

Quantum Statistical Physics

Dr. Ryuichi Shindou¹

September 12, 2015

¹[mail:rshindou@pku.edu.cn](mailto:rshindou@pku.edu.cn)

Contents

1	General Introduction	1
1.1	General Introduction	1
1.2	Linear Harmonic Oscillator	3
1.3	Many Harmonic Oscillator	4
1.4	Field Quantization	6
1.5	System of indistinguishable particles	11
1.6	Creation operator & destruction operator	12
1.7	Hamiltonian and other operators of creation and destruction operators	15
1.8	Degenerate electron gas	15
2	Green's function and field theory at zero temperature	17
2.1	Schrodinger, Heisenberg & interaction picture	17
2.2	Adiabatic switching & Gell-Mann & Low Theorem	17
2.3	Green's function	17
2.4	Wick's theorem	17
2.5	Diagrammatic analysis of perturbation theory (Fermion Case)	17
2.6	(Self-Consist) Hatree-Fock Approximation	17
2.7	Imperfect Fermi Gas	22
2.8	Ring approximation: High density Fermi gas with long-range Coulomb interaction	42
3	Application to electron gas at the zero temperature	71
3.1	Screening effect in a degenerate electron gas	71
3.2	Collective Modes in degenerate electron gas and imperfect Fermi gas – plasma oscillation & zero sound mode	83
4	Field theory at finite temperature	89
4.1	Temperature (Matsubara) Green's function	89
4.2	Temperature Green's function in the interaction picture	89
4.3	Wick's theorem for the temperature Green's function	89
4.4	Feynman's rule for the temperature Green's function	89
4.5	Dyson's equation & Hartree-Fock approximation	89
4.6	Specific heat of an imperfect Fermi gas at low-temperature	89
4.7	Real-time Green's function and generalized Lehmann representation	89
4.8	Linear response at finite temperature	89
4.9	Screening effect & plasma oscillation in an electron gas at finite (or high) temperature	89
5	Application to electron-phonon systems	91
5.1	Non-interacting phonon system and Debye theory of specific heat	91
5.2	Electron-phonon interaction	95

6	Application to superconductivity	97
6.1	Superconducting properties and thermodynamics relations	97
6.2	BCS theory, Cooper pairing and Gap equation	100
6.2.1	BCS(Bardeen Copper Schrieffer) Hamiltonian	100
6.3	Electromagnetic property of superconductors	106
6.4	Microscopic derivation of the Ginzburg Landau equations	106

Disclaimer

This version of textbook is based on Prof. Shindou's lecture notes on Quantum Statistical Physics. The original manuscript is a hand-written one and makes it difficult to search or jump between sections. Hence we several students make a L^AT_EX version.

Definitely, this draft contains numerous deficiency including typing errors or other confusing errors. If any reader of this notes found something incorrect comparing with the original hand-written one, please send us emails to report. We would like to enhance.

Also, this version contains another significant bug (which is waiting for you to fix): for a same notation in different part of the note, it could be represented by differnt way, say x , \mathbf{x} may be same thing but the former one mainly shows in chapter 1 and the later one may be shown up otherwise.

The up-to-date version can be seen at a github page: <https://github.com/laserroger/QSP-Latex>.

QSPLATEX Group
2015 - 09 - 01

Chapter 1

General Introduction

1.1 General Introduction

We begin with a simple math problem.

Suppose that an operator a satisfy the relation:

$$[a, a^\dagger] = 1$$

The problem is to find all eigenvalues and eigenvectors of a Hermitian operator $a^\dagger a$.

We first note that all the eigenvalues are positive.

If $|\alpha\rangle$ is a normalized eigenvector,

$$a^\dagger a |\alpha\rangle = \alpha |\alpha\rangle$$

its eigenvalue α is always positive

$$\alpha = \langle \alpha | a^\dagger a | \alpha \rangle = ||a|\alpha\rangle||^2 \geq 0$$

We next note that

$$\begin{aligned} [a^\dagger a, a] &= [a^\dagger, a] a = -a \\ [a^\dagger a, a^\dagger] &= a^\dagger [a, a^\dagger] = a^\dagger \end{aligned}$$

For this, we see that $a|\alpha\rangle$ is an eigenvector with eigenvalue $\alpha - 1$

$$\begin{aligned} (a^\dagger a) a |\alpha\rangle &= (a a^\dagger a - a) |\alpha\rangle \\ &= (\alpha - 1) a |\alpha\rangle \end{aligned}$$

Similar, $a^\dagger |\alpha\rangle$ is an eigenvector with eigenvalue $\alpha + 1$

$$\begin{aligned} (a^\dagger a) a^\dagger |\alpha\rangle &= (a^\dagger a^\dagger a - a^\dagger) |\alpha\rangle \\ &= (\alpha + 1) a^\dagger |\alpha\rangle \end{aligned}$$

The norm of $a|\alpha\rangle$ is α ,

$$||a|\alpha\rangle||^2 = \langle \alpha | a^\dagger a | \alpha \rangle = \alpha$$

Similar,

$$||a^\dagger|\alpha\rangle||^2 = \langle\alpha|aa^\dagger|\alpha\rangle = \alpha + 1$$

By repeating application of a ‘ n ’ times, we can produce an eigenvector with eigenvalue $\alpha - n$.

$$a^\dagger aa^n|\alpha\rangle = (\alpha - n)a^n|\alpha\rangle$$

But, when n becomes larger than α , the eigenvalue becomes negative, which contradicts the non-negativeness of eigenvalues of $a^\dagger a$.

This means that a given $|\alpha\rangle$ has some integer ‘ n ’, which satisfies

$$a^n|\alpha\rangle \neq 0, \quad a^{n+1}|\alpha\rangle = 0$$

The latter equation means

$$\begin{aligned} 0 &= \langle\alpha|(a^\dagger)^n \cdot a^\dagger a \cdot a^n|\alpha\rangle \\ &= (\alpha - n)\langle\alpha|(a^\dagger)^n a^n|\alpha\rangle \end{aligned}$$

The former means that this:

$$\langle\alpha|(a^\dagger)^n a^n|\alpha\rangle$$

is not zero.

Thus, $\alpha = n$.

This shows that the eigenvalues of $a^\dagger a$ must be non-negative integers and there is a “vacuum state” $|0\rangle$ such that

$$a|0\rangle = 0$$

By repeatedly applying a^\dagger to vacuum state, we see that $(a^\dagger)^n|0\rangle$ is an eigenvector with eigenvalue n .

$$a^\dagger a(a^\dagger)^n|0\rangle = n(a^\dagger)^n|0\rangle$$

The norm of this eigenvector is the factorial of n .

$$\begin{aligned} \langle 0|(a)^n(a^\dagger)^n|0\rangle &= \langle 0|a^{n-1} \cdot \underset{1+a^\dagger a}{aa^\dagger} \cdot (a^\dagger)^{n-1}|0\rangle \\ &= (n - 1 + 1) \cdot \langle 0|(a)^{n-1}(a^\dagger)^{n-1}|0\rangle \\ &= n \cdot (n - 1) \cdot \langle 0|(a)^{n-2}(a^\dagger)^{n-2}|0\rangle \\ &= n \cdot (n - 1) \cdot (n - 2) \cdots 2 \langle 0|aa^\dagger|0\rangle = n! \end{aligned}$$

The normalized eigenvector with eigenvalue ‘ n ’ is

$$|n\rangle = \frac{1}{\sqrt{n!}}(a^\dagger)^n|0\rangle \tag{1.1.1}$$

with this definition, the $|n\rangle$ are orthonormal and satisfy:

$$a^\dagger|n\rangle = \sqrt{n+1}|n+1\rangle \tag{1.1.2}$$

$$\begin{aligned}
a|n\rangle &= \frac{1}{\sqrt{n!}} a(a^\dagger)^n |0\rangle \\
&= \frac{1}{\sqrt{n!}} (a^\dagger a + 1)(a^\dagger)^{n-1} |0\rangle \\
&= \frac{1}{\sqrt{n!}} n(a^\dagger)^{n-1} |0\rangle \\
&= \sqrt{n} \cdot \frac{1}{\sqrt{(n-1)!}} (a^\dagger)^{n-1} |0\rangle \\
&= \sqrt{n} |n-1\rangle \\
a^\dagger a |n\rangle &= n |n\rangle
\end{aligned} \tag{1.1.3}$$

The operator a^\dagger and a are called “raising” and “lowering” operators respectively, because the raise and lower the eigenvalue or $a^\dagger a$. In the later application, a^\dagger and a are called “creation” and “annihilation” operators respectively.

You might be afraid that the ground state “ $|0\rangle$ ” is not necessarily unique, so that the state do not exhaust all the possible eigenstates of $a^\dagger a$.

$$|0\rangle_1, |0\rangle_2, |0\rangle_3, \dots$$

$\rightarrow b$

In such a case, however, we can define some other operators, which commute a and a^\dagger and which connect the ground states.

Thus, we can further classify these ground states according to the eigenvalues of these other operators¹.

1.2 Linear Harmonic Oscillator

Our first application of the general argument is a one-dimensional harmonic oscillator which has a Hamiltonian as this form:

$$H = \frac{p^2}{2m} + \frac{m\omega^2}{2m} x^2 \tag{1.2.1}$$

x and p are the position and the momentum operator of a particle whose mass is ‘ m ’.
 x , p satisfies

$$[x, p] = i\hbar$$

$$\begin{aligned}
H &= \frac{\hbar\omega}{2} \left[\frac{p^2}{m\omega\hbar} + \frac{m\omega}{\hbar} x^2 \right] \\
&\equiv \frac{\hbar\omega}{2} [P^2 + X^2]
\end{aligned}$$

$$[X, P] = i$$

By factorizing an energy unit, the Hamiltonian has a quadratic form of these two dimensionless quantities.

¹like:

$$\begin{aligned}
|n\rangle|\beta\rangle &\stackrel{def}{=} \frac{1}{\sqrt{n!}} (a^\dagger)^n |0\rangle|\beta\rangle \\
b|\beta\rangle &= \beta|\beta\rangle
\end{aligned}$$

They satisfy this commutation relation.

In terms of these two hermitian operators, let us define a following non-hermitian operator:

$$a = \frac{1}{\sqrt{2}}[X + iP]$$

$$a^\dagger = \frac{1}{\sqrt{2}}[X - iP]$$

From this, we see

$$[a, a^\dagger] = 1$$

while the Hamiltonian takes this form:

$$\begin{aligned} H &= \frac{\hbar\omega}{2} \left[\frac{1}{2}(X - iP)(X + iP) + \frac{1}{2}(X + iP)(X - iP) \right] \\ &= \frac{\hbar\omega}{2} [a^\dagger a + a a^\dagger] \\ &= \hbar\omega \left[a^\dagger a + \frac{1}{2} \right] \end{aligned}$$

Now we know how to obtain eigenvalues and eigenvectors of $a^\dagger a$ for any given non-hermitian operator a .

Thus, the eigen energy of H is given by this

$$H|n\rangle = \hbar\omega \left(n + \frac{1}{2} \right) |n\rangle$$

while the eigenstate $|n\rangle$ is obtained from a repeated application of the operator a onto the ‘ground state’.

The ground state $|0\rangle$ satisfies this,

$$\begin{aligned} a|0\rangle &= \frac{1}{\sqrt{2}} \left[XD + i \left(-i \frac{\partial}{\partial X} \right) \right] |0\rangle = \frac{1}{\sqrt{2}} \left[X + \frac{\partial}{\partial X} \right] |0\rangle = 0 \\ |0\rangle &= \frac{1}{N} e^{-X^2/2} \end{aligned}$$

Thus, the eigenstate with eigen energy $(n + 1/2)$ takes this form

$$|n\rangle \propto \left(X - \frac{\partial}{\partial X} \right)^n e^{-\frac{X^2}{2}} = e^{-\frac{X^2}{2}} \times \text{Hermite polynomials}$$

which is expressed in terms of Hermite polynomial with integer index n .

1.3 Many Harmonic Oscillator

Our next application is a system of many harmonic oscillators, which consists of many particles with different mass.

$$H = \sum_i \frac{1}{2m_i} P_i^2 + \sum_{i,j} V_{ij} Q_i Q_j \quad (1.3.1)$$

where Q_i and P_i are canonical coordinates and momenta

$$\begin{cases} [Q_i, Q_j] = [P_i, P_j] = 0 \\ [Q_i, P_j] = i\hbar\delta_{ij} \end{cases}$$

Since these two momentum operators can be exchanged freely, we can regard that V_{ij} is real symmetric

$$V_{ij} = V_{ji} \text{ real}$$

We normalize these coordinates and momenta by differen mass

$$q_i \equiv \sqrt{m_i}Q_i, \quad p_i \equiv \frac{P_i}{\sqrt{m_i}}$$

$$H = \frac{1}{2} \sum_i p_i^2 + \frac{1}{2} \sum_{ij} v_{ij} q_i q_j$$

$$(v_{ij} \equiv \frac{V_{ij}}{\sqrt{m_i m_j}})$$

To obtain eigenstates and eigen energy of H , we first diagonalize this real symmetric matrix by an orthogonal matrix.

$$\mathbb{C}^T \mathbb{V} \mathbb{C} = \mathbb{W}$$

We assume that eigenvalues of \mathbb{V} are all positive

$$c_{\alpha i} c_{\beta j} v_{ij} = \omega_{\alpha}^2 \delta_{\alpha\beta}$$

In terms of this new basis, the Hamiltonian takes the following form:

$$H = \frac{1}{2} \sum_{\alpha} (\tilde{p}_{\alpha}^2 + \omega_{\alpha}^2 \tilde{q}_{\alpha}^2)$$

where \tilde{p}_{α} and \tilde{q}_{α} are canonical momenta and coordinates:

$$\tilde{q}_{\alpha} \equiv \sum_i c_{\alpha i} q_i$$

$$\tilde{p}_{\alpha} \equiv \sum_i c_{\alpha i} p_i$$

$$[\tilde{q}_{\alpha}, \tilde{q}_{\beta}] = [\tilde{p}_{\alpha}, \tilde{p}_{\beta}] = 0$$

$$[\tilde{q}_{\alpha}, \tilde{p}_{\beta}] = i\hbar\delta_{\alpha\beta}$$

Thanks to the basis change, we have a system of decouple linear harmonic oscillators, each of which can be separately solved.

Namely, we form lowering and raising operators for each mode:

$$a_{\alpha} = \frac{1}{\sqrt{2\hbar}} \left[\sqrt{\omega_{\alpha}} \tilde{q}_{\alpha} + \frac{i}{\sqrt{\omega_{\alpha}}} \tilde{p}_{\alpha} \right]$$

$$a_{\alpha}^{\dagger} = \frac{1}{\sqrt{2\hbar}} \left[\sqrt{\omega_{\alpha}} \tilde{q}_{\alpha} - \frac{i}{\sqrt{\omega_{\alpha}}} \tilde{p}_{\alpha} \right]$$

which satisfy

$$\begin{aligned} [a_\alpha, a_\beta] &= [a_\alpha^\dagger, a_\beta^\dagger] = 0 \\ [a_\alpha, a_\beta^\dagger] &= \delta_{\alpha\beta} \end{aligned} \quad (1.3.2)$$

and

$$H = \sum_{\alpha} \hbar\omega_{\alpha} (a_{\alpha}^{\dagger} a_{\alpha} + \frac{1}{2})$$

The eigenstates of H are described by a set of non-negative integers defined for each decoupled mode.

$$H|n_1, n_2, \dots\rangle = \sum_{\alpha} \hbar\omega_{\alpha} (n_{\alpha} + \frac{1}{2})|n_1, n_2, \dots\rangle$$

The integer for a mode α , n_{α} , is an eigenvalue of $a_{\alpha}^{\dagger} a_{\alpha}$.

$$\text{for } \forall \alpha, \quad a_{\alpha}^{\dagger} a_{\alpha} |n_1, n_2, \dots, n_{\alpha}, \dots\rangle = n_{\alpha} |n_1, n_2, \dots, n_{\alpha}, \dots\rangle$$

These eigenstates can be constructed for the ‘ground state’

$$|n_1, n_2, \dots\rangle = \left[\prod_{\alpha} \frac{(a_{\alpha}^{\dagger})^{n_{\alpha}}}{\sqrt{n_{\alpha}!}} \right] |0, 0, \dots\rangle$$

where the ground state is defined by

$$a_{\alpha} |0, 0, \dots\rangle = 0, \text{ for } \forall \alpha$$

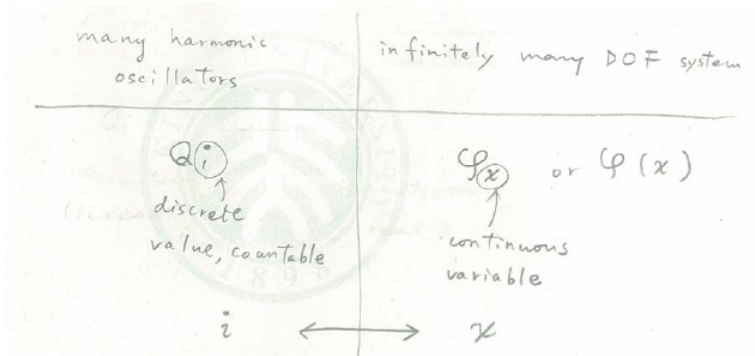
Corresponding to (1.1.2) and (1.1.3), we have

$$a_{\alpha}^{\dagger} |n_1, n_2, \dots, n_{\alpha}, \dots\rangle = \sqrt{n_{\alpha} + 1} |n_1, n_2, \dots, n_{\alpha} + 1, \dots\rangle \quad (1.3.3)$$

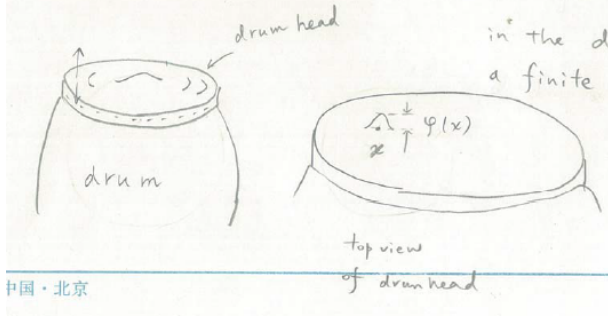
$$a_{\alpha} |n_1, n_2, \dots, n_{\alpha}, \dots\rangle = \sqrt{n_{\alpha}} |n_1, n_2, \dots, n_{\alpha} - 1, \dots\rangle \quad (1.3.4)$$

1.4 Field Quantization

Our final application is a system with infinitely many degrees of freedom. The system is described by what we call a ‘field’ ϕ_x .



An example of the system is a drumhead vibration. Here is a picture of drum:



Every spatial point in the drumhead can have a finite displacement perpendicular to the flat drumhead.

We can define φ of x ($\varphi(x)$) as a displacement at x point.

Namely, a given φ of x specifies how the drumhead is deformed from the flat drum head.

