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Several hypothesis were made around this, some of them said that Standard Solar model was not correct and some others that the Sun ran out of hydrogen, but Bruno Pontecorvo proposed the hypothesis that neutrinos actually have the property to oscillate into different flavors, this theory contradicted the standard model of particle physics because the model theorized that neutrinos were massless particles, but for a neutrino to oscillate into different flavors must have a mass, and in 2001, the Sudbury Neutrino Observatory was able to detect the three neutrino flavors, all of them coming from the Sun, but the Standard Solar Model says that the Sun only produces electron neutrinos, so the conclusion was that neutrino oscillates and actually has a mass value.

## 2 PMNS Matrix

Since the neutrino oscillates and has a mass value, to measure the neutrino oscillations it is used something called the Pontecorvo Maki Nakagawa Sakata Matrix (also known as PMNS matrix, lepton mixing matrix, or neutrino mixing matrix)(eq 1 and 2), where the left part of equation is the expresion of a neutrino type, and the right part is the product of the lepton mixing matrix, which U components are amplitude of the mass eigenstates, and the mass eigenstates of the neutrino, if we apply matrix multiplication we can get a neutrino flavor in terms of the mass eigenstates.

$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{bmatrix} = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{bmatrix} \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{bmatrix} \quad (1)$$

$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{bmatrix} = \begin{bmatrix} U_{e1}\nu_1 + U_{e2}\nu_2 + U_{e3}\nu_3 \\ U_{\mu1}\nu_1 + U_{\mu2}\nu_2 + U_{\mu3}\nu_3 \\ U_{\tau1}\nu_1 + U_{\tau2}\nu_2 + U_{\tau3}\nu_3 \end{bmatrix} \quad (2)$$

$$\nu_e = \sum_{i=1}^3 U_{ei}\nu_i \quad (3)$$

## 3 Neutrinoless double beta decay

Neutrinoless double beta decay is a theoretical type of beta decay that, unlike the standard beta decay  ${}^A_ZX \rightarrow {}^A_{Z+1}X + e^- + \nu_e$  and the not so common double beta decay  ${}^A_ZX \rightarrow {}^A_{Z+2}X + 2e^- + 2\nu_e$ , the neutrinoless double beta decay is characterized for being a beta decay process where the trajectory of the ejected neutrinos collide and annihilate each other (as the name suggests)  ${}^A_ZX \rightarrow {}^A_{Z+2}X + 2e^-$ . Even though the hypothesis is attractive, for this to be possible, neutrino has to be their own antiparticle so it can annihilate with another one, but for technical reasons, instead of theorizing the existence and probability to find a neutrinoless double beta decay, this document will assume

61 it exists and based on that, use the effective mass  $\langle m_{ee} \rangle$  that is derived from  
62 the PMNS matrix, and  $m_{light}$  relationship to find possible mass values for the  
63 neutrino including the possibility of the existence of 1 and 2 sterile neutrinos

## 64 4 Lobster plot

65 Methodology followed to recreate the original lobster of the 3 neutrinos masses  
66 was confirmed that exist  $(m_1, m_2, m_3)$  was using equation 1 and expand it until  
67 we get 3 different mass terms (showed in equation 4, 5, and 6) that are related  
68 to the effective mass.

$$\text{Effective mass } \langle m_{ee} \rangle = \left| \sum U_{ei}^2 m_i \right| \quad (4)$$

$$\langle m_{ee} \rangle = \left| \sum_{i=1}^3 U_{ei}^2 m_i \right| = \left| m_{ee}^{(1)} + m_{ee}^{(2)} e^{2i\alpha} + m_{ee}^{(3)} e^{2i\beta} \right| \quad (5)$$

$$\begin{aligned} \langle m_{ee} \rangle &= |c_{12}^2 c_{13}^2 m_1 + s_{12}^2 c_{13}^2 m_2 e^{i\phi_{12}} + s_{13}^2 m_3 e^{i\phi_{13}}| \\ s_{ij}^2 &= \sin^2 \theta_{ij} \\ c_{ij}^2 &= \cos^2 \theta_{ij} \end{aligned} \quad (6)$$

69  
70  
71 At the end, we got an equation with three terms that each of them is related  
72 to an specific mass term, however, the problem here is that we do not know the  
73 mass values, but we know the absolute value of the square difference of two  
74 masses ( $|\Delta m_{ij}^2|$ ) and the value of  $s_{12}^2$  and  $s_{13}^2$ , so based on that it is possible to  
75 put the mass terms in terms of the lightest mass, but since we know  $|\Delta m_{ij}^2|$ , we  
76 have to consider the case were  $\Delta m_{ij}^2 > 0$  and  $\Delta m_{ij}^2 < 0$  and for doing that, we  
77 have to make a plot considering different scenarios of mass hierarchy, and those  
78 are scenarios are represented as normal hierarchy NH, and inverted hierarchy  
79 IH, since experimental results showed that  $\Delta m_{21}^2 > 0$  (also known as solar mass  
80  $\Delta m_{solar}^2$ ) we know that the lightest mass can only be  $m_1$  or  $m_3$ , and  $\Delta m_{31}^2$  (also  
81 known as atmospheric mass  $\Delta m_A^2$ ) is the one that defines the hierarchy.  
82

Hierarchy name	Hierarchy label	Mass ordering
Normal hierarchy	NH	$m_3 > m_2 > m_1$
Inverted hierarchy	IH	$m_2 > m_1 > m_3$

$$\begin{aligned}
\text{NH: } m_2 &= \sqrt{m_1^2 + \Delta m_{solar}^2} \quad m_3 = \sqrt{m_1^2 + \Delta m_A^2} \\
\text{IH: } m_2 &= \sqrt{m_3^2 + \Delta m_{solar}^2 + \Delta m_A^2} \quad m_1 = \sqrt{m_3^2 + \Delta m_A^2}
\end{aligned} \tag{7}$$

$$\text{where } \Delta m_{solar}^2 = \Delta m_{21}^2 \text{ and } \Delta m_A^2 = \Delta m_{31}^2$$

$$(\Delta m_{31}^2 > 0 \text{ for NH and } \Delta m_{31}^2 < 0 \text{ for IH})$$

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85

## 86 5 Data used for 3 masses

87

Parameter	Best Fit	$3\sigma$
$\Delta m_{21}^2$ ( $10^{-5} \text{ eV}^2$ )	7.53	8.07-6.99
$ \Delta m_{31}^2 $ ( $10^{-3} \text{ eV}^2$ )	2.5283	2.5631-2.4935
$s_{12}^2$	0.307	0.346-0.268
$s_{13}^2$	0.022	0.0241-0.0199

## 88 6 Step by step to make it work for 3 neutrinos

89 So, after having the required materials to mathematically build the plot, next  
90 step is creating the plot using Jupyter Notebook (Details of how the code  
91 was created and how does it works are located [here](#) or go to the next url  
92 <https://github.com/FranciscoTapia61199/Sterile-neutrinos>)

93

94 It is tricky to get this plot, because Python reads it as a 1-dimensional vari-  
95 able (as it is required to do) but there is also the free parameters of  $\phi_{12}$  and  
96  $\phi_{13}$ , and with each different combination, the code will create another curve,  
97 and since there are infinite many combinations of angles, it would be impossible  
98 for the code to run it without burning the PC, so, another strategy is to find  
99 the angles where the boundaries of this function are located and fill each the  
100 space between them to make sure all possible angle combinations are included  
101 without making the PC consume a lot of time and a lot of resources, to calcu-  
102 late the angles that made the boundaries, i used an optimization method using  
103 derivatives to find the max and min curves of the function.

104

$$\frac{\partial(\langle m_{ee} \rangle)}{\partial(\phi_{12})} = 0 \text{ if } \phi_{12} = (0, \pi)$$

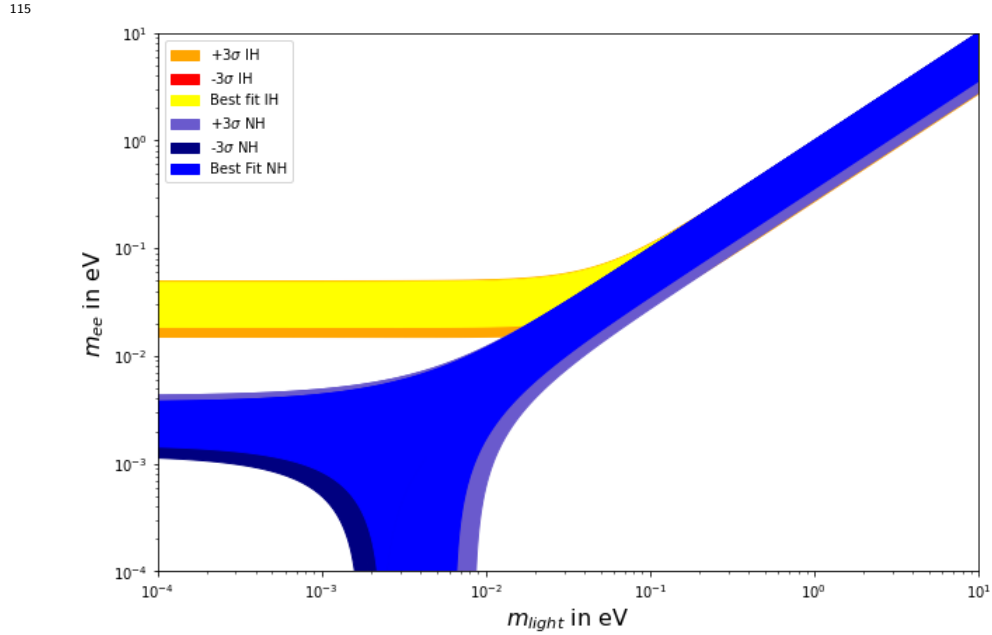
106

$$\frac{\partial(\langle m_{ee} \rangle)}{\partial(\phi_{13})} = 0 \text{ if } \phi_{13} = (0, \pi)$$

upper bound at  $\phi_{12}=0$  and  $\phi_{13}=0$

lower bound at  $\phi_{12} = \pi$  and  $\phi_{13} = \pi$

lower inferior bound (only for NH case) at  $\phi_{12} = \pi$  and  $\phi_{13} = 0$



## 7 Extending the definition for sterile neutrinos

Liquid Scintillator Neutrino Detector (LSND) was a neutrino detector which main purpose was to look for evidence for the neutrino oscillation, and that resulted on the detection of an anomaly where the recorded value  $\Delta m_{41}^2 = 1.78 \text{eV}^2$  but the cosmological limit for neutrino mass is 1 eV, so this anomaly must be something apart for the solar and atmospheric masses.

