimentation as

$$\begin{aligned}\varepsilon_0 &= 8.854 \dots \times 10^{-12} \text{ kg}^{-1} \text{m}^{-3} \text{s}^4 \text{A}^2 \\ \mu_0 &= 1.256 \dots \times 10^{-6} \text{ kg m s}^{-2} \text{A}^{-2}\end{aligned}$$

The speed of light is exactly

$$c = \frac{1}{\sqrt{\mu_0 \varepsilon_0}} = 299\,792\,458 \text{ m s}^{-1}.$$

2 Electrostatics

In a time-independent situation, Maxwell's equations reduce to

$$\begin{aligned}\nabla \cdot \mathbf{E} &= \frac{\rho}{\varepsilon_0} \\ \nabla \cdot \mathbf{B} &= 0 \\ \nabla \times \mathbf{E} &= 0 \\ \nabla \times \mathbf{B} &= \mu_0 \mathbf{J}\end{aligned}$$

Now \mathbf{E} and \mathbf{B} are decoupled so we can study them separately. Electrostatics is the study of the electric field generated by a stationary charge distribution. We'll be looking at

$$\nabla \cdot \mathbf{E} = \frac{\rho}{\varepsilon_0}, \quad \nabla \times \mathbf{E} = 0.$$

2.1 Gauss' Law

Consider a closed surface S enclosing a volume V . Integrate over V and use the divergence theorem to obtain Gauss' law which is

$$\int_S \mathbf{E} \cdot d\mathbf{S} = \frac{Q}{\varepsilon_0},$$

Where

$$Q = \int_V \rho dV$$

is the total charge in V . Gauss' law is the integral version of the first of Maxwell's equations and is valid generally. We get that electric flux is proportional to the total charge enclosed.

In special situations we use Gauss' law together with symmetry to deduce from ρ , by choosing the *Gaussian surface* S appropriately.

2.1.1 Spherical symmetry

Consider a spherically symmetric charge distribution, $\rho(r)$ in spherical polar coordinates with total charge Q contained within an outer radius R . To have spherical symmetry, the electric field should have the form

$$\mathbf{E} = E(r)\mathbf{e}_r.$$

This will satisfy $\nabla \times \mathbf{E} = 0$ as required.

To find $E(r)$ apply Gauss' law to a sphere of radius r . If $r > R$ then we get that

$$\begin{aligned}\int_S \mathbf{E} \cdot d\mathbf{S} &= E(r) \int_S \mathbf{e}_r \cdot d\mathbf{S} \\ &= E(r) \int_S dS \\ &= E(r)4\pi r^2 = \frac{Q}{\epsilon_0}.\end{aligned}$$

Thus

$$\mathbf{E} = \frac{Q}{4\pi\epsilon_0 r^2}\mathbf{e}_r.$$

So the external electric field of a spherically symmetric body depends only on the total charge, and is equivalent to a point charge at the origin with all of the charge. The Lorentz force on a particle of charge q in $r > R$ is

$$\mathbf{F} = q\mathbf{E} = \frac{Qq}{4\pi\epsilon_0 r^2}\mathbf{e}_r.$$

This is the *Coulomb force* between charge particles. The force is repulsive if the charges have the same sign and attractive if the charges have different sign.

In the limit as $R \rightarrow 0$ we obtain the electric field at a *point charge* Q , corresponding to

$$\rho = Q\delta(\mathbf{x}).$$

There is a close analogy between the Coulomb force and the gravitational force between massive particles, recall from IA Dynamics and Relativity that

$$\mathbf{F} = -\frac{GMm}{r^2}\mathbf{e}_r.$$

Both involve an inverse-square law and the product of the charges, however there are some differences.

- (i) While gravity is always attractive, electric forces can be repulsive or attractive;
- (ii) Gravity is very much weaker, due to the much smaller constant of proportionality.

For example if we consider two protons, the ratio of the electric to gravitational force is 10^{36} . On the atom scale, gravity is irrelevant. But the $+$ and $-$ charges balance so accurately, that they cancel on the planetary scale, and gravity is much more dominant.

2.1.2 Cylindrical symmetry

Consider a cylindrically symmetric charge distribution, with $\rho(r)$ in cylindrical polar coordinates with total charge λ per unit length contained within an outer radius R . To have cylindrical symmetry again we have that

$$\mathbf{E} = E(r)\mathbf{e}_r.$$

Again this will satisfy $\nabla \times \mathbf{E} = 0$. To find $E(r)$, apply Gauss' law to a cylinder of radius r arbitrary length L .

If $r > R$ then

$$\begin{aligned}\int_S \mathbf{E} \cdot d\mathbf{S} &= E(r) \int_S \mathbf{e}_r \cdot d\mathbf{S} \\ &= E(r) \int_S dS \\ &= E(r) 2\pi r L = \frac{\lambda L}{\epsilon_0}.\end{aligned}$$

Thus we have that

$$\mathbf{E} = \frac{\lambda}{2\pi\epsilon_0 r} \mathbf{e}_r.$$

In the limit as $R \rightarrow 0$ we obtain the electric field of a line charge λ per unit length, corresponding to $\rho = \lambda\delta(x)\delta(y)$.

2.1.3 Planar symmetry

For a planar charge distribution, we have a charge density of $\rho(z)$ in Cartesian coordinates with total charge σ per unit area contained within a region $-d < z < d$ of thickness $2d$.

We will assume reflective symmetry, so $\rho(z)$ is even.

To have planar symmetry, we have $\mathbf{E} = E(z)\mathbf{e}_z$. Again we have that $\nabla \times \mathbf{E} = 0$. The reflectional symmetry implies that $E(-z) = -E(z)$.

To find $E(z)$ for $z > 0$ apply Gauss' law to a "Gaussian pillbox" of height $2z$ and arbitrary area A . If $z > d$ then

$$\begin{aligned}\int_S \mathbf{E} \cdot d\mathbf{S} &= E(z)A - E(-z)A \\ &= 2E(z)A \\ &= \frac{\sigma A}{\epsilon_0}\end{aligned}$$

Thus we have that

$$\mathbf{E} = \begin{cases} \frac{\sigma}{2\epsilon_0} \mathbf{e}_z & z < d \\ -\frac{\sigma}{2\epsilon_0} \mathbf{e}_z & z < -d \end{cases}.$$

In the limit as $d \rightarrow 0$ we obtain the electric field of a *surface charge* σ per unit area, corresponding to $\rho = \sigma\delta(z)$.

2.1.4 Surface charge and discontinuity

Let \mathbf{n} be a unit vector normal to the charged surface, pointing from region 1 to region 2. In our example we have that $\mathbf{n} = \mathbf{e}_z$. This discontinuity in \mathbf{E} is given by

$$[\mathbf{n} \cdot \mathbf{E}] = \frac{\sigma}{\varepsilon_0}$$

where σ is the surface charge density and

$$[X] = X_2 - X_1$$

denotes a discontinuity between regions 1 and 2.

The tangential components are continuous:

$$[\mathbf{n} \times \mathbf{E}] = 0.$$

And these two equations apply to any surface even if it's curved and non-uniform.

2.2 The electrostatic potential

For a general $\rho(\mathbf{x})$ we cannot determine $\mathbf{E}(\mathbf{x})$ using Gauss' law alone. We'll need to use the Maxwell equation $\nabla \times \mathbf{E} = 0$. This implies that \mathbf{E} is irrotational so it has an *electrostatic* potential $\Phi(\mathbf{x})$, such that

$$\mathbf{E} = -\nabla\Phi.$$

Definition. (Potential difference) The *potential difference* or voltage between two points \mathbf{x}_1 and \mathbf{x}_2 is

$$\Phi(\mathbf{x}_2) - \Phi(\mathbf{x}_1) = \int_{\mathbf{x}_1}^{\mathbf{x}_2} d\Phi = - \int_{\mathbf{x}_1}^{\mathbf{x}_2} \mathbf{E} \cdot d\mathbf{x}$$

and is path independent since $\nabla \times \mathbf{E} = 0$ is zero and the region is simply connected, so the field is conservative .

Definition. (Electric force) The *electric force* on a particle of charge q is

$$\mathbf{F} = q\mathbf{E} = -q\nabla\Phi.$$

Remark. This is a conservative force associated with the potential energy

$$U(\mathbf{x}) = q\Phi(\mathbf{x}).$$

Recall that the first Maxwell equation implies that Φ satisfies Poisson's equation, so

$$-\nabla^2\Phi = \frac{\rho}{\varepsilon_0}.$$

So we have the solution (from IB Methods) as (over all space with boundary conditions that $\Phi \rightarrow 0$ as $|\mathbf{x}| \rightarrow \infty$).

$$\Phi(\mathbf{x}) = \frac{1}{4\pi\varepsilon} \int \frac{\rho(\mathbf{x}')}{|\mathbf{x} - \mathbf{x}'|} d^3\mathbf{x}'.$$

This is the convolution of $\rho(\mathbf{x})$ with the potential of a unit point charge (which relates to our Green's function from IB Methods) $\frac{1}{4\pi\epsilon_0|\mathbf{x}|}$. Namely it is the solution to

$$-\nabla^2\Phi = \frac{\delta(\mathbf{x})}{\epsilon_0}$$

satisfying $\Phi \rightarrow 0$ as $|\mathbf{x}| \rightarrow \infty$. Note that Φ is unaffected if we add an arbitrary constant to Φ (this makes sense since Φ measures a potential difference between two points so increasing the charge uniformly doesn't change). We usually choose this such that $\Phi \rightarrow 0$ as $|\mathbf{x}| \rightarrow \infty$. If $\rho(\mathbf{x})$ does not decay sufficiently rapidly this may not be possible. For example if we have a line charge $E_r \propto \frac{1}{r}$, so we have that $\Phi \propto \log r$ which doesn't go to zero as $r \rightarrow \infty$.

