

Introduction 简介

Test for cite and enumerate and term 同时测试中文

Test for cite and enumerate

Example:

```
[#Kormendy&Ho2013ARA&A, Koss2012ApJ]
[#Kormendy&Ho2013ARA&A, Koss2012ApJ] []
[foo] [#Kormendy&Ho2013ARA&A, Koss2012ApJ]
[foo\]\[bar] [#Kormendy&Ho2013ARA&A, Koss2012ApJ]

[#Kormendy&Ho2013ARA&A, Koss2012ApJ;]
[#Kormendy&Ho2013ARA&A, Koss2012ApJ;] []
[foo] [#Kormendy&Ho2013ARA&A, Koss2012ApJ;]
[foo\]\[bar] [#Kormendy&Ho2013ARA&A, Koss2012ApJ;]
```

Term:

- [1, 2]
 - [1, 2]
- [1, 2, foo]
 - [foo 1, 2, bar]

Enumerate:

1. (author?) [1, 2]
2. (author?) [1, 2]
3. (author?) [1, 2, foo]
4. (author?) [foo 1, foo 2, bar]

Here is not cited().

Here is a enumerate:

1. Type for “1984.”

Here is not:

1984. Type for “1984.”

List of Markers:

\ Backslash0 \ Backslash1

Backtick0 Backtick1

- Asterisk0
- Asterisk1
- plus sign0
- plus sign1
- minus sign0 (hyphen)
- minus sign1 (hyphen)

This is a line. This is a new line.

Cite for intro §Introduction 简介

Cite for cite note (§ 1.1)

Cite for quote §Test for quote

在这里测试中文的显示效果

Test for quote

Test for quote:

- Level 1 Lazy type in for level 1
 - Level 1, must have a empty line for new line, Level 1 (failed new line)
 - Level 2 Lazy type in for level 2, must have a empty type for lazy type
 - Level 2
 - Level 2 (successful new line)
 - Level 3 Lazy type in for level 3, Not lazy type in for level 3

Other test

test for rules

Test for rules:
Some

Thing

Between

different

rules

En...

test for Links

This is **an example**¹ inline link.
This link² has no title attribute.
Also we can add a foot note³
Or like **this**⁴
Or like **id**⁵
I get 10 times more traffic from **Google**⁶ than from **Yahoo**⁷ or **MSN**⁸.

test for emphasis

single asterisks
single underscores
double asterisks
double underscores
this text is surrounded by literal asterisks

test for code

Use the `printf()` function.
There is a literal backtick (`) here.
A single backtick in a code span: `
A backtick-delimited string in a code span: `foo`
Please don’t use any `<blink>` tags.
— is the decimal-encoded equivalent of `—`.

test for equation

Some equations: eq:0 (§ 2.5)

$$x_{1,2} = \frac{-b \pm \sqrt{b^2 - 4ac}}{2a}$$
$$x_{1,2} = \frac{-b \pm \sqrt{b^2 - 4ac}}{2a}$$

(1)

Inline: $x_{1,2} = \frac{-b \pm \sqrt{b^2 - 4ac}}{2a}$, blabla
Inline: $x_{1,2} = \frac{-b \pm \sqrt{b^2 - 4ac}}{2a}$, blabla
公式堆叠测试

$$\text{text}_{i=1}^{i=n} \quad \begin{array}{c} \text{上面} \\ \text{text} \\ \text{下面} \end{array}$$

(2)

eq:1 (Eq. 1) eq3 inline (Eq. 2.5) ￼

$$1 = \Omega_0 + \begin{array}{c} \text{这是} \\ \Omega_\Lambda \\ \text{暗能量} \end{array} + \Omega_k,$$

(3)

And:

$$\Omega_0 = \frac{\rho_0}{\rho_{c0}} = \Omega_m + \Omega_r;$$
$$\Omega_\Lambda = \frac{\rho_\Lambda}{\rho_{c0}};$$
$$\Omega_k = -\frac{kc^2}{a_0^2 H_0^2};$$

(4)

(5)

(6)

test for table

MultiMarkdown table here (Table 1):

表 1: Very very very very very very very very very long caption		
	Grouping	
First Header	Second Header	Third Header
Content	<i>Long Cell</i>	
Content	Cell	Cell
New section	More	Data
And more	And more	