The formulas and techniques used before were defined just for the existence of 3 neutrinos, however, we can actually extend the formulas to include the existence of 4, 5.....n neutrinos (but for the sake of time, coherence and computational resources, here are covered only the cases for 1 and 2 sterile neutrinos, but as is mentioned, in theory, this is can be applied for n neutrinos) and the formula used would be equations 8 and 9.

$$\langle m_{ee} \rangle = \left| \sum_{i=1}^4 U_{ei}^2 m_i \right| = \left| m_{ee}^{(1)} + m_{ee}^{(2)} e^{2i\alpha} + m_{ee}^{(3)} e^{2i\beta} + m_{ee}^{(4)} e^{2i\gamma} \right| \quad (8)$$

$$\langle m_{ee} \rangle = |c_{12}^2 c_{13}^2 c_{14}^2 m_1 + s_{12}^2 c_{13}^2 c_{14}^2 m_2 e^{i\phi_{12}} + s_{13}^2 c_{14}^2 m_3 e^{i\phi_{13}} + s_{14}^2 m_4 e^{i\phi_{14}}| \quad (9)$$

As we can see in the formula, including another mass term to the equation, is not just adding some  $m_4$  times some  $s_{ij}^2 e^{i\phi_{ij}}$  term, it is necessary to multiply each mass term by some  $c_{ij}^2$  (in this case,  $c_{14}^2$ ) and that makes sense, because, if  $\theta_{14} = 0$  we get back the equation with 3 mass terms, so, following that pattern, is possible to expand it for n neutrinos.

Besides including the sterile mass, it is necessary to also include a  $m_{light}$  dependence to  $m_4$ , so, for that, we just repeat the process used previously of the experimental measurement  $|\Delta m_{ij}^2|$  to add the  $m_{light}$  dependence, that in this case we use  $\Delta m_{41}^2$  (also known as Experiment LSND mass  $\Delta m_{LSND}^2$ ), but since the experimental result is an absolute value, besides considering the previous hierarchies, we have to consider the cases were  $m_4$  is the heaviest mass, and were  $m_4$  is the lightest mass, and with all of this in consideration, we can start making the formulas for the plot.

Hierarchy name	Hierarchy label	Mass ordering
Heavy sterile neutrino (Normal hierarchy)	SNH	$m_4 > m_3 > m_2 > m_1$
Heavy sterile neutrino (Inverted hierarchy)	SIH	$m_4 > m_2 > m_1 > m_3$
Light sterile neutrino (Normal hierarchy)	NHS	$m_3 > m_2 > m_1 > m_4$
Light sterile neutrino (Inverted hierarchy)	IHS	$m_2 > m_1 > m_3 > m_4$

Heavy sterile neutrino ( $\Delta m_{41}^2 > 0$ )

$$\text{NH: } m_2 = \sqrt{m_1^2 + \Delta m_{solar}^2} \quad m_3 = \sqrt{m_1^2 + \Delta m_A^2} \quad m_4 = \sqrt{m_1^2 + \Delta m_{LSND}^2}$$

$$\text{IH: } m_2 = \sqrt{m_3^2 + \Delta m_{solar}^2 + \Delta m_A^2} \quad m_1 = \sqrt{m_3^2 + \Delta m_A^2} \quad m_4 = \sqrt{m_3^2 + \Delta m_{LSND}^2 + \Delta m_A^2}$$

$$\text{where } \Delta m_{solar}^2 = \Delta m_{21}^2, \quad \Delta m_A^2 = \Delta m_{31}^2 \text{ and, } \Delta m_{LSND}^2 = \Delta m_{41}^2$$

$$\Delta m_{31}^2 > 0 \text{ for NH and } \Delta m_{31}^2 < 0 \text{ for IH} \quad (10)$$

## 145 8 Data used for 4 masses

146

Parameter	Best Fit	$3\sigma$
$\Delta m_{21}^2$ ( $10^{-5}$ $eV^2$ )	7.53	8.07-6.99
$ \Delta m_{31}^2 $ ( $10^{-3}$ $eV^2$ )	2.5283	2.5631-2.4935
$ \Delta m_{41}^2 $ ( $eV^2$ )	1.78	2.01-1.61
$s_{12}^2$	0.307	0.346-0.268
$s_{13}^2$	0.022	0.0241-0.0199
$s_{14}^2$	0.023	0.04-0.006

## 147 9 Step by step to make it work for 3+1 neutrinos

148

149 To make this plot, we follow exactly the same process as for 3 neutrinos, but  
150 since we are including a new  $\phi_{14}$  dependence, then, the original boundaries  
151 for the 3 neutrino case must be different ones (at least for the NH case), so,  
152 following the same optimization method with derivatives, we found that criti-  
153 cal angles are 0 and  $\pi$  (in fact, even with n neutrinos, the critical values will  
154 always be a mix of 0 and  $\pi$  angles) and based on that, we can construct the plots.

155

$$156 \quad \frac{\partial(\langle m_{ee} \rangle)}{\partial(\phi_{12})} = 0 \text{ if } \phi_{12} = (0, \pi)$$

$$157 \quad \frac{\partial(\langle m_{ee} \rangle)}{\partial(\phi_{13})} = 0 \text{ if } \phi_{13} = (0, \pi)$$

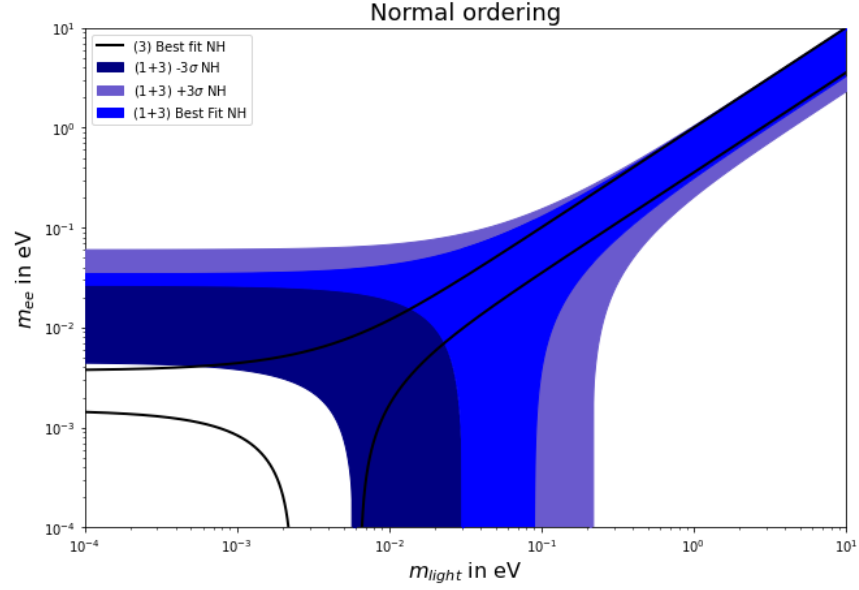
$$158 \quad \frac{\partial(\langle m_{ee} \rangle)}{\partial(\phi_{14})} = 0 \text{ if } \phi_{14} = (0, \pi)$$

159 upper bound at  $\phi_{12}=0$ ,  $\phi_{13}=0$ , and  $\phi_{14}=0$

160 lower bound at  $\phi_{12} = \pi$ ,  $\phi_{13} = \pi$ , and  $\phi_{14} = \pi$

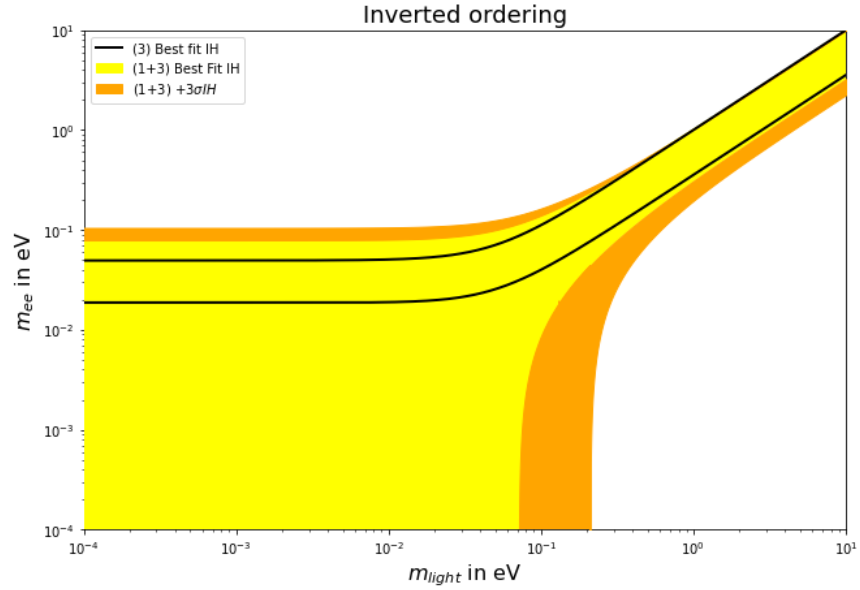
161 lower inferior bound (only for SNH case) at  $\phi_{12} = 0$ ,  $\phi_{13} = 0$ , and  $\phi_{14} = \pi$

## 162 9.1 Heavy sterile neutrino (SNH)



163

## 164 9.2 Heavy sterile neutrino (SIH)



165

166 We also have to include the case where  $m_4$  is the lightest mass, however,  
 167 according to the previous hierarchies,  $m_{light}$  in case for the active neutrinos



168 (the ones we know that actually exist) is  $m_1$  or  $m_3$ , but neither of them are the  
 169 lightest mass anymore, so, for experimental purposes, we are going to include  
 170 2 types of plots, one where we put everything in terms of the lightest mass,  
 171 regardless if is active or sterile, and one where we put everything in terms  
 172 of the lightest active mass, as mentioned before, we include one of the active  
 173 neutrino mass separated to be able to see the possibilities to find the sterile  
 174 neutrino via experimental measurements. (Details of how the plots were made  
 175 in <https://github.com/FranciscoTapia61199/Sterile-neutrinos>).