Let us consider φ of x whose motion is described by the following Lagrangian

$$L(\varphi, \dot{\varphi}) = \frac{1}{2} \int d^2x \dot{\varphi}(x) \dot{\varphi}(x) - \frac{1}{2} \int d^2x \int d^2x' K(x-x') \varphi(x) \varphi(x') \quad \text{with } K(x-x') = K(x'-x) \quad (1.4.1)$$

The corresponding equation of motion is

$$0 = \frac{\partial}{\partial t} \frac{\delta L}{\delta \dot{\varphi}(x)} - \frac{\delta \dot{L}}{\delta \varphi(x)} = \ddot{\varphi}(x) + \int d^2x' K(x-x') \varphi(x') \quad (1.4.2)$$

When the interaction between fields is spatially local like this

$$K(x-x') = -c^2 \Delta^2 \delta(x-x'),$$

the equation of motion becomes the usual wave equation

$$\nabla^2 \varphi(x) - \frac{1}{c^2} \ddot{\varphi}(x) = 0$$

If we regard that $\varphi(x)$ is a coordinate variable for each x , the conjugate momentum to φ is

$$\Pi(x) = \frac{\delta L}{\delta \dot{\varphi}(x)} = \dot{\varphi}(x)$$

The Hamiltonian is then

$$H = \int d^2x \Pi(x) \dot{\varphi}(x) - L = \frac{1}{2} \int d^2x \Pi(x) \Pi(x) + \frac{1}{2} \int d^2x \int d^2x' K(x-x') \varphi(x) \varphi(x') \quad (1.4.3)$$

To quantize the system, we suppose that $\varphi(x)$ and $\Pi(x)$ are Hermitian operators, which satisfy

$$\begin{aligned} [\varphi(x), \varphi(x')] &= [\Pi(x), \Pi(x')] = 0 \\ [\varphi(x), \Pi(x')] &= i\hbar \delta^2(x-x') \end{aligned} \quad (1.4.4)$$

In such a case, it is helpful to express the field in the momentum representation,

$$\tilde{\varphi}(k) = \int d^2x \varphi(x) e^{-ikx}$$

$$\tilde{\Pi}(x) = \int d^2x \Pi(x) e^{-ikx}$$

where the inverse transformation is

$$\varphi(x) = \int \frac{d^2k}{(2\pi)^2} \tilde{\varphi}(k) e^{ikx}$$

with

$$\int \frac{d^2k}{(2\pi)^2} e^{-ikx} = \delta^2(x) \quad (1.4.5)$$

Since $\varphi(x)$ and $\Pi(x)$ are hermitian, we have

$$\varphi^\dagger(k) = \varphi(-k)$$

$$\Pi^\dagger(k) = \Pi(-k)$$

The commutation relations are

$$\begin{aligned} [\tilde{\varphi}(k), \tilde{\varphi}(k')] &= [\tilde{\Pi}(k), \tilde{\Pi}(k')] = 0 \\ [\tilde{\varphi}(k), \tilde{\Pi}(k')] &= i\hbar(2\pi)^2 \delta^2(k + k') \end{aligned} \quad (1.4.6)$$

Now let us introduce

$$\omega^2(k) = \int d^2x K(x) e^{-ikx}$$

Since $K(x) = K(-x)$, $\omega(k)$ is real and even function in k .

$$\omega^*(k) = \omega(k) = \omega(-k)$$

Let us again assume that $\omega(k)$ are non-negative for all k , so that we replace this by $\omega^2(k)$. Rewriting the Hamiltonian in terms of these fields and functions, we obtain

$$H = \frac{1}{2} \int \frac{d^2k}{(2\pi)^2} \left[\tilde{\Pi}(-k) \tilde{\varphi}(k) + \omega^2(k) \tilde{\varphi}(-k) \tilde{\varphi}(k) \right] = \frac{1}{2} \int \frac{d^2k}{(2\pi)^2} \left[\tilde{\Pi}^\dagger(k) \tilde{\varphi}(k) + \omega^2(k) \tilde{\varphi}^\dagger(k) \tilde{\varphi}(k) \right] \quad (1.4.7)$$

We next define annihilation and creation operators as follows,

$$\begin{aligned} a(k) &= \frac{1}{\sqrt{2}} \left[\sqrt{\frac{\omega(k)}{\hbar}} \tilde{\varphi}(k) + \frac{i}{\sqrt{\omega(k)\hbar}} \tilde{\Pi}(k) \right] \\ a^\dagger(k) &= \frac{1}{\sqrt{2}} \left[\sqrt{\frac{\omega(k)}{\hbar}} \tilde{\varphi}(k) - \frac{i}{\sqrt{\omega(k)\hbar}} \tilde{\Pi}(k) \right] \end{aligned} \quad (1.4.8)$$

Their commutation relations are calculated as follows,

$$\begin{aligned}
[a(k), a(k')] &= \frac{1}{2\hbar} \left[\sqrt{\omega_k} \tilde{\varphi}(k) + \frac{i}{\sqrt{\omega_k}} \tilde{\Pi}(k), \sqrt{\omega_{k'}} \tilde{\varphi}(k') + \frac{i}{\sqrt{\omega_{k'}}} \tilde{\Pi}(k') \right] \\
&= \frac{1}{2\hbar} \sqrt{\frac{\omega_k}{\omega_{k'}}} i [\tilde{\varphi}(k), \tilde{\Pi}(k')] + \frac{1}{2\hbar} \sqrt{\frac{\omega_{k'}}{\omega_k}} i [\tilde{\Pi}(k), \tilde{\varphi}(k')] \\
&= (-1)(2\pi)^2 \delta^2(k + k') + (2\pi)^2 \delta^2(k + k') = 0 \\
[a^\dagger(k), a^\dagger(k')] &= 0
\end{aligned}$$

$$\begin{aligned}
[a(k), a^\dagger(k')] &= \frac{1}{2\hbar} \left[\sqrt{\omega_k} \tilde{\varphi}(k) + \frac{i}{\sqrt{\omega_k}} \tilde{\Pi}(k), \sqrt{\omega_{k'}} \tilde{\varphi}(-k') - \frac{i}{\sqrt{\omega_{k'}}} \tilde{\Pi}(-k') \right] \\
&= \frac{-i}{2\hbar} \sqrt{\frac{\omega_k}{\omega_{k'}}} [\tilde{\varphi}(k), \tilde{\Pi}(-k')] + \frac{i}{2\hbar} [\tilde{\Pi}(k), \tilde{\varphi}(-k')] \\
&= (2\pi)^2 \delta^2(k - k')
\end{aligned} \tag{1.4.9}$$

If we write H in terms of a & a^\dagger , we obtain

$$H = \frac{1}{2} \int \frac{d^2k}{(2\pi)^2} \hbar \omega(k) [a^\dagger(k) a(k) + a(k) a^\dagger(k)] = \frac{1}{2} \int \frac{d^2k}{(2\pi)^2} \hbar \omega(k) a^\dagger(k) a(k) + C \tag{1.4.10}$$

Thus, eigenstate of H is given by the repeated application of $a^\dagger(k)$ onto the ground state¹

$$\begin{aligned}
|k_1\rangle &= a^\dagger(k_1)|0\rangle, \quad (E = \omega(k_1)) \\
|k_1, k_2\rangle &= a^\dagger(k_1) a^\dagger(k_2)|0\rangle, \quad (E = \omega(k_1) + \omega(k_2))
\end{aligned}$$

while the ground state satisfies

$$a(k)|0\rangle = 0, \quad \text{for } \forall k$$

To characterize these multi-particle states, let us consider the following operators,

$$\mathbb{P} \equiv \int \frac{d^2k}{(2\pi)^2} \hbar k a^\dagger(k) a(k) \tag{1.4.11}$$

which satisfies

$$\begin{aligned}
[\mathbb{P}, a^\dagger(k)] &= \hbar k a^\dagger(k) \\
[\mathbb{P}, a(k)] &= -\hbar k a(k)
\end{aligned}$$

so that

$$\begin{aligned}
\mathbb{P}|k_1, k_2, \dots\rangle &= \mathbb{P} a^\dagger(k_1) a^\dagger(k_2) \dots |0\rangle = a^\dagger(k_1) (\hbar k_1 + \mathbb{P}) a^\dagger(k_2) \dots |0\rangle \\
&= \hbar k_1 |k_1, k_2, \dots\rangle + a^\dagger(k_1) \mathbb{P} a^\dagger(k_2) \dots |0\rangle \\
&= \hbar k_1 |k_1, k_2, \dots\rangle + a^\dagger(k_1) a^\dagger(k_2) (\hbar k_2 + \mathbb{P}) \dots |0\rangle \\
&= \dots \\
&= \hbar(k_1 + k_2 + \dots) |k_1, k_2, \dots\rangle
\end{aligned}$$

¹Regarding these energy units as particles respectively, we can think of these excited states as one-particle state, two-particle state, and so on.

we can also see that

$$\begin{aligned}
[\mathbb{P}, \varphi(x)] &= \left[\mathbb{P}, \int \frac{d^2k}{(2\pi)^2} \sqrt{\frac{\hbar}{2\omega(k)}} (a(k)e^{ik \cdot x} + a^\dagger(k)e^{-ik \cdot x}) \right] \\
&= \int \frac{d^2k}{(2\pi)^2} \sqrt{\frac{\hbar}{2\omega(k)}} \{ (-\hbar k) a(k) e^{ik \cdot x} + (\hbar k) a^\dagger(k) e^{-ik \cdot x} \} \\
&= i\hbar \frac{\partial}{\partial x} \left\{ \int \frac{d^2k}{(2\pi)^2} \sqrt{\frac{\hbar}{2\omega(k)}} (a(k)e^{ik \cdot x} + a^\dagger(k)e^{-ik \cdot x}) \right\} \\
&= i\hbar \frac{\partial}{\partial x} \varphi(x)
\end{aligned}$$

Similarly

$$[\mathbb{P}, \Pi(x)] = i\hbar \cdot \frac{\partial}{\partial x} \Pi(x)$$

Using these relations, one can show that

$$e^{a \cdot \mathbb{P} / i\hbar} \varphi(x) e^{-a \cdot \mathbb{P} / i\hbar} = \varphi(x + a)$$

$$e^{a \cdot \mathbb{P} / i\hbar} \Pi(x) e^{-a \cdot \mathbb{P} / i\hbar} = \Pi(x + a)$$

Namely, taking a derivative of r.h.s with respect to a , we have

$$\begin{aligned}
\frac{\partial}{\partial a} \left[e^{a \cdot \mathbb{P} / i\hbar} \varphi(x) e^{-a \cdot \mathbb{P} / i\hbar} \right] &= \frac{1}{i\hbar} e^{a \cdot \mathbb{P} / i\hbar} [\mathbb{P}, \varphi(x)] e^{-a \cdot \mathbb{P} / i\hbar} \\
&= e^{a \cdot \mathbb{P} / i\hbar} \frac{\partial}{\partial x} \varphi(x) e^{-a \cdot \mathbb{P} / i\hbar} \\
&= \frac{\partial}{\partial x} \left\{ e^{a \cdot \mathbb{P} / i\hbar} \varphi(x) e^{-a \cdot \mathbb{P} / i\hbar} \right\}
\end{aligned}$$

This means that

$$e^{a \cdot \mathbb{P} / i\hbar} \varphi(x) e^{-a \cdot \mathbb{P} / i\hbar} = f(x + a)$$

while f should be ϕ when $a = 0$, thus we have

$$e^{a \cdot \mathbb{P} / i\hbar} \varphi(x) e^{-a \cdot \mathbb{P} / i\hbar} = \varphi(x + a) \tag{1.4.12}$$

Similar we have

$$e^{a \cdot \mathbb{P} / i\hbar} \Pi(x) e^{-a \cdot \mathbb{P} / i\hbar} = \Pi(x + a)$$

These two relations means that \mathbb{P} is a generator of the spatial translation and is therefore the momentum operator.

1.5 System of indistinguishable particles

In the preceding sections, we consider the quantum states of an oscillator system as multiparticle states, where each particle has an energy unit and is often called a “phonon”. In this section, we follow a different line of reasoning. We will start with a space of states describing a single particle, which is either Bose particle or Fermi particle. We then construct the multiple-particle states based on standard methods.

For the Bose particle case, we will arrive at a system of states and operators, that is mathematically equivalent to that found previously for a system of oscillators. Thereby, this shows that the interpretation of oscillator states as many-phonon states is consistent with the usual description of many-particle systems with Bose statistics.

We will develop a formalism with Fermi-Dirac statistics, for which the states do not resemble those of a harmonic-oscillator system.

Consider first the case of distinguishable particles.

A n -particle state is given by a simple product of one-particle states

$$|\psi\rangle = |\psi_1\rangle|\psi_2\rangle \cdots |\psi_n\rangle$$

You can also write this in the coordinate representation as

$$\psi(x_1, \cdots, x_n) = \psi_1(x_1)\psi_2(x_2) \cdots \psi_n(x_n)$$

where x_j denotes the coordinate of j -th particle and $\langle x_j|\psi_j\rangle = \psi_j(x_j)$.

Please notice that these n -particles are all distinguishable from one another.

Namely, each particle has its own coordinate x , and we can distinguish j -th particle from m -th particle, because they belong to different one-particle states

$$\psi_j(x) \neq \psi_m(x), \quad \text{for } j \neq m$$

Let us next consider indistinguishable particles.

To make up a many-particle wavefunction for indistinguishable particles out of this, we need to make any pair of particles cannot be distinguished from each other.

There is two-ways of doing this, one way is to “symmetrize” this with respect to all the possible permutations of n particles:

$$\psi_1(x_1) \cdot \psi_2(x_2) \cdots \psi_n(x_n) \xrightarrow{\text{“symmetrize”}} \frac{1}{\sqrt{n!}} \sum_P \psi_{P(1)}(x_1) \cdot \psi_{P(2)}(x_2) \cdots \psi_{P(n)}(x_n)$$

Here P denotes a permutation among n numbers, $\{1, 2, \cdots, n\}$

$$P \equiv \begin{pmatrix} 1 & 2 & \cdots & n \\ P(1) & P(2) & \cdots & P(n) \end{pmatrix}$$

into $\{P(1), P(2), \cdots, P(n)\}$

we take the sum over all possible permutations here.

The total number of possible permutations among n distinguishable objects are given by a factorial of n . To normalize this many-particle wavefunction, I put 1 over the square root of the factorial of n . We can see that the exchange between the j -th particle coordinate and the m -th particle coordinate does not change this many-particle wavefunction.

$$\begin{aligned}
\psi(x_2, x_1, x_3, \dots, x_n) &= \frac{1}{\sqrt{n!}} \sum_P \psi_{P(1)}(x_2) \psi_{P(1)}(x_2) \psi_{P(3)}(x_3) \cdots \psi_{P(n)}(x_n) \\
&= \frac{1}{\sqrt{n}} \sum_P \psi_{P(1)}(x_2) \psi_{P(1)}(x_2) \psi_{P(3)}(x_3) \cdots \psi_{P(n)}(x_n) \\
&\quad \left\{ P' = P \cdot \begin{pmatrix} 1 & 2 & 3 & \cdots & n \\ 2 & 1 & 3 & \cdots & n \end{pmatrix} \right\} \\
&= \frac{1}{\sqrt{n}} \sum_{P'} \psi_{P'(1)}(x_1) \psi_{P'(2)}(x_2) \psi_{P'(3)}(x_3) \cdots \psi_{P'(n)}(x_n) \\
&= \psi(x_1, x_2, x_3, \dots, x_n)
\end{aligned}$$

Here P' is a product between a permutation P and the permutation which exchange 1 and 2, while remains other indices intact.

This product can be regarded as a mapping from P to P' and this mapping is one-to-one mapping. Namely, a set of all possible permutations is unchanged under that mapping.

Therefore, the right hand side is identical to the original wavefunctions.

This argument so far holds true not only for the exchange between arbitrary pair of two particles but also for an arbitrary permutations among n particles.

$$\psi(x_{Q(1)}, x_{Q(2)}, \dots, x_{Q(n)}) = \cdots = \psi(x_1, x_2, \dots, x_n)$$

Because of this symmetrized property, n -particles in this wavefunction are indistinguishable from one another.

An indistinguishable particle in a symmetrized many-particle wavefunction is called as “boson”.

The other way of making wavefunctions for indistinguishable particles is “antesymmetrize” the wavefunction for distinguishable particles.

1.6 Creation operator & destruction operator

So far, we have constructed a complete orthonormal set of states in the multiparticle space. With these constructions in hand, we are now ready to define creation and destruction operators. These operators change the number of particles, so that they connect states of different numbers of particles. We will see that all the physical operators are expressed in terms of creation and destruction operators.

Let $|\varphi\rangle$ be any one-particle state. We define a^\dagger of φ to be that linear operator which satisfies this, for any n -particle states $|\psi_1, \dots, \psi_n\rangle$.

$$\begin{aligned}
a^\dagger \varphi |\psi_1, \dots, \psi_n\rangle &= |\varphi, \psi_1, \dots, \psi_n\rangle & (e) \\
&\text{for any } |\psi_1, \dots, \psi_n\rangle & (1.6.1)
\end{aligned}$$

For $n = 0$, this simply means that

$$a^\dagger(\varphi)|vac\rangle = |\varphi\rangle$$

We call $a^\dagger(\varphi)$ as the creation operator for the state $|\varphi\rangle$, and its hermitian conjugate $a(\varphi)$ as the destruction or annihilation operator.

A destruction operator converts an n -particle states into a $(n - 1)$ -particle states.

To find the effect of $a(\varphi)$ on an n -particle state $|\psi_1, \psi_2, \dots, \psi_n\rangle$, we multiply on the left by an arbitrary $(n - 1)$ -particle state, which is this

$$\begin{aligned}
& \langle x_1, \dots, x_{n-1} | a(\varphi) | \psi_1, \dots, \psi_n \rangle \\
&= (\langle \psi_1, \dots, \psi_n | a^\dagger(\varphi) | x_1, \dots, x_{n-1} \rangle)^* \\
&= (\langle \psi_1, \dots, \psi_n | \varphi, x_1, \dots, x_{n-1} \rangle)^* \\
&= \left| \begin{array}{cccc} \langle \psi_1 | \phi \rangle & \langle \psi_1 | x_1 \rangle & \cdots & \langle \psi_1 | x_{n-1} \rangle \\ \langle \psi_2 | \phi \rangle & \langle \psi_2 | x_1 \rangle & \cdots & \langle \psi_2 | x_{n-1} \rangle \\ \vdots & \vdots & \ddots & \vdots \\ \langle \psi_n | \phi \rangle & \langle \psi_n | x_1 \rangle & \cdots & \langle \psi_n | x_{n-1} \rangle \end{array} \right|_\xi^* \\
&= \left\{ \sum_{k=1}^n \xi^{k-1} \langle \psi_k | \varphi \rangle \left| \begin{array}{ccc} \langle \psi_1 | x_1 \rangle & \cdots & \langle \psi_1 | x_{n-1} \rangle \\ \langle \psi_2 | x_1 \rangle & \cdots & \langle \psi_2 | x_{n-1} \rangle \\ \vdots & \ddots & \vdots \\ \langle \psi_n | x_1 \rangle & \cdots & \langle \psi_n | x_{n-1} \rangle \end{array} \right|_\xi \right\}^* \\
&= \sum_{k=1}^n \xi^{k-1} \langle \varphi | \psi_k \rangle \langle x_1, \dots, x_{n-1} | \psi_1, \dots, \psi_{k-1}, \psi_{k+1}, \dots, \psi_n \rangle
\end{aligned}$$

Since this relation holds true for any arbitrary x_1 and so on, we finally have

$$a(\varphi) | \psi_1, \dots, \psi_n \rangle = \sum_{k=1}^n \xi^{k-1} \langle \varphi | \psi_k \rangle | \psi_1, \dots, \psi_{k-1}, \psi_{k+1}, \dots, \psi_n \rangle \quad (1.6.2)$$

We can derive

$$a^\dagger(\varphi_1) a^\dagger(\varphi_2) = \xi a^\dagger(\varphi_2) a^\dagger(\varphi_1)$$

Namely, the following relations holds true for arbitrary ψ_1, \dots, ψ_n

$$a^\dagger(\varphi_1) a^\dagger(\varphi_2) | \psi_1, \dots, \psi_n \rangle = \xi a^\dagger(\varphi_2) a^\dagger(\varphi_1) | \psi_1, \dots, \psi_n \rangle$$

This leads to the commutation relations for boson and fermion.

$$[a^\dagger(\varphi_1), a^\dagger(\varphi_2)]_{-\xi} = 0 \quad (1.6.3)$$

where

$$[A, B]_{-\xi} = AB - \xi BA = \begin{cases} AB - BA & \text{boson} \\ AB + BA & \text{fermion} \end{cases}$$

Taking the hermite conjugate of this, we have

$$[a(\varphi_1), a(\varphi_2)]_{-\xi} = 0 \quad (1.6.4)$$

To find the commutation relations between the creation operator and destruction operator, we first calculate

$$\begin{aligned}
& a(\varphi_1) a^\dagger(\varphi_2) | \psi_1, \dots, \psi_n \rangle \\
&= a(\varphi_1) | \varphi_2, \psi_1, \dots, \psi_n \rangle \\
&= \langle \varphi_1 | \varphi_2 \rangle | \psi_1, \dots, \psi_n \rangle + \sum_{k=1}^n \xi^k \langle \varphi_1 | \psi_k \rangle | \varphi_2, \psi_1, \dots, \psi_{k-1}, \psi_{k+1}, \dots, \psi_n \rangle
\end{aligned}$$

and

$$\begin{aligned}
& a^\dagger(\varphi_2)a(\varphi_1)|\psi_1, \dots, \psi_n\rangle \\
&= \sum_{k=1}^n \xi^{k-1} \langle \varphi_1 | \psi_k \rangle a^\dagger(\varphi_2) |\psi_k\rangle |\varphi_2, \psi_1, \dots, \psi_{k-1}, \psi_{k+1}, \dots, \psi_n\rangle \\
&= \sum_{k=1}^n \xi^{k-1} \langle \varphi_1 | \psi_k \rangle |\varphi_2, \psi_k\rangle |\varphi_2, \psi_1, \dots, \psi_{k-1}, \psi_{k+1}, \dots, \psi_n\rangle
\end{aligned}$$

Multiply this by ξ and subtract it from the former one, we see that

$$[a(\varphi_1)a^\dagger(\varphi_2) - \xi a^\dagger(\varphi_2)a(\varphi_2)]|\psi_1, \dots, \psi_n\rangle = \langle \varphi_1 | \varphi_2 \rangle |\psi_1, \dots, \psi_n\rangle$$

Equivalently, we have

$$[a(\varphi_1), a^\dagger(\varphi_2)]_{-\xi} = \langle \varphi_1 | \varphi_2 \rangle \quad (1.6.5)$$

These three relations are the fundamental “commutation relations” for creation & destruction operators. When these operators are introduced in terms of an orthonormal basis, this relation becomes

$$[a(\alpha), a^\dagger(\beta)]_{-\xi} = \langle \alpha | \beta \rangle = \delta_{\alpha\beta} \quad (1.6.6)$$

Where one-particle states $|\alpha\rangle$ and $|\beta\rangle$ are from a complete orthonormal set of one-particle states, $\{|1\rangle, |2\rangle, \dots\}$

Let us next compare these commutation relations with raising and lowering operators for linear harmonic oscillators.

