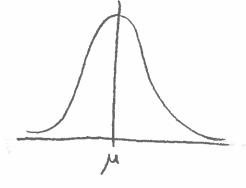
1) The ever-popular normal (or Gaussian) distribution looks

like this:



and is defined like this:

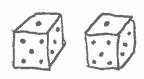
$$f(x|\mu,\sigma^2) = \frac{-(x-\mu)^2}{\sqrt{2\pi\sigma^2}}$$



But why? Where does this ugly looking expression come from?

2) Well, we can start by reminding ourselves that probability distributions. Sum to 1. So we can take any function g(x) over a discrete set:





3) And we can turn it into a probability distribution P(x) by dividing the function by a constant Z such that it sums to 1.

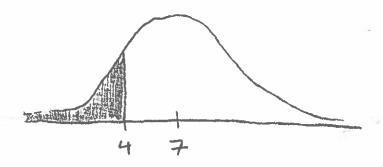
want
$$\Sigma P(x) = 1 \Rightarrow \sum_{x} g(x) = 1$$

$$\Rightarrow Z = \sum_{x} g(x)$$

(4) Now we can answer a question like "what is the probability that $x \le 4?$ " by summing the probability $P(x) \left| \frac{1}{36} + \frac{2}{36} \right|$

mass less than or equal to 4.

5) Now what if the function g(x) is over a continuous set like the real numbers?



If we can find a constant Z s.t.

$$\int_{Z}^{\infty} g(x) dx = 1$$

then we can define the probability that $x \le 4$ as the area under the curve shown about.

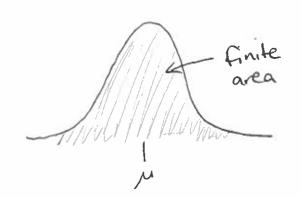
© Note that if this is possible at all, then $Z = \int_{-\infty}^{\infty} g(x) dx$

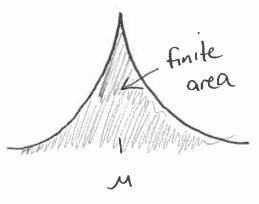
Thus it's only possible if Sg(x)dx is finite.

3 So let's say we want to build a continuous distribution from scratch.

We already know that we want $\int_{-\infty}^{\infty} g(x) dx$ to be finite.

What else might we want? Let's make it so that it has a unique maximum (mode), u, and that the Probability of x gets smaller as x gets further away from the mode M. J. So something like:

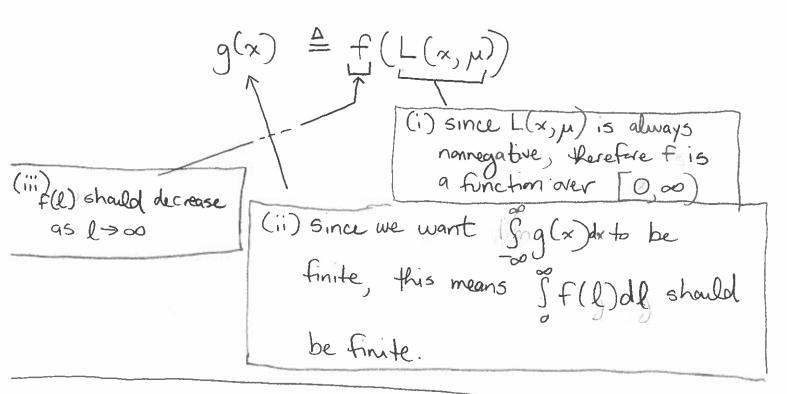




⁽³⁾ In other words, we want g(x) to be a function of some absolute distance between x and μ .

- let $L(x,\mu)$ be the nonnegative distance - let $g(x) \triangleq f(L(x,\mu))$ what, now fluis two more functions?

9) There are a cauple implications here:



10) So we need to find a function f(l) over $[0,\infty)$ such that Sf(l) all is finite. It'd also be supernice if f(l) was easy to integrate. What about $f(l) = e^{-l}$?

$$\int_{0}^{\infty} f(l) dl = \int_{0}^{\infty} e^{-l} dl$$

$$= \lim_{k \to \infty} \int_{0}^{\infty} e^{-l} dl$$

$$= \lim_{k \to \infty} \left[-e^{-l} \right]_{0}^{k}$$

$$= \lim_{k \to \infty} \left[-e^{-k} - (-e^{0}) \right]$$

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12) To recap, if we define:

$$g(x) = e^{-L(x,\mu)}$$

where $L(x,\mu)$ is a nonnegative distance between x and a mode μ , then we get lots of nice properties:

- Sg(x) is finite
- g(x) is maximal at a unique mode M.
- = g(x) decreases as x gets further away from u.