Light sterile neutrino ( $\Delta m_{41}^2 < 0$ )

$$\text{NH: } m_2 = \sqrt{m_4^2 + \Delta m_{solar}^2 + \Delta m_{\text{LSND}}^2} \quad m_3 = \sqrt{m_4^2 + \Delta m_A^2 + \Delta m_{\text{LSND}}^2} \quad m_1 = \sqrt{m_4^2 + \Delta m_{\text{LSND}}^2}$$

$$\text{IH: } m_2 = \sqrt{m_4^2 + \Delta m_{solar}^2 + \Delta m_{\text{LSND}}^2} \quad m_3 = \sqrt{m_4^2 + \Delta m_A^2 - \Delta m_{\text{LSND}}^2} \quad m_1 = \sqrt{m_4^2 + \Delta m_{\text{LSND}}^2}$$

$$\text{where } \Delta m_{solar}^2 = \Delta m_{21}^2, \Delta m_A^2 = \Delta m_{31}^2 \text{ and, } \Delta m_{\text{LSND}}^2 = \Delta m_{41}^2$$

$$\Delta m_{31}^2 > 0 \text{ for NH and } \Delta m_{31}^2 < 0 \text{ for IH} \quad (11)$$

(active) Light sterile neutrino ( $\Delta m_{41}^2 < 0$ )

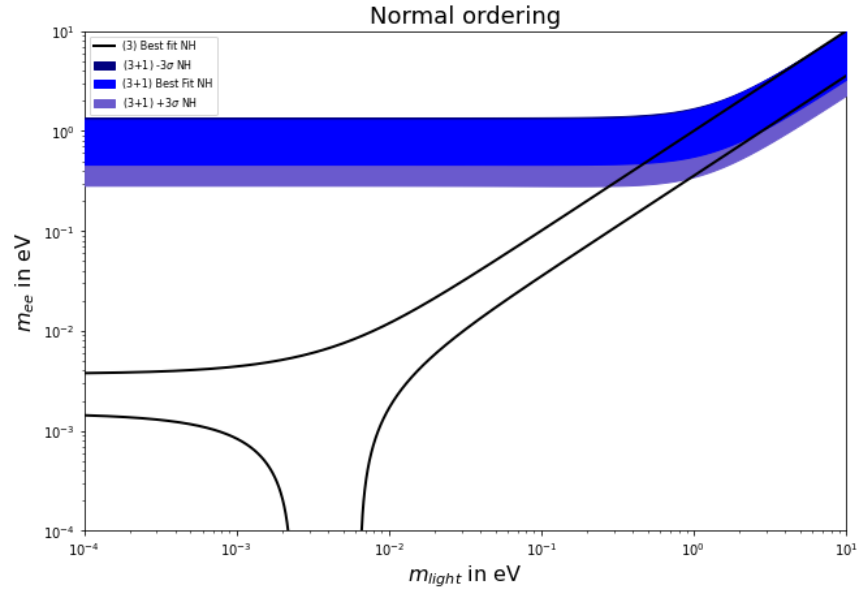
$$\text{NH: } m_2 = \sqrt{m_1^2 + \Delta m_{solar}^2} \quad m_3 = \sqrt{m_1^2 + \Delta m_A^2} \quad m_4 = \sqrt{m_1^2 - \Delta m_{\text{LSND}}^2}$$

$$\text{IH: } m_2 = \sqrt{m_3^2 + \Delta m_{solar}^2 + \Delta m_A^2} \quad m_1 = \sqrt{m_3^2 + \Delta m_A^2} \quad m_4 = \sqrt{m_3^2 + \Delta m_A^2 - \Delta m_{\text{LSND}}^2}$$

$$\text{where } \Delta m_{solar}^2 = \Delta m_{21}^2, \Delta m_A^2 = \Delta m_{31}^2 \text{ and, } \Delta m_{\text{LSND}}^2 = \Delta m_{41}^2$$

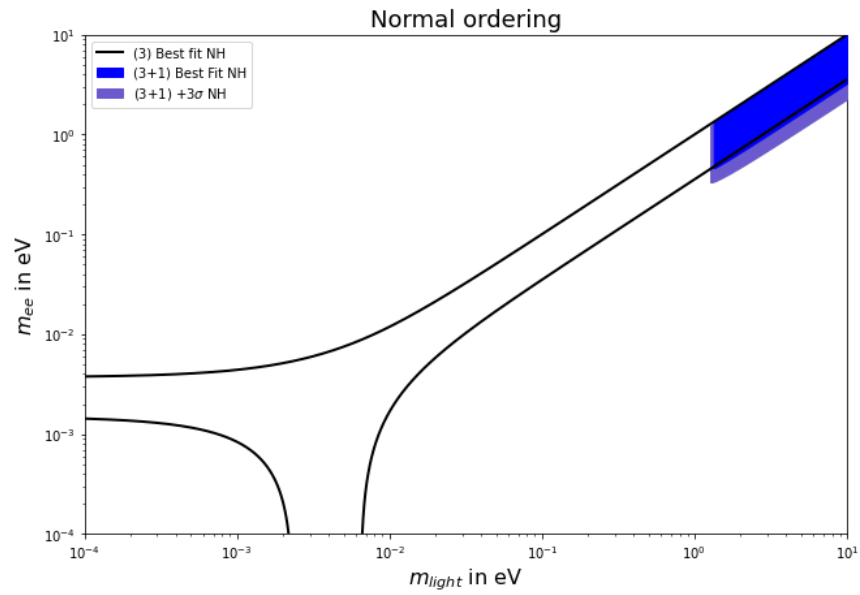
$$\Delta m_{31}^2 > 0 \text{ for NH and } \Delta m_{31}^2 < 0 \text{ for IH} \quad (12)$$

176 **9.3 Light sterile neutrino (NHS)**



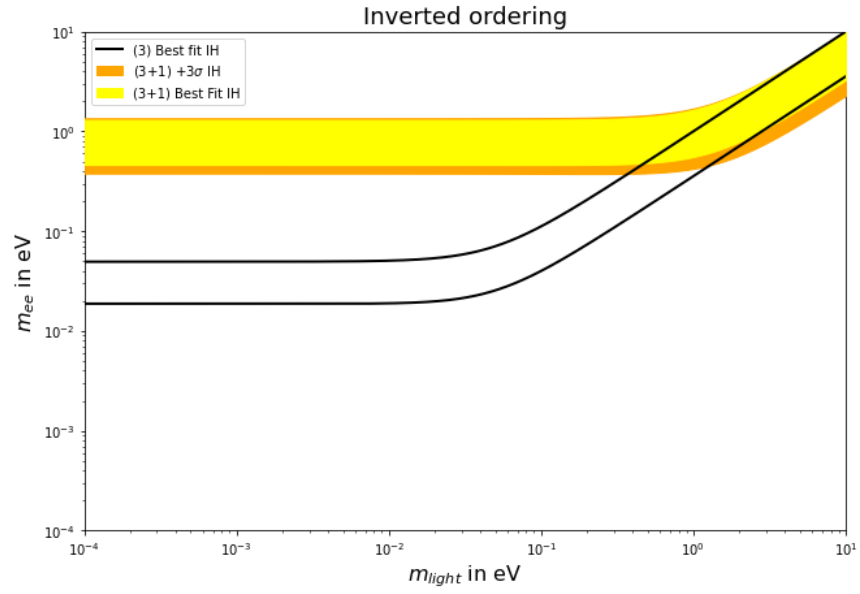
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178 **9.4 (active) Light sterile neutrino (NHS)**



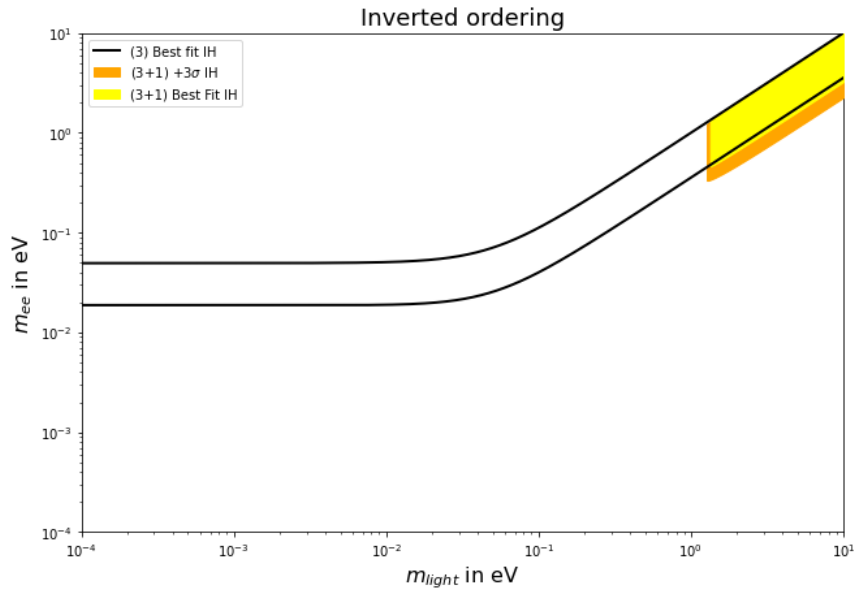
179

180 **9.5 Light sterile neutrino (IHS)**



181

182 **9.6 (active) Light sterile neutrino (IHS)**



183

## 10 2 sterile neutrinos

Again, following the same pattern explained for 1 sterile neutrino case, we can expand  $\langle m_{ee} \rangle$  for 2 sterile neutrinos, again, including the  $s_{15}^2$  and  $m_5$  terms.