For boson case, we further define an orthonormal basis for the multiparticle space as

$$|n_1, n_2, \dots\rangle = \frac{|1, \dots, 1, 2, \dots, 2, \dots\rangle}{\sqrt{n_1! n_2! \dots}} \quad (1.6.7)$$

Where we use $\{|1\rangle, |2\rangle, \dots\}$ as a complete orthonormal set of one-particle states and n_α is the number of times α appear in the ket state in the right hand side. In terms of this orthonormal basis, eqs (1.6.1), (1.6.2) reduce to the followings:

$$a^\dagger(\alpha)|n_1, n_2, \dots, n_\alpha, \dots\rangle = \sqrt{n_\alpha + 1}|n_1, n_2, \dots, n_\alpha + 1, \dots\rangle \quad (1.6.8)$$

$$a(\alpha)|n_1, n_2, \dots, n_\alpha, \dots\rangle = \sqrt{n_\alpha}|n_1, n_2, \dots, n_\alpha - 1, \dots\rangle \quad (1.6.9)$$

while the commutation relations are

$$\begin{aligned}
[a(\alpha), a(\beta)] &= [a^\dagger(\alpha), a^\dagger(\beta)] = 0 \\
[a(\alpha), a^\dagger(\beta)] &= \delta_{\alpha\beta}
\end{aligned} \quad (1.6.10)$$

These three equations are identical to eqs (1.3.2), (1.3.3), (1.3.4)

This means that creation & destruction operators for a system of Bose particles looks exactly like those for a system of linear harmonic oscillators.

For fermions case, we have

$$\begin{aligned}
a_{\alpha}^{\dagger}|\alpha_1, \dots, \alpha_n\rangle &= |\alpha, \alpha_1, \dots, \alpha_n\rangle \\
a_{\alpha}|\alpha_1, \dots, \alpha_n\rangle &= \sum_{k=1}^n (-1)^{k-1} \delta_{\alpha\alpha_k} |\alpha_1, \dots, \alpha_{k-1}, \alpha_{k+1}, \dots, \alpha_n\rangle
\end{aligned} \tag{1.6.11}$$

$$\begin{cases} [a_{\alpha}, a_{\beta}]_{+} = [a_{\alpha}^{\dagger}, a_{\beta}^{\dagger}] = 0 \\ [a_{\alpha}, a_{\beta}^{\dagger}]_{+} = \delta_{\alpha\beta} \end{cases} \tag{1.6.12}$$

Contrary to the boson case, there is no a priori reason for postulating these anticommutation relations. These relations can be derived only from the antisymmetrization postulate for fermions.

Let us put some remarks on these general commutation relations.

One advantage of deriving them in such a general form is that we are not tied down to a particular basis of one-particle states. Suppose we use a basis of momentum eigenstates, $|\mathbf{p}\rangle$ with $\langle\mathbf{p}|\mathbf{p}'\rangle = (2\pi)^3\delta^3(\mathbf{p} - \mathbf{p}')$. Then we have

$$\begin{aligned}
[a(\mathbf{p}), a^{\dagger}(\mathbf{p}')]_{-\xi} &= (2\pi)^3\delta^3(\mathbf{p} - \mathbf{p}') \\
[a(\mathbf{p}), a(\mathbf{p}')]_{-\xi} &= [a^{\dagger}(\mathbf{p}), a^{\dagger}(\mathbf{p}')]_{-\xi} = 0
\end{aligned} \tag{1.6.13}$$

A complete orthogonal sets of states in the multiparticle space can be constructed from the vacuum state

$$|\mathbf{p}_1, \mathbf{p}_2, \dots, \mathbf{p}_n\rangle = a^{\dagger}(\mathbf{p}_1) \cdots a^{\dagger}(\mathbf{p}_n)|vac\rangle$$

If we use a basis of position eigenstates $|\mathbf{x}\rangle$

1.7 Hamiltonian and other operators of creation and destruction operators

1.8 Degenerate electron gas

Chapter 2

Green's function and field theory at zero temperature

2.1 Schrodinger, Heisenberg & interaction picture

2.2 Adiabatic switching & Gell-Mann & Low Theorem

2.3 Green's function

2.4 Wick's theorem

2.5 Diagrammatic analysis of perturbation theory (Fermion Case)

2.6 (Self-Consist) Hartree-Fock Approximation

The first approximation scheme is the self-consistent Hartree-Fock approximation, which begins with the 1st order approximation for the proper selfenergy.

The diagram shows a fermion line (a vertical line with diagonal hatching) labeled with x_i at the top and x_i' at the bottom. This is followed by an approximation symbol \approx , then a summation $\sum_{(1)} \star$ (where \star is a star symbol). This is followed by an equals sign, then two terms separated by a plus sign. The first term is a fermion line with a self-energy loop (a circle with a wavy line inside). The second term is a fermion line with a self-energy loop (a circle with a wavy line inside).

In this approximation, a particle moves in a single particle potential that is created by all the other particles by way of interaction potential. But, in this approximation, these internal fermion lines are just non-interacting Green's function, so that background particles which create an effective single-particle potential are treated as non-interacting. But, in reality, the background particles also move under the effective single-particle potential. Therefore, it is much more natural to replace these non-interacting Green's functions by the interacting Green's function itself. Namely,



where these bold lines themselves are given by this proper selfenergy via Dyson equation,



We will study these equations in details.

To this end, we will consider a spin= $\frac{1}{2}$ fermion system in a static spin-independent potential $U(\mathbf{x})$.

$$\begin{aligned}\hat{H}_0 &= \int d^3x \psi_\alpha^\dagger(x) \left[-\frac{\hbar^2 \nabla^2}{2m} + U(\mathbf{x}) \right] \psi_\alpha(x) \\ \hat{H}_1 &= \frac{1}{2} \int d^3x d^3x' \psi_\alpha^\dagger(x) \psi_\beta^\dagger(x') V(x-x') \psi_\beta(x') \psi_\alpha(x)\end{aligned}$$

Here we assumed the spin-independent interaction potential for simplicity.

Within the self-consistent Hatree-Fock approximation, the Dyson equation is given by these two equations:

$$\left\{ \begin{array}{l} G(x, y) = G^0(x, y) + \int d^4x_1 d^4x'_1 G^0(x, x_1) \Sigma^\star(x_1, x'_1) G(x'_1, y) \\ \hbar \Sigma^\star(x_1, x'_1) = -i\delta(t_1 - t'_1) \times \left[\delta(\mathbf{x}_1 - \mathbf{x}'_1) 2 \int d^3x_2 \left(\begin{array}{l} G(\mathbf{x}_2, t_2; \mathbf{x}_2, t_2 + 0) V(\mathbf{x}_1 - \mathbf{x}_2) \\ -G(\mathbf{x}_1, t_1; \mathbf{x}'_1, t_1 + 0) V(\mathbf{x}_1 - \mathbf{x}'_1) \end{array} \right) \right] \end{array} \right.$$

Where the factor 2 comes from the spin summations.

Now, it is convenient to use a Fourier representation.

$$\begin{aligned}G(\mathbf{x}, t; \mathbf{x}', t') &= \frac{1}{2\pi} \int d\omega e^{-i\omega(t-t')} G(\mathbf{x}, \mathbf{x}'; \omega) \\ G^0(\mathbf{x}, t; \mathbf{x}', t') &= \frac{1}{2\pi} \int d\omega e^{-i\omega(t-t')} G^0(\mathbf{x}, \mathbf{x}'; \omega) \\ \Sigma^\star(\mathbf{x}, t; \mathbf{x}', t') &= \delta(t - t') \Sigma^\star(\mathbf{x}, \mathbf{x}') \\ &= \frac{1}{2\pi} \int d\omega e^{-i\omega(t-t')} \Sigma^\star(\mathbf{x}, \mathbf{x}')\end{aligned}$$

$$G(\mathbf{x}, \mathbf{y}; \omega) = G^0(\mathbf{x}, \mathbf{y}; \omega) + \int d^3\mathbf{x}_1 d^3\mathbf{x}'_1 G^0(\mathbf{x}, \mathbf{x}_1; \omega) \times \Sigma^\star(\mathbf{x}_1, \mathbf{x}'_1) G(\mathbf{x}'_1, \mathbf{y}, \omega) \quad (2.6.1)$$

$$\begin{aligned}\hbar \Sigma^\star(\mathbf{x}_1, \mathbf{x}'_1) &= -i2\delta(\mathbf{x}_1 - \mathbf{x}'_1) \int d^3\mathbf{x}_2 V(\mathbf{x}_1 - \mathbf{x}_2) \times \frac{1}{2\pi} \int d\omega e^{i\omega 0t} G(\mathbf{x}_2, \mathbf{x}_2; \omega) \\ &\quad + iV(\mathbf{x}_1 - \mathbf{x}'_1) \frac{1}{2\pi} \int d\omega e^{i\omega 0t} G(\mathbf{x}_1, \mathbf{x}'_1; \omega)\end{aligned} \quad (2.6.2)$$

Let us also introduce the orthonormal eigenfunctions of H_0 .

$$H_0 \varphi_j^0(\mathbf{x}) \equiv \left[-\frac{\hbar^2 \nabla^2}{2m} + U(\mathbf{x}) \right] \varphi_j^0(\mathbf{x}) = \epsilon_j \varphi_j^0(\mathbf{x})$$

In terms of this eigenfunctions, the field operator in the interaction representation can be expanded as

$$\begin{cases} \psi(\mathbf{x}, t) = \sum_j \varphi_j(\mathbf{x}) c_j e^{-i \frac{\epsilon_j t}{\hbar}} \\ \psi^\dagger(\mathbf{x}, t) = \sum_j \varphi_j^*(\mathbf{x}) c_j^\dagger e^{+i \frac{\epsilon_j t}{\hbar}} \end{cases}$$

where we used

$$e^{+i \frac{H_0 t}{\hbar}} c_j e^{-i \frac{H_0 t}{\hbar}} = e^{-i \frac{\epsilon_j t}{\hbar}} c_j$$

In terms of them, the non-interacting Green's function can be calculated as

$$\begin{aligned} iG^0(\mathbf{x}, t; \mathbf{x}', t') &= \langle \Phi_0 | \Gamma[\psi_I(\mathbf{x}, t) \psi_I^\dagger(\mathbf{x}', t')] | \Phi_0 \rangle \\ &= \theta(t - t') \langle \Phi_0 | \psi_I(\mathbf{x}, t) \psi_I^\dagger(\mathbf{x}', t') | \Phi_0 \rangle \\ &\quad - \theta(t' - t) \langle \Phi_0 | \psi_I^\dagger(\mathbf{x}', t') \psi_I(\mathbf{x}, t) | \Phi_0 \rangle \\ &= \theta(t - t') \sum_j \varphi_j^0(\mathbf{x}) \varphi_j^{0*}(\mathbf{x}') e^{-i \frac{\epsilon_j}{\hbar}(t-t')} \times \langle \Phi_0 | c_j c_j^\dagger | \Phi_0 \rangle \\ &\quad - \theta(t' - t) \sum_j \varphi_j^0(\mathbf{x}) \varphi_j^{0*}(\mathbf{x}') e^{-i \frac{\epsilon_j}{\hbar}(t-t')} \times \langle \Phi_0 | c_j^\dagger c_j | \Phi_0 \rangle \\ &= \sum_j \varphi_j^0(\mathbf{x}) \varphi_j^{0*}(\mathbf{x}') e^{-i \frac{\epsilon_j}{\hbar}(t-t')} \times \left\{ \theta(t - t') \langle \Phi_0 | c_j c_j^\dagger | \Phi_0 \rangle - \theta(t' - t) \langle \Phi_0 | c_j^\dagger c_j | \Phi_0 \rangle \right\} \\ &= \sum_j \varphi_j^0(\mathbf{x}) \varphi_j^{0*}(\mathbf{x}') e^{-i \frac{\epsilon_j}{\hbar}(t-t')} \times \left\{ \theta(t - t') \theta(\epsilon_j^0 - \epsilon_F^0) - \theta(t' - t) \theta(-\epsilon_j^0 + \epsilon_F^0) \right\} \end{aligned}$$

where ϵ_F^0 is the energy of the highest filled state.

The Fourier transformation leads to

$$G^0(\mathbf{x}, \mathbf{x}'; \omega) = \sum_j \varphi_j^0(\mathbf{x}) \varphi_j^{0*}(\mathbf{x}') \times \left[\frac{\theta(\epsilon_j^0 - \epsilon_F^0)}{\omega - \hbar^{-1} \epsilon_j^0 + i\eta} + \frac{\theta(\epsilon_F^0 - \epsilon_j^0)}{\omega - \hbar^{-1} \epsilon_j^0 - i\eta} \right] \quad (2.6.3)$$

The particle density at \mathbf{x} in the unperturbed ground state is estimated as

$$\begin{aligned} n^0(\mathbf{x}) &= -i 2 \frac{1}{2\pi} \int d\omega e^{i\omega 0t} G^0(\mathbf{x}, \mathbf{x}; \omega) \\ &= 2 \sum_j |\varphi_j^0(\mathbf{x})|^2 \theta(\epsilon_F^0 - \epsilon_j^0) \end{aligned}$$

where the factor 2 comes from spin.

The total number of fermions is given by its spatial integral:

$$N = \int d^3\mathbf{x} n^0(\mathbf{x}) = 2 \sum_j \theta(\epsilon_F^0 - \epsilon_j^0)$$

where we assume the normalization of $\varphi_j^0(\mathbf{x})$

These three equations (eq.(2.6.1), eq.(2.6.2), eq.(2.6.3)) comprise a set of coupled equations for the Green's function G .

To solve these coupled equations in favor of G , let us apply a following differential operator onto eq.(2.6.1)

$$L_1 = \hbar\omega + \frac{\hbar^2 \nabla^2}{2m} - V(\mathbf{x})$$

If we apply this onto $G^0(\mathbf{x}, \mathbf{y}; \omega)$, we obtain the delta function in space,

$$\begin{aligned} L_1 G^0(\mathbf{x}, \mathbf{y}; \omega) &= \sum_j (\hbar\omega - \epsilon_j^0) \varphi_j^0(\mathbf{x}) \varphi_j^{0*}(\mathbf{y}) \times \left[\frac{\theta(\epsilon_j^0 - \epsilon_F^0)}{\omega - \hbar^{-1}\epsilon_j^0 + i\eta} + \frac{\theta(\epsilon_F^0 - \epsilon_j^0)}{\omega - \hbar^{-1}\epsilon_j^0 - i\eta} \right] \\ &= \hbar \sum_j \varphi_j^0(\mathbf{x}) \varphi_j^{0*}(\mathbf{y}) = \hbar \delta^3(\mathbf{x} - \mathbf{y}) \end{aligned}$$

Thus, an application of L_1 onto eq.(2.6.1) leads to

$$L_1 G(\mathbf{x}, \mathbf{y}; \omega) = \hbar \delta^3(\mathbf{x} - \mathbf{y}) + \int d^3 \mathbf{x}' \hbar \Sigma^\star(\mathbf{x}, \mathbf{x}') G(\mathbf{x}', \mathbf{y}; \omega)$$

or equivalently

$$\left[\hbar\omega + \frac{\hbar^2 \nabla_{\mathbf{x}}^2}{2m} - V(\mathbf{x}) \right] G(\mathbf{x}, \mathbf{y}; \omega) - \int d^3 \mathbf{x}' \hbar \Sigma^\star(\mathbf{x}, \mathbf{x}') G(\mathbf{x}', \mathbf{y}; \omega) = \hbar \delta^3(\mathbf{x} - \mathbf{y}) \quad (2.6.4)$$

Generally, such a linearized integral equation can be solved in terms of following set of eigenstates $\varphi_j(\mathbf{x})$

$$\left[-\frac{\hbar^2 \nabla_{\mathbf{x}}^2}{2m} + V(\mathbf{x}) \right] \varphi_j(\mathbf{x}) + \int d^3 \mathbf{x}' \hbar \Sigma^\star(\mathbf{x}, \mathbf{x}') \varphi_j(\mathbf{x}') = \epsilon_j \varphi_j(\mathbf{x}) \quad (2.6.5)$$

$$\begin{aligned} iG_{\alpha\beta}(\mathbf{x}, t; \mathbf{x}', t') &= \sum_n \{ \theta(t - t') e^{\frac{-i(E_n - E)(t - t')}{\hbar}} \times \langle \Psi_0 | \psi_\alpha(\mathbf{x}) | \Psi_n \rangle \langle \Psi_n | \psi_\beta^\dagger(\mathbf{x}') | \Psi_0 \rangle \\ &\quad - \theta(t' - t) e^{\frac{-i(E_n - E)(t - t')}{\hbar}} \times \langle \Psi_0 | \psi_\beta^\dagger(\mathbf{x}') | \Psi_n \rangle \langle \Psi_n | \psi_\alpha(\mathbf{x}) | \Psi_0 \rangle \} \end{aligned}$$

$$\begin{aligned} G_{\alpha\beta}(\mathbf{x}, \mathbf{x}'; \omega) &= \sum_n \left[\frac{\langle \Psi_0 | \psi_\alpha(\mathbf{x}) | \Psi_n \rangle \langle \Psi_n | \psi_\beta^\dagger(\mathbf{x}') | \Psi_0 \rangle}{\omega - \hbar^{-1}(E_n - E) + i\eta} \right. \\ &\quad \left. + \frac{\langle \Psi_0 | \psi_\beta^\dagger(\mathbf{x}') | \Psi_n \rangle \langle \Psi_n | \psi_\alpha(\mathbf{x}) | \Psi_0 \rangle}{\omega - \hbar^{-1}(E_n - E) - i\eta} \right] \end{aligned}$$

$$\begin{aligned} \frac{1}{2\pi} \int_{-\infty}^{+\infty} d\omega G_{\alpha\beta}(\mathbf{x}, \mathbf{x}'; \omega) e^{i\omega 0t} &= i \sum_n \langle \Psi_0 | \psi_\beta^\dagger(\mathbf{x}') | \Psi_n \rangle \langle \Psi_n | \psi_\alpha(\mathbf{x}) | \Psi_0 \rangle \\ \left(i \int_{-\infty}^{+\infty} d\omega G(\mathbf{x}, \mathbf{x}'; \omega) e^{i\omega 0t} \right)^* &= i \int_{-\infty}^{+\infty} d\omega G(\mathbf{x}', \mathbf{x}; \omega) e^{i\omega 0t} \end{aligned}$$

where $\hbar \Sigma^\star$ play role of a static but no-local potential.

Since Σ^\star is hermitian,

$$\Sigma^{\star\star}(\mathbf{x}, \mathbf{x}') = \Sigma^\star(\mathbf{x}', \mathbf{x})$$

$\{\varphi_j\}$ comprises complete set of orthonormal eigenfunctions with

$$\sum_j \varphi_j(\mathbf{x}) \varphi_j^*(\mathbf{x}') = \delta^3(\mathbf{x} - \mathbf{x}')$$

$$\int d^3 \mathbf{x} \varphi_j^*(\mathbf{x}) \varphi_m(\mathbf{x}) = \delta_{jm}$$

and ϵ_j is real.

In terms of this eigenstates, the Green's function satisfying eq.(2.6.4) is roughly given in a following form:

$$G(\mathbf{x}, \mathbf{y}; \omega) \sim \sum_j \varphi_j(\mathbf{x}) \varphi_j^*(\mathbf{y}) \cdot \frac{1}{\omega - \hbar^{-1} \epsilon_j}$$

According to eq.(??), the particle density and total number of the particles are given by the frequency integral of this Green's function.

$$n(\mathbf{x}) = -i \cdot 2 \cdot \frac{1}{2\pi} \int_{-\infty}^{+\infty} d\omega e^{i\omega 0t} G(\mathbf{x}, \mathbf{x}; \omega)$$

$$N = -i \cdot 2 \cdot \frac{1}{2\pi} \int d^3\mathbf{x} \int_{-\infty}^{+\infty} d\omega e^{i\omega 0t} G(\mathbf{x}, \mathbf{x}; \omega)$$

But, since the perturbation \hat{H}_1 conserves the total number of particles, the number of particles thus calculated should be equal to the number of particles in the unperturbed ground state. To make it possible, the denominator in the r.h.s. of this equation should acquire an infinitesimally small imaginary part like this:

$$G(\mathbf{x}, \mathbf{y}; \omega) = \sum_j \varphi_j(\mathbf{x}) \varphi_j^*(\mathbf{y}) \left[\frac{\theta(\epsilon_j - \epsilon_F)}{\omega - \hbar^{-1} \epsilon_j + i\eta} + \frac{\theta(\epsilon_F - \epsilon_j)}{\omega - \hbar^{-1} \epsilon_j - i\eta} \right] \quad (\eta > 0) \quad (2.6.6)$$

ϵ_F here denote the energy of the highest filled state of these orthonormal eigenstates, which satisfies

$$\sum_j \theta(\epsilon_F - \epsilon_j) = N \equiv \sum_j \theta(\epsilon_F^0 - \epsilon_j^0)$$

substituting this solution back into eq.(2.6.2), we obtain the self-energy itself, which is given by these orthonormal eigenstates.

$$\begin{aligned} \hbar \Sigma^\star(\mathbf{x}_1, \mathbf{x}_1') &= 2 \delta(\mathbf{x}_1 - \mathbf{x}_1') \int d^3\mathbf{x}_2 V(\mathbf{x}_1 - \mathbf{x}_2) \times \sum_j |\varphi_j(\mathbf{x}_2)|^2 \theta(\epsilon_F - \epsilon_j) \\ &\quad - V(\mathbf{x}_1 - \mathbf{x}_1') \sum_j \varphi_j(\mathbf{x}_1) \varphi_j^*(\mathbf{x}_1') \theta(\epsilon_F - \epsilon_j) \\ &= \delta(\mathbf{x}_1 - \mathbf{x}_1') \int d^3\mathbf{x}_2 V(\mathbf{x}_1 - \mathbf{x}_2) n(\mathbf{x}_2) \quad (\text{Fock term}) \\ &\quad - V(\mathbf{x}_1 - \mathbf{x}_1') \sum_j \varphi_j(\mathbf{x}_1) \varphi_j^*(\mathbf{x}_1') \theta(\epsilon_F - \epsilon_j) \quad (\text{Hartree term}) \end{aligned} \quad (2.6.7)$$

Within this HF approximation, the self-energy consists of two terms. One is a local term propotional to particle density, and the other is non-local exchange term. These two are often called as Hartree term and Fock term respectively.