2.2.1 Point charge

The potential due to a point charge q at the origin is

$$\Phi(\mathbf{x}) = \frac{q}{4\pi\epsilon_0|\mathbf{x}|} = \frac{q}{4\pi\epsilon_0 r}.$$

2.2.2 Electric dipole

Two equal and opposite charges at different positions. Without loss of generality consider charges $-q$ at $\mathbf{x} = 0$ and $+q$ at $\mathbf{x} = \mathbf{d}$. The potential due to the dipole is

$$\Phi(\mathbf{x}) = \frac{q}{4\pi\epsilon_0} \left(-\frac{1}{|\mathbf{x}|} + \frac{1}{|\mathbf{x} - \mathbf{d}|} \right)$$

Apply Taylor's theorem for a scalar field,

$$f(\mathbf{x} + \mathbf{h}) = f(\mathbf{x}) + (\mathbf{h} \cdot \nabla)f(\mathbf{x}) + \frac{1}{2}(\mathbf{h} \cdot \nabla)^2 f(\mathbf{x}) + O(|\mathbf{h}|^3).$$

So we get that

$$\Phi(\mathbf{x}) = \frac{q}{4\pi\epsilon_0} \left(-\frac{1}{r} + \frac{1}{r} - (\mathbf{d} \cdot \nabla)\frac{1}{r} + O(|\mathbf{d}|^2) \right) = \frac{q\mathbf{d} \cdot \mathbf{x}}{4\pi\epsilon_0|\mathbf{x}|^3} + O(|\mathbf{d}|^2).$$

In the limit as $|\mathbf{d}| \rightarrow 0$ with $q\mathbf{d}$ finite, we obtain a *point dipole* with *electric dipole moment*

$$\mathbf{p} = q\mathbf{d}.$$

which has potential

$$\Phi(\mathbf{x}) = \frac{\mathbf{p} \cdot \mathbf{x}}{4\pi\epsilon_0|\mathbf{x}|^3}$$

and electric field

$$\mathbf{E} = -\nabla\Phi = \frac{3(\mathbf{p} \cdot \mathbf{x})\mathbf{x} - |\mathbf{x}|^2\mathbf{p}}{4\pi\epsilon_0|\mathbf{x}|^5}.$$

In spherical polar coordinates aligned with $\mathbf{p} = p\mathbf{e}_z$. So

$$\Phi = \frac{p \cos \theta}{4\pi\epsilon_0 r^2}.$$

Then we get that

$$E_r = -\frac{\partial\Phi}{\partial r} = \frac{2p\cos(\theta)}{4\pi\epsilon_0 r^3}$$

and

$$E_\theta = -\frac{1}{r} \frac{\partial\Phi}{\partial\theta} = \frac{p\sin\theta}{4\pi\epsilon_0 r^3}.$$

From our alignment we have that $E_\phi = 0$.

Remark. Note that

- (i) Φ and \mathbf{E} are not spherically symmetric.
- (ii) They decrease more rapidly with r than a point charge since the dipole are nearly cancelling each other out.

A point dipole \mathbf{p} at the origin corresponds to

$$\rho(\mathbf{x}) = -\mathbf{p} \cdot \nabla \delta(\mathbf{x}),$$

So we can find the associated potential Φ as

$$\Phi(\mathbf{x}) = \mathbf{p} \cdot \nabla \left(\frac{1}{4\pi\epsilon_0 |\mathbf{x}|} \right).$$

2.2.3 Field lines and equipotentials

Electric field lines are the integral curves of \mathbf{E} being tangent to \mathbf{E} everywhere. Since we have that $\nabla \cdot \mathbf{E} = \frac{\rho}{\epsilon_0}$, field lines begin on positive charges and end on negative charges. In electrostatics, $\mathbf{E} = -\nabla\Phi$, so field lines are perpendicular to the equipotential surfaces of which Φ are constant.

2.2.4 Dipole in an external field

Consider a dipole \mathbf{p} in an external field $\mathbf{E}_{\text{external}} = -\nabla\Phi$ generated by distinct charges. With $-q$ at \mathbf{x} and $+q$ and $\mathbf{x} + \mathbf{d}$, the potential energy at the dipole due to the external field is

$$\begin{aligned} U &= -q\Phi(\mathbf{x}) + Q\Phi(\mathbf{x} + \mathbf{d}) \\ &= q(\mathbf{d} \cdot \nabla)\Phi(\mathbf{x}) + O(|\mathbf{d}|^2) \end{aligned}$$

In the limit at the point dipole,

$$U = \mathbf{p} \cdot \nabla\Phi = -\mathbf{p} \cdot \mathbf{E}_{\text{external}}$$

and is minimised when \mathbf{p} is aligned with $\mathbf{E}_{\text{external}}$.

2.2.5 Multipole expansion

For a general charge distribution $\rho(\mathbf{x})$ confined to a ball $\{V : |\mathbf{x}| < R\}$,

$$\Phi(\mathbf{x}) = \frac{1}{4\pi\epsilon_0} \int_V \frac{\rho(\mathbf{x}')}{|\mathbf{x} - \mathbf{x}'|} d^3\mathbf{x}'.$$

We'll look at the external potential at \mathbf{x} with $\mathbf{x} \notin V$. Expand

$$\frac{1}{|\mathbf{x} - \mathbf{x}'|} = \frac{1}{r} - (\mathbf{x}' \cdot \nabla) \frac{1}{r} + \frac{1}{2} (\mathbf{x}' \cdot \nabla)^2 \frac{1}{r} + O(|\mathbf{x}'|^3).$$

Which is

$$= \frac{1}{r} \left[1 + \frac{\mathbf{x}' \cdot \mathbf{x}}{r^2} + \frac{3(\mathbf{x}' \cdot \mathbf{x})^2 - |\mathbf{x}'|^2 |\mathbf{x}|^2}{2r^4} + O\left(\frac{R^3}{r^3}\right) \right]$$

This leads to the *multipole expansion* of the potential,

$$\Phi(\mathbf{x}) = \frac{1}{4\pi\epsilon_0} \left(\frac{Q}{r} + \frac{\mathbf{p} \cdot \mathbf{x}}{r^2} + \frac{1}{2} \frac{Q_{ij} x_i x_j}{r^5} + \dots \right).$$

The first three multipole moments:

- (i) The total charge, $Q = \int_V \rho(\mathbf{x}) d^3\mathbf{x}$.
- (ii) The electric dipole moment $\mathbf{p} = \int_V \mathbf{x} \rho(\mathbf{x}) d^3\mathbf{x}$.
- (iii) The electric quadrupole moment. This is a second order tensor which is traceless and symmetric,

$$Q_{ij} = \int_V (3x_i x_j - |\mathbf{x}|^2 \delta_{ij}) \rho(\mathbf{x}) d^3\mathbf{x}.$$

For $\gg R$, Φ and \mathbf{E} look increasingly like those of a point charge Q , unless $Q = 0$, in which case they look like those of a point dipole, unless $\mathbf{p} = 0$, etc.

2.3 Electrostatic energy

The work done against the electric force, $\mathbf{F} = q\mathbf{E}$, in bringing in a particle of charge q from infinity (where we assume that $\Phi = 0$ at infinity) to \mathbf{x} is

$$- \int_{\infty}^{\mathbf{x}} \mathbf{F} \cdot d\mathbf{x} = +q \int_{\infty}^{\mathbf{x}} \nabla \Phi \cdot d\mathbf{x} = q\Phi(\mathbf{x}).$$

Consider assembling a configuration of N point charges one by one. Particle i of charge q_i is brought from ∞ to \mathbf{x}_i while the previous particles remain fixed. For the first particle no work is involved, $W_1 = 0$. For the second particle

$$W_2 = q_2 \left(\frac{q_1}{4\pi\epsilon_0 |\mathbf{x}_2 - \mathbf{x}_1|} \right)$$

and for the third particle

$$W_3 = q_3 \left(\frac{q_1}{4\pi\epsilon_0 |\mathbf{x}_3 - \mathbf{x}_1|} + \frac{q_2}{4\pi\epsilon_0 |\mathbf{x}_3 - \mathbf{x}_2|} \right)$$

So the total work done is

$$U = \sum_{i=1}^N W_i = \sum_{i=2}^N \sum_{j=1}^{i-1} \frac{q_i q_j}{4\pi\epsilon_0 |\mathbf{x}_i - \mathbf{x}_j|}.$$

This can be rewritten as

$$U = \frac{1}{2} \sum_{i=1}^N \sum_{j \neq i} \frac{q_i q_j}{4\pi\epsilon_0 |\mathbf{x}_i - \mathbf{x}_j|}.$$

or

$$U = \frac{1}{2} \sum_{i=1}^N q_i \Phi(\mathbf{x}_i).$$

We can generalise to a continuous charge distribution $\rho(\mathbf{x})$ occupying a finite volume V .

$$U = \frac{1}{2} \int_V \rho(\mathbf{x}) \Phi(\mathbf{x}) d^3\mathbf{x}.$$

Using the first Maxwell equation we get that

$$\begin{aligned} U &= \frac{1}{2} \int_V (\epsilon_0 \nabla \cdot \mathbf{E}) \Phi dV \\ &= \frac{\epsilon_0}{2} \int_V (\nabla \cdot (\Phi \mathbf{E}) - \mathbf{E} \cdot \nabla \Phi) dV \\ &= \frac{\epsilon_0}{2} \int_S \Phi \mathbf{E} \cdot d\mathbf{S} + \int_V \frac{\epsilon_0 |\mathbf{E}|^2}{2} dV. \end{aligned}$$

Let $S = \partial V$ be a sphere of radius $R \rightarrow \infty$. Then $\Phi = O(R^{-1})$ and $\mathbf{E} = O(R^{-2})$ on S while the area of S is $O(R^2)$, so \int_S is $O(R^{-1})$ and $\rightarrow 0$ as $R \rightarrow \infty$. Then

$$U = \int \frac{\epsilon_0 |\mathbf{E}|^2}{2} dV$$

where the integral is taken over all space, not just the volume where the charges are contained.