$$\langle m_{ee} \rangle = \left| \sum_{i=1}^5 U_{ei}^2 m_i \right| = \left| \left| m_{ee}^{(1)} \right| + \left| m_{ee}^{(2)} \right| e^{2i\alpha} + \left| m_{ee}^{(3)} \right| e^{2i\beta} + \left| m_{ee}^{(4)} \right| e^{2i\gamma} + \left| m_{ee}^{(5)} \right| e^{2i\delta} \right| \quad (13)$$

$$\langle m_{ee} \rangle = \left| c_{12}^2 c_{13}^2 c_{14}^2 c_{15}^2 m_1 + s_{12}^2 c_{13}^2 c_{14}^2 c_{15}^2 m_2 e^{i\phi_{12}} + s_{13}^2 c_{14}^2 c_{15}^2 m_3 e^{i\phi_{13}} + s_{14}^2 c_{15}^2 m_4 e^{i\phi_{14}} + s_{15}^2 m_5 e^{i\phi_{15}} \right| \quad (14)$$

And again, similar to the equation of 4 neutrinos, if we set  $\theta_{15} = 0$  all  $s_{15}^2$  become 0, and all  $c_{15}^2$  become 1 and we get the 4 neutrinos equation back.

Since we are including a new mass  $m_5$ , we have to put it in terms of  $m_{light}$ , for that, we use the experimental value  $|\Delta m_{51}^2|$  (named  $\Delta m_{New}^2$  on this note.), and since is an absolute value, we have to consider again the different hierarchies SSN, SSI, SNSa, SISa, SNSb, SISb, NSS, and ISS.

Hierarchy name	Hierarchy label	Mass ordering
2 Heavy steriles (Normal hierarchy)	SSN	$m_5 > m_4 > m_3 > m_2 > m_1$
2 Heavy steriles (Inverted hierarchy)	SSI	$m_5 > m_4 > m_2 > m_1 > m_3$
Heavy $m_5$ light $m_4$ (Normal hierarchy)	SNS[m5]	$m_5 > m_3 > m_2 > m_1 > m_4$
Heavy $m_5$ light $m_4$ (Inverted hierarchy)	SIS[m5]	$m_5 > m_2 > m_1 > m_3 > m_4$
Heavy $m_4$ light $m_5$ (Normal hierarchy)	SNS[m4]	$m_4 > m_3 > m_2 > m_1 > m_5$
Heavy $m_4$ light $m_5$ (Inverted hierarchy)	SIS[m4]	$m_4 > m_2 > m_1 > m_3 > m_5$
2 Light steriles (Normal hierarchy)	NSS	$m_3 > m_2 > m_1 > m_5 > m_4$
2 Light steriles (Inverted hierarchy)	ISS	$m_2 > m_1 > m_3 > m_5 > m_4$

2 Heavy sterile neutrinos ( $\Delta m_{41}^2 > 0$ , and  $\Delta m_{51}^2 > 0$ )

$$\begin{aligned}
\text{NH: } m_2 &= \sqrt{m_1^2 + \Delta m_{solar}^2} \quad m_3 = \sqrt{m_1^2 + \Delta m_A^2} \\
m_4 &= \sqrt{m_1^2 + \Delta m_{LSND}^2} \quad m_5 = \sqrt{m_1^2 + \Delta m_{New}^2} \\
\\ 
\text{IH: } m_2 &= \sqrt{m_3^2 + \Delta m_{solar}^2 + \Delta m_A^2} \quad m_1 = \sqrt{m_3^2 + \Delta m_A^2} \\
m_4 &= \sqrt{m_3^2 + \Delta m_{LSND}^2 + \Delta m_A^2} \quad m_5 = \sqrt{m_3^2 + \Delta m_{New}^2 + \Delta m_A^2}
\end{aligned} \tag{15}$$

where  $\Delta m_{solar}^2 = \Delta m_{21}^2$ ,  $\Delta m_A^2 = \Delta m_{31}^2$  and,  $\Delta m_{LSND}^2 = \Delta m_{41}^2$

$\Delta m_{31}^2 > 0$  for NH and  $\Delta m_{31}^2 < 0$  for IH

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198

$m_5$  Heavy  $m_4$  Light neutrinos ( $\Delta m_{41}^2 < 0$ , and  $\Delta m_{51}^2 > 0$ )

$$\begin{aligned}
\text{NH: } m_1 &= \sqrt{m_4^2 + \Delta m_{LSND}^2} \quad m_2 = \sqrt{m_4^2 + \Delta m_{LSND}^2 + \Delta m_{solar}^2} \\
m_3 &= \sqrt{m_4^2 + \Delta m_{LSND}^2 + \Delta m_A^2} \quad m_5 = \sqrt{m_4^2 + \Delta m_{LSND}^2 + \Delta m_{New}^2} \\
\\ 
\text{IH: } m_1 &= \sqrt{m_4^2 + \Delta m_{LSND}^2} \quad m_2 = \sqrt{m_4^2 + \Delta m_{LSND}^2 + \Delta m_{solar}^2} \\
m_3 &= \sqrt{m_4^2 + \Delta m_{LSND}^2 - \Delta m_A^2} \quad m_5 = \sqrt{m_4^2 + \Delta m_{LSND}^2 + \Delta m_{New}^2}
\end{aligned} \tag{16}$$

where  $\Delta m_{solar}^2 = \Delta m_{21}^2$ ,  $\Delta m_A^2 = \Delta m_{31}^2$  and,  $\Delta m_{LSND}^2 = \Delta m_{41}^2$

$\Delta m_{31}^2 > 0$  for NH and  $\Delta m_{31}^2 < 0$  for IH

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(active)  $m_5$  Heavy  $m_4$  Light neutrinos ( $\Delta m_{41}^2 < 0$ , and  $\Delta m_{51}^2 > 0$ )

$$\begin{aligned}
\text{NH: } m_2 &= \sqrt{m_1^2 + \Delta m_{solar}^2} \quad m_3 = \sqrt{m_1^2 + \Delta m_A^2} \\
m_4 &= \sqrt{m_1^2 - \Delta m_{LSND}^2} \quad m_5 = \sqrt{m_1^2 + \Delta m_{New}^2} \\
\text{IH: } m_2 &= \sqrt{m_3^2 + \Delta m_{solar}^2 + \Delta m_A^2} \quad m_1 = \sqrt{m_3^2 + \Delta m_A^2} \\
m_4 &= \sqrt{m_3^2 + \Delta m_{LSND}^2 + \Delta m_A^2} \quad m_5 = \sqrt{m_3^2 + \Delta m_{New}^2 + \Delta m_A^2}
\end{aligned} \tag{17}$$

where  $\Delta m_{solar}^2 = \Delta m_{21}^2$ ,  $\Delta m_A^2 = \Delta m_{31}^2$  and,  $\Delta m_{LSND}^2 = \Delta m_{41}^2$

$\Delta m_{31}^2 > 0$  for NH and  $\Delta m_{31}^2 < 0$  for IH

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202

$m_4$  Heavy  $m_5$  Light neutrinos ( $\Delta m_{41}^2 > 0$ , and  $\Delta m_{51}^2 < 0$ )

$$\begin{aligned}
\text{NH: } m_1 &= \sqrt{m_5^2 + \Delta m_{New}^2} \quad m_2 = \sqrt{m_5^2 + \Delta m_{New}^2 + \Delta m_{solar}^2} \\
m_3 &= \sqrt{m_5^2 + \Delta m_{New}^2 + \Delta m_A^2} \quad m_4 = \sqrt{m_5^2 + \Delta m_{LSND}^2 + \Delta m_{New}^2} \\
\text{IH: } m_1 &= \sqrt{m_5^2 + \Delta m_{New}^2} \quad m_2 = \sqrt{m_5^2 + \Delta m_{New}^2 + \Delta m_{solar}^2} \\
m_3 &= \sqrt{m_5^2 + \Delta m_{New}^2 - \Delta m_A^2} \quad m_4 = \sqrt{m_5^2 + \Delta m_{LSND}^2 + \Delta m_{New}^2}
\end{aligned} \tag{18}$$

where  $\Delta m_{solar}^2 = \Delta m_{21}^2$ ,  $\Delta m_A^2 = \Delta m_{31}^2$  and,  $\Delta m_{LSND}^2 = \Delta m_{41}^2$

$\Delta m_{31}^2 > 0$  for NH and  $\Delta m_{31}^2 < 0$  for IH

203

204

(active)  $m_4$  Heavy  $m_5$  Light neutrinos ( $\Delta m_{41}^2 > 0$ , and  $\Delta m_{51}^2 < 0$ )

$$\begin{aligned}
\text{NH: } m_2 &= \sqrt{m_1^2 + \Delta m_{solar}^2} \quad m_3 = \sqrt{m_1^2 + \Delta m_A^2} \\
m_4 &= \sqrt{m_1^2 + \Delta m_{LSND}^2} \quad m_5 = \sqrt{m_1^2 - \Delta m_{New}^2} \\
\text{IH: } m_2 &= \sqrt{m_3^2 + \Delta m_{solar}^2 + \Delta m_A^2} \quad m_1 = \sqrt{m_3^2 + \Delta m_A^2} \\
m_4 &= \sqrt{m_3^2 + \Delta m_{LSND}^2 + \Delta m_A^2} \quad m_5 = \sqrt{m_3^2 - \Delta m_{New}^2 + \Delta m_A^2}
\end{aligned} \tag{19}$$

where  $\Delta m_{solar}^2 = \Delta m_{21}^2$ ,  $\Delta m_A^2 = \Delta m_{31}^2$  and,  $\Delta m_{LSND}^2 = \Delta m_{41}^2$

