

Gravitational Algorithm for Anomaly Detection

Javier Fumanal Idocin¹, Alfonso Induráin¹, María Minárová²,
Mikel Ferrero Jaurrieta¹, Humberto Bustince¹

3rd February, 2022

¹ Public University of Navarre

² Slovak University of Technology in Bratislava

16th International Conference on Fuzzy Set Theory and Applications
Liptovský Ján, Slovak Republic



Table of Contents

1. Gravitation Clustering
2. Anomaly Detection and Gravitational Clustering
3. Modification of the original Gravitational Clustering
4. Experimental results

Gravitational clustering

What is cluster analysis?

- maximal mutual similarity inside a cluster
- minimal similarity between clusters

Gravitational clustering algorithm (Wright, 1977)

alternating of two steps:

- (A) **Attracting** - driven by Newton gravitational law
controlling parameter: δ - maximal shift (of the fastest particle) within one time step
- (C) **Clustering**
controlling parameter: ϵ - limiting distance of particles
 - Setting:
 - initial time to zero
 - influencing parameters ϵ and $\delta, \delta < \epsilon$
 - initial mass of each particle to 1

Algorithm

- Repeating (i)-(vi) until 1 particle remains
 - (i) checking the distances of particles. If less than ϵ , unification in new one: mass = sum of masses, position in the center of gravity
 - (ii) movement driving function $\forall p_i, \forall [t, t + dt]$

$$\vec{g}_i(t) = \frac{1}{2} G \sum_{j \neq i} \frac{1}{m_j(t)} \frac{\vec{s}_j(t) - \vec{s}_i(t)}{|\vec{s}_j(t) - \vec{s}_i(t)|^3} dt^2, \quad G > 0$$

- (iii) the fastest particle: $l = \arg(\max_i \{|\vec{g}_i|(t)\})$, p_l
- (iv) current time step length computation: $|\vec{g}_l(t)| = \delta \Rightarrow dt(t)$
- (v) new position of each particle p_i : $\vec{s}_i(t + dt(t)) = \vec{s}_i(t) + \vec{g}_i(t)$
- (vi) $t \leftarrow t + dt(t)$

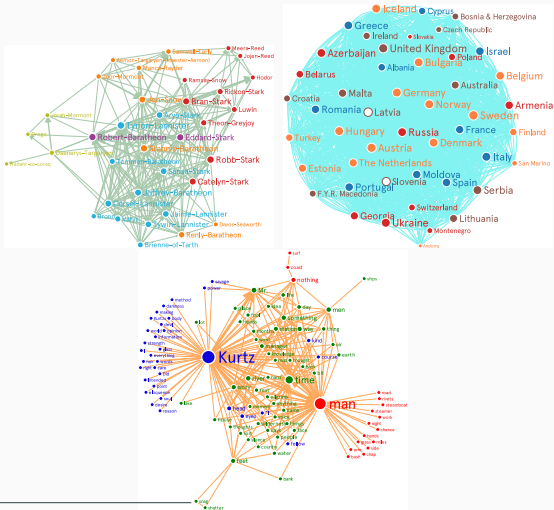
Anomaly detection and Gravitational Clustering

Borgia Clustering.

1. Different natural affinities between parties → affinity functions.
2. Balance of power → particle mass.
3. Scaling size: small communities differ from big communities → change the behaviour while communities grow.

Fumanal-Idocin, J., Alonso-Betanzos, A., Cordón, O., Bustince, H., & Minárová, M. (2020). Community detection and social network analysis based on the Italian wars of the 15th century. *Future Generation Computer Systems*, 113, 25-40.

Borgia Clustering

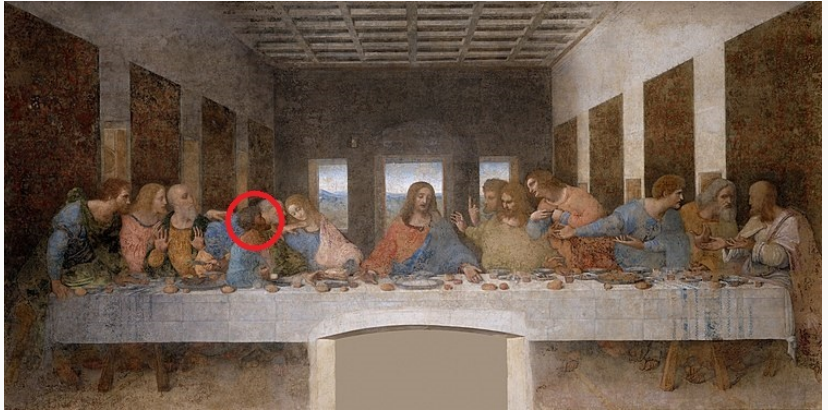


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There is always the weird one...

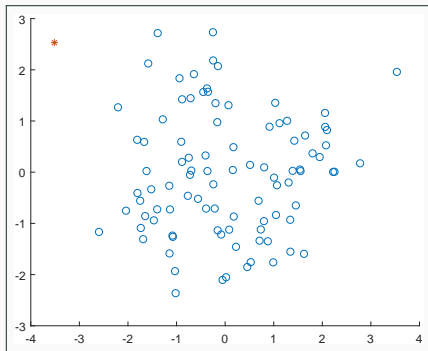


There is always the weird one...



A challenge... and an opportunity

Sometimes particles are **too far** away... bad results!