Remark. This implies that energy is stored in the electric field, even in a vacuum.

Any of expression for U suggests that the self-energy of a point charge is infinite, hence for U to be useful, we discard all self-energies since it is unchanging and causes no force.

2.4 Conductors

In a *conductor* such as a metal, some charges can move freely. In electrostatics we require

$$\mathbf{E} = 0, \quad \Phi = \text{constant}$$

inside a conductor, hence $\rho = 0$. Otherwise free charges would move in a response to the electric force and a current would flow.

However a surface charge density σ can exist on the surface of a conductor, which is an equipotential.

Taking \mathbf{n} to point out of the conductor, the condition,

$$\mathbf{n} \cdot \mathbf{E} = \frac{\sigma}{\epsilon_0}$$

becomes

$$\mathbf{n} \cdot \mathbf{E} = \frac{\sigma}{\epsilon_0} \quad \text{immediately outside the conductor.}$$

The constant potential of a conductor can be set by connecting it to a battery or another conductor.

Definition. (Earthed/Grounded conductor) An *earthed* or *grounded* conductor is connected to the ground, usually taken as $\Phi = 0$.

To find $\Phi(\mathbf{x})$ and $\mathbf{E}(\mathbf{x})$ due to the charge distribution $\rho(\mathbf{x})$ in the presence of conductors with surface S_i and potentials Φ_i we solve Poisson's equation

$$-\nabla^2 \Phi = \frac{\rho}{\varepsilon_0}$$

with Dirichlet boundary conditions

$$\Phi = \Phi_i \quad \text{on } S_i.$$

The solution depends linearly on ρ and $\{\Phi_i\}$.

Let's see an example. Take a point charge q at position $(0, 0, h)$ in a half space ($z > 0$) bounded by an earthed conducting wall. Hence we have the boundary condition $\Phi = 0$ on $z = 0$. By the method of images, the solution in $z > 0$ is identical to that of a dipole, with image charge $-q$ placed at $(0, 0, -h)$. The wall coincides with an equipotential of the dipole, namely the line with $\Phi = 0$ which is the same as our boundary condition. The induced surface charge density on the wall can be worked out from

$$\frac{\sigma}{\varepsilon_0} = \mathbf{n} \cdot \mathbf{E} = E_z = -\frac{2qh}{4\pi\varepsilon_0(r^2 + h^2)^{3/2}}.$$

The total induced surface charge is

$$\begin{aligned} \int_0^\infty \sigma \, 2\pi r \, dr &= -qh \int_0^\infty \frac{r \, dr}{(r^2 + h^2)^{3/2}} \\ &= -q \end{aligned}$$

which is equal to the image charge.

Definition. (Capacitor) A simple *capacitor* consists of two separated conductors carrying charges $\pm Q$. If the potential difference between them is V , then the capacitance is defined by

$$C = \frac{Q}{V}$$

and depends only on the geometry, because Φ depends linearly on Q .

For example, consider two infinite parallel plates separated by some distance d . Let the plate surfaces at $z = 0, z = d$ have surface charge densities $\pm\sigma$. Then $\mathbf{E} = E\mathbf{e}_z$ with $E = \sigma/\varepsilon_0 = \text{const.}$ for $0 < z < d$, with $\mathbf{E} = 0$ elsewhere. So

$$\Phi = -Ez + \text{const}, \quad \text{and} \quad V = Ed.$$

The same solution holds approximately for parallel plates of $A \gg d^2$ if end effects are neglected. So

$$C = \frac{Q}{V} \approx \frac{\sigma A}{Ed} \approx \frac{\varepsilon_0 A}{d}.$$

The electrostatic energy stored in a capacitor is

$$\begin{aligned} U &= \int \frac{\varepsilon_0 |\mathbf{E}|^2}{2} dV \\ &\approx \frac{\varepsilon_0 E^2}{2} Ad \\ &\approx \frac{1}{2} CV^2. \end{aligned}$$

In general,

$$U = \frac{1}{2} CV^2 = \frac{Q^2}{2C}.$$

The work done in moving an element of charge δQ from one plate to another is $\delta W = V \delta Q$, so the total work done is

$$\int_0^Q \frac{Q'}{C} dQ' = \frac{Q^2}{2C}$$

for any geometry of the capacitor. Or we can use

$$\begin{aligned} U &= \frac{1}{2} \int \rho \Phi dV \\ &= \frac{1}{2} Q \Phi_+ - \frac{1}{2} Q \Phi_- \\ &= \frac{1}{2} QV. \end{aligned}$$

3 Magnetostatics

Magnetostatics is the study of the magnetic field generated by a stationary current distribution.

$$\begin{aligned} \nabla \times \mathbf{B} &= \mu_0 \mathbf{J} \\ \nabla \cdot \mathbf{B} &= 0, \end{aligned}$$

and the first equation here implies that $\nabla \cdot \mathbf{J} = 0$, the time-independent equation of charge conservation.

3.1 Ampere's law

Consider a closed curve C that is the boundary of an open surface S . Integrate the fourth Maxwell equation over S and apply Stokes' theorem to obtain Ampere's law.

$$\int_C \mathbf{B} \cdot d\mathbf{x} = \mu_0 I$$

where

$$I = \int_S \mathbf{J} \cdot d\mathbf{S}.$$

I is the total current through the surface S . Since $\nabla \cdot \mathbf{J} = 0$, the same current I flows through *any* open surface S such that $\partial S = C$. Ampere's law is the integral version of the fourth Maxwell equation and is valid provided that $\frac{\partial \mathbf{E}}{\partial t} = 0$. Ampere's law is saying that

circulation of magnetic field around loop \propto total current through loop.

In special situations we can use Ampere's law together with symmetry to deduce \mathbf{B} from \mathbf{J} .

A cylindrically symmetric situation could involve

- An axial current distribution

$$J_z(r)\mathbf{e}_z.$$

- An azimuthal current distribution

$$J_\phi(r)\mathbf{e}_\phi$$

or a combination. (In fact $\nabla \cdot \mathbf{J} = 0$ excludes a radial current.)

The same applies to \mathbf{B} . The curl in the fourth Maxwell equation implies that B_ϕ is linearly related to J_z and B_z is linearly related to J_ϕ . Let's consider a cylindrical wire of radius R which carries a total current I parallel to its axis. To find $B_\phi(r)$ generated by $J_z(r)$, apply Ampere's law to a circle C of radius R .

If $r > R$ then

$$\begin{aligned}\int_C \mathbf{B} \cdot d\mathbf{x} &= B_\phi(r) \int_C \mathbf{e}_\phi \cdot d\mathbf{x} \\ &= B_\phi(r) \int_C d\ell \\ &= B_\phi(r) 2\pi r = \mu_0 I.\end{aligned}$$

Thus, outside the wire,

$$\mathbf{B} = \frac{\mu_0 I}{2\pi r} \mathbf{e}_\phi.$$

Now for a solenoid. A thin wire is coiled around a cylindrical tube of radius R . An *ideal solenoid* is infinitely long and tightly wound, having cylindrical symmetry and purely azimuthal current. The wire carries current I and has N turns per unit length of the tube.

To find $B_z(r)$ generated by $J_\phi(r)$, apply Ampere's law to a rectangular loop C . Taking $a < b < R$ or $R < a < b$ gives that

$$L(B_z(a) - B_z(b)) = 0.$$

and taking $a < R < b$ gives

$$L(B_z(a) - B_z(b)) = \mu_0 N L I.$$

Assuming that $B_z(r) \rightarrow 0$ as $r \rightarrow \infty$, we deduce that

$$B_z(r) = \begin{cases} \mu_0 N I & r < R \\ 0 & r > R \end{cases}$$

The ideal solenoid is an example of a surface current, here of the form

$$J_\phi(r) = K_\phi \delta(r - R)$$

with $K_\phi = NI$.