$\Delta m_{31}^2 > 0$  for NH and  $\Delta m_{31}^2 < 0$  for IH

205

206

2 Light sterile neutrinos ( $\Delta m_{41}^2 < 0$ , and  $\Delta m_{51}^2 < 0$ )

$$\begin{aligned}
\text{NH: } m_1 &= \sqrt{m_4^2 + \Delta m_{LSND}^2} \quad m_2 = \sqrt{m_4^2 + \Delta m_{LSND}^2 + \Delta m_{solar}^2} \\
m_3 &= \sqrt{m_4^2 + \Delta m_{LSND}^2 + \Delta m_A^2} \quad m_5 = \sqrt{m_4^2 + \Delta m_{LSND}^2 - \Delta m_{New}^2} \\
\text{IH: } m_1 &= \sqrt{m_4^2 + \Delta m_{LSND}^2} \quad m_2 = \sqrt{m_4^2 + \Delta m_{LSND}^2 + \Delta m_{solar}^2} \\
m_3 &= \sqrt{m_4^2 + \Delta m_{LSND}^2 - \Delta m_A^2} \quad m_5 = \sqrt{m_4^2 + \Delta m_{LSND}^2 - \Delta m_{New}^2}
\end{aligned} \tag{20}$$

where  $\Delta m_{solar}^2 = \Delta m_{21}^2$ ,  $\Delta m_A^2 = \Delta m_{31}^2$  and,  $\Delta m_{LSND}^2 = \Delta m_{41}^2$

$\Delta m_{31}^2 > 0$  for NH and  $\Delta m_{31}^2 < 0$  for IH

207

208

(active) 2 Light sterile neutrinos ( $\Delta m_{41}^2 < 0$ , and  $\Delta m_{51}^2 < 0$ )

$$\begin{aligned}
\text{NH: } m_2 &= \sqrt{m_1^2 + \Delta m_{solar}^2} \quad m_3 = \sqrt{m_1^2 + \Delta m_A^2} \\
m_4 &= \sqrt{m_1^2 - \Delta m_{\text{LSND}}^2} \quad m_5 = \sqrt{m_1^2 - \Delta m_{New}^2} \\
\\ 
\text{IH: } m_2 &= \sqrt{m_3^2 + \Delta m_{solar}^2 + \Delta m_A^2} \quad m_1 = \sqrt{m_3^2 + \Delta m_A^2} \\
m_4 &= \sqrt{m_3^2 - \Delta m_{\text{LSND}}^2 + \Delta m_A^2} \quad m_5 = \sqrt{m_3^2 - \Delta m_{New}^2 + \Delta m_A^2}
\end{aligned} \tag{21}$$

where  $\Delta m_{solar}^2 = \Delta m_{21}^2$ ,  $\Delta m_A^2 = \Delta m_{31}^2$  and,  $\Delta m_{\text{LSND}}^2 = \Delta m_{41}^2$

$\Delta m_{31}^2 > 0$  for NH and  $\Delta m_{31}^2 < 0$  for IH

209

210

## 211 11 Data used for 5 masses

212

Parameter	Best Fit	$3\sigma$
$\Delta m_{21}^2$ ( $10^{-5} \text{ eV}^2$ )	7.53	8.07-6.99
$ \Delta m_{31}^2 $ ( $10^{-3} \text{ eV}^2$ )	2.5283	2.5631-2.4935
$ \Delta m_{41}^2 $ ( $\text{eV}^2$ )	1.78	2.01-1.61
$ \Delta m_{51}^2 $ ( $\text{eV}^2$ )	0.87	0.97-0.77
$s_{12}^2$	0.307	0.346-0.268
$s_{13}^2$	0.022	0.0241-0.0199
$s_{14}^2$	0.023	0.04-0.006
$s_{15}^2$	0.02	0.035-0.005



## 213 12 Step by step to make it work for 3+2 neu- 214 trinos

215 Basically to plot the graph for 2 sterile neutrinos we use the same approach as  
216 the previous plots, however, as we have new mass terms, then, the boundaries  
217 of the plot change again, fortunately, we just have to use the same optimization  
218 technique but including a  $\phi_{15}$  angle.

$$219 \quad \frac{\partial(\langle m_{ee} \rangle)}{\partial(\phi_{12})} = 0 \text{ if } \phi_{12} = (0, \pi)$$

$$221 \quad \frac{\partial(\langle m_{ee} \rangle)}{\partial(\phi_{13})} = 0 \text{ if } \phi_{13} = (0, \pi)$$

$$222 \quad \frac{\partial(\langle m_{ee} \rangle)}{\partial(\phi_{14})} = 0 \text{ if } \phi_{14} = (0, \pi)$$

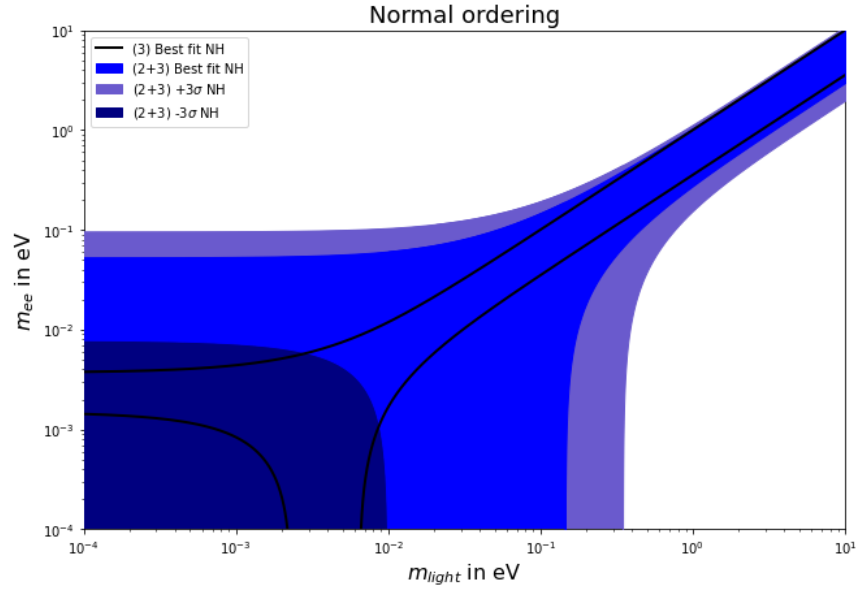
$$223 \quad \frac{\partial(\langle m_{ee} \rangle)}{\partial(\phi_{15})} = 0 \text{ if } \phi_{15} = (0, \pi)$$

224  
225 upper bound at  $\phi_{12}=0, \phi_{13}=0, \phi_{14}=0$ , and  $\phi_{15}=0$

226 lower bound at  $\phi_{12} = \pi, \phi_{13} = \pi, \phi_{14} = \pi$ , and  $\phi_{15} = \pi$

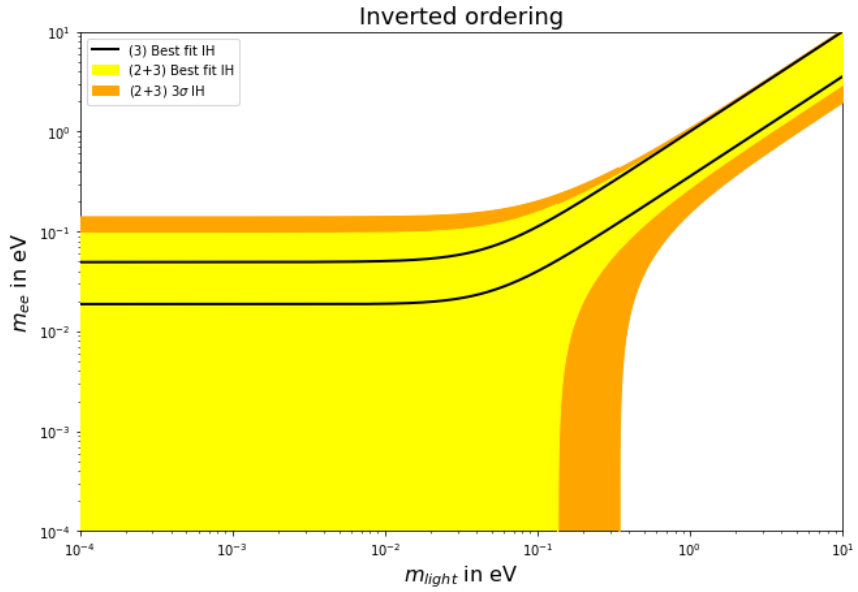
227 lower inferior bound (only for SSN case) at  $\phi_{12} = 0, \phi_{13} = 0, \phi_{14} = \pi$ , and  
228  $\phi_{15} = 0$

229 **12.1 2 Heavy steriles (SSN)**



230

231 **12.2 2 Heavy steriles (SSI)**

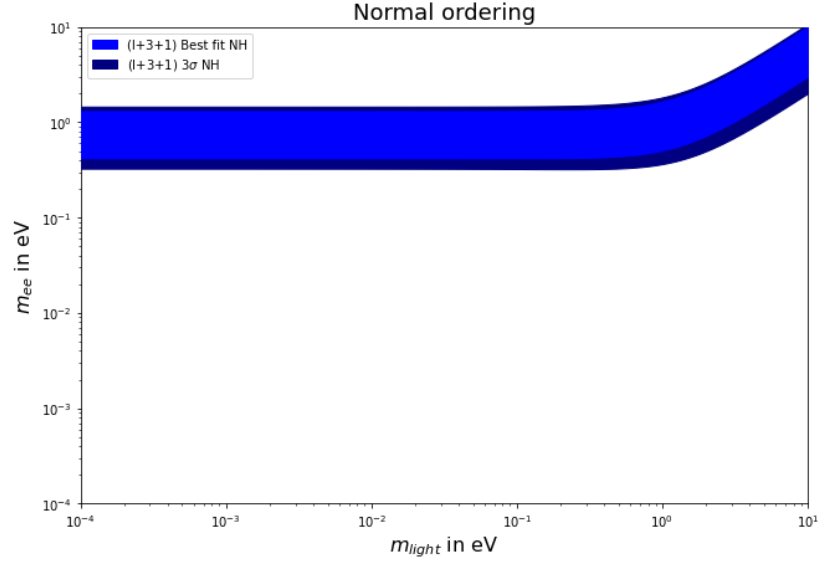


232

233 And like the 1 sterile neutrino case, even when the lightest mass for these 3  
 234 hierarchies are  $m_4$  and  $m_5$ , for experimental purposes, we also include separated