Large Table:large (Table 2)

表 2: Photometric Parameters

Target name	Aperture	m_Y	m_U	M_z	M_u	$u - z$	M_*
(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)
SDSS J1108+0659	18''	16.24±0.05	18.48±0.26	-23.12	-21.37	1.76±0.26	11.14
SDSS J1131−0204	12''	15.81±0.01	18.94±0.09	-23.04	-20.39	2.65±0.10	11.24
SDSS J1146+5110	18''	15.89±0.03	18.40±0.14	-22.68	-20.65	2.02±0.14	11.00
SDSS J1332+0606	12''	16.82±0.04	19.29±0.16	-22.86	-20.88	1.98±0.17	11.07

Col. 2: the aperture sizes that are large enough to enclose more than 95% flux of the galaxies; Col. 3 & 4: the apparent magnitudes in Y and U bands measured with the apertures correspondingly. The uncertainties are estimated from 1 σ variation of the sky background; Col. 5 & 6: the SDSS z - and u -band absolute magnitudes of the targets, transformed from m_Y and m_U correspondingly, assuming a flat local spectra ($f_\lambda \sim const.$) around the relevant frequencies; Col. 7: color calculated from the u and z magnitudes; Col. 8: the stellar mass in solar unit of the targets estimated from the z - and u -band magnitudes (see the text for details, §??). Col. 9: the bulge-to-total mass ratios. The bulge mass is obtained by adding up the two bulge components in GALFIT decomposition results for each galaxy.

Test for figs

Include some figure:

Image one (image with given size, 200x400): image fixed (Figure 4)

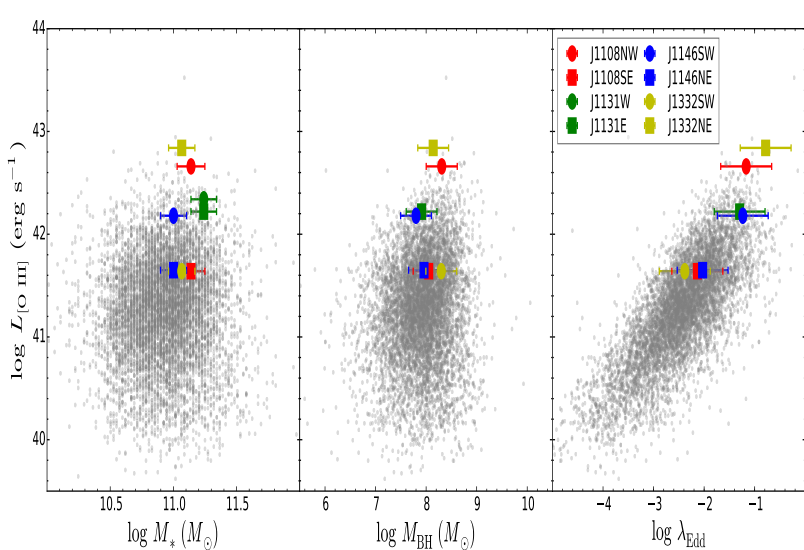


图 1: Fig with give size, 200x400

Image two (image direct): image direct, no way to ref it (??)