Generally, a surface current density, \mathbf{K} , produces a discontinuity in the tangential magnetic field:

$$[\mathbf{n} \times \mathbf{B}] = \mu_0 \mathbf{K}.$$

It follows from Ampere's law that the normal component is continuous, i.e.

$$[\mathbf{n} \cdot \mathbf{B}] = 0.$$

3.2 The magnetic vector potential

The second Maxwell equation implies that \mathbf{B} can be written in terms of a *magnetic vector potential*.

Definition. (Magnetic vector potential) For a magnetic field \mathbf{B} , the *magnetic vector potential* is the vector $\mathbf{A}(\mathbf{x})$ such that

$$\mathbf{B} = \nabla \times \mathbf{A}.$$

Remark. \mathbf{A} is *not* unique. If we make a *gauge transformation*, replacing \mathbf{A} with

$$\tilde{\mathbf{A}} = \mathbf{A} + \nabla\chi,$$

where $\chi(\mathbf{x})$ is an arbitrary scalar field, then \mathbf{B} is unchanged since $\mathbf{B} = \nabla \times \mathbf{A} = \nabla \times \tilde{\mathbf{A}}$.

A convenient gauge for many calculations is the *Coulomb gauge* in which $\nabla \cdot \mathbf{A} = 0$. We can assume this condition *wlog*. If $\nabla \cdot \mathbf{A} \neq 0$, then we can make a gauge transformation such that $\nabla \cdot \tilde{\mathbf{A}} = 0$ by choosing χ to be the solution of Poisson's equation

$$-\nabla^2 \chi = \nabla \cdot \mathbf{A}.$$

In terms of \mathbf{A} the fourth Maxwell equation becomes

$$\nabla \times (\nabla \times \mathbf{A}) = \mu_0 \mathbf{J}.$$

Using the identity,

$$\nabla \times (\nabla \times \mathbf{A}) = \nabla(\nabla \cdot \mathbf{A}) - \nabla^2 \mathbf{A}$$

and assuming Coulomb gauge with $\nabla \cdot \mathbf{A} = 0$, we obtain Poisson's equation in vector form,

$$-\nabla^2 \mathbf{A} = \mu_0 \mathbf{J}$$

3.3 The Biot-Savart law

The solution of Poisson's equation is

$$\mathbf{A}(\mathbf{x}) = \frac{\mu_0}{4\pi} \int \frac{\mathbf{J}(\mathbf{x}')}{|\mathbf{x} - \mathbf{x}'|} d^3\mathbf{x}'.$$

We should check that the solution satisfies the assumed Coulomb gauge condition

$$\begin{aligned} \nabla \cdot \mathbf{A}(\mathbf{x}) &= \frac{\mu_0}{4\pi} \int_V \nabla \cdot \frac{\mathbf{J}(\mathbf{x}')}{|\mathbf{x} - \mathbf{x}'|} d^3\mathbf{x}' \\ &= \frac{\mu_0}{4\pi} \int_V \mathbf{J}(\mathbf{x}') \cdot \nabla \left(\frac{1}{|\mathbf{x} - \mathbf{x}'|} \right) d^3\mathbf{x}' \\ &= -\frac{\mu_0}{4\pi} \int_V \mathbf{J}(\mathbf{x}') \cdot \nabla' \left(\frac{1}{|\mathbf{x} - \mathbf{x}'|} \right) d^3\mathbf{x}' \\ &= -\frac{\mu_0}{4\pi} \int_V \nabla' \cdot \left(\frac{\mathbf{J}(\mathbf{x}')}{|\mathbf{x} - \mathbf{x}'|} \right) d^3\mathbf{x}' \\ &= -\frac{\mu_0}{4\pi} \int_{\partial V} \frac{\mathbf{J}(\mathbf{x}') \cdot d\mathbf{S}'}{|\mathbf{x} - \mathbf{x}'|}. \end{aligned}$$

Thus $\nabla \cdot \mathbf{A} = 0$, as assumed. if the current is contained in some finite volume and we take V to be at least as large, or if \mathbf{J} decays sufficiently as $|\mathbf{x}| \rightarrow \infty$.

To find the magnetic field, derive $\mathbf{B} = \nabla \times \mathbf{A}$. This gives the following law.

Theorem. (Biot-Savart law)

$$\mathbf{B}(\mathbf{x}) = \frac{\mu_0}{4\pi} \int \frac{\mathbf{J}(\mathbf{x}') \times (\mathbf{x} - \mathbf{x}')}{|\mathbf{x} - \mathbf{x}'|^3} d^3\mathbf{x}'$$

A special case is when the current is restricted to a thin wire in the form of a curve C . Then the current element $\mathbf{J}d^3\mathbf{x}$ can be replaced by $I d\mathbf{x}$. By charge conservation, we get that I is constant along the wire, hence we can take it outside of the integral, so the Biot-Savart law for a current carrying wire is

$$\mathbf{B}(\mathbf{x}) = \frac{\mu_0 I}{4\pi} \int_C \frac{d\mathbf{x}' \times (\mathbf{x} - \mathbf{x}')}{|\mathbf{x} - \mathbf{x}'|^3}.$$

Alternatively, the thin wire current density can be represented as

$$\mathbf{J}(\mathbf{x}) = I \int_C \delta(\mathbf{x} - \mathbf{x}') d^3\mathbf{x}',$$

which gives the same equation if substituted in. Charge conservation takes the form

$$\begin{aligned} \nabla \cdot \mathbf{J}(\mathbf{x}) &= I \int_C \nabla \delta(\mathbf{x} - \mathbf{x}') \cdot d\mathbf{x}' \\ &= -I \int_C -C \nabla \delta(\mathbf{x} - \mathbf{x}') \cdot d\mathbf{x}' \\ &= -I[\delta(\mathbf{x} - \mathbf{x}_2) - \delta(\mathbf{x} - \mathbf{x}_1)] \end{aligned}$$

where C runs from \mathbf{x}_1 to \mathbf{x}_2 . If C is closed then $\mathbf{x}_2 = \mathbf{x}_1$ and $\nabla \cdot \mathbf{J} = 0$ as expected. If C is infinite then $\nabla \cdot \mathbf{J} = 0$ for any finite \mathbf{x} . Let's check the thin-wire version of Biot-Savart's law gives the same result as Ampere's law for a long straight thin wire along the z axis of cylindrical polars. We have that $\mathbf{x} = r\mathbf{e}_r$ (taking $z = 0$ wlog) and $\mathbf{x}' = z'\mathbf{e}_z$. So $\mathbf{x} - \mathbf{x}' = r\mathbf{e}_r - z'\mathbf{e}_z$. We also have that $d\mathbf{x}' = z'\mathbf{e}_z$. This gives that

$$\mathbf{B}(\mathbf{x}) = \frac{\mu_0 I}{4\pi} \mathbf{e}_\phi \int_{-\infty}^{\infty} \frac{r dz'}{(r^2 + z'^2)^{3/2}} = \frac{\mu_0 I}{2\pi r} \mathbf{e}_\phi$$

as expected.

3.4 Magnetic dipoles

For a general current distribution $\mathbf{J}(\mathbf{x})$ confined to a ball $\{V : |\mathbf{x}| < R\}$,

$$\mathbf{A}(\mathbf{x}) = \frac{\mu_0}{4\pi} \int_V \frac{\mathbf{J}(\mathbf{x}')}{|\mathbf{x} - \mathbf{x}'|} d^3\mathbf{x}'.$$

The external field for $|\mathbf{x}| = r > R$ can be evaluated by expanding

$$\frac{1}{|\mathbf{x} - \mathbf{x}'|} = \frac{1}{r} \left(1 + \frac{\mathbf{x}' \cdot \mathbf{x}}{r^2} + O\left(\frac{R^2}{r^2}\right) \right),$$

leading to a multipole expansion as before. We need to calculate the moments of the current distribution.

Since $\mathbf{J} = 0$, on ∂V and $\nabla \cdot \mathbf{J} = 0$ the divergence theorem implies that

$$\begin{aligned} 0 &= \int_{\partial V} x_i J_j dS_j = \int_V \partial_j (x_i J_j) d^3 \mathbf{x} \\ &= \int_V (\delta_{ij} J_j + x_i \partial_j J_j) d^3 \mathbf{x} \\ &= \int_V J_i d^3 \mathbf{x} \end{aligned}$$

so the zeroth moment vanishes. Similarly,

$$\begin{aligned} 0 &= \int_{\partial V} x_i x_j J_k dS_k = \int_V \partial_k (x_i x_j J_k) d^3 \mathbf{x} \\ &= \int_V (\delta_{ik} x_j J_k + x_i \delta_{jk} J_k + x_i x_j \partial_k J_k) d^3 \mathbf{x} \\ &= \int_V x_j J_i d^3 \mathbf{x} + \int_V x_i J_j d^3 \mathbf{x}. \end{aligned}$$

So the first moment is an antisymmetric tensor. Like the electric dipole moment we define:

Definition. (Magnetic dipole moment)

$$\mathbf{m} = \frac{1}{2} \int_V \mathbf{x} \times \mathbf{J} d^3 \mathbf{x}$$

or in component form

$$m_i = \frac{1}{2} \varepsilon_{ijk} \int_V x_j J_k d^3 \mathbf{x}.$$

This is a vector related to the antisymmetric matrix by

$$\int_V x_i J_j d^3 \mathbf{x} = \varepsilon_{ijk} m_k.$$

Returning to the multipole expansion for \mathbf{A} , we have

$$A_i(\mathbf{x}) = \frac{\mu_0}{4\pi|\mathbf{x}|} \left(\int_V J_i(\mathbf{x}') d^3 \mathbf{x}' + \dots \right)$$

we know the first term is zero so

$$\mathbf{A}(\mathbf{x}) \approx \mathbf{A}_{\text{dipole}}(\mathbf{x}) = \frac{\mu_0}{4\pi} \frac{\mathbf{m} \times \mathbf{x}}{|\mathbf{x}|^3}$$

which is the vector potential due to a point dipole \mathbf{m} at the origin. The corresponding magnetic field is

$$\mathbf{B}_{\text{dipole}} = \nabla \times \mathbf{A}_{\text{dipole}} = \frac{\mu_0}{4\pi} \left(\frac{3(\mathbf{m} \cdot \mathbf{x})\mathbf{x} - |\mathbf{x}|^2 \mathbf{m}}{|\mathbf{x}|^5} \right).$$

A point dipole \mathbf{m} at the origin corresponds to the current density and vector potential

$$\mathbf{J} = \nabla \times (\mathbf{m}\delta(\mathbf{x})), \quad \mathbf{A} = \nabla \times \left(\frac{\mu_0 \mathbf{m}}{4\pi|\mathbf{x}|} \right).$$

The magnetic dipole moment of a thin wire carrying current I around a closed curve C is

$$\mathbf{m} = \frac{I}{2} \int_C \mathbf{x} \times d\mathbf{x}.$$

To evaluate this, let \mathbf{a} be any constant vector. Then by Stokes' theorem,

$$\begin{aligned} \mathbf{a} \cdot \mathbf{m} &= \frac{I}{2} \int_C \mathbf{a} \cdot (\mathbf{x} \times d\mathbf{x}) \\ &= \frac{I}{2} \int_C (\mathbf{a} \times \mathbf{x}) \cdot d\mathbf{x} \\ &= \frac{I}{2} \int_S (\nabla \times (\mathbf{a} \times \mathbf{x})) \cdot d\mathbf{S} \\ &= I \int_S \mathbf{a} \cdot d\mathbf{S} \end{aligned}$$

where S is an open surface with boundary C and we use

$$\nabla \times (\mathbf{a} \times \mathbf{x}) = \mathbf{x} \cdot \nabla \mathbf{a} - \mathbf{a} \cdot \nabla \mathbf{x} + (\nabla \cdot \mathbf{x})\mathbf{a} - (\nabla \cdot \mathbf{a})\mathbf{x} = 2\mathbf{a}.$$

Since \mathbf{a} is arbitrary we conclude that

$$\mathbf{m} = I\mathbf{S}$$

where

$$\mathbf{S} = \int_S d\mathbf{S}$$

is the vector area of the surface S which depends only on C .

The simplest example is a circular loop, for example $x^2 + y^2 = a^2$ for which $\mathbf{m} = I\pi a^2 \mathbf{e}_z$. On the z -axis the dipole approximation gives that

$$B_z = \frac{\mu_0}{4\pi} \left(\frac{3m_z z^2 - z^2 m_z}{|z|^5} \right) = \frac{\mu_0 I a^2}{2|z|^3}$$

while the exact solution is

$$B_z = \frac{\mu_0 I a^2}{2(z^2 + a^2)^{3/2}}.$$

Magnetic field lines are the integral curves of \mathbf{B} . Since $\nabla \cdot \mathbf{B} = 0$, they are continuous.

Definition. (Permanent magnets) A bar magnet has north and south poles and a dipole moment. This comes from the superposition of aligned dipoles on the atomic scale. Atoms contain electrons which are spinning charges particles with a magnetic dipole moment.

A classical model of a particle is a spinning charged sphere which is a current loop with a magnetic dipole moment proportional to its charge and spin. As far as we know, there is no magnetic charges (monopoles).

The liquid iron outer core of the Earth is a conducting fluid in convective motion and supports electric currents that generate a magnetic field. At the Earth's surface this resembles a dipole field.

3.5 Magnetic forces

Recall that the Lorentz force on a particle of charge q_i at position $\mathbf{x}_i(t)$ is

$$\mathbf{F} = q_i(\mathbf{E} + \dot{\mathbf{x}}_i \times \mathbf{B})$$

where both \mathbf{E} and \mathbf{B} are evaluated at $\mathbf{x}_i(t)$.

In the limit of continuous charge and current densities, the Lorentz force per unit volume is then

$$\rho\mathbf{E} + \mathbf{J} \times \mathbf{B}.$$

We can recover the discrete version of the Lorentz force by substituting

$$\begin{aligned}\rho &= \sum_i q_i \delta(\mathbf{x} - \mathbf{x}_i(t)) \\ \mathbf{J} &= \sum_i q_i \dot{\mathbf{x}}_i(t) \delta(\mathbf{x} - \mathbf{x}_i(t)).\end{aligned}$$

Let's look at the force between thin wires. Consider two or more thin wires with current I_i along a curve C_i . The total magnetic field is $\mathbf{B} = \sum_i \mathbf{B}_i$, where

$$\mathbf{B}_i(\mathbf{x}) = \frac{\mu_0 I_i}{4\pi} \int_{C_i} \frac{d\mathbf{x}_i \times (\mathbf{x} - \mathbf{x}_i)}{|\mathbf{x} - \mathbf{x}_i|^3}$$

is the magnetic field due to wire i . The current density is $\mathbf{J} = \sum_i \mathbf{J}_i$, where

$$\mathbf{J}_i(\mathbf{x}) = I_i \int_{C_i} \delta(\mathbf{x} - \mathbf{x}_i) d\mathbf{x}_i.$$

The total magnetic field acting on a volume V is

$$\mathbf{F} = \int_V \mathbf{J} \times \mathbf{B} dV.$$

The force acting on wire i is

$$\begin{aligned}\mathbf{F}_i &= \int \mathbf{J}_i(\mathbf{x}) \times \mathbf{B}(\mathbf{x}) d^3\mathbf{x} \\ &= I_i \int_{C_i} d\mathbf{x}_i \times \mathbf{B}(\mathbf{x}_i).\end{aligned}$$

Since $\mathbf{B} = \sum_i \mathbf{B}_i$, we have that

$$\mathbf{F}_i = \sum_j \mathbf{F}_{ij}$$

where

$$\mathbf{F}_{ij} = I_i \int_{C_i} d\mathbf{x}_i \times \mathbf{B}_j(\mathbf{x}_i)$$

is the force on wire i due to wire j . Using the Biot-Savart law,

$$\mathbf{F}_{ij} = \frac{\mu_0 I_i I_j}{4\pi} \int_{C_i} \int_{C_j} d\mathbf{x}_i \times \left(\frac{d\mathbf{x}_j \times (\mathbf{x}_i - \mathbf{x}_j)}{|\mathbf{x}_i - \mathbf{x}_j|^3} \right).$$

This can be rewritten (Example Sheet 3) in a manifestly antisymmetric way that shows that

$$\mathbf{F}_{ji} = -\mathbf{F}_{ij}$$

as expected from Newton's third law. The self force \mathbf{F}_{ii} vanishes, although the thin-wire integral is singular when $i = j$ and it is better to treat the case of thick wires.

Let's consider two infinitely long parallel thin wires separated by a distance r . Use cylindrical polars centred on the second wire. We have that $\mathbf{B}_2 = \frac{\mu_0 I_2}{2\pi r} \mathbf{e}_\phi$.

$$\mathbf{F}_{12} = I_1 \int_{-\infty}^{\infty} dz \mathbf{e}_z \times \mathbf{B}_2.$$

The total force is infinite. The force per unit length is

$$I_1 \mathbf{e}_z \times \mathbf{B}_2 = -\frac{\mu_0 I_1 I_2}{2\pi r} \mathbf{e}_r.$$

This is directed towards wire 2 if $I_1 I_2 > 0$ so the force is attractive if the currents are aligned and repulsive otherwise.

Let's consider the force and torque on a magnetic dipole. Consider a localised current distribution confined to a ball $\{V : |\mathbf{x}| < R\}$. Place this in an external magnetic field $\mathbf{B}(\mathbf{x})$ that varies slowly over the length scale R . The magnetic torque about the origin on the current loop is

$$\begin{aligned} \boldsymbol{\tau} &= \int_V \mathbf{x} \times (\mathbf{J}(\mathbf{x}) \times \mathbf{B}(\mathbf{x})) d^3\mathbf{x} \\ &= \int_V ((\mathbf{x} \cdot \mathbf{B}(\mathbf{x}))\mathbf{J}(\mathbf{x}) - (\mathbf{x} \cdot \mathbf{J}(\mathbf{x}))\mathbf{B}(\mathbf{x})) d^3\mathbf{x}. \end{aligned}$$

within V , $\mathbf{B}(\mathbf{x})$ can be expanded as a Taylor series,

$$\mathbf{B}_i(\mathbf{x}) = \mathbf{B}_i(0) + x_j \partial_j \mathbf{B}_i(0) + \dots$$

retaining only the zero order term (uniform field), we have that

$$\tau_i \approx B_j(0) \int_V x_j J_i d^3\mathbf{x} - B_i(0) \int_V x_j J_j d^3\mathbf{x}.$$

Recall the first moments of the current distribution:

$$\int_V x_i J_j d^3\mathbf{x} = \varepsilon_{ijk} m_k.$$

Thus

$$\tau_i \approx B_j(0) \varepsilon_{jik} m_k.$$

In general,

$$\boldsymbol{\tau} \approx \mathbf{m} \times \mathbf{B},$$

where \mathbf{B} is evaluated at the dipole's location and $\boldsymbol{\tau}$ is measured about this point.