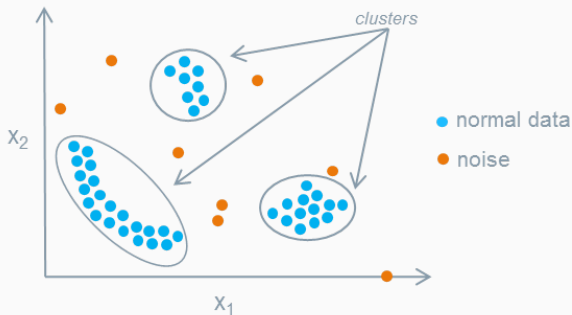
A challenge... and an opportunity

Can we exploit this?

Modifications to the original Gravitational Clustering

Anomaly detection

Anomaly. A rare item or observation which deviate significantly from the majority of the data (Hawkins, 1980).



Anomaly clusters

If a particle/cluster is “too far away” is an anomaly!

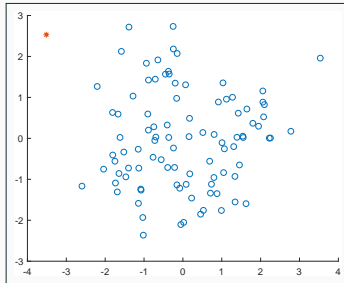
Two checks:

1. How **far**: percentage of the simulation that took place before it joined the rest of the observations
2. How **big**: if a lot of observations are close, then they are not anomalies. (Cluster size)

Setting. P = Maximum anomaly cluster size. T = coupling time.
If a particle/cluster takes up more than T to join other cluster can be considered an anomaly.

Distances and particles

- Anomalies can affect the intuitive process of aggregation forces: a distant particle can create incorrect clustering configurations.



- Solution: use a more general expression for the aggregation of forces to mitigate extreme values.

Generalizing the distance aggregation of forces

From

$$g_i^d(t) = \sum_{j \neq i} \frac{1}{m_i(t)} \frac{s_j^d(t) - s_i^d(t)}{|\vec{s}_j(t) - \vec{s}_i(t)|^3} dt^2(t)$$

to

$$g_i^d(t) = F_{j \neq i} \frac{1}{m_i(t)} \frac{s_j^d(t) - s_i^d(t)}{|\vec{s}_j(t) - \vec{s}_i(t)|^3} dt^2(t)$$

The discrete Sugeno integral with respect to the fuzzy measure m is defined as a function $S_m : [0, 1]^n \rightarrow [0, 1]$, given for every $\mathbf{x} = (x_1, \dots, x_n)$ (Bardozzo et al., 2020)

$$S_m(\mathbf{x}) = F\{G(x_{\sigma(i)}, m(A_i)) | i = 1, \dots, n\} \quad (1)$$

where \mathbf{x}_σ is an increasing permutation of \mathbf{x} such that $0 \leq x_{\sigma(1)} \leq \dots \leq x_{\sigma(n)}$. With the convention that $x_{\sigma(0)} = 0$, $A_i = \{(i), (i+1), \dots, (n)\}$, F is a n -ary function and G is a bi-variant function.

For our experimentation, we take as F the summation and G as the product.

Ad-hoc solution for the Sugeno integral

- We are aggregating vectors of forces (not numbers).
 1. We sort the vector forces based on their moduli.
 2. We multiply each vector to the correspondent measure.
 3. We sum the resulting vectors.

Experimental results

Datasets utilized

Table 1: Imabalance ratio, features and samples for each dataset studied.

Dataset	Imb. Ratio	Features	Samples
Ecoli1	3.36	7	336
Ecoli3	8.6	7	336
Glass6	6.38	9	214
Wisconsin	1.86	8	683
Yeast3	8.1	8	1484

Keel repository: <https://sci2s.ugr.es/keel/imbalanced.php>

Target metric. F1 score

Table 2: F1 score for the proposed algorithm using different aggregation functions for the attraction forces.

Dataset	Aggregation	
	Sum	Ad-hoc FS-Sugeno like
Ecoli1	0.8468	0.8350
Ecoli3	0.9285	0.9432
Glass6	0.9254	0.9057
Wisconsin	0.8062	0.7954
Yeast3	0.9324	0.9324

- Different aggregations \rightarrow different solutions to combine.
- In order to add more diversity we can use some well-known tactics:
 1. Sampling only a % of the observations.
 2. Using only some features.
- We generate k different models using randomly chosen aggregation, features and samples.
- Majority vote as final decision.

Results for decision making

Table 3: F1 score for the anomaly-Grav algorithm using different numbers of models.

Datasets	Number of Models			
	5	10	15	20
Ecoli1	0.8841	0.8900	0.8947	0.8952
Ecoli3	0.8887	0.8918	0.8943	0.8949
Glass6	0.8925	0.8925	0.8938	0.8944
Wisconsin	0.8767	0.8904	0.8956	0.8968
Yeast3	0.8884	0.8897	0.8916	0.8925

Conclusions and Future lines

- We proposed an adaptation of the gravitational clustering to perform anomaly detection.
- Performance measured using F1 score is good!
- Different aggregations than the sum can lead to better results.
- Decision making schemes seems to be very successful.

- Use more aggregation functions.
- Develop further the decision making scheme.

Thanks for your attention

Any questions?

javier.fumanal@unavarra.es