To summarize, eq.(2.6.5) and eq.(2.6.7) comprise a self-consistent equations for the orthonormal eigenstates φ_j and its eigenvalue ϵ_j .

Namely, once φ_j and ϵ_j are given, the self-energy Σ^\star are calculated by eq.(2.6.7). In terms of this self-energy, the orthonormal eigenstates φ_j and ϵ_j are recalculated. This process is repeated, until a self-consistent solution for $\{\varphi_j\}$ and $\{\epsilon_j\}$ is obtained.

Once the self-consistent solution is obtained, we can calculate Green's function from eq.(2.6.6), from which we can calculate various physical quantities such as ground state energy and ground state expectation value of single particle operators.

This is all about the Hartree-Fock approximation, which is quite commonly used in condensed matter physics research.

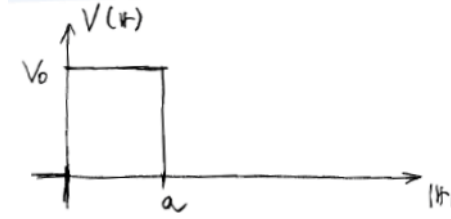
In the following two sections, I will describe two other standard approximation scheme commonly used in the contemporary research.

2.7 Imperfect Fermi Gas

The first system considered is a dilute Fermi gas with strong short-range repulsive interaction, which is sometimes called as “imperfect Fermi gas”.

We assume that the repulsive interaction is s-wave like, so that the potential depends only on a spatial distance between two particles.

To be specific, we assume that the repulsive interaction is described by a square-well shape potential.



where positive V_0 means repulsive and “ a ” represents an interaction range.

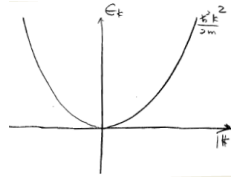
“strong short-range repulsive interaction” means that “ V_0 ” will be taken to be sufficiently large, but “ a ” will be taken to be small.

$$\begin{cases} V_0 \rightarrow \text{large} \\ a \rightarrow \text{small} \end{cases}$$

For the kinetic energy part, we simply assume the free electron Hamiltonian

$$H_0 = \sum_{\alpha} \int d^3\mathbf{x} \psi_{\alpha}^{\dagger}(\mathbf{x}) \left[-\frac{\hbar^2 \nabla_{\mathbf{x}}^2}{2m} \right] \psi_{\alpha}(\mathbf{x})$$

whose dispersion take a parabolic form.



“dilute Fermi gas” means that the chemical potential is chosen so that the Fermi momentum is sufficiently small.

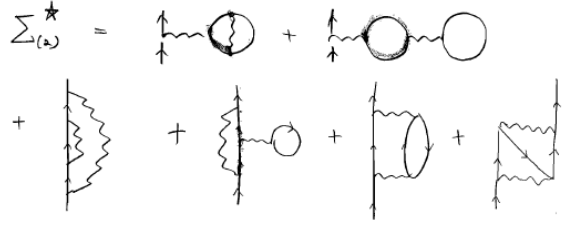
$$k_F \rightarrow \text{small}$$

Now $k_F a$ comprises a dimensionless quantity. The natural coupling constant for “imperfect Fermi gas” is “ $k_F a$ ”. Namely, we are going to analyze the “imperfect Fermi gas” by treating “ $k_F a$ ” as a small parameter. Thereby, physical quantities such as ground state energy, chemical potential and quasi-particle spectrum are calculated, perturbatively in small $k_F a$.

$$\text{e.g.) } \frac{E}{N} = \frac{\hbar^2 k_F^2}{2m} \left[A + B(k_F a) + C(k_F a)^2 + \dots \right]$$

To do this perturbation expansion, systematically, let us begin with the 2nd order contribution to the proper self energy.

According to eq.(2.6.4), the 2nd order contribution can be enumerated as follows:



For the uniform electron gas case, these four term exactly reduce to zero.

For example, consider this diagram.

The self energy part can be calculated as this.



$$\left(\frac{i}{\hbar}\right)^2 \frac{1}{(2\pi)^8} \int d^3\mathbf{q} \int d\epsilon \int d^3\mathbf{q}' \int d\epsilon' \times \left\{ V(\mathbf{q} - \mathbf{q}') V(\mathbf{k} - \mathbf{q}) e^{i\epsilon' 0t} G_0(\mathbf{q}', \epsilon') \times [G_0(\mathbf{q}, \epsilon)]^2 \right\}$$

where

$$G_0(\mathbf{q}, \epsilon) \equiv \frac{\theta(|\mathbf{q}| - k_F)}{\epsilon - \epsilon_{\mathbf{q}} + i\eta} + \frac{\theta(k_F - |\mathbf{q}|)}{\epsilon - \epsilon_{\mathbf{q}} - i\eta}$$

Putting this into here, this part becomes like this:

$$\int \frac{d\epsilon'}{2\pi} \left[\frac{\theta(|\mathbf{q}| - k_F)}{\epsilon' - \epsilon_{\mathbf{q}'} + i\eta} + \frac{\theta(k_F - |\mathbf{q}|)}{\epsilon' - \epsilon_{\mathbf{q}'} - i\eta} \right]^2$$

Since

$$\theta(|\mathbf{q}| - k_F) \theta(k_F - |\mathbf{q}|) = 0$$

the cross term vanishes, so that we have

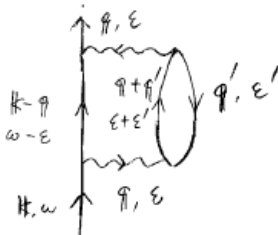
$$\int \frac{d\epsilon'}{2\pi} \left[\frac{\theta(|\mathbf{q}| - k_F)}{(\epsilon' - \epsilon_{\mathbf{q}'} + i\eta)^2} + \frac{\theta(k_F - |\mathbf{q}|)}{(\epsilon' - \epsilon_{\mathbf{q}'} - i\eta)^2} \right]$$

Now we can integrate this, by closing the contour either in the upper half plane or in the lower half plane (By using the method of contour integration).

But, since both of these two are the second order pole, the contour integration always reduce to zero, irrespective of whether the integration is carried out in the upper-half plane or in the lower-half plane.

Therefore, this contribution always reduce to zero. Similarly, we can immediately see that all of these also reduce to zero, because these parts give the same integration as this.

Thus, we have only to calculate these two contributions,



$$-\left(\frac{i}{\hbar}\right)^2 \frac{1}{(2\pi)^8} \int d^3\mathbf{q} \int d\epsilon \int d^3\mathbf{q}' \int d\epsilon' \times \left\{ V^2(\mathbf{q}) G_0(\mathbf{k} - \mathbf{q}, \omega - \epsilon) \times G_0(\mathbf{q} + \mathbf{q}', \epsilon + \epsilon') G_0(\mathbf{q}', \epsilon') \right\}$$

Let me first integrate over these internal frequencies by using the method of contour integration.

$$\begin{aligned} & \int \frac{d\epsilon}{2\pi} \int \frac{d\epsilon'}{2\pi} \left[\frac{\theta(|\mathbf{k}-\mathbf{q}|-k_F)}{\omega-\epsilon-\epsilon_{\mathbf{k}-\mathbf{q}}+\eta} + \frac{\theta(k_F-|\mathbf{k}-\mathbf{q}|)}{\omega-\epsilon-\epsilon_{\mathbf{k}-\mathbf{q}}-\eta} \right] \\ & \times \left[\frac{\theta(|\mathbf{q}+\mathbf{q}'|-k_F)}{\epsilon+\epsilon'-\epsilon_{\mathbf{q}+\mathbf{q}'}+\eta} + \frac{\theta(k_F-|\mathbf{q}+\mathbf{q}'|)}{\epsilon+\epsilon'-\epsilon_{\mathbf{q}+\mathbf{q}'}-\eta} \right] \\ & \times \left[\frac{\theta(|\mathbf{q}'|-k_F)}{\epsilon'-\epsilon_{\mathbf{q}'}+\eta} + \frac{\theta(k_F-|\mathbf{q}'|)}{\epsilon'-\epsilon_{\mathbf{q}'}-\eta} \right] \end{aligned}$$

I will first integrate over ϵ , whose integrand is the product between these two:

$$\left[\frac{\theta(|\mathbf{k}-\mathbf{q}|-k_F)}{\omega-\epsilon-\epsilon_{\mathbf{k}-\mathbf{q}}+\eta} + \frac{\theta(k_F-|\mathbf{k}-\mathbf{q}|)}{\omega-\epsilon-\epsilon_{\mathbf{k}-\mathbf{q}}-\eta} \right] \times \left[\frac{\theta(|\mathbf{q}+\mathbf{q}'|-k_F)}{\epsilon+\epsilon'-\epsilon_{\mathbf{q}+\mathbf{q}'}+\eta} + \frac{\theta(k_F-|\mathbf{q}+\mathbf{q}'|)}{\epsilon+\epsilon'-\epsilon_{\mathbf{q}+\mathbf{q}'}-\eta} \right]$$

From this product, we have four terms.

But these two reduce to zero, when the contour integration over ϵ is carried out.

$$\frac{\dots}{\omega-\epsilon-\epsilon_{\mathbf{k}-\mathbf{q}}\pm\eta} \cdot \frac{\dots}{\epsilon+\epsilon'-\epsilon_{\mathbf{q}+\mathbf{q}'}\mp\eta}$$

Namely these two have a pole in the upper half plane, while these two have a pole in the lower half plane. Thus, for this term, we can close the contour in the lower half plane, while for this term, we can close the contour in the upper half plane, which reduce them to zero under the integration.

On the one hand, these combinations always have a pole both in the upper half plane and in the lower half plane.

$$\frac{\dots}{\omega-\epsilon-\epsilon_{\mathbf{k}-\mathbf{q}}\pm\eta} \cdot \frac{\dots}{\epsilon+\epsilon'-\epsilon_{\mathbf{q}+\mathbf{q}'}\pm\eta}$$

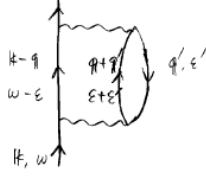
Thus, irrespective of how the contour is closed, these two combinations give finite contribution.

$$\begin{aligned} & \int \frac{d\epsilon}{2\pi} \left\{ \frac{\theta(|\mathbf{k}-\mathbf{q}|-k_F)}{\omega-\epsilon-\epsilon_{\mathbf{k}-\mathbf{q}}+\eta} \cdot \frac{\theta(|\mathbf{q}+\mathbf{q}'|-k_F)}{\epsilon+\epsilon'-\epsilon_{\mathbf{q}+\mathbf{q}'}+\eta} + \frac{\theta(k_F-|\mathbf{k}-\mathbf{q}|)}{\omega-\epsilon-\epsilon_{\mathbf{k}-\mathbf{q}}-\eta} \cdot \frac{\theta(k_F-|\mathbf{q}+\mathbf{q}'|)}{\epsilon+\epsilon'-\epsilon_{\mathbf{q}+\mathbf{q}'}-\eta} \right\} \\ & = -i \cdot \frac{\theta(|\mathbf{k}-\mathbf{q}|-k_F)\theta(|\mathbf{q}+\mathbf{q}'|-k_F)}{\omega-\epsilon_{\mathbf{k}-\mathbf{q}}+\epsilon'-\epsilon_{\mathbf{q}+\mathbf{q}'}+2\eta} + i \cdot \frac{\theta(k_F-|\mathbf{k}-\mathbf{q}|)\theta(k_F-|\mathbf{q}+\mathbf{q}'|)}{\omega-\epsilon_{\mathbf{k}-\mathbf{q}}+\epsilon'-\epsilon_{\mathbf{q}+\mathbf{q}'}-2\eta} \end{aligned}$$

Let me next integrate over ϵ' , by substituting this into here. From this product, we again have four terms.

Now these two have a pole in the upper half plane, while these two have a pole in the lower half plane. Due to the same reason, only these two combinations give us a finite contribution.

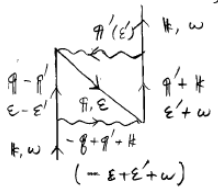
$$\begin{aligned} & (-i) \left\{ \int \frac{d\epsilon'}{2\pi} \frac{\theta(|\mathbf{k}-\mathbf{q}|-k_F)\theta(|\mathbf{q}+\mathbf{q}'|-k_F)}{\epsilon'+\omega-\epsilon_{\mathbf{k}-\mathbf{q}}-\epsilon_{\mathbf{q}+\mathbf{q}'}+2\eta} \cdot \frac{\theta(k_F-|\mathbf{q}'|)}{\epsilon'-\epsilon_{\mathbf{q}'}-\eta} \right. \\ & \quad \left. - \int \frac{d\epsilon'}{2\pi} \frac{\theta(k_F-|\mathbf{k}-\mathbf{q}|)\theta(k_F-|\mathbf{q}+\mathbf{q}'|)}{\epsilon'+\omega-\epsilon_{\mathbf{k}-\mathbf{q}}-\epsilon_{\mathbf{q}+\mathbf{q}'}-2\eta} \cdot \frac{\theta(|\mathbf{q}'|-k_F)}{\epsilon'-\epsilon_{\mathbf{q}'}+\eta} \right\} \\ & = (-i) \left\{ i \frac{\theta(|\mathbf{k}-\mathbf{q}|-k_F)\theta(|\mathbf{q}+\mathbf{q}'|-k_F)\theta(k_F-|\mathbf{q}'|)}{\omega-\epsilon_{\mathbf{k}-\mathbf{q}}-\epsilon_{\mathbf{q}+\mathbf{q}'}+\epsilon_{\mathbf{q}'}+3\eta} \right. \\ & \quad \left. + i \frac{\theta(k_F-|\mathbf{k}-\mathbf{q}|)\theta(k_F-|\mathbf{q}+\mathbf{q}'|)\theta(|\mathbf{q}'|-k_F)}{\omega-\epsilon_{\mathbf{k}-\mathbf{q}}-\epsilon_{\mathbf{q}+\mathbf{q}'}+\epsilon_{\mathbf{q}'}-3\eta} \right\} \end{aligned}$$



By substituting this into the original expression we obtain this:

$$\begin{aligned}
 &= \frac{1}{\hbar^2} \int \frac{d\mathbf{q}}{(2\pi)^3} \int \frac{d\mathbf{q}'}{(2\pi)^3} \\
 &\quad \times V^2(\mathbf{q}) \left\{ \frac{\theta(|\mathbf{k} - \mathbf{q}| - k_F) \theta(|\mathbf{q} + \mathbf{q}'| - k_F) \theta(k_F - |\mathbf{q}'|)}{\omega - \epsilon_{\mathbf{k}-\mathbf{q}} - \epsilon_{\mathbf{q}+\mathbf{q}'} + \epsilon_{\mathbf{q}'}} \right. \\
 &\quad \left. + \frac{\theta(k_F - |\mathbf{k} - \mathbf{q}|) \theta(k_F - |\mathbf{q} + \mathbf{q}'|) \theta(|\mathbf{q}'| - k_F)}{\omega - \epsilon_{\mathbf{k}-\mathbf{q}} - \epsilon_{\mathbf{q}+\mathbf{q}'} + \epsilon_{\mathbf{q}'}} \right\} \quad (2.7.1)
 \end{aligned}$$

Similarly we can calculate the other 2nd order contribution to the proper self-energy.



$$\begin{aligned}
 &\left(\frac{i}{\hbar}\right)^2 \frac{1}{(2\pi)^8} \int d^3\mathbf{q} \int d\epsilon \int d^3\mathbf{q}' \int d\epsilon' \\
 &\quad \times \left\{ V(\mathbf{q}') V(-\mathbf{q} + \mathbf{q}' + \mathbf{k}) \times G_0(\mathbf{q} - \mathbf{q}', \epsilon - \epsilon') G_0(\mathbf{q}, \epsilon) G_0(\mathbf{q}' + \mathbf{k}, \epsilon' + \omega) \right\}
 \end{aligned}$$

$$\begin{aligned}
 &\int \frac{d\epsilon}{2\pi} \int \frac{d\epsilon'}{2\pi} \left[\frac{\theta(|\mathbf{q} - \mathbf{q}'| - k_F)}{\epsilon - \epsilon' - \epsilon_{\mathbf{q}-\mathbf{q}'} + \eta} + \frac{\theta(k_F - |\mathbf{q} - \mathbf{q}'|)}{\epsilon - \epsilon' - \epsilon_{\mathbf{q}-\mathbf{q}'} - \eta} \right] \\
 &\quad \times \left[\frac{\theta(|\mathbf{q}| - k_F)}{\epsilon - \epsilon_{\mathbf{q}} + \eta} + \frac{\theta(k_F - |\mathbf{q}|)}{\epsilon - \epsilon_{\mathbf{q}} - \eta} \right] \\
 &\quad \times \left[\frac{\theta(|\mathbf{q}' + \mathbf{k}| - k_F)}{\epsilon' + \omega - \epsilon_{\mathbf{q}'+\mathbf{k}} + \eta} + \frac{\theta(k_F - |\mathbf{q}' + \mathbf{k}|)}{\epsilon' + \omega - \epsilon_{\mathbf{q}'+\mathbf{k}} - \eta} \right] \\
 &= \int \frac{d\epsilon}{2\pi} \left\{ \frac{\theta(|\mathbf{q}| - k_F)}{\epsilon - \epsilon_{\mathbf{q}} + \eta} + \frac{\theta(k_F - |\mathbf{q}|)}{\epsilon - \epsilon_{\mathbf{q}} - \eta} \right\} \\
 &\quad \times (-i) \left\{ \frac{\theta(|\mathbf{q} - \mathbf{q}'| - k_F) \theta(|\mathbf{q}' + \mathbf{k}| - k_F)}{\epsilon + \omega - \epsilon_{\mathbf{q}'+\mathbf{k}} - \epsilon_{\mathbf{q}-\mathbf{q}'} + 2\eta} \right. \\
 &\quad \left. + \frac{\theta(k_F - |\mathbf{q} - \mathbf{q}'|) \theta(k_F - |\mathbf{q}' + \mathbf{k}|)}{\epsilon + \omega - \epsilon_{\mathbf{q}'+\mathbf{k}} - \epsilon_{\mathbf{q}-\mathbf{q}'} - 2\eta} \right\} \\
 &= - \left\{ \frac{\theta(|\mathbf{q}| - k_F) \theta(k_F - |\mathbf{q} - \mathbf{q}'|) \theta(k_F - |\mathbf{q}' + \mathbf{k}|)}{\omega + \epsilon_{\mathbf{q}} - \epsilon_{\mathbf{q}'+\mathbf{k}} - \epsilon_{\mathbf{q}-\mathbf{q}'} - 3\eta} \right. \\
 &\quad \left. + \frac{\theta(k_F - |\mathbf{q}|) \theta(|\mathbf{q} - \mathbf{q}'| - k_F) \theta(|\mathbf{q}' + \mathbf{k}| - k_F)}{\omega + \epsilon_{\mathbf{q}} - \epsilon_{\mathbf{q}'+\mathbf{k}} - \epsilon_{\mathbf{q}-\mathbf{q}'} + 3\eta} \right\} \\
 &\rightarrow + \int \frac{d^3\mathbf{q}}{(2\pi)^3} \int \frac{d^3\mathbf{q}'}{(2\pi)^3} \times \left[\frac{\theta(k_F - |\mathbf{q}|) \theta(|\mathbf{q} - \mathbf{q}'| - k_F) \theta(|\mathbf{q}' + \mathbf{k}| - k_F)}{\omega + \epsilon_{\mathbf{q}} - \epsilon_{\mathbf{q}'+\mathbf{k}} - \epsilon_{\mathbf{q}-\mathbf{q}'}} \right. \\
 &\quad \left. + \frac{\theta(|\mathbf{q}| - k_F) \theta(k_F - |\mathbf{q} - \mathbf{q}'|) \theta(k_F - |\mathbf{q}' + \mathbf{k}|)}{\omega + \epsilon_{\mathbf{q}} - \epsilon_{\mathbf{q}'+\mathbf{k}} - \epsilon_{\mathbf{q}-\mathbf{q}'}} \right] \quad (2.7.2)
 \end{aligned}$$

Observing these two expressions, notice that the integrands always contain the step function

$$\theta(x) = \begin{cases} 1 & x > 0 \\ 0 & x < 0 \end{cases}$$

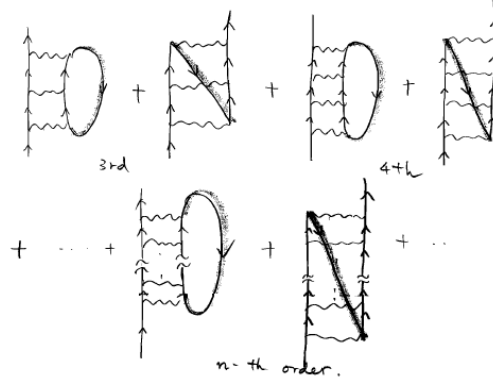
which is nothing but the fermion distribution function at the zero temperature. Thus, this step function represents whether the intermediate state is particle-like excited state or hole-like excited state.