For the force, we need to go to the first order of the Taylor expansion of \mathbf{B} .

$$\mathbf{F} = \int_V \mathbf{J}(\mathbf{x}) \times \mathbf{B}(\mathbf{x}) d^3\mathbf{x}$$

and in components

$$\begin{aligned}
F_i &\approx \int_V \varepsilon_{ijk} J_j(\mathbf{x}) (B_k(0) + x_\ell \partial_\ell B_k(0)) d^3\mathbf{x} \\
&= 0 + \varepsilon_{ijk} \partial_\ell B_k(0) \varepsilon_{\ell j n} m_n \\
&= \partial_i B_k(0) m_k - \partial_k B_k(0) m_i \\
&= \partial_i (m_k B_k)(0) \quad \text{since } \nabla \cdot \mathbf{B} = 0
\end{aligned}$$

In general, $\mathbf{F} \approx \nabla(\mathbf{m} \cdot \mathbf{B})$. Which can be written as $\mathbf{F} = -\nabla U$, where

$$U = -\mathbf{m} \cdot \mathbf{B}$$

is the potential energy of a magnetic dipole in an external field. As in the electric case this is minimised when \mathbf{m} is aligned with \mathbf{B} .

4 Electrodynamics

4.1 Faraday's law of induction

We're interested in the third Maxwell equation,

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$

which implies that a time-dependent magnetic field, must be accompanied by an electric field. This can induce a current to flow in a conductor. This is called *electromagnetic induction*.

First let's look at a static circuit. Consider a closed curve C that is the boundary of a time-independent open surface S . Integrate the third Maxwell equation over S and use Stokes' theorem.

$$\begin{aligned}
\int_C \mathbf{E} \cdot d\mathbf{x} &= - \int_S \frac{\partial \mathbf{B}}{\partial t} \cdot d\mathbf{S} \\
&= \frac{d}{dt} \int_S \mathbf{B} \cdot d\mathbf{S}
\end{aligned}$$

This is Faraday's law of induction for a static circuit.

$$\mathcal{E} = -\frac{d\mathcal{F}}{dt}.$$

where

$$\mathcal{E} = \int_C \mathbf{E} \cdot d\mathbf{x}$$

is the *electromotive force* (emf) around C and

$$\mathcal{F} = \int_S \mathbf{B} \cdot d\mathbf{S}$$

is the *magnetic flux* through S .

Since $\nabla \cdot \mathbf{B} = 0$, the flux \mathcal{F} is the same through any surface S such that $\partial S = C$, so it can be regarded as the magnetic flux through C .

Using $\mathbf{B} = \nabla \times \mathbf{A}$ and Stokes' theorem we can write $\mathcal{F} = \int_C \mathbf{A} \cdot d\mathbf{x}$, which is invariant under a gauge transformation. The emf is not actually a force. It is the line integral of the Lorentz force on a particle of unit charge confined to C .

$$\mathcal{E} = \frac{1}{q} \int_C \mathbf{F} \cdot d\mathbf{x} = \int_C (\mathbf{E} + \dot{\mathbf{x}} \times \mathbf{B}) \cdot d\mathbf{x} = \int_C \mathbf{E} \cdot d\mathbf{x}$$

since $\dot{\mathbf{x}}$ is tangent to C for a particle confined to a time-independent curve C . We will see later that if C coincides with a thin wire of resistance R , then the current induced in the wire is $I = \mathcal{E}/R$.

There are several ways in which the magnetic flux through C could change with time.

- A magnet is moved near C ;
- A current-carrying circuit is moved near C ;
- The current in a nearby circuit is changed.

All these will induce an emf around C and cause a current to flow.

Now let's look at Faraday's law for a moving circuit. Let $C(t)$ be time-dependent closed curve that is the boundary of an open surface $S(t)$. How does the magnetic flux through S ,

$$\mathcal{F} = \int_S \mathbf{B} \cdot d\mathbf{S}$$

change in time? We have

$$\begin{aligned} \mathcal{F}(t + \delta t) - \mathcal{F}(t) &= \int_{S(t+\delta t)} \mathbf{B}(\mathbf{x}, t + \delta t) \cdot d\mathbf{S} - \int_{S(t)} \mathbf{B}(\mathbf{x}, t) \cdot d\mathbf{S} \\ &= \int_{S(t+\delta t)} \left(\mathbf{B}(\mathbf{x}, t) + \frac{\partial \mathbf{B}}{\partial t} \delta t + O(\delta t^2) \right) \cdot d\mathbf{S} - \int_{S(t)} \mathbf{B}(\mathbf{x}, t) \cdot d\mathbf{S} \\ &= \int_{S(t+\delta t) - S(t)} \mathbf{B}(\mathbf{x}, t) \cdot d\mathbf{S} + \int_{S(t)} \frac{\partial \mathbf{B}}{\partial t} \cdot d\mathbf{S} \delta t + O(\delta t^2) \end{aligned}$$

Let δV be the volume swept out by $S(t)$ in the time interval δt . Its boundary is the closed surface $S(t + \delta t) - S(t) + \Sigma$, where Σ is the surface swept out by $C(t)$ in the time δt .

Looking at a line element $d\mathbf{x}$ on $C(t)$ which sweeps out a surface of height $\mathbf{v} \delta t$ with normal $d\mathbf{S}$. By the second Maxwell equation and the divergence theorem,

$$\begin{aligned} 0 &= \int_{\delta V} (\nabla \cdot \mathbf{B}) dV \\ &= \int_{S(t+\delta t) - S(t)} \mathbf{B} \cdot d\mathbf{S} + \int_{\Sigma} \mathbf{B} \cdot d\mathbf{S}. \end{aligned}$$

To evaluate the last term, parametrise C as $\mathbf{x} = \mathbf{x}(\lambda, t)$, where λ is a parameter around C . An element of C is

$$d\mathbf{x} = \frac{\partial \mathbf{x}}{\partial \lambda} d\lambda$$

and has velocity

$$\mathbf{v} = \frac{\partial \mathbf{x}}{\partial t}.$$

In time δt it sweeps out the vector area element

$$d\mathbf{S} = d\mathbf{x} \times (\mathbf{v}\delta t)$$

which points out of δV as we require. Thus

$$\int_{\sigma} \mathbf{B} \cdot d\mathbf{S} = \int_C \mathbf{B} \cdot (d\mathbf{x} \times \mathbf{v}) \delta t + O(\delta t^2).$$

We then have

$$\mathcal{F}(t + \delta t) - \mathcal{F}(t) = - \int_C (\mathbf{v} \times \mathbf{B}) \cdot d\mathbf{x} \delta t + \int_S \frac{\partial \mathbf{B}}{\partial t} \cdot d\mathbf{S} \delta t + O(\delta t^2)$$

from which

$$\begin{aligned} \frac{d\mathcal{F}}{dt} &= - \int_C (\mathbf{v} \times \mathbf{B}) \cdot d\mathbf{x} + \int_S \frac{\partial \mathbf{B}}{\partial t} \cdot d\mathbf{S} \\ &= - \int_C (\mathbf{v} \times \mathbf{B}) \cdot d\mathbf{x} - \int_S (\nabla \times \mathbf{E}) \cdot d\mathbf{S} \\ &= - \int_C (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot d\mathbf{x}. \end{aligned}$$

We recover Faraday's law,

$$\mathcal{E} = - \frac{d\mathcal{F}}{dt}$$

with the redefined emf

$$\mathcal{E} = \int_C (\mathbf{E} + \mathbf{v}' \times \mathbf{B}) \cdot d\mathbf{x}.$$

This is again the line integral around C of the Lorentz force on a particle of unit charge confined to C . (For which the perpendicular components of $\dot{\mathbf{x}}$ must agree with those of the curve velocity \mathbf{v} .)

Law. (Lenz's law) The direction of the induced current is always such as to produce a magnetic field that opposes the change in flux that causes the emf.

Example. A circular wire in the $x - y$ -plane. If B_z inside the loop increases in time, then

$$\mathcal{E} = - \frac{d\mathcal{F}}{dt} < 0.$$

This induces a clockwise current ($I < 0$) that generates a magnetic field $B_z < 0$ inside the loop. Hence the minus sign in Faraday's law. This avoids an unstable situation in which the flux grows indefinitely.

Definition. (Inductance) If a current I around a circuit C generates a magnetic field

with flux \mathcal{F} then the *inductance* of the circuit is defined by

$$L = \frac{\mathcal{F}}{I}$$

and depends only on the geometry.

Example. An ideal solenoid with cross-sectional area A and N turns per unit length. The uniform field $B = \mu_0 NI$ inside the solenoid produces a flux BA per turn, so the inductance per unit length of the solenoid is $\mu_0 N^2 A$.

