

TRANSITION TO COHERENCE IN DENSE CORES USING THE GREEN BANK AMMONIA SURVEY (GAS)

JAIME E. PINEDA¹, RACHEL K. FRIESEN², ERIK ROSOLOWSKY³, FELIPE ALVES¹, ANA CHACÓN-TANARRO¹, HOPE HOW-HUAN CHEN⁴, MICHAEL CHUN-YUAN CHEN⁵, JAMES DI FRANCESCO^{5,6}, JARED KEOWN⁵, HELEN KIRK⁶, ANNA PUNANOVA¹, YOUNG MIN SEO^{7,8}, YANCY SHIRLEY⁸, ADAM GINSBURG⁹, CHRISTINE HALL¹⁰, STELLA S. R. OFFNER¹¹, AYUSHI SINGH¹², HÉCTOR G. ARCE¹³, PAOLA CASELLI¹, ALYSSA A. GOODMAN⁴, PETER G. MARTIN¹⁴, CHRISTOPHER MATZNER¹², PHILIP C. MYERS⁴, ELENA REDAELLI¹

¹Max-Planck-Institut für extraterrestrische Physik, Giessenbachstrasse 1, 85748 Garching, Germany

²Dunlap Institute for Astronomy & Astrophysics, University of Toronto, 50 St. George Street, Toronto, Ontario, Canada M5S 3H4

³Department of Physics, University of Alberta, Edmonton, AB, Canada

⁴Harvard-Smithsonian Center for Astrophysics, 60 Garden St., Cambridge, MA 02138, USA

⁵Department of Physics and Astronomy, University of Victoria, 3800 Finnerty Road, Victoria, BC, Canada V8P 5C2

⁶Herzberg Astronomy and Astrophysics, National Research Council of Canada, 5071 West Saanich Road, Victoria, BC, V9E 2E7, Canada

⁷Jet Propulsion Laboratory, NASA, 4800 Oak Grove Dr, Pasadena, CA 91109, USA

⁸Steward Observatory, 933 North Cherry Avenue, Tucson, AZ 85721, USA

⁹National Radio Astronomy Observatory, Socorro, NM 87801, USA

¹⁰Department of Physics, Engineering Physics & Astronomy, Queen's University, Kingston, Ontario, Canada K7L 3N6

¹¹Department of Astronomy, University of Massachusetts, Amherst, MA 01003, USA

¹²Department of Astronomy & Astrophysics, University of Toronto, 50 St. George Street, Toronto, Ontario, Canada M5S 3H4

¹³Department of Astronomy, Yale University, P.O. Box 208101, New Haven, CT 06520-8101, USA

¹⁴Canadian Institute for Theoretical Astrophysics, University of Toronto, 60 St. George St., Toronto, Ontario, Canada, M5S 3H8

ABSTRACT

We use the first data release (DR1) of the Green Bank Ammonia Survey (GAS). GAS is an ambitious Large Program at the Green Bank Telescope to map all regions within the northern hemisphere Gould Belt star-forming regions with $A_v \gtrsim 7$ in emission from NH₃ and other key molecular tracers. This first release includes the data for four regions in Gould Belt clouds: L1688 in Ophiuchus, Orion A North in Orion, NGC 1333 in Perseus, and B18 in Taurus. We study the velocity dispersion obtained towards all these regions and identify sharp transitions between super-sonic and sub-sonic turbulence in molecular clouds.

Keywords: ISM: clouds — stars: formation — ISM: molecules — ISM: individual (Perseus Molecular Complex, L1451, HH211, IRAS03282)

1. INTRODUCTION

2. RESULTS

3. SUMMARY

JEP, AP and PC acknowledge the financial support of the European Research Council (ERC; project PALs 320620). RKF is a Dunlap Fellow at the Dunlap Institute for Astronomy & Astrophysics. The Dunlap In-

stitute is funded through an endowment established by the David Dunlap family and the University of Toronto. EWR, PGM and CDM are supported by Discovery Grants from NSERC of Canada. SSRO acknowledges support from NSF grant AST-1510021. The National Radio Astronomy Observatory is a facility of the National Science Foundation operated under cooperative agreement by Associated Universities, Inc.

Facility: Green Bank Telescope

Software: Astropy (?), Matplotlib (?), pyspeckit (?)