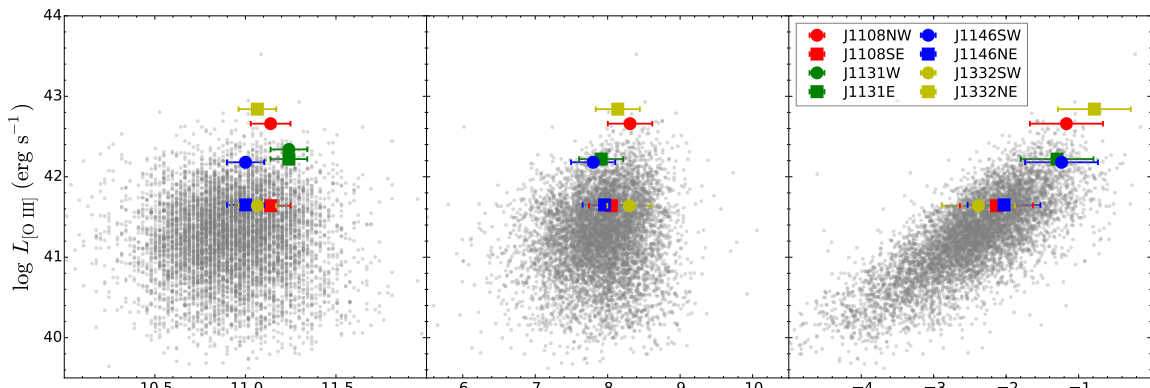


图 2: Direct figure title

Image three (image: normal): Figure 3

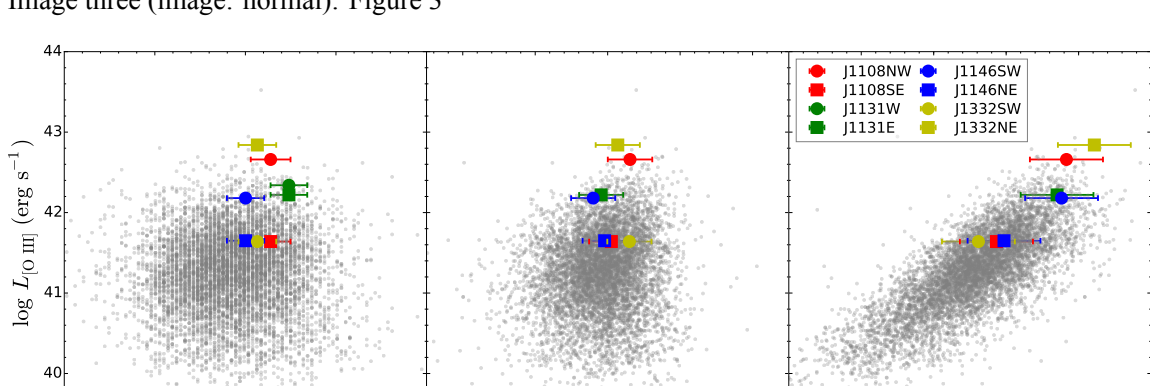


图 3: Here we have our long long blablaFigblablaFigblablaFigblabla FigtitleblablaFigblablaF igblablaFigblablaFigti tleblablaFigblablaFig, A demonstration of our near- to far-IR SEDs of PG quasars, produced by [4]., I try to ref intro (§ 1)

Image Four (image:small):Figure 4

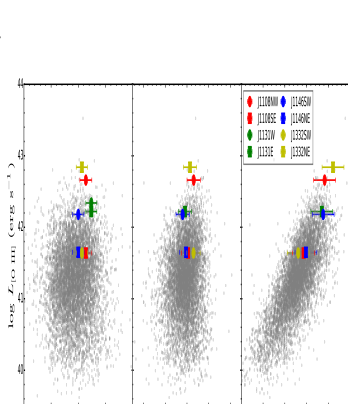


图 4: 0.3 \textwidth and 0.25 \textheight, compare the size with Figure 3

Good include: cmp (Figure 5)