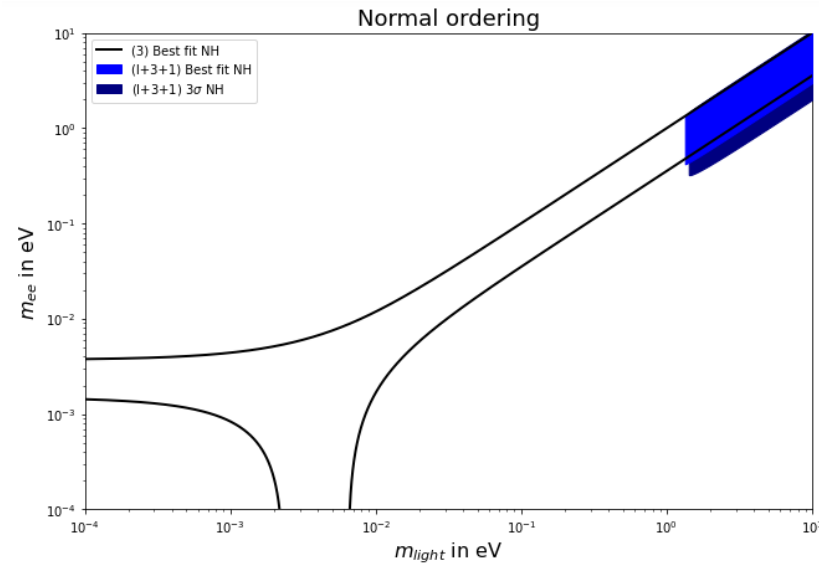
235 plots in terms of  $m_1$  for NH and  $m_3$  for IH (Details of how the plots were  
 236 made in <https://github.com/FranciscoTapia61199/Sterile-neutrinos> and go to  
 237 "2 steriles" folder).