Due to this fermi distribution function, the momentum integral regions associated with these internal lines are highly restricted in the low density limit.

For example, this step function $\theta(k_F - |\mathbf{q}|)$ restrict this momentum (\mathbf{q}) integral within a small sphere whose radius is k_F . While these two step function $\theta(k_F - |\mathbf{q} - \mathbf{q}'|)\theta(k_F - |\mathbf{q} + \mathbf{k}|)$ restrict these two momentum (\mathbf{q} and \mathbf{q}') variables within a sphere of radius k_F .

Accordingly, in the low density limit, the first term vanishes on the order of k_F , while the 2nd term vanishes on the order of k_F^2 . Thus, we can regard that the first term dominates over the 2nd term, and these two contribution gives a quantity of the order of $k_F a$ in the low-density limit.

When generalizing this argument into higher order contribution to the proper self energy, we see that the leading order contribution in the low density limit comes only from the so-called ladder-type diagrams depicted here:



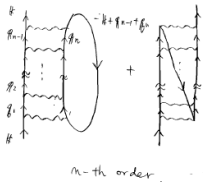
$$(2.7.3)$$

A basic feature of these ladder-type diagrams is as follows.

When all the interaction lines are put along the horizontal direction, any ladder-type diagrams contain only one internal line which runs in the reverse direction to the other internal lines.

In these diagrams, these are the ones which run in the reverse direction.

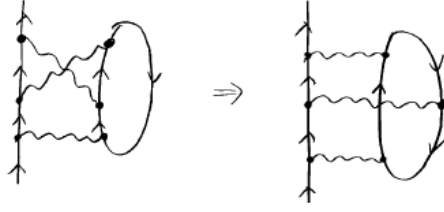
Due to this feature, the integrand of these diagram contains $(2n - 1)$ -numbers of fermi distribution function for particle times one fermi distribution function for hole.



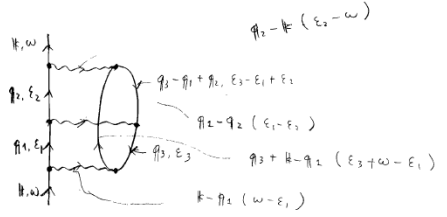
$$\begin{aligned}
 &= \int \frac{d^3 \mathbf{q}_1}{(2\pi)^3} \int \frac{d^3 \mathbf{q}_2}{(2\pi)^3} \cdots \int \frac{d^3 \mathbf{q}_n}{(2\pi)^3} V^n \\
 &\quad \underbrace{(2n - 1) \text{ number for particle lines}} \quad \underbrace{(1) \text{ hole line}} \\
 &\times \left\{ \frac{\theta(|\mathbf{q}_1| - k_F) \theta(|\mathbf{q}_2| - k_F) \cdots \theta(|\mathbf{q}_n| - k_F) \theta(k_F - |-\mathbf{k} + \mathbf{q}_{n-1} + \mathbf{q}_n|)}{(\text{energy denominator})} \right. \\
 &\quad \left. + o((k_F a)^2) \right\}
 \end{aligned}$$

Any Feynman diagrams other than these ladder-type diagram contain at least two internal lines which run in the reverse direction to the other internal lines. Therefore, their leading order term always contain at least two hole lines, so that they are at most the 2nd order of $k_F a$.

For example, consider a following 3rd order perturbation to the proper self energy term.



When all the interaction lines are put in the horizontal way, this Feynman diagram contain two internal lines which run in the reverse direction to the other.

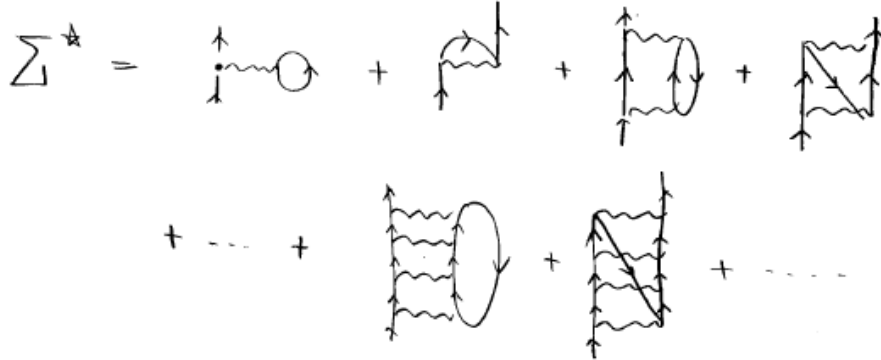


$$\begin{aligned}
 & - \left(\frac{i}{\hbar}\right)^3 \frac{1}{(2\pi)^{12}} \int d^3\mathbf{q}_1 \int d\epsilon_1 \int d^3\mathbf{q}_2 \int d\epsilon_2 \int d^3\mathbf{q}_3 \int d\epsilon_3 \\
 & \times V(\mathbf{k} - \mathbf{q}_1) V(\mathbf{q}_1 - \mathbf{q}_2) V(\mathbf{q}_2 - \mathbf{k}) \\
 & \times G_0(\mathbf{q}_1, \epsilon_1) G_0(\mathbf{q}_2, \epsilon_2) G_0(\mathbf{q}_3, \epsilon_3) \\
 & \times G_0(\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{q}_2, \epsilon_3 - \epsilon_1 + \epsilon_2) G_0(\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{k}, \epsilon_3 - \epsilon_1 + \omega)
 \end{aligned}$$

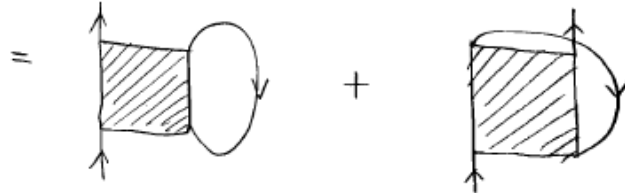
$$\begin{aligned}
 & \int \frac{d\epsilon_1}{2\pi} \int \frac{d\epsilon_2}{2\pi} \int \frac{d\epsilon_3}{2\pi} \left[\frac{\theta(|\mathbf{q}_1| - k_F)}{\epsilon_1 - \epsilon_{\mathbf{q}_1} + \eta} + \frac{\theta(k_F - |\mathbf{q}_1|)}{\epsilon_1 - \epsilon_{\mathbf{q}_1} - \eta} \right] \\
 & \times \left[\frac{\theta(|\mathbf{q}_2| - k_F)}{\epsilon_2 - \epsilon_{\mathbf{q}_2} + \eta} + \frac{\theta(k_F - |\mathbf{q}_2|)}{\epsilon_2 - \epsilon_{\mathbf{q}_2} - \eta} \right] \times \left[\frac{\theta(|\mathbf{q}_3| - k_F)}{\epsilon_3 - \epsilon_{\mathbf{q}_3} + \eta} + \frac{\theta(k_F - |\mathbf{q}_3|)}{\epsilon_3 - \epsilon_{\mathbf{q}_3} - \eta} \right] \\
 & \times \left[\frac{\theta(|\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{q}_2| - k_F)}{\epsilon_3 - \epsilon_1 + \epsilon_2 - \epsilon_{\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{q}_2} + \eta} + \frac{\theta(k_F - |\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{q}_2|)}{\epsilon_3 - \epsilon_1 + \epsilon_2 - \epsilon_{\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{q}_2} - \eta} \right] \\
 & \times \left[\frac{\theta(|\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{k}| - k_F)}{\epsilon_3 - \epsilon_1 + \omega - \epsilon_{\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{k}} + \eta} + \frac{\theta(k_F - |\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{k}|)}{\epsilon_3 - \epsilon_1 + \omega - \epsilon_{\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{k}} - \eta} \right] \\
 & = \int \frac{d\epsilon_1}{2\pi} \int \frac{d\epsilon_2}{2\pi} \int \frac{d\epsilon_3}{2\pi} \left[\frac{\theta(|\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{q}_2| - k_F) \theta(k_F - |\mathbf{q}_2|)}{\epsilon_3 - \epsilon_1 + \epsilon_{\mathbf{q}_2} - \epsilon_{\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{q}_2} + 2\eta} \right. \\
 & \quad \left. - \frac{\theta(k_F - |\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{q}_2|) \theta(|\mathbf{q}_2| - k_F)}{\epsilon_3 - \epsilon_1 + \epsilon_{\mathbf{q}_2} - \epsilon_{\mathbf{q}_3 - \mathbf{q}_1 + \mathbf{q}_2} - 2\eta} \right] \times [\dots]
 \end{aligned}$$

$$\begin{aligned}
& \frac{\epsilon_3 - \epsilon_1 \rightarrow \epsilon_3}{\mathbf{q}_3 - \mathbf{q}_2 \rightarrow \mathbf{q}_3} \ell \int \frac{d\epsilon_1}{2\pi} \int \frac{d\epsilon_3}{2\pi} \left[\frac{\theta(|\mathbf{q}_3 + \mathbf{q}_2| - k_F) \theta(k_F - |\mathbf{q}_2|)}{\epsilon_3 + \epsilon_{\mathbf{q}_2} - \epsilon_{\mathbf{q}_3 + \mathbf{q}_2} + 2i\eta} \right. \\
& \quad \left. - \frac{\theta(k_F - |\mathbf{q}_3 + \mathbf{q}_2|) \theta(|\mathbf{q}_2| - k_F)}{\epsilon_3 + \epsilon_{\mathbf{q}_2} - \epsilon_{\mathbf{q}_3 + \mathbf{q}_2} - 2i\eta} \right] \\
& \quad \times \left[\frac{\theta(|\mathbf{q}_1| - k_F)}{\epsilon_1 - \epsilon_{\mathbf{q}_1} + i\eta} + \frac{\theta(k_F - |\mathbf{q}_1|)}{\epsilon_1 - \epsilon_{\mathbf{q}_1} - i\eta} \right] \\
& \quad \times \left[\frac{\theta(|\mathbf{q}_3 + \mathbf{k}| - k_F)}{\epsilon_3 + \omega - \epsilon_{\mathbf{q}_3 + \mathbf{k}} + i\eta} + \frac{\theta(k_F - |\mathbf{q}_3 + \mathbf{k}|)}{\epsilon_3 + \omega - \epsilon_{\mathbf{q}_3 + \mathbf{k}} - i\eta} \right] \\
& \quad \times \left[\frac{\theta(|\mathbf{q}_3 + \mathbf{q}_1| - k_F)}{\epsilon_3 + \epsilon_1 - \epsilon_{\mathbf{q}_3 + \mathbf{q}_1} + i\eta} + \frac{\theta(k_F - |\mathbf{q}_3 + \mathbf{q}_1|)}{\epsilon_3 + \epsilon_1 - \epsilon_{\mathbf{q}_3 + \mathbf{q}_1} - i\eta} \right] \\
& = (i)^2 \int \frac{d\epsilon_3}{2\pi} \left[\frac{\theta(|\mathbf{q}_3 + \mathbf{q}_2| - k_F) \theta(k_F - |\mathbf{q}_2|)}{\epsilon_3 + \epsilon_{\mathbf{q}_2} - \epsilon_{\mathbf{q}_3 + \mathbf{q}_2} + 2i\eta} \right. \\
& \quad \left. - \frac{\theta(k_F - |\mathbf{q}_3 + \mathbf{q}_2|) \theta(|\mathbf{q}_2| - k_F)}{\epsilon_3 + \epsilon_{\mathbf{q}_2} - \epsilon_{\mathbf{q}_3 + \mathbf{q}_2} - 2i\eta} \right] \\
& \quad \times \left[-\frac{\theta(|\mathbf{q}_1| - k_F) \theta(k_F - |\mathbf{q}_3 + \mathbf{q}_1|)}{\epsilon_3 - \epsilon_{\mathbf{q}_3 + \mathbf{q}_1} + \epsilon_{\mathbf{q}_1} - 2i\eta} + \frac{\theta(k_F - |\mathbf{q}_1|) \theta(|\mathbf{q}_3 + \mathbf{q}_1| - k_F)}{\epsilon_3 - \epsilon_{\mathbf{q}_3 + \mathbf{q}_1} + \epsilon_{\mathbf{q}_1} + 2i\eta} \right] \\
& \quad \times \left[\frac{\theta(|\mathbf{q}_3 + \mathbf{k}| - k_F)}{\epsilon_3 + \omega - \epsilon_{\mathbf{q}_3 + \mathbf{k}} + i\eta} + \frac{\theta(k_F - |\mathbf{q}_3 + \mathbf{k}|)}{\epsilon_3 + \omega - \epsilon_{\mathbf{q}_3 + \mathbf{k}} - i\eta} \right]
\end{aligned}$$

Due to these reasonings, the leading order diagrams in the low density limit are given by these ladder-type diagrams.

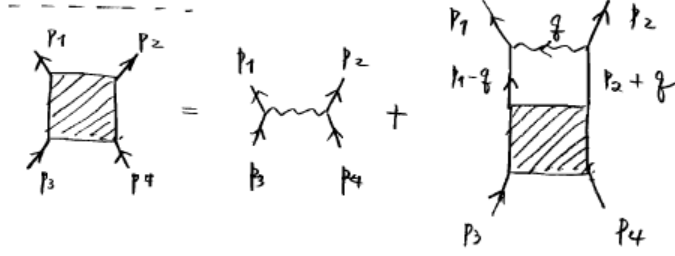


These ladder-type of diagram can be given by so-called the effective two-particle interaction.



(2.7.4)

Where the effective two-particle interaction is given by this:



(2.7.5)

The corresponding analytic expression are given as follows:

$$\hbar \Sigma^\star(p) = -2i \frac{1}{(2\pi)^4} \int d^4k G^0(k) \Gamma(p, k; p, k) + i \frac{1}{(2\pi)^4} \int d^4k G^0(k) \Gamma(k, p; p, k) \quad (2.7.4')$$

$$\Gamma(p_1, p_2; p_3, p_4) = v_0(p_1 - p_3) + i\hbar^{-1} \frac{1}{(2\pi)^4} \int d^4q v_0(q) \times G^0(p_1 - q) G^0(p_2 + q) \Gamma(p_1 - q, p_2 + q; p_3, p_4) \quad (2.7.6)$$

with $v_0(q) = V_0(\mathbf{q})$

This factor “2” comes from the summation over spin index associated with this closed loop, while the corresponding spin index here is fixed by the external line so that there is no factor “2” in the 2nd term. The minus sign here comes from this closed loop, while there is no closed loop in the 2nd term.

This self consistent equation for the two-particle interaction is called Bethe-Salpeter equation, for the effective interaction.

Now, the calculation of the proper self energy reduces to solving the Bethe-Salpeter equation in favor for the effective interaction.

To do this, let us define an “effective wavefunction” Q out of the effective interaction.

$$\Gamma(p_1, p_2; p_3, p_4) \equiv \frac{1}{(2\pi)^4} \int d^4q v_0(q) Q(p_1 - q, p_2 + q; p_3, p_4) \quad (2.7.7)$$

The B.S. equation for the effective wavefunction is given by¹:

$$Q(p_1, p_2; p_3, p_4) = (2\pi)^4 \delta^4(p_1 - p_3) + \left(\frac{i}{\hbar}\right) G^0(p_1) G^0(p_2) \times \int \frac{d^4q}{(2\pi)^4} v_0(q) \cdot Q(p_1 - q, p_2 + q; p_3, p_4) \quad (2.7.8)$$

To study this B.S. equation, let us first introduce the total center of mass wave vector as

$$P \equiv p_1 + p_2 = p_3 + p_4 \quad (2.7.9)$$

where the last equality comes from the conservation of total four-momentum.

We also introduce the relative wave vectors as

$$p \equiv \frac{1}{2}(p_1 - p_2), p' \equiv \frac{1}{2}(p_3 - p_4) \quad (2.7.10)$$

1

$$\begin{aligned} \Gamma(p_1, p_2; p_3, p_4) &= (2\pi)^{-4} \int d^4q v_0(q) (2\pi)^4 \delta(p_1 - q - p_3) \\ &\quad + (2\pi)^{-4} \int d^4q v_0(q) \frac{i}{\hbar} G^0(p_1 - q) G^0(p_2 + q) \Gamma(p_1 - q, p_2 + q; p_3, p_4) \end{aligned}$$

In terms of these three wave vectors, p_1, p_2, p_3, p_4 are given as follows.

$$\begin{cases} p_1 = \frac{1}{2}P + p, p_2 = \frac{1}{2}P - p \\ p_3 = \frac{1}{2}P + p', p_4 = \frac{1}{2}P - p' \end{cases} \quad (2.7.11)$$

Let us integrate $Q(p_1, p_2; p_3, p_4)$ over the frequency component of the former relative wave vector (p):¹

$$\begin{aligned} \frac{1}{2\pi} \int dp_0 Q(p_1, p_2; p_3, p_4) &= \frac{1}{2\pi} \int dp_0 Q\left(\frac{1}{2}P + p, \frac{1}{2}P - p; \frac{1}{2}P + p', \frac{1}{2}P - p'\right) \\ &\equiv \chi(\mathbf{p}, p', P) \end{aligned} \quad (2.7.12)$$

which we call as $\chi(\mathbf{p}, p', P)$.

Now that $v_0(q)$ is independent of the frequency component of q : $v_0(q) \equiv V_0(\mathbf{q})$, the B.S. equation for Q can be made into that for χ .