Exercise. Show that the magnetic flux through a thin wire C_i due to a current I_j around another thin wire C_j is $\mathcal{F}_{ij} = L_{ij} I_j$ where the *mutual inductance* is

$$L_{ij} = \frac{\mu_0}{4\pi} \int_{C_i} \int_{C_j} \frac{d\mathbf{x}_i \cdot \mathbf{x}_j}{|\mathbf{x}_i - \mathbf{x}_j|} = L_{ji}.$$

When the current I around a circuit C is varies, an emf

$$\mathcal{E} = -\frac{d\mathcal{F}}{dt} = -L \frac{dI}{dt}$$

is induced. In a small time interval δt , a charge $\delta Q = I \delta t$ flows around C and the work done on it by the Lorentz force is

$$\delta W = \mathcal{E} \delta Q = -LI \frac{dI}{dt} \delta t.$$

So the rate at which work is done by the current on the EM field is

$$-\frac{dW}{dt} = LI \frac{dI}{dt} = \frac{d}{dt} \left(\frac{1}{2} LI^2 \right).$$

Consider reaching a magnetostatic state by building up the current from 0 to I The energy stored is

$$\begin{aligned} U &= \frac{1}{2} LI^2 \\ &= \frac{1}{2} I \mathcal{F} \\ &= \frac{1}{2} I \int_C \mathbf{A} \cdot d\mathbf{x} \\ &= \frac{1}{2} \int_V \mathbf{J} \cdot \mathbf{A} dV. \end{aligned}$$

This is analogous to $U = \frac{1}{2} \int \rho \Phi dV$ in electrostatics. Now using the third Maxwell equation we have

$$U = \frac{1}{2\mu_0} \int (\nabla \times \mathbf{B}) \cdot \mathbf{A} dV$$

and $(\nabla \times \mathbf{B}) \cdot \mathbf{A} = \nabla \cdot (\mathbf{B} \times \mathbf{A}) + \mathbf{B} \cdot (\nabla \times \mathbf{A})$. If we take the integral over all space then the first term gives zero by the divergence theorem, since $|\mathbf{B}| = O\left(\frac{1}{|\mathbf{x}|^3}\right)$ and $|\mathbf{A}| = O\left(\frac{1}{|\mathbf{x}|^2}\right)$ as $|\mathbf{x}| \rightarrow \infty$ for a finite current distribution, leaving

$$U = \int \frac{|\mathbf{B}|^2}{2\mu_0} dV$$

as the energy stored in the magnetic field.

4.2 Ohm's law

In a stationary conductor

$$\mathbf{J} = \sigma \mathbf{E}$$

where σ is the electrical conductivity. This is not a fundamental physical law but a constitutive relation, a macroscopic property of a material. The inverse relation

$$\mathbf{E} = \sigma^{-1} \mathbf{J}$$

where σ^{-1} is the *resistivity* (usually denoted by ρ which are both in conflict with the notation for charge densities). In some materials σ is not a scalar, but instead is a tensor, hence we denote the resistivity by σ^{-1} to show that sometimes we need to invert a tensor if the material is not isotropic.

Definition. (Perfect conductor) A *perfect conductor* corresponds to $\sigma \rightarrow \infty$ so $\mathbf{E} = 0$.

Definition. (Perfect insulator) A *perfect insulator* corresponds to $\sigma \rightarrow 0$ so $\mathbf{J} = 0$.

Consider a straight wire of length L in the direction of the unit vector \mathbf{n} and with uniform cross-sectional area A and conductivity σ . If the electric field is $\mathbf{E} = E\mathbf{n}$ where E is constant then $\mathbf{J} = \sigma E\mathbf{n}$ and the total current is $I = \sigma EA$. The potential difference (voltage) along the wire is

$$V = \int \mathbf{E} \cdot d\mathbf{x} = EL = \frac{IL}{\sigma A}.$$

So $V = IR$ where $R = \frac{L}{\sigma A}$ is the *resistance* of the wire.

Accompanying the resistance of a wire is *Joule heating* which is the conversion of EM energy into heat at rate $I^2 R$. If the voltage V is maintained by a battery then $VI = I^2 R$ is the rate at which the emf of the battery ($\mathcal{E} = V$) does work to maintain the current I .

4.3 Time-dependent electric fields

In electrodynamics we can no longer write

$$\mathbf{E} = -\nabla\Phi$$

since \mathbf{E} is no longer irrotational. But we still have the second Maxwell equation, so \mathbf{B} is still divergence free so we can still write

$$\mathbf{B} = \nabla \times \mathbf{A}$$

and the third Maxwell equation gives that

$$\nabla \times \left(\mathbf{E} + \frac{\partial \mathbf{A}}{\partial t} \right) = 0$$

which allows us to write

$$\mathbf{E} = -\nabla\Phi - \frac{\partial \mathbf{A}}{\partial t}$$

which generalises the electrostatic expression. Let's check this is well-defined by ensuring \mathbf{E} is unchanged by a time-dependent gauge transformation. Suppose that

$$\tilde{\mathbf{A}} = \mathbf{A} + \nabla\chi, \quad \tilde{\Phi} = \Phi - \frac{\partial\chi}{\partial t}$$

where $\chi(\mathbf{x}, t)$ is any scalar field, then both \mathbf{E} and \mathbf{B} are unchanged.

In magnetostatics we used Ampere's law

$$\int_C \mathbf{B} \cdot d\mathbf{x} = \mu_0 \int_S \mathbf{J} \cdot d\mathbf{S} = \mu_0 I$$

or its differential form

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{J}.$$

For a time-dependent situation we have to use the whole of the fourth Maxwell equation,

$$\nabla \times \mathbf{B} = \mu_0 \left(\mathbf{J} + \varepsilon_0 \frac{\partial \mathbf{E}}{\partial t} \right)$$

which contains an extra term, the *displacement current*.

Why is this extra term needed? Without it we would have $\nabla \cdot \mathbf{J} = 0$, which describes charge conservation in a situation where ρ is constrained to remain constant. But suppose we place free particles of positive charge in some localised region. Repulsive Coulomb forces cause the particles to separate implying that $\nabla \cdot \mathbf{J} > 0$.

We have seen that charge conservation in the correct form

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{J} = 0$$

follows from Maxwell's equations, including the displacement current.

4.4 Electromagnetic waves

We will consider freely evolving electric and magnetic fields in a vacuum, in the absence of charges and currents. We have that

$$\begin{aligned} \nabla \cdot \mathbf{E} &= 0 \\ \nabla \cdot \mathbf{B} &= 0 \\ \nabla \times \mathbf{E} &= -\frac{\partial \mathbf{B}}{\partial t} \\ \nabla \times \mathbf{B} &= \mu_0 \varepsilon_0 \frac{\partial \mathbf{E}}{\partial t}. \end{aligned}$$

Let's eliminate \mathbf{B} by taking the time derivative of the fourth Maxwell equation and substituting in the third equation. We get

$$\begin{aligned} \mu_0 \varepsilon_0 \frac{\partial^2 \mathbf{E}}{\partial t^2} &= \frac{\partial}{\partial t} (\nabla \times \mathbf{B}) \\ &= \nabla \times \frac{\partial \mathbf{B}}{\partial t} \\ &= -\nabla \times (\nabla \times \mathbf{E}) \\ &= \nabla^2 \mathbf{E} \end{aligned}$$

where we use the identity (and the first Maxwell equation)

$$\nabla \times (\nabla \times \mathbf{E}) = \nabla(\nabla \cdot \mathbf{E}) - \nabla^2 \mathbf{E}.$$

Alternatively, eliminate \mathbf{E} by taking the time derivative of the third Maxwell equation and substituting in the fourth.

$$\begin{aligned}\frac{\partial^2 \mathbf{B}}{\partial t^2} &= -\frac{\partial}{\partial t}(\nabla \times \mathbf{E}) \\ &= -\nabla \times \frac{\partial \mathbf{E}}{\partial t} \\ &= -\frac{1}{\mu_0 \varepsilon_0} \nabla \times (\nabla \times \mathbf{E}) \\ &= \frac{1}{\mu_0 \varepsilon_0} \nabla^2 \mathbf{B}.\end{aligned}$$

So each (Cartesian) component of \mathbf{E} and \mathbf{B} satisfies the wave equation

$$\frac{\partial^2 u}{\partial t^2} = c^2 \nabla^2 u$$

with wave speed

$$c = \frac{1}{\sqrt{\mu_0 \varepsilon_0}}$$

which is the speed of light (in a vacuum). We expect this since light is an electromagnetic wave involving oscillations of \mathbf{E} and \mathbf{B} .

Depending on the wavelength, EM waves can be radio waves, microwaves, infrared, visible light, ultraviolet, x-rays, and gamma rays.

Example. Consider a plane wave in which \mathbf{E} and \mathbf{B} depend only on (x, t) and not on (y, z) . A simple example is

$$\mathbf{E} = E(x, t) \mathbf{e}_y$$

where $E(x, t)$ satisfies the 1D wave equation

$$\frac{\partial^2 E}{\partial t^2} = c^2 \frac{\partial^2 E}{\partial x^2}.$$

The general solution (solving using method of characteristics) is

$$E(x, t) = f(x - ct) + g(x + ct)$$

which can be seen as the sum of a wave travelling without change of form in the positive x direction and another travelling in the negative x direction. We can use Maxwell's equation to find the corresponding magnetic field \mathbf{B} . We have

$$\begin{aligned}\frac{\partial \mathbf{B}}{\partial t} &= -\nabla \times \mathbf{E} \\ &= -\frac{\partial E}{\partial x} \mathbf{e}_z \\ &= (-f'(x - ct) - g'(x + ct)) \mathbf{e}_z\end{aligned}$$

and so

$$\mathbf{B} = B(x, t) \mathbf{e}_z$$

with $B(x, t) = \frac{1}{c}(f(x - ct) - g(x + ct))$. This also satisfies $\nabla \times \mathbf{B} = \mu_0 \varepsilon_0 \frac{\partial \mathbf{E}}{\partial t}$.