### 238 12.3 Heavy $m_5$ light $m_4$ (SNSa)



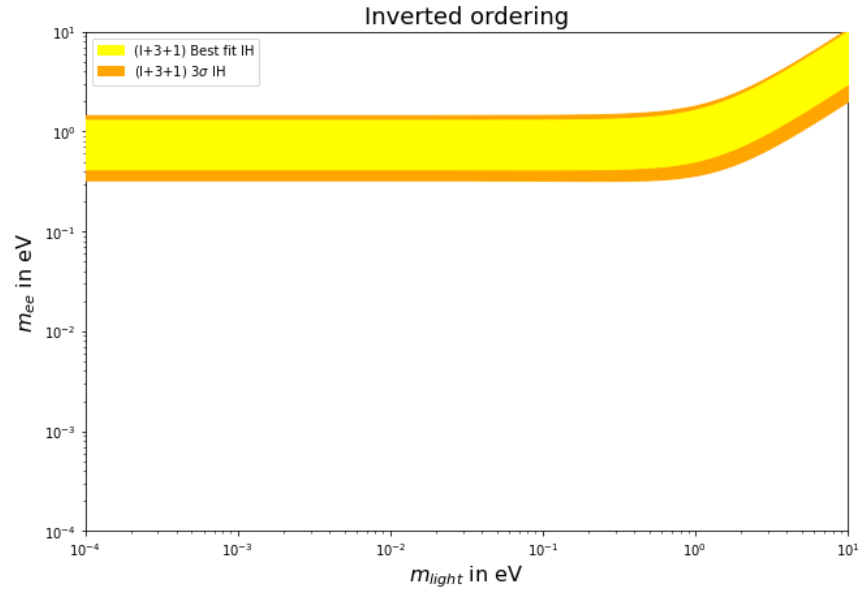
239

### 240 12.4 (active) Heavy $m_5$ light $m_4$ (SNSa)



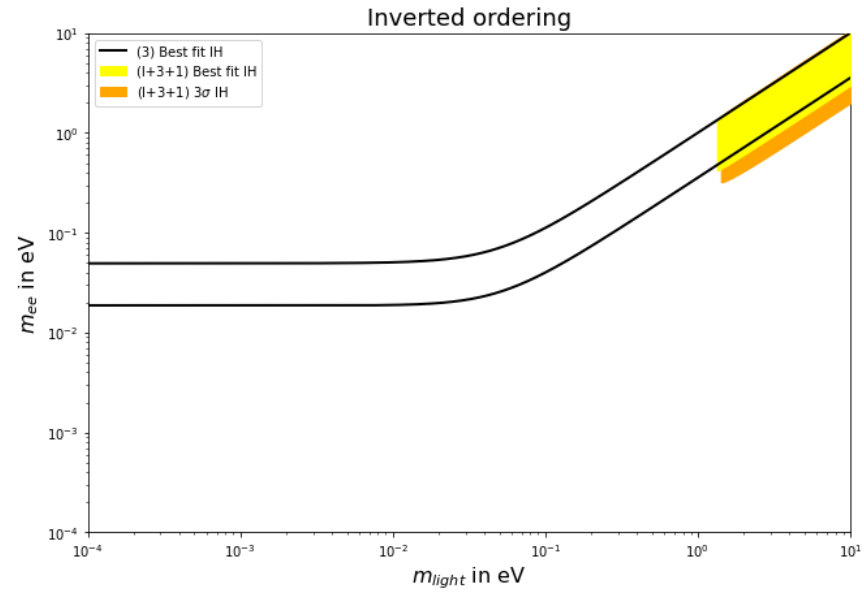
241

242 **12.5 Heavy  $m_5$  light  $m_4$  (SISa)**



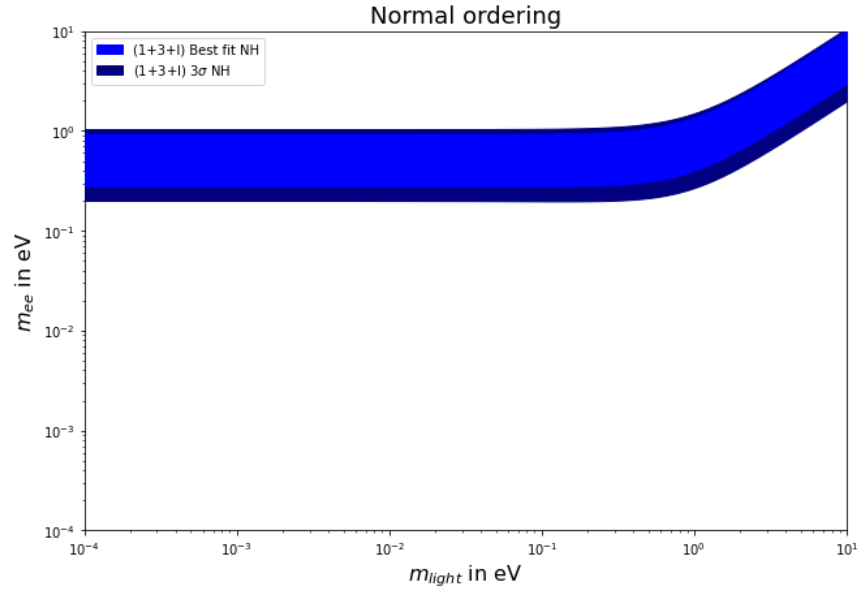
243

244 **12.6 (active) Heavy  $m_5$  light  $m_4$  (SISa)**



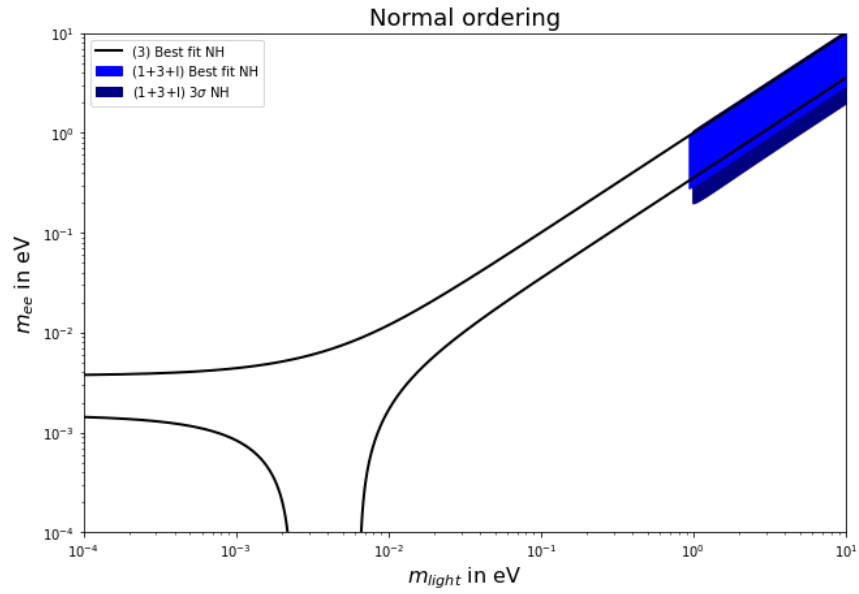
245

246 **12.7 Heavy  $m_4$  light  $m_5$  (SNSb)**



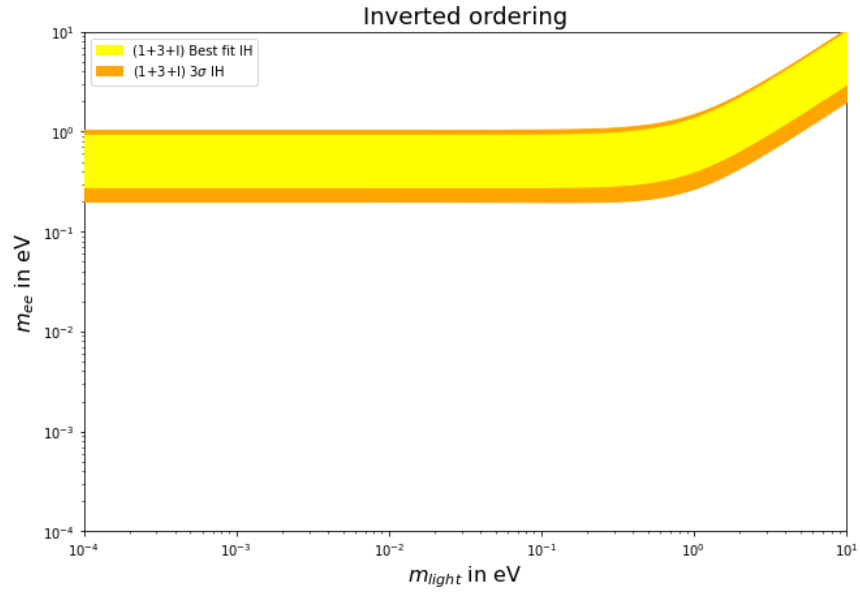
247

248 **12.8 (active) Heavy  $m_4$  light  $m_5$  (SNSb)**



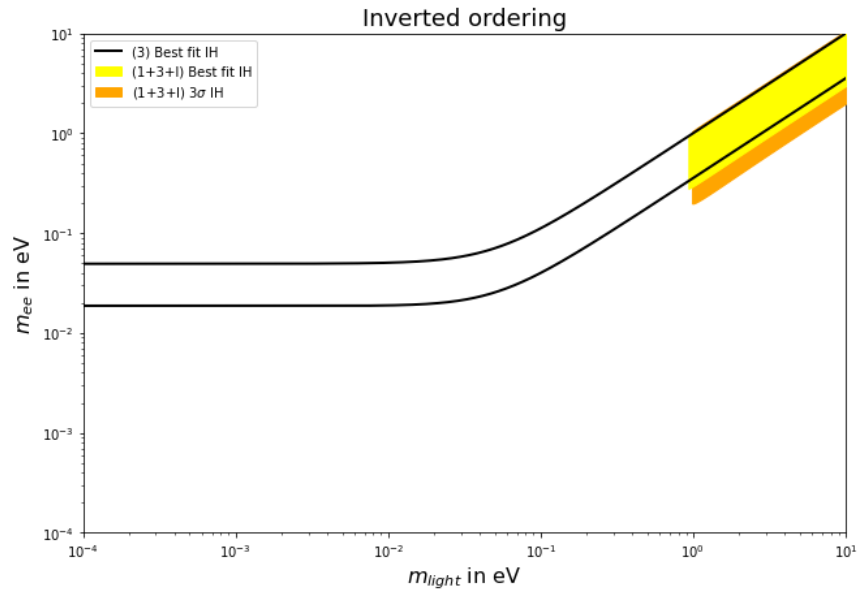
249

250 **12.9 Heavy  $m_4$  light  $m_5$  (SISb)**



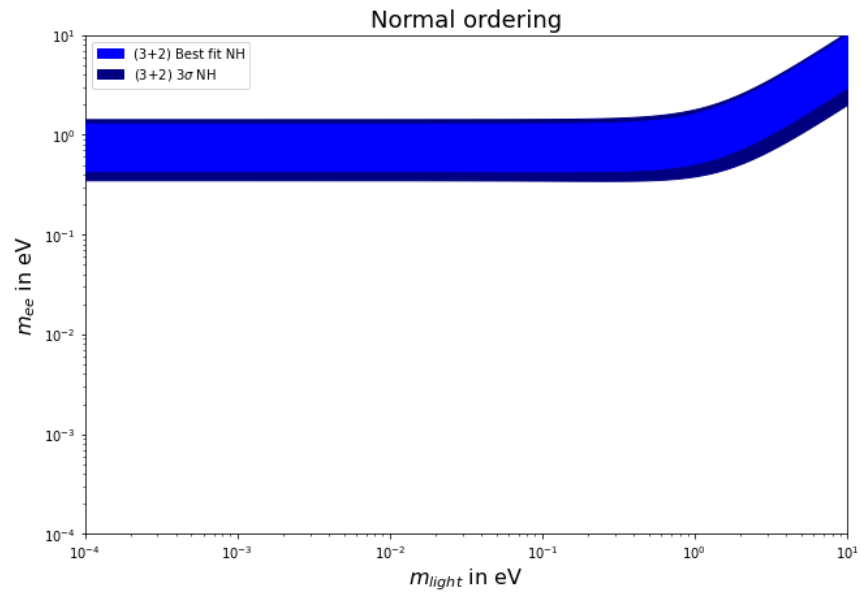
251

252 **12.10 (active) Heavy  $m_4$  light  $m_5$  (SISb)**

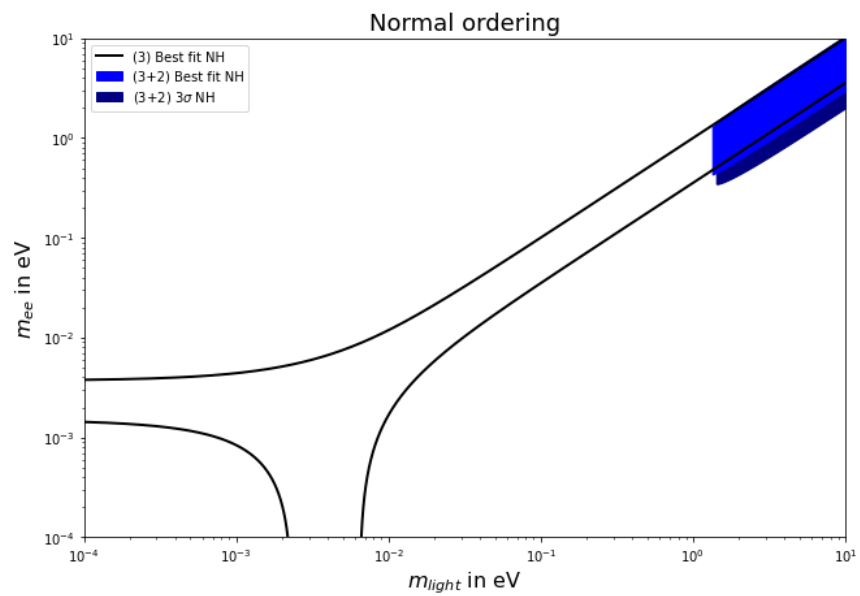


253

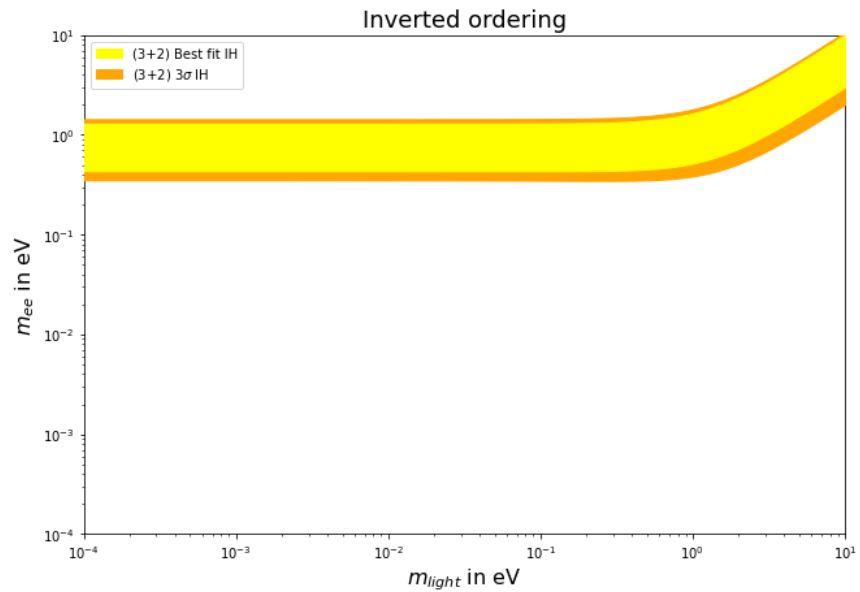
254 **12.11 2 Light steriles (NSS)**



256 **12.12 (active) 2 Light steriles (NSS)**

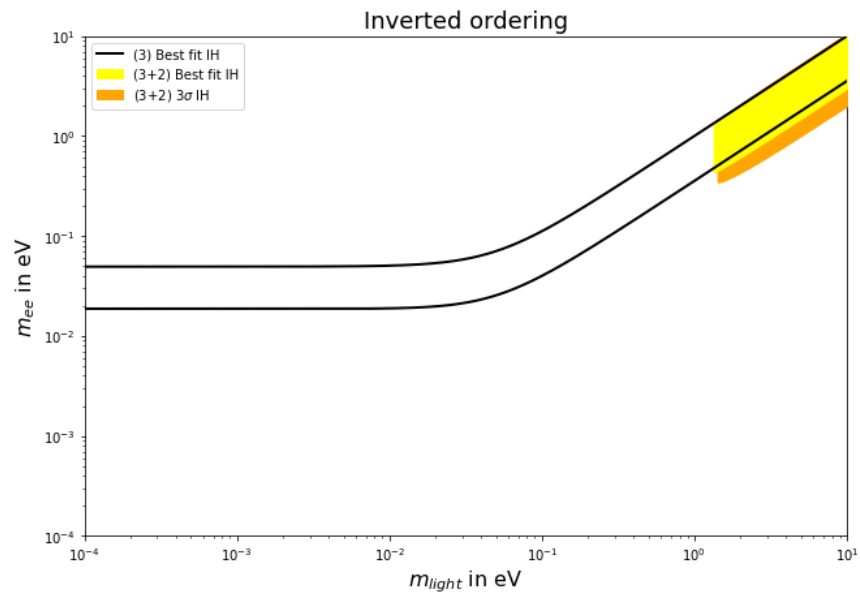


258 **12.13 2 Light steriles (ISS)**



259

260 **12.14 (active) 2 Light steriles (ISS)**



261



### 13 Dynamic plots and how to make them

In addition to the previously shown plots, we include also animations that shows how the plots change with respect to  $\theta_{14}$  for the case of 1 sterile neutrino and  $\theta_{15}$  for the case of 2 sterile neutrinos, (For simplicity reasons, the animations made are first, the ones that transforms the 3 neutrinos model to the 3+1 neutrino model, and then the ones that transforms the 3+1 neutrino models to 3+2 neutrino models.) the purpose of why making the dynamic plots in function of those angles is because that, if those angles become zero, we get back the previous plot (for example, if  $\theta_{14} = 0$  we get the 3 neutrino plot, and if  $\theta_{15} = 0$  we get back the 3+1 neutrino plot.)