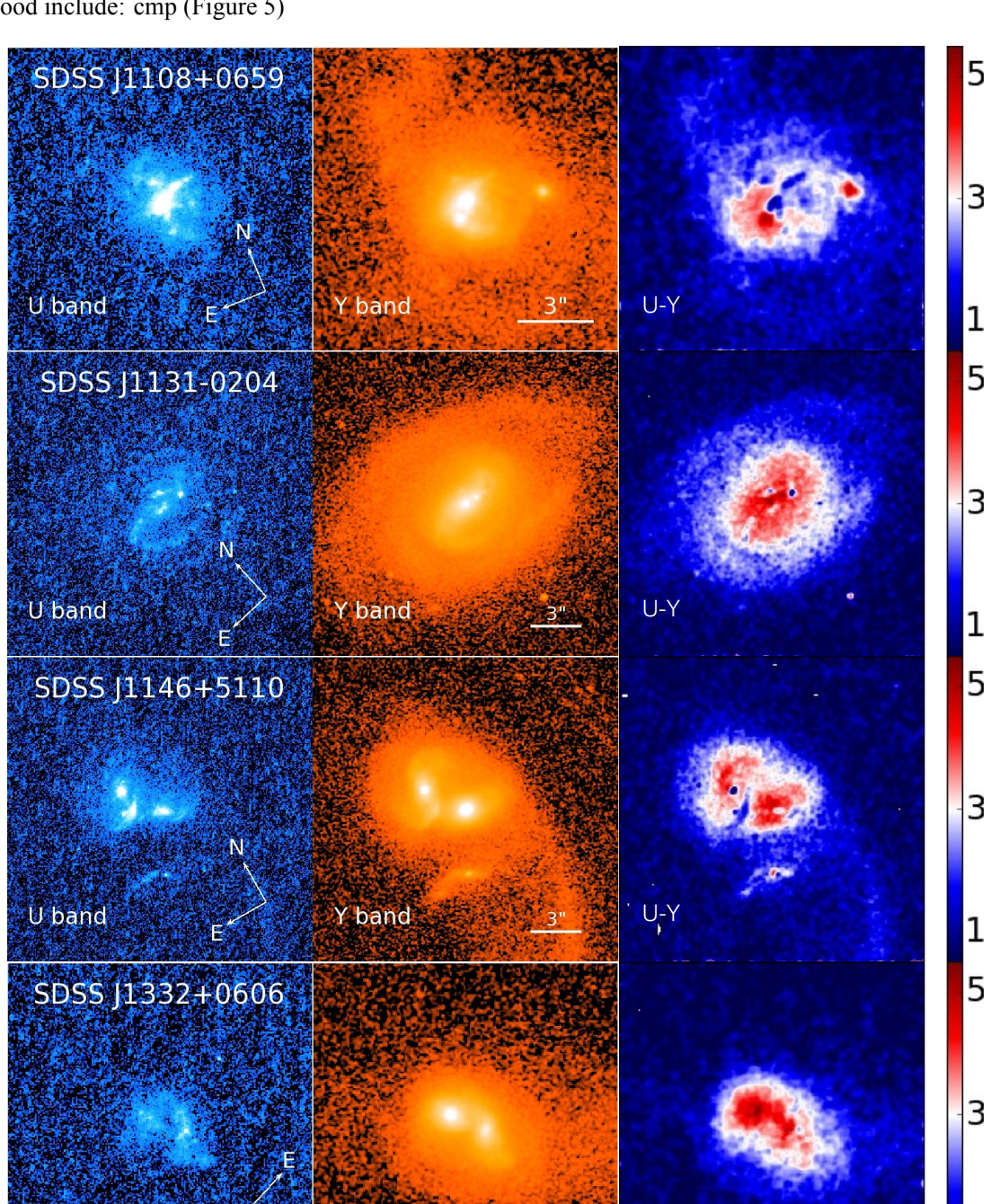


图 5: Images observed through U- and Y-band filters are shown in *left* and *middle* columns. The color maps are in the *right column*. The FOV and orientation of the images in each row are the same. The compass and 3'' scale bar are on the lower right corner of *left* and *middle* panels in each row.