Namely, this part reduces to this²:

$$\int \frac{d^4 q}{(2\pi)^4} v_0(q) Q(p_1 - q, p_2 + q; p_3, p_4) = \int \frac{d^3 \mathbf{q}}{(2\pi)^3} V_0(\mathbf{q}) \chi(\mathbf{p} - \mathbf{q}, p', P)$$

So that, by integrating over p_0 , we can reduce this B.S. equation into this:

$$\begin{aligned} &\int \frac{dp_0}{2\pi} Q\left(\frac{1}{2}P + p, \frac{1}{2}P - p; \frac{1}{2}P + p', \frac{1}{2}P - p'\right) \\ &= \int \frac{dp_0}{2\pi} (2\pi)^4 \delta^4(p - p') + i\hbar^{-1} \int \frac{dp_0}{2\pi} G^0\left(\frac{1}{2}P + p\right) G^0\left(\frac{1}{2}P - p\right) \times \int \frac{d^3 \mathbf{q}}{(2\pi)^3} V_0(\mathbf{q}) \chi(\mathbf{p} - \mathbf{q}, p', P) \\ &= (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') + i\hbar^{-1} \int \frac{dp_0}{2\pi} G^0\left(\frac{1}{2}P + p\right) G^0\left(\frac{1}{2}P - p\right) \times \int \frac{d^3 \mathbf{q}}{(2\pi)^3} V_0(\mathbf{q}) \chi(\mathbf{p} - \mathbf{q}, p', P) \end{aligned}$$

Namely, we obtain the B.S. equation for χ :

$$\chi(\mathbf{p}, p', P) = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') + \frac{i}{\hbar} \int \frac{dp_0}{2\pi} G^0\left(\frac{1}{2}P + p\right) G^0\left(\frac{1}{2}P - p\right) \times \int \frac{d^3 \mathbf{q}}{(2\pi)^3} V_0(\mathbf{q}) \chi(\mathbf{p} - \mathbf{q}, p', P) \quad (2.7.13)$$

When solving this equation iteratively, one can see that $\chi(\mathbf{p}, p', P)$ thus given is actually independent from the frequency component of p :

$$\begin{aligned} \chi(\mathbf{p}, p', P) &= (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') + \frac{i}{\hbar} \int \frac{dp_0}{2\pi} G^0\left(\frac{1}{2}P + p\right) G^0\left(\frac{1}{2}P - p\right) V_0(\mathbf{p} - \mathbf{p}') \\ &\quad + \left(\frac{i}{\hbar}\right)^2 \int \frac{dp_0}{2\pi} G^0\left(\frac{1}{2}P + p\right) G^0\left(\frac{1}{2}P - p\right) \\ &\quad \times \frac{1}{(2\pi)^3} \int d^3 \mathbf{q} V_0(\mathbf{q}) \int \frac{d(p_0 - q_0)}{2\pi} G^0\left(\frac{1}{2}P + p - q\right) G^0\left(\frac{1}{2}P - p + q\right) V_0(\mathbf{p} - \mathbf{q} - \mathbf{p}') \\ &\quad + \dots \end{aligned}$$

Thus, we just call $\chi(\mathbf{p}, p', P)$ as $\chi(\mathbf{p}, \mathbf{p}', P)$.

$$\chi(\mathbf{p}, \mathbf{p}', P) = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') + \frac{i}{\hbar} \int \frac{dp_0}{2\pi} G^0\left(\frac{P}{2} + p\right) G^0\left(\frac{P}{2} - p\right) \times \int \frac{d^3 \mathbf{q}}{(2\pi)^3} V_0(\mathbf{q}) \chi(\mathbf{p} - \mathbf{q}, \mathbf{p}', P) \quad (2.7.14)$$

¹ $p \equiv (\mathbf{p}, p_0), p' \equiv (\mathbf{p}', p'_0)$
² $q \equiv (\mathbf{q}, q_0)$

It is now easy to evaluate the underlined part:

$$\begin{aligned}
&= \frac{i}{\hbar} \int \frac{dp_0}{2\pi} \left[\frac{\theta(|\frac{\mathbf{P}}{2} + \mathbf{p}| - k_F)}{\hbar(\frac{P_0}{2} + p_0) - \epsilon_{\frac{\mathbf{P}}{2} + \mathbf{p}} + i\eta} + \frac{\theta(k_F - |\frac{\mathbf{P}}{2} + \mathbf{p}|)}{\hbar(\frac{P_0}{2} + p_0) - \epsilon_{\frac{\mathbf{P}}{2} + \mathbf{p}} - i\eta} \right] \\
&\quad \times \left[\frac{\theta(|\frac{\mathbf{P}}{2} - \mathbf{p}| - k_F)}{\hbar(\frac{P_0}{2} - p_0) - \epsilon_{\frac{\mathbf{P}}{2} - \mathbf{p}} + i\eta} + \frac{\theta(k_F - |\frac{\mathbf{P}}{2} - \mathbf{p}|)}{\hbar(\frac{P_0}{2} - p_0) - \epsilon_{\frac{\mathbf{P}}{2} - \mathbf{p}} - i\eta} \right] \\
&= i(-i) \left[\frac{\theta(|\frac{\mathbf{P}}{2} + \mathbf{p}| - k_F)\theta(|\frac{\mathbf{P}}{2} - \mathbf{p}| - k_F)}{\hbar P_0 - \epsilon_{\frac{\mathbf{P}}{2} + \mathbf{p}} - \epsilon_{\frac{\mathbf{P}}{2} - \mathbf{p}} + 2i\eta} - \frac{\theta(k_F - |\frac{\mathbf{P}}{2} + \mathbf{p}|)\theta(k_F - |\frac{\mathbf{P}}{2} - \mathbf{p}|)}{\hbar P_0 - \epsilon_{\frac{\mathbf{P}}{2} + \mathbf{p}} - \epsilon_{\frac{\mathbf{P}}{2} - \mathbf{p}} - 2i\eta} \right]
\end{aligned}$$

Now that $\epsilon_{\mathbf{k}} \equiv \frac{\hbar^2 \mathbf{k}^2}{2m}$, we have

$$\hbar P_0 - \epsilon_{\frac{1}{2}\mathbf{P} + \mathbf{p}} - \epsilon_{\frac{1}{2}\mathbf{P} - \mathbf{p}} = \hbar P_0 - \frac{\hbar^2 \mathbf{P}^2}{4m} - \frac{\hbar^2 \mathbf{p}^2}{m} = E - \frac{\hbar^2 \mathbf{p}^2}{m} \quad (2.7.15')$$

Using also

$$N(\mathbf{P}, \mathbf{p}) = 1 - n_{\frac{\mathbf{P}}{2} + \mathbf{p}}^0 - n_{\frac{\mathbf{P}}{2} - \mathbf{p}}^0 \quad (2.7.15)$$

with $n_{\mathbf{p}}^0 \equiv \theta(k_F - |\mathbf{p}|)$, the underlined part can be simplified into this:

$$\frac{i}{\hbar} \int \frac{dp_0}{2\pi} G^0(\frac{P}{2} + p) G^0(\frac{P}{2} - p) = \frac{N(\mathbf{P}, \mathbf{p})}{E - \frac{\hbar^2 \mathbf{p}^2}{m} + i\eta N(\mathbf{P}, \mathbf{p})}$$

and B.S. equation can be given as follows:

$$\chi(\mathbf{p}, \mathbf{p}', P) = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') + \frac{N(\mathbf{P}, \mathbf{p})}{E - \frac{\hbar^2 \mathbf{p}^2}{m} + i\eta N(\mathbf{P}, \mathbf{p})} \int \frac{d^3 \mathbf{q}}{(2\pi)^3} \times V_0(\mathbf{q}) \chi(\mathbf{p} - \mathbf{q}, \mathbf{p}', P) \quad (2.7.16)$$

In terms of χ thus given, the effective two-particle interaction is given as follows:

$$\Gamma(\mathbf{p}, \mathbf{p}', P) = \frac{1}{(2\pi)^3} \int d^3 \mathbf{q} V_0(\mathbf{q}) \chi(\mathbf{p} - \mathbf{q}, \mathbf{p}', P) \quad (2.7.17)$$

To solve this Bethe-Salpeter equation, let us first consider the low density limit, in which $N(\mathbf{P}, \mathbf{p})$ can be replaced by 1.

$$\chi_0(\mathbf{p}, \mathbf{p}', P) = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') + \frac{1}{E - \frac{\hbar^2 \mathbf{p}^2}{m} + i\eta} \times \int \frac{d^3 \mathbf{q}}{(2\pi)^3} V_0(\mathbf{q}) \chi_0(\mathbf{p} - \mathbf{q}, \mathbf{p}', P) \quad (2.7.18)$$

By scaling E and V_0 by $\frac{\hbar^2}{m}$,

$$\epsilon = \frac{mE}{\hbar^2}, v_0(\mathbf{q}) = \frac{mV_0(\mathbf{q})}{\hbar^2} \quad (2.7.19)$$

we have:

$$\chi_0(\mathbf{p}, \mathbf{p}', P) = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') + \frac{1}{\epsilon - \mathbf{p}^2 + i\eta} \times \int \frac{d^3 \mathbf{q}}{(2\pi)^3} v_0(\mathbf{q}) \chi_0(\mathbf{p} - \mathbf{q}, \mathbf{p}', P) \quad (2.7.18')$$

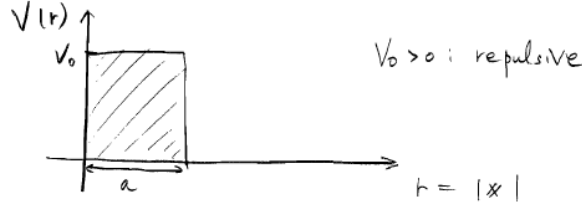
This equation can be solved closely in parallel to the scattering problem of one electron by a spherical hard core potential.

I suppose that the scattering problem are usually given in the course of Quantum Mechanics, so that I will not go into the details of it. But, to make this class to be self-contained, I will briefly summarize the scattering problem of single electron by the hard core potential.

Consider that an incident electron wave of momentum \mathbf{k} is injected to a spherical hard core potential:

$$\left[-\frac{\hbar^2 \nabla^2}{2m} + V(\mathbf{x}) \right] \psi(\mathbf{x}) = E\psi(\mathbf{x}) \quad (2.7.A.1)$$

where $V(\mathbf{x})$ represent the hard core potential depending only on the radial coordinate.



Since the incident wave have momentum \mathbf{k} , we want find an eigenstate of this Hamiltonian which have energy of $\frac{\hbar^2 \mathbf{k}^2}{2m}$.

With a proper scaling, this reduces to

$$[\nabla^2 + |\mathbf{k}|^2] \psi(\mathbf{x}) = v(\mathbf{x})\psi(\mathbf{x}) \quad \left(v(\mathbf{x}) = \frac{2mV(\mathbf{x})}{\hbar^2} \right) \quad (2.7.A.1')$$

When \mathbf{x} is far from the origin at which the hard core scattering potential is located, the eigenwavefunction become a plane-wave with the momentum \mathbf{k} .

$$\psi(\mathbf{x}) = e^{i\mathbf{k} \cdot \mathbf{x}} + \dots$$

Since this requirement is nothing but the boundary condition, we put \mathbf{k} as a subscript of ψ :

$$\psi_{\mathbf{k}}(\mathbf{x}) = e^{i\mathbf{k} \cdot \mathbf{x}} + \dots$$

When \mathbf{x} becomes close to the scattering potential, we need correction to the eigenwavefunction, which can be given like this:

$$\psi_{\mathbf{k}}(\mathbf{x}) = e^{i\mathbf{k} \cdot \mathbf{x}} - \int d^3\mathbf{y} G^{(+)}(\mathbf{x} - \mathbf{y}) v(\mathbf{y}) \psi_{\mathbf{k}}(\mathbf{y}) \quad (2.7.A.2)$$

An “out-going-wave” Green’s function $G^{(+)}$ which satisfies this:

$$\left(\nabla_{\mathbf{x}}^2 + |\mathbf{k}|^2 \right) G^{(+)}(\mathbf{x} - \mathbf{y}) = -\delta^3(\mathbf{x} - \mathbf{y}) \quad (2.7.A.3)$$

is given by this:

$$G^{(+)}(\mathbf{x} - \mathbf{y}) = \int \frac{d^3\mathbf{p}}{(2\pi)^3} \frac{e^{i\mathbf{p} \cdot (\mathbf{x} - \mathbf{y})}}{p^2 - k^2 - i\eta} = \frac{1}{4\pi} \frac{e^{ik|\mathbf{x} - \mathbf{y}|}}{|\mathbf{x} - \mathbf{y}|} \quad (2.7.A.4)$$

You can check that this set of solutions indeed satisfies this Schrödinger equation.

If the scattering potential has no bound state, then this eigenstate support a complete set of states:

$$\int \frac{d^3\mathbf{k}}{(2\pi)^3} \psi_{\mathbf{k}}(\mathbf{x}) \psi_{\mathbf{k}'}^*(\mathbf{x}') = \delta^3(\mathbf{x} - \mathbf{x}') \quad (2.7.A.5)$$

In the momentum representation,

$$\begin{cases} \psi_{\mathbf{k}}(\mathbf{p}) \equiv \int d^3\mathbf{x} e^{-i\mathbf{p} \cdot \mathbf{x}} \psi_{\mathbf{k}}(\mathbf{x}) \\ v(\mathbf{p}) \equiv \int d^3\mathbf{x} e^{-i\mathbf{p} \cdot \mathbf{x}} v(\mathbf{x}) \end{cases} \quad (2.7.A.6)$$

This equation can be rewritten as

$$\psi_{\mathbf{k}}(\mathbf{p}) = (2\pi)^3 \delta^3(\mathbf{k} - \mathbf{p}) - \frac{1}{\mathbf{p}^2 - \mathbf{k}^2 - i\eta} \int \frac{d^3\mathbf{q}}{(2\pi)^3} v(\mathbf{q}) \psi_{\mathbf{k}}(\mathbf{p} - \mathbf{q}) \quad (2.7.A.7)$$

Then the scattering amplitude for a transition from an incident wave vector \mathbf{k} to a final wave vector \mathbf{p} is defined by this part:

$$\tilde{f}(\mathbf{p}, \mathbf{k}) \equiv -4\pi f(\mathbf{p}, \mathbf{k}) \equiv \int \frac{d\mathbf{q}}{(2\pi)^3} v(\mathbf{q}) \psi_{\mathbf{k}}(\mathbf{p} - \mathbf{q}) \quad (2.7.A.8)$$

In terms of the scattering amplitude, this reduces to this:

$$\psi_{\mathbf{k}}(\mathbf{p}) = (2\pi)^3 \delta^3(\mathbf{k} - \mathbf{p}) + \frac{\tilde{f}(\mathbf{p}, \mathbf{k})}{k^2 - p^2 + i\eta} \quad (2.7.A.7')$$

According to this completeness relation, this satisfies the corresponding completeness relation, which is given like this:

$$\int \frac{d^3\mathbf{k}}{(2\pi)^3} \psi_{\mathbf{k}}(\mathbf{p}) \psi_{\mathbf{k}'}^*(\mathbf{p}') = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') \quad (2.7.A.9)$$

When the norm of the outgoing-wave momentum ($|\mathbf{k}| = |\mathbf{k}'|$) and that of the incident wave is same, the scattering amplitude is calculated by “the method of partial wave”:¹

$$f(\mathbf{k}, \mathbf{k}') = \sum_{l=0}^{\infty} \frac{2l+1}{k} e^{i\delta_l} \sin(\delta_l) P_l(\cos\theta) \quad (2.7.A.10)$$

where θ is angle subtended by \mathbf{k} and \mathbf{k}' and $k = |\mathbf{k}| = |\mathbf{k}'|$.

δ_l is called a “phase shift” associated with l -th partial wave.

In the limit of hard core ($V_0 \rightarrow +\infty$), the phase shift can be expanded with respect to small k_a , as:²

$$\tan(\delta_l) = -\frac{(k_a)^{2l+1}}{(2l+1)[(2l-1)!!]^2} + o((k_a)^{2l+2}) \quad (2.7.A.11)$$

These are the summary of “scattering problem of a single electron by a spherical hard core potential”, which becomes very useful for solving the Bethe-Salpeter equation in the low-density limit.

$$\chi_0(\mathbf{p}, \mathbf{p}', P) = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') + \frac{1}{\epsilon - \mathbf{p}^2 + i\eta} \cdot \int \frac{d^3\mathbf{q}}{(2\pi)^3} v_0(\mathbf{q}) \chi_0(\mathbf{p} - \mathbf{q}, \mathbf{p}', P) \quad (2.7.18')$$

To see this, let us rewrite this as follows

$$\begin{aligned} (\epsilon - \mathbf{p}^2 + i\eta) \chi_0(\mathbf{p}, \mathbf{p}', P) - \int \frac{d^3\mathbf{q}}{(2\pi)^3} v_0(\mathbf{q}) \chi_0(\mathbf{p} - \mathbf{q}, \mathbf{p}', P) &= (2\pi)^3 (\epsilon - \mathbf{p}^2 + i\eta) \delta^3(\mathbf{p} - \mathbf{p}') \\ &= (2\pi)^3 (\epsilon - \mathbf{p}'^2 + i\eta) \delta^3(\mathbf{p} - \mathbf{p}') \end{aligned} \quad (2.7.18'')$$

This equation can be satisfied by

$$\chi_0(\mathbf{p}, \mathbf{p}', P) = (\epsilon - \mathbf{p}'^2 + i\eta) \int \frac{d\mathbf{k}}{(2\pi)^3} \frac{\psi_{\mathbf{k}}(\mathbf{p}) \psi_{\mathbf{k}'}^*(\mathbf{p}')}{\epsilon - \mathbf{k}^2 + i\eta} \quad (2.7.20)$$

¹ $P_l(x)$ is the Legendre polynomial.

²see eq. (19.21) in the text book by L.Schiff “Quantum Mechanics” (3rd edition).

where $\psi_{\mathbf{k}}(\mathbf{p})$ is given by eq.(2.7.A.7), a solution of scattering problem of a single electron by hard core spherical potential.

Here the scattering potential $V(\mathbf{x})$ is replaced by the interaction potential between two particles $V_0(\mathbf{x})$

$$\begin{cases} V(\mathbf{x}) \rightarrow V_0(\mathbf{x}) \\ v(\mathbf{q}) \rightarrow v_0(\mathbf{q}) \end{cases} \quad (2.7.21)$$

With this replacement in mind, we then recycle all the results obtained in the scattering problem here.

By the way, one can easily see this:

$$\begin{aligned} (\epsilon - \mathbf{p}^2 + \imath\eta)\chi_0(\mathbf{p}, \mathbf{p}', P) &= (\epsilon - \mathbf{p}'^2 + \imath\eta) \int \frac{d^3\mathbf{k}}{(2\pi)^3} \frac{(\epsilon - \mathbf{p}^2 + \imath\eta)\psi_{\mathbf{k}}(\mathbf{p})\psi_{\mathbf{k}}^*(\mathbf{p}')}{\epsilon - \mathbf{k}^2 + \imath\eta} \\ &= (\epsilon - \mathbf{p}'^2 + \imath\eta) \int \frac{d^3\mathbf{k}}{(2\pi)^3} \psi_{\mathbf{k}}(\mathbf{p})\psi_{\mathbf{k}}^*(\mathbf{p}') \\ &\quad + (\epsilon - \mathbf{p}'^2 + \imath\eta) \int \frac{d^3\mathbf{k}}{(2\pi)^3} \frac{(\mathbf{k}^2 - \mathbf{p}^2 + \imath\eta)\psi_{\mathbf{k}}(\mathbf{p})\psi_{\mathbf{k}}^*(\mathbf{p}')}{\epsilon - \mathbf{k}^2 + \imath\eta} \\ &= (\epsilon - \mathbf{p}'^2 + \imath\eta)\delta^3(\mathbf{p} - \mathbf{p}') \\ &\quad + (\epsilon - \mathbf{p}'^2 + \imath\eta) \int \frac{d^3\mathbf{k}}{(2\pi)^3} \int \frac{d\mathbf{q}}{(2\pi)^3} \frac{v(\mathbf{q})\psi_{\mathbf{k}}(\mathbf{p} - \mathbf{q})\psi_{\mathbf{k}}^*(\mathbf{p}')}{\epsilon - \mathbf{k}^2 + \imath\eta} \\ &= (\epsilon - \mathbf{p}'^2 + \imath\eta)\delta^3(\mathbf{p} - \mathbf{p}') + \int \frac{d\mathbf{q}}{(2\pi)^3} N(\mathbf{q})\chi_0(\mathbf{p} - \mathbf{q}, \mathbf{p}', P) \end{aligned}$$

Using (2.7.A.7'), this can be rewritten as:

$$\begin{aligned} \chi_0(\mathbf{p}, \mathbf{p}', P) &= (\epsilon - \mathbf{p}'^2 + \imath\eta) \int \frac{d^3\mathbf{k}}{(2\pi)^3} \frac{\psi_{\mathbf{k}}(\mathbf{p})\psi_{\mathbf{k}}^*(\mathbf{p}')}{\epsilon - \mathbf{k}^2 + \imath\eta} \\ &= (\epsilon - \mathbf{p}'^2 + \imath\eta) \int \frac{d^3\mathbf{k}}{(2\pi)^3} \frac{\psi_{\mathbf{k}}(\mathbf{p})}{\epsilon - \mathbf{k}^2 + \imath\eta} \times \left\{ (2\pi)^3 \delta^3(\mathbf{p}' - \mathbf{k}) + \frac{\tilde{f}^*(\mathbf{p}', \mathbf{k})}{\mathbf{k}^2 - \mathbf{p}'^2 - \imath\eta} \right\} \\ &= \psi_{\mathbf{p}'}(\mathbf{p}) + \int \frac{d^3\mathbf{k}}{(2\pi)^3} \psi_{\mathbf{k}}(\mathbf{p}) \times \frac{\epsilon - \mathbf{p}'^2 + \imath\eta}{[\epsilon - \mathbf{k}^2 + \imath\eta][\mathbf{k}^2 - \mathbf{p}'^2 - \imath\eta]} \tilde{f}^*(\mathbf{p}', \mathbf{k}) \\ &= \psi_{\mathbf{p}'}(\mathbf{p}) + \int \frac{d^3\mathbf{k}}{(2\pi)^3} \psi_{\mathbf{k}}(\mathbf{p}) \times \left[\frac{1}{\epsilon - \mathbf{k}^2 + 2\imath\eta} + \frac{1}{\mathbf{k}^2 - \mathbf{p}'^2 - \imath\eta} \right] \tilde{f}^*(\mathbf{p}', \mathbf{k}) \end{aligned} \quad (2.7.22)$$

substituting this into eq.(2.7.17), we finally obtain:

$$\begin{aligned} \Gamma_0(\mathbf{p}, \mathbf{p}', P) &= \frac{1}{(2\pi)^3} \frac{\hbar^2}{m} \int d^3\mathbf{q} v_0(\mathbf{q}) \chi_0(\mathbf{p} - \mathbf{q}, \mathbf{p}', P) \\ \frac{m}{\hbar^2} \Gamma_0(\mathbf{p}, \mathbf{p}', P) &\stackrel{2.7.22}{=} \frac{1}{(2\pi)^3} \int d^3\mathbf{q} v_0(\mathbf{q}) \psi_{\mathbf{p}'}(\mathbf{p} - \mathbf{q}) \\ &\quad + \frac{1}{(2\pi)^3} \int d^3\mathbf{q} v_0(\mathbf{q}) \int \frac{d^3\mathbf{k}}{(2\pi)^3} \psi_{\mathbf{k}}(\mathbf{p} - \mathbf{q}) \left[\frac{1}{\epsilon - \mathbf{k}^2 + \imath\eta} + \frac{1}{\mathbf{k}^2 - \mathbf{p}'^2 - \imath\eta} \right] \tilde{f}^*(\mathbf{p}', \mathbf{k}) \\ &\stackrel{2.7.A.8}{=} \tilde{f}(\mathbf{p}, \mathbf{p}') + \int \frac{d^3\mathbf{k}}{(2\pi)^3} \tilde{f}(\mathbf{p}, \mathbf{k}) \left[\frac{1}{\epsilon - \mathbf{k}^2 + \imath\eta} + \frac{1}{\mathbf{k}^2 - \mathbf{p}'^2 - \imath\eta} \right] \times \tilde{f}^*(\mathbf{p}', \mathbf{k}) \end{aligned} \quad (2.7.23)$$

This expression suggests that the effective two particle interaction in the low density limit can be expressed only by the scattering amplitude $\tilde{f}(\mathbf{k}, \mathbf{k}')$, which is calculated from the scattering problem of a single electron.