$$\langle m_{ee} \rangle = |c_{12}^2 c_{13}^2 c_{14}^2 c_{15}^2 m_1 + s_{12}^2 c_{13}^2 c_{14}^2 c_{15}^2 m_2 e^{i\phi_{12}} + s_{13}^2 c_{14}^2 c_{15}^2 m_3 e^{i\phi_{13}} + s_{14}^2 c_{15}^2 m_4 e^{i\phi_{14}} + s_{15}^2 m_5 e^{i\phi_{15}}| \quad (22)$$

if  $\theta_{15} = 0$

$$\langle m_{ee} \rangle = |c_{12}^2 c_{13}^2 c_{14}^2 m_1 + s_{12}^2 c_{13}^2 c_{14}^2 m_2 e^{i\phi_{12}} + s_{13}^2 c_{14}^2 m_3 e^{i\phi_{13}} + s_{14}^2 m_4 e^{i\phi_{14}}| \quad (23)$$

if  $\theta_{14} = 0$

$$\langle m_{ee} \rangle = |c_{12}^2 c_{13}^2 m_1 + s_{12}^2 c_{13}^2 m_2 e^{i\phi_{12}} + s_{13}^2 m_3 e^{i\phi_{13}}| \quad (24)$$

The first step to make an animation is choose a "time dependent" value (more than time dependent, in this case would be angle dependent, but for technical reasons we assume that those are angles that change with time.) and as showed in the previous equations 19-21, a good choice are  $\theta_{14}$  and  $\theta_{15}$ .

Next, we have to see how those angles will change with respect of time;  $\theta_{14}$  and  $\theta_{15}$  have fix values, but we are also interested in see how the animation change with respect of those angles, so, a clever way to solve this could be to make each frame of the animation a fractional part of those angles (or in this case, the  $s_{ij}^2$  and  $c_{ij}^2$ ), we can represent that like this:

$$s_{ijframe}^2 = (\frac{currentframe}{totalframes - 1}) s_{ij}^2 \quad (25)$$

$$c_{ijframe}^2 = 1 - s_{ijframe}^2 \quad (26)$$

For example, for these animations, the total frames used are 100, and since we want to start on a point where we do not have the  $\theta_{ij}$ , the first frame would be zero, (that is why in the equation 25 we substract 1 from the total frames, to include the zero frame on the top).

Now that we know how the animation will change, another important step is to preserve the boundaries of each graph to include all possible angles, but at the same time keep it consistent with how the animation change frame by frame, considering the fact that the boundaries of the plots could change dramatically with the change of  $\theta_{ij}$ , it is challenging to achieve that. Surprisingly in the case of all IH plots and NH that are not SNH and SSN, the boundaries actually preserved and if the angles of  $\theta_{ij}$  change, the boundaries were still able to contain and include all possible combinations (Note, in fact the inferior boundaries actually changed a little bit, however since the boundaries were beyond the graph y-scale, it was not necessary to calculate them frame by frame.)

However, the most challenging ones to actually be able to do where the SNH model, due that the inferior left boundary changes, and the transition from one model to the new model was everything but smooth and precise. Fortunately, there was a method that works not only for 4, 5, but potentially for n neutrinos.

## 14 Boundary generator

To locate and calculate the boudaries for the SNH model, the methodology used was the following, first was use the equations 19-21 (Depending on which model was required to transform) and separate each mass term, but instead of making them dependent to the lightest mass, those were made dependent to the current frame in where there angles are, and for the lightest mass, we pick the smallest value possible for the plot, in this case 0.0001, as it shows here:

### 14.1 3 neutrinos to 3+1 neutrinos

$$s_{14frame}^2 = \left( \frac{currentframe}{totalframes - 1} \right) s_{14}^2 \quad (27)$$

$$c_{14frame}^2 = 1 - s_{14frame}^2 \quad (28)$$

$$\begin{aligned} m_1 &= 0.0001 \\ m_2 &= \sqrt{m_1^2 + \Delta m_{solar}^2} \\ m_3 &= \sqrt{m_1^2 + \Delta m_A^2} \\ m_4 &= \sqrt{m_1^2 + \Delta m_{LSND}^2} \end{aligned}$$

$$\begin{aligned} m_{ee}^{(1)} &= c_{12}^2 c_{13}^2 c_{14}^2 m_1 \\ m_{ee}^{(2)} &= s_{12}^2 c_{13}^2 c_{14}^2 m_2 \\ m_{ee}^{(3)} &= s_{13}^2 c_{14}^2 m_3 \\ m_{ee}^{(4)} &= s_{14}^2 m_4 \end{aligned}$$

(Note, we are not including the  $e^{i\phi_{ij}}$  because we know from previous pages that the critical values are 0 and  $\pi$ , but that is the equivalent to just add or subtract the  $m_{ee}^{(i)}$  in different combinations that we will show now.)

Since we fix the lightest mass to be 0.001, that means those mass terms are in terms of the current frame, but we are not interested in the mass terms by themselves, but interested in how those mass terms add and subtract on this way.

$$\begin{aligned}
m0 &= |m_{ee}^{(1)} + m_{ee}^{(2)} + m_{ee}^{(3)} + m_{ee}^{(4)}| \\
m1 &= |m_{ee}^{(1)} + m_{ee}^{(2)} + m_{ee}^{(3)} - m_{ee}^{(4)}| \\
m2 &= |m_{ee}^{(1)} + m_{ee}^{(2)} - m_{ee}^{(3)} + m_{ee}^{(4)}| \\
m3 &= |m_{ee}^{(1)} + m_{ee}^{(2)} - m_{ee}^{(3)} - m_{ee}^{(4)}| \\
m4 &= |m_{ee}^{(1)} - m_{ee}^{(2)} + m_{ee}^{(3)} + m_{ee}^{(4)}| \\
m5 &= |m_{ee}^{(1)} - m_{ee}^{(2)} + m_{ee}^{(3)} - m_{ee}^{(4)}| \\
m6 &= |m_{ee}^{(1)} - m_{ee}^{(2)} - m_{ee}^{(3)} + m_{ee}^{(4)}| \\
m7 &= |m_{ee}^{(1)} - m_{ee}^{(2)} - m_{ee}^{(3)} - m_{ee}^{(4)}|
\end{aligned}$$

Now that we have this, what we actually want to know is which combination of masses will give the minimum value for each frame, and by looking at which combination of masses gives us the minimum value for certain specific frame, we can then assign specific boundaries for specific frames, making sure the animation is smooth and contains all possible angle combinations (For more details on how this work, here is attached a code named "Boundary generator" at <https://github.com/FranciscoTapia61199/Sterile-neutrinos> that explains how it works.)

These method was used exclusively to calculate the lower left boundary and how it changes with respect to the frame for the SNH model, however, it can be applied for all models and potentially for n neutrinos, but due that those other models keep the same boundaries even with the change of frames, this was not necessary.

And in this repository <https://github.com/FranciscoTapia61199/Sterile-neutrinos> at the animation section are the results of how the plots change with respect of angle (Note, all plots regardless of what is the actual lightest mass are for NH in terms of  $m_1$  and for IH in terms of  $m_3$  for experimental purposes, and all plots runs at 15 fps.)

## 15 Remarkable Observations

The evolution of the plots in comparison to the original 3 neutrino plots is...interesting, the first remarkable observation to made is that the plot, depending on which hierarchy is used, is extremely sensitive, small changes to the mixing angles and mass values would result into changes in shape of the whole graph that are extremely visible for the transition from the 3 neutrino model to the heavy sterile models, it could be with 1 or 2 steriles, the changes are visible for NH and IH cases, each time we are including a heavy mass term, it is hard to distinguish if a neutrino mass is from a NH or IH, but even that is almost indistinguishable, due that that region is located on a low energy area ( $10^{-2}$  eV -  $10^{-5}$  eV) is extremely hard to detect with actual equipment.

Thinks are also not looking well for the case where the 1 sterile model transforms to 2 sterile model, if a big change is visible from 3 neutrinos to 1 sterile neutrino model, the inclusion of another sterile makes just small changes to the plot, basically detection regions are the same for NH and IH.

For the case where 1 sterile neutrino was heavy and the other one light, we see little to no visible change in the graph, and also the results are very similar with the case of 2 light sterile neutrinos, regardless of the hierarchy used.

The main observation of this project is that, the most accurate model to predict the mass of the neutrino is the 3 neutrino model due that NH and IH are enough separated to be distinguished and the fact that the hierarchy that is more likely to detect is IH, the moment we are sure of the actual mass value of the active neutrinos, then, it is expected to formulate better techniques to detect the sterile neutrino.