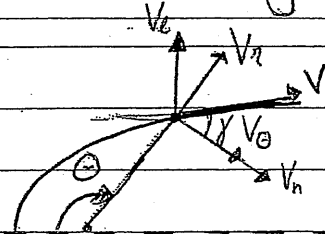


Question 2: Velocity components

august 29'02

$$a) \quad r = \frac{a(1-e^2)}{1+e\cos\theta} = \frac{H^2/\mu}{1+e\cos\theta}$$

$$\begin{aligned} \Rightarrow \dot{r} &= \frac{H^2/\mu}{(1+e\cos\theta)^2} e \sin\theta \dot{\theta} \cdot \frac{r^2}{r^2} \\ &= r^2 \frac{H^2/\mu}{(1+e\cos\theta)^2} e \sin\theta \dot{\theta} \frac{(1+e\cos\theta)^2}{H^4/\mu^2} \\ &= r^2 \frac{\mu}{H^2} e \sin\theta \dot{\theta} \\ &= \frac{\mu}{H} e \sin\theta \end{aligned}$$



$$\begin{cases} \dot{r} = V \sin\gamma \\ r\dot{\theta} = V \cos\gamma \\ H = r V \cos\gamma \end{cases}$$

$$r = \frac{H^2/\mu}{1+e\cos\theta} = \frac{H/\mu}{1+e\cos\theta} [r^2 \dot{\theta}]$$

$$\Rightarrow r\dot{\theta} = \frac{1+e\cos\theta}{H/\mu} = \frac{\mu}{H} (1+e\cos\theta)$$

$$\Rightarrow V_l = \frac{\dot{r}}{\sin(180-\theta)} = \frac{\dot{r}}{\sin\theta} = \frac{\dot{r}}{rH/\mu e} = \frac{\mu e}{H}$$

$$V_n = r\dot{\theta} + \frac{\dot{r}}{\tan(180-\theta)} = r\dot{\theta} + \frac{\dot{r}}{\tan\theta}$$

$$\Rightarrow \cancel{r\dot{\theta} + \frac{\dot{r}}{\sin\theta} \cos\theta} = \cancel{r\dot{\theta} + \frac{\mu}{H} e \cdot \frac{1}{e} \left[ \frac{r^2 \dot{\theta}}{H} - 1 \right]}$$

$$\cancel{= r\dot{\theta} + \frac{\mu}{H}}$$

$$= \frac{\mu}{H} (1+e\cos\theta) + \frac{\dot{r}}{\sin\theta} \cos\theta = \frac{\mu}{H} (1+e\cos\theta) - \frac{\mu}{H} e \cos\theta$$

$$= \frac{\mu}{H}$$

$$c) E - e \sin E = M$$

$$dE - e \cos E dE = dM$$

$$(1 - e \cos E) = \frac{dM}{dE} = \left( \frac{dE}{dM} \right)^{-1}$$

Substitution in

$$r = a(1 - e \cos E) = a \frac{dM}{dE}$$

So

$$\left( \frac{r}{a} \right) = \left( \frac{dE}{dM} \right)^{-1}$$

$$d) E = M + e \sin M + e^2 \frac{1}{2} \sin(2M)$$

$$dE = dM + e \cos M dM - \frac{1}{2} e^2 \cos(2M) dM$$

$$\frac{dE}{dM} = 1 + e \cos M - e^2 \cos(2M) + \dots$$

$$(1+y)^{-1} = 1 - y + y^2 - \dots \quad y = e \cos M - e^2 \cos(2M)$$

$$(1 + e \cos M - e^2 \cos(2M))^{-1} = 1 - e \cos M + e^2 \cos(2M) + e^2 \cos^2 M$$

1/

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Question 3 Keplers equationaug 29 '62  
aug 24 '61

a)  $E - e \sin E = n(t - T) = M$

$$r = a(1 - e \cos E)$$

$E$ : eccentric anomaly

$e$ : anomaly

$n$ : mean angular velocity  $n = \sqrt{\frac{\mu}{a^3}}$

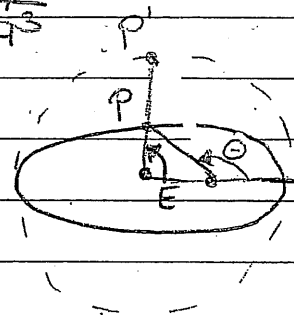
$t$ : time

$T$ : time of last perigee

$r$ : distance

$a$ : semi-major axis

$M$ : mean anomaly



$$2a$$

$$a(1+e)$$

b)  $x = E - M = e \sin E = e \sin(x + M)$

$$= e \sin x \cos M + e \cos x \sin M$$

$$= e \cos M \left( x - \frac{1}{6} x^3 + \dots \right) + e \sin M \left( 1 - \frac{1}{2} x^2 + \dots \right)$$

$$x_1 e + x_2 e^2 + x_3 e^3 + \dots = \sin M e + \cos M x_1 e^2 + O(e^3)$$

$$x_1 = \sin M$$

$$x_2 = x_1 \cos M = \sin M \cos M = \frac{1}{2} \sin(2M)$$

$$E = x + M = M + e \sin M + e^2 \frac{1}{2} \sin(2M) + O(e^3)$$

relat eff

## Question 3 Relativistic effects 29 august 02

$$a) \frac{d^2 u}{d\varphi^2} + u = \frac{\mu}{H^2}$$

$u$ : distance satellite to gravitational body  
 $u = 1/r$

$\varphi$ : angular velocity

$\mu$ : gravitational constant of body

$H$ : angular velocity

$$\frac{d^2 u}{d\varphi^2} + u = \frac{\mu}{H^2} + 3 \frac{\mu}{c^2} u^2$$

$c$  = speed of light

$$b) = \frac{\mu}{H^2} \left[ 1 + 3 \frac{H^2 u^2}{c^2} \right] = \frac{\mu}{H^2} \left[ 1 + 3 \frac{V_\varphi^2}{c^2} \right]$$

$$c^2 \gg V_\varphi^2$$

$$c) \quad u = u_0 + \alpha \frac{\mu^2}{H^4} \left[ 1 + \frac{1}{2} e^2 + \underbrace{e\varphi \sin(\varphi - \omega)}_{\text{always increasing}} - \frac{1}{6} e^2 \cos 2(\varphi - \omega) \right]$$

- increase  $u$  with constant term
- fluctuating term growing constantly with  $\varphi$
- pure oscillation with constant amplitude

d) only taking long term in account

$$u = \frac{\mu}{H^2} \left[ 1 + e \cos(\varphi - \omega) + \alpha \frac{\mu}{H^2} e \varphi \sin(\varphi - \omega) \right]$$

$$u = \frac{\mu}{H^2} \left[ 1 + e \cos(\varphi - \omega) + \beta e \varphi \sin(\varphi - \omega) \right]$$