We now try to solve eq.(2.7.16) for finite $N(\mathbf{P}, \mathbf{p})$.

$$\chi(\mathbf{p}, \mathbf{p}', P) = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') + \frac{N(\mathbf{P}, \mathbf{p})}{\epsilon - p^2 + i\eta N(\mathbf{P}, \mathbf{p})} \int \frac{d^3 \mathbf{q}}{(2\pi)^3} v(\mathbf{q}) \chi(\mathbf{p} - \mathbf{q}, \mathbf{p}', P) \quad (2.7.16')$$

$$\int \frac{d^3 \mathbf{p}'}{(2\pi)^3} \left[(2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') - \frac{v(\mathbf{p} - \mathbf{p}')}{\epsilon - \mathbf{p}^2 + i\eta} \right] \chi_0(\mathbf{p}', \mathbf{p}'', P) = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}')$$

$$\int d^3 \mathbf{p}' \chi_0^{-1}(\mathbf{p}, \mathbf{p}', P) \chi_0(\mathbf{p}', \mathbf{p}'', P) = (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}'')$$

Thus

$$\chi_0^{-1}(\mathbf{p}, \mathbf{p}', P) = \delta^3(\mathbf{p} - \mathbf{p}') - \frac{N(\mathbf{p} - \mathbf{p}')}{\epsilon - \mathbf{p}^2 + i\eta} \frac{1}{(2\pi)^3}$$

$$\int d^3 \mathbf{p} \chi_0(\mathbf{p}'', \mathbf{p}, P) \chi_0^{-1}(\mathbf{p}, \mathbf{p}', P) = (2\pi)^3 \delta^3(\mathbf{p}'' - \mathbf{p}')$$

We first subtract the eq in the low density limit:

$$\chi(\mathbf{p}, \mathbf{p}', P) - \frac{1}{\epsilon - \mathbf{p}^2 + i\eta} \int \frac{d^3 \mathbf{q}}{(2\pi)^3} v(\mathbf{q}) \chi(\mathbf{p} - \mathbf{q}, \mathbf{p}', P)$$

$$= (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') + \left(\frac{N(\mathbf{P}, \mathbf{p})}{\epsilon - \mathbf{p}^2 + i\eta N(\mathbf{P}, \mathbf{p})} - \frac{1}{\epsilon - \mathbf{p}^2 + i\eta} \right) \times \int \frac{d^3 \mathbf{q}}{(2\pi)^3} v(\mathbf{q}) \chi(\mathbf{p} - \mathbf{q}, \mathbf{p}', P)$$

where the underlined part can be replaced by the effective two-particle interaction $\frac{m}{\hbar^2} \Gamma(\mathbf{p}, \mathbf{p}', P)$.

$$\int d^3 \mathbf{p}'' \chi_0^{-1}(\mathbf{p}, \mathbf{p}'', P) \chi(\mathbf{p}'', \mathbf{p}', P)$$

$$= (2\pi)^3 \delta^3(\mathbf{p} - \mathbf{p}') + \left(\frac{N(\mathbf{P}, \mathbf{p})}{\epsilon - \mathbf{p}^2 + i\eta N(\mathbf{P}, \mathbf{p})} - \frac{1}{\epsilon - \mathbf{p}^2 + i\eta} \right) \times \frac{m}{\hbar^2} \Gamma(\mathbf{p}, \mathbf{p}', P)$$

$$\int d^3 \mathbf{q} \chi_0(\mathbf{p}, \mathbf{q}, P) \int d^3 \mathbf{p}'' \chi_0^{-1}(\mathbf{q}, \mathbf{p}'', P) \chi(\mathbf{p}'', \mathbf{p}', P)$$

$$= \int d^3 \mathbf{q} \chi_0(\mathbf{p}, \mathbf{q}, P) (2\pi)^3 \delta^3(\mathbf{q} - \mathbf{p}')$$

$$+ \int d^3 \mathbf{q} \chi_0(\mathbf{p}, \mathbf{q}, P) \left[\frac{N(\mathbf{P}, \mathbf{q})}{\epsilon - \mathbf{q}^2 + i\eta N(\mathbf{P}, \mathbf{q})} - \frac{1}{\epsilon - \mathbf{q}^2 + i\eta} \right] \frac{m}{\hbar^2} \Gamma(\mathbf{q}, \mathbf{p}', P)$$

$$\Leftrightarrow (2\pi)^3 \chi(\mathbf{p}, \mathbf{p}', P) = (2\pi)^3 \chi_0(\mathbf{p}, \mathbf{p}', P)$$

$$+ \int d^3 \mathbf{p}'' \chi_0(\mathbf{p}, \mathbf{p}'', P) \left[\frac{N(\mathbf{P}, \mathbf{p}'')}{\epsilon - \mathbf{p}''^2 + i\eta N(\mathbf{P}, \mathbf{p}'')} - \frac{1}{\epsilon - \mathbf{p}''^2 + i\eta} \right] \frac{m}{\hbar^2} \Gamma(\mathbf{p}'', \mathbf{p}', P)$$

$$\chi(\mathbf{p}, \mathbf{p}', P) = \chi_0(\mathbf{p}, \mathbf{p}', P)$$

$$+ \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \chi_0(\mathbf{p}, \mathbf{k}, P) \times \left[\frac{N(\mathbf{P}, \mathbf{k})}{\epsilon - \mathbf{k}^2 + i\eta N(\mathbf{P}, \mathbf{k})} - \frac{1}{\epsilon - \mathbf{k}^2 + i\eta} \right] \frac{m}{\hbar^2} \Gamma(\mathbf{k}, \mathbf{p}', P)$$

Then, multiplying $v(\mathbf{q})$ and replacing \mathbf{p} by $\mathbf{p} - \mathbf{q}$ and intergrating over \mathbf{q} , we can replace this by Γ and these two by Γ_0 respectively.

$$\Gamma(\mathbf{p}, \mathbf{p}', P) = \Gamma_0(\mathbf{p}, \mathbf{p}', P)$$

$$+ \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \Gamma_0(\mathbf{p}, \mathbf{k}, P) \times \left[\frac{N(\mathbf{P}, \mathbf{k})}{\epsilon - \mathbf{k}^2 + i\eta N(\mathbf{P}, \mathbf{k})} - \frac{1}{\epsilon - \mathbf{k}^2 + i\eta} \right] \frac{m}{\hbar^2} \Gamma(\mathbf{k}, \mathbf{p}', P) \quad (2.7.24)$$

A set of these two equations gives effective two-particle potential Γ in terms of the scattering amplitude \tilde{f} .

Importantly, this scattering amplitude \tilde{f} can be calculated from a single-particle scattering problem due to a hard core spherical scattering potential, which corresponds to our short-range s-wave repulsive interaction potential.

This set of equation are called as Galitskii's integral equation.

For a low density Fermi gas, eq.(2.7.24) can be solved iteratively as a power series in small $k_F a (\ll 1)$. Namely, the integrand of the 2nd term vanishes, when the vectors $\frac{1}{2}\mathbf{P} + \mathbf{k}$ and $\frac{1}{2}\mathbf{P} - \mathbf{k}$ are both outside the Fermi sea.

In a 'small-Fermi-sea' limit, the integral region of the momentum \mathbf{k} is restricted, so that the 2nd term can be roughly estimated as

$$\Gamma_0 \cdot k_F \cdot m \Gamma \frac{1}{\hbar^2}$$

Thus, the correction to Γ due to the 2nd term can be shown to be subleading order:

$$\frac{\Gamma - \Gamma_0}{\Gamma} = \frac{k_F m \Gamma_0}{\hbar} = k_F \tilde{f} \stackrel{2.7.A.10 \& 2.7.A.11}{=} k_F a \ll 1$$

Before solving eq.(2.7.24) iteratively, please also notice that, according to eq.(2.7.4'), the full set of variable $\mathbf{p}, \mathbf{p}', P$ is not necessary.

Namely, all that is necessary is the proper self-energy, which is given by this:

$$\hbar \Sigma^\star(p) = -2i \int \frac{d^4 k}{(2\pi)^4} G^0(k) \Gamma(p, k; p, k) + i \int \frac{d^4 k}{(2\pi)^4} G^0(k) \Gamma(k, p; p, k)$$

Using

$$\left. \begin{aligned} P &= p_1 + p_2 = p_3 + p_4 \\ p &= \frac{1}{2}(p_1 - p_2) \\ p' &= \frac{1}{2}(p_3 - p_4) \end{aligned} \right\} \text{2.7.9 \& 2.7.10 \& 2.7.11}$$

we have

$$\Gamma(p_1, p_2; p_3, p_4) \rightarrow \Gamma(p, p', P)$$

Thus

$$\Gamma(p, k; p, k) \rightarrow \Gamma(p - k, p - k, \frac{p + k}{2}) = \Gamma(q, q, P)$$

where

$$\begin{aligned} \frac{1}{2}(p - k) &= q \\ p + k &= P \end{aligned}$$

and

$$\Gamma(p, k; k, p) \rightarrow \Gamma(q, -q, P)$$

According to eq.(2.7.17), Γ does not depend on the frequency component of first two four-dimensional momentum vector, so that, we replace these q by three-dimensional momentum \mathbf{q} .

$$\frac{1}{2}(\mathbf{p} - \mathbf{k}) = \mathbf{q}, \mathbf{p} + \mathbf{k} = \mathbf{P}, p_0 + k_0 = P_0$$

$$\hbar \Sigma^\star(p) = -2i \int \frac{d^4 k}{(2\pi)^4} G^0(k) \Gamma(\mathbf{q}, \mathbf{q}, P) + i \int \frac{d^4 k}{(2\pi)^4} G^0(k) \Gamma(\mathbf{q}, -\mathbf{q}, P) \quad (2.7.25)$$

Many of the physical quantities can be calculated from self-energy, which is now given by two-particle effective interaction with a particular set of its variables.

Namely, \mathbf{q} , \mathbf{q} and P ; or \mathbf{q} , $-\mathbf{q}$ and P .

Thus, we will calculate Γ with these particular set of variables, iteratively as a power of small $k_F a$.

To be more specific, we will obtain Γ to the 2nd order of $k_F a$.

Using eq.(2.7.23) and eqs. (2.7.A.10) & (2.7.A.11), we first obtain Γ_0 to the 2nd order in $k_F a$.

From eqs. (2.7.A.10) & (2.7.A.11), the scattering amplitude can be expanded in small $k_F a$ as this, for the case with $|\mathbf{k}| = |\mathbf{k}'|$.¹

$$\begin{aligned} f(\mathbf{k}, \mathbf{k}') &= \sum_{l=0}^{+\infty} \frac{2l+1}{k} e^{i\delta_l} \sin \delta_l P_l(\cos \theta) \\ &= k^{-1} (1 + i\delta_0) \delta_0 + \underline{o(k^{-1} \cdot (ka)^3)} \\ &\stackrel{\underline{2.7.A.11}}{=} -a + ika^2 + o(k^2 a^3) \\ \tilde{f}(\mathbf{k}, \mathbf{k}') &= 4\pi a - 4\pi i a^2 k + o(k^2 a^3) \end{aligned}$$

Since \tilde{f} is already at most 1st order in small a , we can replace \tilde{f} in the 2nd term of eq.(2.7.23) by its first order contribution, while replace f in the first term of eq.(2.7.23) by its expression up to the 2nd order.

Such treatment gives us Γ_0 to the 2nd order in small a :

$$\begin{aligned} \frac{m}{\hbar^2} \Gamma_0(\mathbf{q}, \mathbf{q}, P) &= 4\pi a - 4\pi i q a^2 + \int \frac{d^3 k'}{(2\pi)^3} 4\pi a \left(\frac{1}{\epsilon - k'^2 + i\eta} + \frac{1}{k'^2 - q^2 - i\eta} \right) 4\pi a + o(a^3) \\ &= 4\pi a - 4\pi i q a^2 + (4\pi a)^2 \times \int \frac{d^3 k'}{(2\pi)^3} \left(\frac{1}{\epsilon - k'^2 + i\eta} + \frac{1}{k'^2 - q^2 - i\eta} \right) + \dots \\ &= 4\pi a - 4\pi i q a^2 + (4\pi a)^2 \times \\ &\quad \int \frac{d^3 k'}{(2\pi)^3} \left(\frac{1}{\epsilon - k'^2 + i\eta} + \frac{\mathcal{P}}{k'^2 - q^2} + i\pi \frac{1}{2q} \delta(|\mathbf{k}'| - |\mathbf{q}|) \right) + \dots \\ \underline{\underline{\int \frac{d^3 k'}{(2\pi)^3} \rightarrow 4\pi |\mathbf{k}'|^2 d|\mathbf{k}'|}} &= 4\pi a - 4\pi i q a^2 + i(4\pi a)^2 \times \frac{q}{4\pi} \\ &\quad + (4\pi a)^2 \int \frac{d^3 k'}{(2\pi)^3} \left(\frac{1}{\epsilon - k'^2 + i\eta} + \frac{\mathcal{P}}{k'^2 - q^2} \right) + \dots \\ &= 4\pi a + (4\pi a)^2 \int \frac{d^3 k'}{(2\pi)^3} \left(\frac{1}{\epsilon - k'^2 + i\eta} + \frac{\mathcal{P}}{k'^2 - q^2} \right) + \dots \end{aligned}$$

substituting this solution to eq.(2.7.24), we obtain Γ up to the second order in small a :

$$\begin{aligned} \frac{m}{\hbar^2} \Gamma(\mathbf{q}, \mathbf{q}, P) &= 4\pi a + (4\pi a)^2 \int \frac{d^3 \mathbf{k}'}{(2\pi)^3} \left(\frac{1}{\epsilon - \mathbf{k}'^2 + i\eta} + \frac{\mathcal{P}}{\mathbf{k}'^2 - \mathbf{q}^2} \right) \\ &\quad + (4\pi a)^2 \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \left[\frac{N(\mathbf{P}, \mathbf{k})}{\epsilon - \mathbf{k}^2 + i\eta N(\mathbf{P}, \mathbf{k})} - \frac{1}{\epsilon - \mathbf{k}^2 + i\eta} \right] \\ &= 4\pi a + (4\pi a)^2 \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \left[\frac{N(\mathbf{P}, \mathbf{k})}{\epsilon - \mathbf{k}^2 + i\eta N(\mathbf{P}, \mathbf{k})} + \frac{\mathcal{P}}{\mathbf{k}^2 - \mathbf{q}^2} \right] + o(a^3) \end{aligned} \tag{2.7.26}$$

To the 2nd order in small a , we can also see that

$$\Gamma(-\mathbf{q}, \mathbf{q}, P) = \Gamma(\mathbf{q}, \mathbf{q}, P) + o(a^3) \tag{2.7.26'}$$

¹the underlined part: see eq. (2.7.A.11).

Thus, substituting these solution into eq.(2.7.25), we finally obtain:

$$\begin{aligned}\hbar\Sigma^\star(p) &= -\imath \int \frac{d^4k}{(2\pi)^4} G^0(k) \Gamma(\mathbf{q}, \mathbf{q}, P) \\ &\equiv \hbar\Sigma_1^\star(p) + \hbar\Sigma_2^\star(p) + o((k_F a)^3)\end{aligned}\quad (2.7.27)$$

$$\hbar\Sigma_1^\star(p) = -\imath \int \frac{d^4k}{(2\pi)^4} G^0(k) e^{\imath k_0 0+} \cdot \frac{4\pi a \hbar^2}{m} \quad (2.7.27')$$

$$\hbar\Sigma_2^\star(p) = -\imath \int \frac{d^4k}{(2\pi)^4} G^0(k) e^{\imath k_0 0+} \times (4\pi a)^2 \int \frac{d^3\mathbf{k}'}{(2\pi)^3} \left[\frac{N(\mathbf{P}, \mathbf{k}')}{\epsilon - \mathbf{k}'^2 + \imath \eta N(\mathbf{P}, \mathbf{k}')} + \frac{\mathcal{P}}{\mathbf{k}'^2 - \mathbf{q}^2} \right] \quad (2.7.27'')$$

where $\Sigma_1^\star(p)$ is the first order in $k_F a$, while $\Sigma_2^\star(p)$ is the 2nd order in $k_F a$.

We put $e^{\imath k_0 0+}$ in order to this frequency integral to be well-defined.

The reason for choosing positive small number instead of negative small number is essentially same as that explained in the item ⑦ of Feynman's rule in the momentum space.

With this phase factor in mind, one can easily see that Σ_1^\star is indeed on the first order in $k_F a$.

$$\begin{aligned}\hbar\Sigma_1^\star(p) &= -\imath \int \frac{d^3\mathbf{k}}{(2\pi)^3} \frac{4\pi a \hbar^2}{m} \int \frac{dk_0}{2\pi} e^{\imath k_0 0+} \left[\frac{\theta(k - k_F)}{k_0 - \omega_k + \imath \eta} + \frac{\theta(k_F - k)}{k_0 - \omega_k - \imath \eta} \right] \\ &= \frac{4\pi a \hbar^2}{m} \int \frac{d^3\mathbf{k}}{(2\pi)^3} \theta(k_F - k) \\ &= \frac{4\pi a \hbar^2}{m} \cdot \frac{4\pi}{(2\pi)^3} \frac{k_F^3}{3} = \frac{\hbar^2 k_F^2}{m} \cdot \frac{2}{3\pi} k_F a\end{aligned}\quad (2.7.28)$$

To evaluate Σ_2^\star , let us remember that ϵ here depends on the frequency component k , i.e. k_0 . Namely, according to eqs.(2.7.15') & (2.7.19) ϵ is given like this:

$$\epsilon = \frac{mP_0}{\hbar} - \frac{\mathbf{P}^2}{4} = \frac{m}{\hbar} (p_0 + k_0) - \frac{\mathbf{P}^2}{4}$$

Thus we obtain Σ_2 as

$$\begin{aligned}\hbar\Sigma_2^\star(p) &= -\imath \int \frac{d^3\mathbf{k}}{(2\pi)^3} \int \frac{dk_0}{2\pi} e^{\imath k_0 0+} \\ &\quad \times \left[\frac{\theta(|\mathbf{k}| - k_F)}{k_0 - \omega_{\mathbf{k}} + \imath \eta} + \frac{\theta(k_F - |\mathbf{k}|)}{k_0 - \omega_{\mathbf{k}} - \imath \eta} \right] \\ &\quad \times (4\pi a)^2 \int \frac{d^3\mathbf{k}'}{(2\pi)^3} \left[\frac{N(\mathbf{P}, \mathbf{k}')}{\frac{m}{\hbar} p_0 + \frac{m}{\hbar} k_0 - \frac{\mathbf{P}^2}{4} - \mathbf{k}'^2 + \imath \eta N(\mathbf{P}, \mathbf{k}')} + \frac{\mathcal{P}}{\mathbf{k}'^2 - \mathbf{q}^2} \right] \\ &= -\imath (4\pi a)^2 \int \frac{d^3\mathbf{k}}{(2\pi)^3} \int \frac{d^3\mathbf{k}'}{(2\pi)^3} (-2\pi\imath) \frac{1}{2\pi} \\ &\quad \left\{ \frac{\theta(|\mathbf{k}| - k_F) \delta_{N,-1} N}{\frac{m}{\hbar} p_0 + \frac{m}{\hbar} \omega_{\mathbf{k}} - \frac{\mathbf{P}^2}{4} - \mathbf{k}'^2 - \imath \eta} - \frac{\theta(k_F - |\mathbf{k}|) \delta_{N,1} N}{\frac{m}{\hbar} p_0 + \frac{m}{\hbar} \omega_{\mathbf{k}} - \frac{\mathbf{P}^2}{4} - \mathbf{k}'^2 + \imath \eta} - \frac{\mathcal{P}}{\mathbf{k}'^2 - \mathbf{q}^2} \theta(k_F - |\mathbf{k}|) \right\}\end{aligned}$$

using

$$\begin{aligned}\frac{m}{\hbar} \omega_{\mathbf{k}} &= m\epsilon_{\mathbf{k}} = \frac{\mathbf{k}^2}{2m} m = \frac{\mathbf{k}^2}{2} \\ \frac{1}{2} \mathbf{k}^2 - \frac{\mathbf{P}^2}{4} &= \frac{1}{2} \mathbf{k}^2 - \frac{1}{4} (\mathbf{p} + \mathbf{k})^2 = -\frac{1}{2} \mathbf{P}^2 + \frac{1}{4} (\mathbf{p} - \mathbf{k})^2 = -\frac{1}{2} \mathbf{p}^2 + \mathbf{q}^2\end{aligned}$$

$$\begin{aligned}
\hbar\Sigma_2^\star(p) = & (4\pi a)^2 \int \frac{d^3\mathbf{k}}{(2\pi)^3} \int \frac{d^3\mathbf{k}'}{(2\pi)^3} \\
& \times \left\{ \frac{\theta(k_F - |\mathbf{k}|)\theta(|\frac{1}{2}\mathbf{P} + \mathbf{k}'| - k_F)\theta(|\frac{1}{2}\mathbf{P} - \mathbf{k}'| - k_F)}{\frac{m}{\hbar}p_0 - \frac{1}{2}\mathbf{P}^2 + \mathbf{Q}^2 - \mathbf{k}'^2 + i\eta} \right. \\
& + \frac{\theta(|\mathbf{k}| - k_F)\theta(k_F - |\frac{1}{2}\mathbf{P} + \mathbf{k}'|)\theta(k_F - |\frac{1}{2}\mathbf{P} - \mathbf{k}'|)}{\frac{m}{\hbar}p_0 - \frac{1}{2}\mathbf{P}^2 + \mathbf{Q}^2 - \mathbf{k}'^2 - i\eta} \\
& \left. + \frac{\mathcal{P}}{\mathbf{k}'^2 - \mathbf{Q}^2} \theta(k_F - |\mathbf{k}|) \right\}
\end{aligned} \tag{2.7.29}$$

In terms of these proper self energy, the single-particle Green's function takes a following form

$$G(\mathbf{p}, p_0) = \left[p_0 - \frac{\hbar\mathbf{p}^2}{2m} - \Sigma_1^\star(\mathbf{p}, p_0) - \Sigma_2^\star(\mathbf{p}, p_0) \right]^{-1}$$

where the one-particle spectrum can be calculated as a pole of this Green's function. The pole of this function is generally complex value:

$$p_0 = k^{-1}\epsilon_{\mathbf{p}} + i\gamma_{\mathbf{p}}, \tag{2.7.30}$$

which satisfies this equation:

$$p_0 - \frac{\hbar\mathbf{p}^2}{2m} - \Sigma_1^\star(\mathbf{p}, p_0) - \Sigma_2^\star(\mathbf{p}, p_0) = 0 \tag{2.7.31}$$

The imaginary part $\gamma_{\mathbf{p}}$ stands for the inverse of a life-time of one-particle excitation $a_{\mathbf{p}}|F.S.\rangle$, while the real part $\epsilon_{\mathbf{p}}$ describes a “renormalized” one-particle spectrum, which includes the effect of particle-particle interactions.