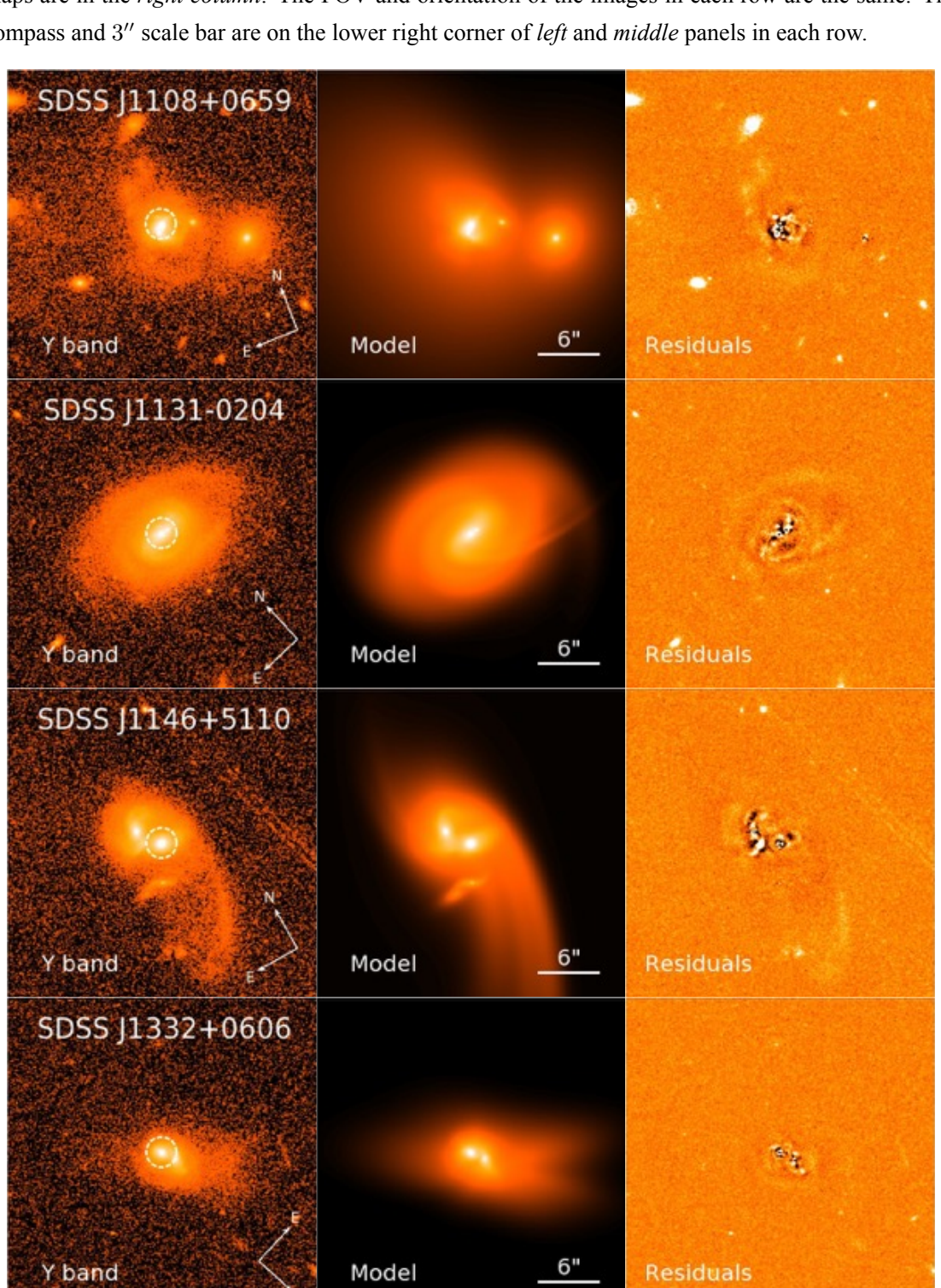


图 6: Model fits of the *HST* Y-band images of the four sources. *Left column* is the Y-band image of each target. The white dashed circles demonstrate the coverage of SDSS fibers (3'' in diameter). The compass of each image is shown in the lower right corner. *Middle column* is GALFIT best fit of each target. The scale bar is shown in the lower-right corner. The FOVs of three panels in each row are the same. The pixel brightness scale is logarithmic in the first two columns. The residuals are show in *Right column*, where the pixel brightness scale is linear. We simultaneously fit the close companions of J1108 and J1146.

References

- [1] J. Kormendy and L. C. Ho, “Coevolution (Or Not) of Supermassive Black Holes and Host Galaxies,” *ARAA*, vol. 51, pp. 511–653, Aug. 2013.
- [2] M. Koss, R. Mushotzky, E. Treister, S. Veilleux, R. Vasudevan, and M. Trippe, “Understanding Dual Active Galactic Nucleus Activation in the nearby Universe,” *APJL*, vol. 746, p. L22, Feb. 2012.
- [3] M. Pović, M. Sánchez-Portal, A. M. Pérez García, A. Bongiovanni, J. Cepa, M. Huertas-Company, M. A. Lara-López, M. Fernández Lorenzo, A. Ederoclite, E. Alfaro, H. Castañeda, J. Gallego, J. I. González-Serrano, and J. J. González, “AGN-host galaxy connection: morphology and colours of X-ray selected AGN at $z < 2$,” *AAP*, vol. 541, p. A118, May 2012.
- [4] J. Kormendy and D. Richstone, “Inward Bound—The Search For Supermassive Black Holes In Galactic Nuclei,” *ARAA*, vol. 33, p. 581, 1995.