Of particular importance is a *monochromatic wave* of a singular angular frequency ω so

$$E = E_0 \cos(kx - \omega t)$$

$$B = \frac{E_0}{c} \cos(kx - \omega t)$$

where E_0 is a constant amplitude and $k = \frac{\omega}{c}$ is the *wavenumber* related to the wavelength λ by $k = \frac{2\pi}{\lambda}$.

Remark. Let's make some notes about these monochromatic waves

- The (angular) frequency and wavenumber are related by the *dispersion relation*

$$\omega^2 = c^2 k^2$$

so $\omega = \pm ck$.

- The oscillations of \mathbf{E} and \mathbf{B} are in phase but in orthogonal directions.
- The waves are *transverse*. The oscillatory fields are orthogonal to the direction in which the wave propagates.

Because Maxwell's equations are linear, Electromagnetic waves of different amplitudes, frequencies and directions can be superposed.

A more general approach to plane EM waves is to seek solutions of Maxwell's equations of the form

$$\mathbf{E} = \text{Re}(\mathbf{E}_0 e^{i(\mathbf{k} \cdot \mathbf{x} - \omega t)})$$

$$\mathbf{B} = \text{Re}(\mathbf{B}_0 e^{i(\mathbf{k} \cdot \mathbf{x} - \omega t)})$$

where \mathbf{E}_0 and \mathbf{B}_0 are complex constant vector amplitudes, \mathbf{k} is the real, constant *wavevector* and ω is the real constant angular frequency. The wavenumber is $k = |\mathbf{k}|$. The wave equation is satisfied by \mathbf{E} and \mathbf{B} if ω and k satisfy the dispersion relation

$$\omega^2 = c^2 k^2.$$

The individual Maxwell equations reduce to algebraic conditions.

$$\begin{aligned} \nabla \cdot \mathbf{E} &\implies \mathbf{k} \cdot \mathbf{E}_0 = 0 \\ \nabla \cdot \mathbf{B} &\implies \mathbf{k} \cdot \mathbf{B}_0 = 0 \\ \nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t} &\implies \mathbf{k} \times \mathbf{E}_0 = \omega \mathbf{B}_0 \\ \nabla \times \mathbf{B} = \mu_0 \varepsilon_0 \frac{\partial \mathbf{E}}{\partial t} &\implies \mathbf{k} \times \mathbf{B}_0 = -\frac{\omega}{c^2} \mathbf{E}_0 \end{aligned}$$

The fourth equation is redundant because the third and third together with the dispersion relation imply

$$\mathbf{k} \times \mathbf{B}_0 = \frac{1}{\omega} \mathbf{k} \times (\mathbf{k} \times \mathbf{E}_0) = -\frac{k^2}{\omega} \mathbf{E}_0 = -\frac{\omega}{c^2} \mathbf{E}_0.$$

Suppose that \mathbf{E}_0 is real. Then \mathbf{B}_0 is also real and the vectors $\mathbf{k}, \mathbf{E}_0, \mathbf{B}_0$ form an arothogonal triad. So \mathbf{E} and \mathbf{B} oscillate in fixed directions, with are perpendicular to each other and to the

direction of propagation. This is similar to the 1D wave considered previously and corresponds to a linearly polarised wave. Polarisation describes the orientation of the electric field.

Now suppose that \mathbf{E}_0 is complex and of the form

$$\mathbf{E}_0 = \mathbf{a} - i\mathbf{b}$$

where \mathbf{a}, \mathbf{b} are linearly independent real vectors in the plane perpendicular to \mathbf{k} . Then the direction of

$$\mathbf{E} = \mathbf{a} \cos(\mathbf{k} \cdot \mathbf{x} - \omega t) + \mathbf{b} \sin(\mathbf{k} \cdot \mathbf{x} - \omega t)$$

is not fixed but rotates in the plane perpendicular to \mathbf{k} . In general it traces out an ellipse and the wave is said to be *elliptically polarised*.

The special case of *circular polarisation* occurs when $|\mathbf{a}| = |\mathbf{b}|$ and $\mathbf{a} \cdot \mathbf{b} = 0$. Then \mathbf{E} traces out a circle.

The polarisation is *right-handed* if $\mathbf{b} = \hat{\mathbf{k}} \times \mathbf{a}$ and *left-handed* if $\mathbf{b} = \mathbf{a} \times \hat{\mathbf{k}}$.

Suppose a perfect conductor occupies the half-space $x > 0$. An *incident wave* approaching the surface $x = 0$ in the $+x$ direction has the form

$$\mathbf{E}_{\text{inc}} = \text{Re} \left(E_0 e^{i(kx - \omega t)} \mathbf{e}_y \right)$$

where we take $k > 0$, $\omega = ck > 0$ and $E_0 \in \mathbb{R}$.

Inside the conductor we have $\mathbf{E} = 0$. Since the tangential components of \mathbf{E} are continuous at the interface ($[\mathbf{n} \times \mathbf{E}] = 0$) we require $E_y = 0$ at $x = 0$. This can be satisfied by adding a *reflected wave*

$$\mathbf{E}_{\text{ref}} = \text{Re} \left(-E_0 e^{i(-kx - \omega t)} \mathbf{e}_y \right)$$

propagating in the $-x$ direction. The combined solution is

$$\begin{aligned} \mathbf{E} &= \mathbf{E}_{\text{inc}} + \mathbf{E}_{\text{ref}} \\ &= \text{Re} \left(2iE_0 \sin(kx) e^{-i\omega t} \mathbf{e}_y \right) \\ &= 2E_0 \sin(kt) \sin(\omega t) \mathbf{e}_y \end{aligned}$$

which is a standing wave.

Exercise. Show that \mathbf{B} is discontinuous at $x = 0$ and work out the surface current density $\mathbf{K}(t)$.

4.5 Electromagnetic energy

In electrostatics we found that the energy per unit volume in \mathbf{E} is $\frac{1}{2}\epsilon_0|\mathbf{E}|^2$. In magnetostatics we found the energy per unit volume was $\frac{1}{2\mu_0}|\mathbf{B}|^2$.

Now derive a general conservation law from Maxwell's equations. Consider the rate of change of the electric and magnetic energy per unit volume.

$$\begin{aligned} \frac{\partial}{\partial t} \left(\frac{1}{2}\epsilon|\mathbf{E}|^2 + \frac{1}{2\mu_0}|\mathbf{B}|^2 \right) &= \mathbf{E} \cdot \left(\epsilon_0 \frac{\partial \mathbf{E}}{\partial t} \right) + \mathbf{B} \cdot \left(\frac{1}{\mu_0} \frac{\partial \mathbf{B}}{\partial t} \right) \\ &= \mathbf{E} \cdot \left(\frac{1}{\mu_0} \nabla \times \mathbf{B} - \mathbf{J} \right) + \mathbf{B} \cdot \left(-\frac{1}{\mu_0} \nabla \times \mathbf{E} \right). \end{aligned}$$

An identity from vector calculus:

$$\nabla \cdot (\mathbf{E} \times \mathbf{B}) = \mathbf{B} \cdot (\nabla \times \mathbf{E}) - \mathbf{E} \cdot (\nabla \times \mathbf{B}).$$

Divide by μ_0 and add to the previous equation

$$\frac{\partial}{\partial t} \left(\frac{1}{2} \varepsilon_0 |\mathbf{E}|^2 + \frac{1}{2\mu_0} |\mathbf{B}|^2 \right) + \nabla \cdot \left(\frac{\mathbf{E} \times \mathbf{B}}{\mu_0} \right) = -\mathbf{E} \cdot \mathbf{J}.$$

This is the energy equation for the EM field. Write it as

$$\frac{\partial w}{\partial t} + \nabla \cdot \mathbf{S} = -\mathbf{E} \cdot \mathbf{J}$$

where

$$w = \frac{1}{2} \varepsilon_0 |\mathbf{E}|^2 + \frac{1}{2\mu_0} |\mathbf{B}|^2$$

is the *energy density* and

$$\mathbf{S} = \frac{\mathbf{E} \times \mathbf{B}}{\mu_0}$$

is the *energy flux density* and it known as the *Poynting vector*.

In the absense of currents this becomes

$$\frac{\partial w}{\partial t} + \nabla \cdot \mathbf{S} = 0$$

and expresses the conservation of electromagnetic energy. Integrate over some time independent volume V with $\partial V = S$ and use the divergence theorem

$$\frac{d}{dt} \int_V w dV + \int_S \mathbf{S} \cdot d\mathbf{S} = 0.$$

So the rate of change of energy in V is equal to minus the flux of energy through the surface S .

If we do however have $\mathbf{J} \neq 0$, we can see that $\mathbf{E} \cdot \mathbf{J}$ is the rate at which the EM field loses energy (per unit volume) by doing work on charged particles via the Lorentz force.