13) Now we just have to decide on a nonnegative distance $L(x, \mu)$. Let's start with just the absolute difference between x and μ .

$$L_{1}(x, \mu) = \frac{|x - \mu|}{b}$$

We include a "scalling" factor b that allows us to specify what "far" means. Maybe we're talking about distances between cities, and we want 10 miles to be "close" and 1000 miles to be "far". We could use b=10, which gives us

$$L_1(10,0)=1$$
 $L_1(1000,0)=100$

Or maybe we're talking about in terplanetary distances, and so 240,000 miles (from the Earth to the moon) is near and 33,000,000 miles (from the Earth to Mas) is "far". We could use b=240000, which gives us: $L_1(240000,0)=1$ $L_1(33000000,0)=137.5$

This choice of distance gives us:
$$g(x) = e^{-\frac{|x-\mu|}{b}}$$

Remember, to turn this into a probability function, all we need to do is divide by $Z = \int_{-\infty}^{\infty} g(x) dx$.

(5)
$$\int_{-\infty}^{\infty} e^{-\frac{|x-\mu|}{b}} dx = \int_{-\infty}^{\infty} e^{\frac{|x-\mu|}{b}} dx + \int_{\mu}^{\infty} e^{\frac{|x-\mu|}{b}} dx$$

$$= \left(\lim_{k \to -\infty} \int_{k}^{\infty} e^{\frac{|x-\mu|}{b}} dx\right) + \left(\lim_{k \to \infty} \int_{\mu}^{\infty} e^{\frac{|x-\mu|}{b}} dx\right)$$

$$= \left(\lim_{k \to -\infty} \int_{k}^{\infty} e^{\frac{|x-\mu|}{b}} dx\right) + \left(\lim_{k \to \infty} \left[-e^{\frac{|x-\mu|}{b}} \right]_{\mu}^{k}\right)$$

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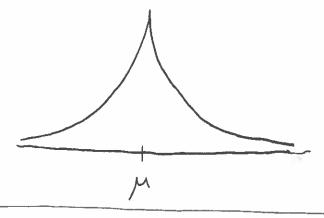
$$= \left(\lim_{k \to -\infty} \int_{k}^{\infty} e^{\frac{|x-\mu|}{b}} dx\right) + \left(\lim_{k \to -\infty} \int_{k$$

(6) So we have a distribution!
$$p(x) = g(x) = \frac{-|x-\mu|}{b}$$

$$\int_{-\infty}^{-1} g(x) dx = \frac{-|x-\mu|}{2b}$$

That's not the normal distribution

No, it's not the normal distribution, but it's nevertheless a fin and sometimes handy distribution called the Laplace distribution.



(7) What if we try a different distance function L(x, u)? For instance, we could try the Squared distance (also with a scaling parameter b): $L_2(x_{>}\mu) = (x-\mu)^2$

Note this is also nonneaptive, just like L, (2, 11).

(18) This choice of distance gives us:
$$g(x) = e^{-\frac{(x-\mu)^2}{b}}$$

To turn this into a distribution, we need to divide g(x) by $Z = \int_{-\infty}^{\infty} g(x) dx$.

We'll just look this one up: $\int_{-\infty}^{\infty} e^{-(x-\mu)^2} dx = \sqrt{b\pi}$

19) Our resulting distribution is:
$$P(x) = \frac{q(x)}{s} = \frac{1-(x-(x-w)^2)}{s}$$
That's not the normal dist... oh wait.

If we substitute the parameter $\sigma^2 = \frac{b}{2}$, then we get $p(x) = \frac{1}{1 - (x - w^2)} = \frac{-(x - w^2)}{2}$

which is indeed our friendly neighborhood normal distribution.

20) We write:

to mean

$$p(x) = \frac{1}{\sqrt{2\pi\sigma^2}} e^{\frac{-(x-\mu)^2}{2\sigma^2}K}$$

and we write

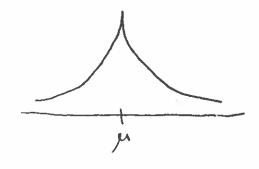
to mean

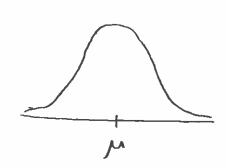
$$p(x) = \frac{1}{2b} e^{-\frac{|x-\mu|}{b}}$$

we substituted or for b because it's a more interpretable parameter (it's the variance of the distribution)



1 It's worth examining when you might want to use Laplace (µ,b) vs Normal (µ, 02)





Laplace can be a little annoying to use because it's not differentiable (i.e. spiky) at μ .

DISTRIBUTIONS

(2) On the other hand, the Laplace has so-called "heavier tails" than a normal distribution of similar variance.

Compare Normal (O, 1) with Laplace (O, 1), which each have variance !.

variance is the expected squared difference of a drawn sample x from the mean u

Suppose we randomly take a single sample from each distribution, once per day. How often should I expect to get a sample...

Naplace | Normal

≥3?

> 5?

>7?

4 per year / levery two years

H per decade | every 10000 years (once in recorded history)

3 per century every 2 billion years (twice in the history of Earth)

²³⁾ Basically the normal distribution doesn't give much allowance to rare events (sometimes called "black swan events") - they should basically never happen, according to the normal distribution.

For some phenomena (many actually), the normal distribution is a great choice. But others, like say the stock market, are subject to fluke events too often to be effectively explained by a normal distribution.