Now that Σ_1^\star is free from \mathbf{p} and p_0 , and it is of the first order of small $k_F a$, we can solving this equation in favor for the pole, perturbatively in small $k_F a$.

Namely,

$$\begin{aligned}
\text{0th: } p_0 &= \frac{\hbar\mathbf{p}^2}{2m} \\
\text{1st: } p_0 &= \frac{\hbar\mathbf{p}^2}{2m} + \frac{\hbar k_F^2}{m} \cdot \frac{2k_F a}{3\pi}
\end{aligned}$$

To the 2nd order in small $k_F a$, we have only to substitute the zero-th order solution of p_0 into the argument of Σ_2^\star in this equation:

$$p_0 - \frac{\hbar\mathbf{p}^2}{2m} - \frac{\hbar k_F^2}{m} \cdot \frac{2k_F a}{3\pi} - \Sigma_2^\star(\mathbf{p}, \frac{\hbar\mathbf{p}^2}{2m}) = 0$$

or

$$\begin{aligned}
p_0 &= \frac{\hbar\mathbf{p}^2}{2m} + \frac{\hbar k_F^2}{m} \cdot \frac{2k_F a}{3\pi} + \Sigma_2^\star(\mathbf{p}, \frac{\hbar\mathbf{p}^2}{2m}) \\
&= \frac{\hbar\mathbf{p}^2}{2m} + \frac{\hbar k_F^2}{m} \cdot \frac{2k_F a}{3\pi} + \frac{\hbar}{m} 16\pi^2 a^2 \int \frac{d^3\mathbf{k} d^3\mathbf{k}'}{(2\pi)^6} \\
&\quad \left[\frac{\theta(k_F - |\mathbf{k}|)\theta(|\frac{1}{2}\mathbf{P} + \mathbf{k}| - k_F)\theta(|\frac{1}{2}\mathbf{P} - \mathbf{k}| - k_F)}{\frac{1}{2}\mathbf{P}^2 - \frac{1}{2}\mathbf{P}^2 + \mathbf{Q}^2 - \mathbf{k}'^2 + i\eta} \right. \\
&\quad + \frac{\theta(|\mathbf{k}| - k_F)\theta(k_F - |\frac{1}{2}\mathbf{P} + \mathbf{k}|)\theta(k_F - |\frac{1}{2}\mathbf{P} - \mathbf{k}|)}{\frac{1}{2}\mathbf{P}^2 - \frac{1}{2}\mathbf{P}^2 + \mathbf{Q}^2 - \mathbf{k}'^2 - i\eta} \\
&\quad \left. - \frac{\mathcal{P}}{\mathbf{Q}^2 - \mathbf{k}'^2} \theta(k_F - |\mathbf{k}|) \right]
\end{aligned} \tag{2.7.32}$$

$$\begin{aligned}
&= \frac{\hbar \mathbf{p}^2}{2m} + \frac{\hbar^2 k_F^2}{m} \left\{ \frac{2k_F a}{3\pi} + 16\pi^2 (k_F a)^2 \times \right. \\
&\quad \int \frac{d^3 \mathbf{k} d^3 \mathbf{k}'}{(2\pi)^6} \left[\frac{\theta(1 - |\mathbf{k}|) \theta(|\frac{1}{2}\mathbf{P} + \mathbf{k}| - 1) \theta(|\frac{1}{2}\mathbf{P} - \mathbf{k}| - 1)}{\mathbf{q}^2 - \mathbf{k}'^2 + i\eta} \right. \\
&\quad + \frac{\theta(|\mathbf{k}| - 1) \theta(1 - |\frac{1}{2}\mathbf{P} + \mathbf{k}|) \theta(1 - |\frac{1}{2}\mathbf{P} - \mathbf{k}|)}{\mathbf{q}^2 - \mathbf{k}'^2 - i\eta} \\
&\quad \left. \left. - \frac{\mathcal{P}}{\mathbf{q}^2 - \mathbf{k}'^2} \theta(1 - |\mathbf{k}|) \right] + o((k_F a)^3) \right\}
\end{aligned} \tag{2.7.33}$$

where $\mathbf{P} = \mathbf{p} + \mathbf{k}$, $\mathbf{q} = \frac{1}{2}(\mathbf{p} - \mathbf{k})$.

Note also that, within this integral, all the momentum variables are rescaled by k_F , so that this integral itself is on the order of unit.

This momentum integrals require an involved mathematics, so that I'm not going to calculate this here. But, the integrals were carried out by Victor Galitskii in 1958 in this paper: Sov. Phys. JETP **7**, 104, (1958).

To put his result first, the inverse of the life time is given like this:

$$\gamma_{\mathbf{p}} = \frac{\hbar k_F^2}{2m} \left\{ \frac{2}{\pi} (k_F a)^2 \left(\frac{k_F - |\mathbf{p}|}{k_F} \right)^2 \text{sgn}(k_F - |\mathbf{p}|) + o((k_F a)^3) \right\} \tag{2.7.34}$$

which is valid when $||\mathbf{p}| - k_F|$ is much smaller than k_F itself.

Thus when $|\mathbf{p}| \rightarrow k_F$, the life time of the one-particle excited state becomes infinite.

The long-lived single-particle excitations are often called as quasi-particles.

His calculation also dictates that $\gamma_{\mathbf{p}}$ changes the sign at k_F , which is the Fermi momentum in the non-interacting electron gas. This implies that the ground state in interacting system remains a Fermi sea filled up the wavenumber k_F , but a different dispersion relation for quasi-particle spectrum $\epsilon_{\mathbf{p}}$.

Since $\gamma_{\mathbf{p}}$ changes sign at k_F , the preceding general argument based on Lehmann representation concludes that the chemical potential is equal to a quasi particle energy evaluated at k_F .

$$\mu = \epsilon_{k_F}$$

Galitskii's evaluation of eq.(2.7.33) shows that the right hand side is calculated as follows

$$\mu = \frac{\hbar^2 k_F^2}{2m} \left[1 + \frac{4}{3\pi} k_F a + \frac{4}{15\pi^2} (11 - 2\ln 2) (k_F a)^2 \right] \tag{2.7.35}$$

up to the 2nd order in small $k_F a$,

Near the Fermi surface, the quasi-particle spectrum can be expanded in small $\mathbf{p} - |k_F|$:

$$\begin{aligned}
\epsilon_{|\mathbf{p}|} &= \epsilon_{k_F} + \left. \frac{\partial \epsilon_{|\mathbf{p}|}}{\partial |\mathbf{p}|} \right|_{k_F} (p - k_F) + \dots \\
&\equiv \epsilon_{k_F} + \frac{\hbar^2 k_F}{m^*} (p - k_F) + \dots
\end{aligned} \tag{2.7.36}$$

From which we can define effective mass at the Fermi surface.

A detailed calculation with eq.(2.7.24) shows that the effective mass thus defined is given as follow to the 2nd order in small $k_F a$.

$$\frac{m^*}{m} = 1 + \frac{8}{15\pi^2} (7\ln 2 - 1) (k_F a)^2 \tag{2.7.37}$$

m^* has no 1st order correction in $k_F a$, which is simply because Σ_1^\star has no momentum dependence.

The effective mass of quasi-particles determines the heat capacity of the system in the Low-temperature limit as

$$\frac{C_V}{V} \rightarrow \frac{k_B^2 T m^* k_F}{3\hbar^2} \quad (T \rightarrow 0) \quad (2.7.38)$$

Another physical quantity which we can calculate along this line of argument is the ground state energy.

According to standard theory of thermodynamics ($dE = TdS - PdV + \mu dN$), the chemical potential is given by a partial derivative of the internal energy E with respect to the total number of particles with fixed volume V and entropy S .

$$\mu = \left(\frac{\partial E}{\partial N} \right)_{S,V} \quad (2.7.39)$$

As discussed above, the ground state remains a Fermi sea filled up to the wavenumber k_F , which has no degeneracy in general. Thus, as far as physical quantities are evaluated at ground state, we can regard that the entropy is fixed to be zero.

Thus, we can replace this by

$$\mu = \left(\frac{\partial E}{\partial N} \right)_V$$

Now that the chemical potential is evaluated as a polynomial of small $k_F a$, where k_F depends on V and N as

$$k_F = \left(\frac{3\pi^2 N}{V} \right)^{\frac{1}{3}} \quad (2.7.40)$$

As such, the ground state energy at given particle, N can be calculated from the integral of μ over the particle number up to N .¹

$$\begin{aligned} E &= \int_0^N dN' \mu(V, N') \\ &= \int_0^N dN' \frac{\hbar^2 k_F^2}{2m} \times \left[1 + \frac{4}{3\pi} k_F a + \frac{4}{15\pi^2} (11 - 2\ln 2) (k_F a)^2 \right] \\ &= \frac{\hbar^2 k_F^2}{2m} \cdot N \cdot \left[\frac{3}{2+3} + \frac{3}{3+3} \cdot \frac{4}{3\pi} k_F a + \frac{3}{4+3} \cdot \frac{4}{15\pi^2} (11 - 2\ln 2) (k_F a)^2 \right] \end{aligned}$$

From which we obtain

$$\frac{E}{N} = \frac{\hbar^2 k_F^2}{2m} \left[\frac{3}{5} + \frac{2}{3\pi} k_F a + \frac{4}{35\pi^2} (11 - 2\ln 2) (k_F a)^2 + \dots \right] \quad (2.7.41)$$

This is the ground state energy given perturbatively in small $k_F a$. Clearly the first term is just a ground state energy in the non-interacting limit, while the next two terms are correction to this.

¹

$$\begin{aligned} \int_0^N dN' (k_F(N'))^\lambda &= \int_0^N dN' \left(\frac{3\pi^2}{V} \right)^{\frac{\lambda}{3}} (N')^{\frac{\lambda}{3}} \\ &= \left(\frac{3\pi^2}{V} \right)^{\frac{\lambda}{3}} \frac{1}{\frac{\lambda}{3} + 1} \left[(N')^{\frac{\lambda}{3} + 1} \right]_0^N \\ &= \left(\frac{3\pi^2}{V} \right)^{\frac{\lambda}{3}} \frac{3}{\lambda + 3} N^{\frac{\lambda}{3} + 1} \\ &= \frac{3}{\lambda + 3} k_F(N) \cdot N \end{aligned}$$

Especially, the 2nd order correction was originally obtained by Huang and Yang and Lee and Yang, prior to Galitskii.

$$\begin{cases} \text{K.Huang \& C.N.Yang Phy.Rev.}\underline{105} \text{ 767 (1957)} \\ \text{T.D.Lee \& C.N.Yang Phy.Rev.}\underline{105} \text{ 1119(1957)} \end{cases}$$

2.8 Ring approximation: High density Fermi gas with long-range Coulomb interaction

The 2nd system considered is a degenerate, high-density electron gas with long-range Coulomb interaction. The most straightforward approximation approach is to sum up all the leading order terms in the proper self-energy.

In the case of the low-density imperfect Fermi gas, the leading terms summed up is the ladder-type diagrams. In the case of the high-density Fermi gas with long-range Coulomb interaction, the leading terms summed up is the ring-type diagrams.

In the following, I will describe how the ground state energy of high-density Fermi gas with long-range Coulomb interaction can be calculated from the ring-type diagrams.

In the previous section, we have obtained ground state energy, by calculating the self-energy. In this section, however, we will employ an alternative way of calculating the ground state energy.

We begin with a general Hamiltonian for interacting many-particle systems

$$\hat{H}(\lambda) = \hat{H}_0 + \lambda \hat{H}_1$$

where \hat{H}_0 denote the kinetic energy part and \hat{H}_1 denote the interaction part.

We put λ as an overall coefficient of the interaction part, so that $\hat{H}(\lambda)$ interpolates between non-interacting Hamiltonian and interacting Hamiltonian.

$$\hat{H}(\lambda = 0) = \hat{H}_0, \hat{H}(\lambda = 1) = \hat{H} = \hat{H}_0 + \hat{H}_1$$

Suppose that the ground state of $\hat{H}(\lambda)$ is non-degenerate and is given by $|\Psi_0(\lambda)\rangle$:

$$\hat{H}(\lambda)|\Psi_0(\lambda)\rangle = E_0(\lambda)|\Psi_0(\lambda)\rangle$$

with $\langle\Psi_0(\lambda)|\Psi_0(\lambda)\rangle = 1$, we have

$$E_0(\lambda) = \langle\Psi_0(\lambda)|\hat{H}(\lambda)|\Psi_0(\lambda)\rangle$$

so that the derivative of $E_0(\lambda)$ with respect to λ is given by

$$\begin{aligned} \frac{d}{d\lambda} E_0(\lambda) &= \left\langle \frac{\partial \Psi_0(\lambda)}{\partial \lambda} \right| \hat{H}(\lambda) | \Psi_0(\lambda) \rangle + \langle \Psi_0(\lambda) | \hat{H}(\lambda) \left| \frac{\partial \Psi_0(\lambda)}{\partial \lambda} \right\rangle + \langle \Psi_0(\lambda) | \hat{H}_1(\lambda) | \Psi_0(\lambda) \rangle \\ &= E_0(\lambda) \cdot \frac{\partial}{\partial \lambda} \{ \langle \Psi_0(\lambda) | \Psi_0(\lambda) \rangle \} + \langle \Psi_0(\lambda) | \hat{H}_1(\lambda) | \Psi_0(\lambda) \rangle \\ &= \langle \Psi_0(\lambda) | \hat{H}_1(\lambda) | \Psi_0(\lambda) \rangle \end{aligned}$$

Now, we want the ground state energy at $\lambda = 1$, which is given by a integral of the r.h.s with respect to λ :

$$E_0(\lambda = 1) = E_0(\lambda = 0) + \int_{\lambda=0}^{\lambda=1} \frac{d\lambda}{\lambda} \langle \Psi_0(\lambda) | \lambda \hat{H}_1 | \Psi_0(\lambda) \rangle \quad (2.8.1)$$

where the ground state energy in the non-interacting limit, $E_0(\lambda = 0)$, is already known.

As such we have only calculate the integrand as a function of λ .

To do this explicitly, let us restrict the system studied here to a spatially homogeneous Fermi system with a spin-independent static interaction potential. Namely,

$$\begin{cases} \hat{H}_0 = \sum_{\alpha} \int d^3\mathbf{x} \psi_{\alpha}^{\dagger}(\mathbf{x}) \left[-\frac{\hbar^2 \nabla_{\mathbf{x}}^2}{2m} \right] \psi_{\alpha}(\mathbf{x}) \\ \hat{H}_1 = \frac{1}{2} \sum_{\alpha, \beta} \int d^3\mathbf{x} \int d^3\mathbf{x}' V(\mathbf{x} - \mathbf{x}') \psi_{\alpha}^{\dagger}(\mathbf{x}) \psi_{\beta}^{\dagger}(\mathbf{x}') \times \psi_{\beta}(\mathbf{x}') \psi_{\alpha}(\mathbf{x}) \end{cases}$$

$$\begin{aligned} \langle \Psi_0(\lambda) | \lambda \hat{H}_1(\lambda) | \Psi_0(\lambda) \rangle &\equiv \langle \hat{H}_1 \rangle_{\lambda} \\ &= \frac{\lambda}{2} \int d^3\mathbf{x} d^3\mathbf{x}' V(\mathbf{x} - \mathbf{x}') \sum_{\alpha, \beta} \langle \psi_{\alpha}^{\dagger}(\mathbf{x}) \psi_{\beta}^{\dagger}(\mathbf{x}') \psi_{\beta}(\mathbf{x}') \psi_{\alpha}(\mathbf{x}) \rangle_{\lambda} \\ &= \frac{1}{2} \int d^3\mathbf{x} d^3\mathbf{x}' V(\mathbf{x} - \mathbf{x}') \times \left[\langle \hat{n}(\mathbf{x}) \hat{n}(\mathbf{x}') \rangle_{\lambda} - \delta^3(\mathbf{x} - \mathbf{x}') \langle \hat{n}(\mathbf{x}) \rangle_{\lambda} \right] \end{aligned}$$

In a spatially homogeneous system, this can be replaced by a constant, which is identical to $\frac{N}{V}$ irrespective of λ .

$$= \frac{1}{2} \int d^3\mathbf{x} d^3\mathbf{x}' V(\mathbf{x} - \mathbf{x}') \times \left[\langle \hat{n}(\mathbf{x}) \hat{n}(\mathbf{x}') \rangle_{\lambda} - \delta^3(\mathbf{x} - \mathbf{x}') n \right]$$

For later convenience, we introduce deviation of the density from its average

$$\tilde{n}(\mathbf{x}) \equiv \hat{n}(\mathbf{x}) - \langle \hat{n}(\mathbf{x}) \rangle = \hat{n}(\mathbf{x}) - n$$

In terms of this operator, the r.h.s. can be further calculated as

$$= \frac{\lambda}{2} \int d^3\mathbf{x} d^3\mathbf{x}' V(\mathbf{x} - \mathbf{x}') \times \left[\langle \tilde{n}(\mathbf{x}) \tilde{n}(\mathbf{x}') \rangle_{\lambda} + n^2 - \delta^3(\mathbf{x} - \mathbf{x}') n \right]$$

To make use of the diagrammatic analysis, let us introduce a time-ordered density-density correlation function.

$$\imath D^{\lambda}(x, x') \equiv \frac{\langle \Psi_0(\lambda) | \Gamma[\tilde{n}_{H(\lambda)}(x) \tilde{n}_{H(\lambda)}(x')] | \Psi_0(\lambda) \rangle}{\langle \Psi_0(\lambda) | \Psi_0(\lambda) \rangle} \quad (2.8.2)$$

where both x and x' are four-component coordinate vector $x \equiv (\mathbf{x}, t)$, $x' \equiv (\mathbf{x}', t')$.

In terms of this, the r.h.s. can be rewritten as:

$$= \frac{\lambda}{2} \int d^3\mathbf{x} d^3\mathbf{x}' V(\mathbf{x} - \mathbf{x}') \times \left[\imath D^{\lambda}(\mathbf{x}, t; \mathbf{x}', t) + n^2 - \delta^3(\mathbf{x} - \mathbf{x}') n \right]$$

where this is just an equal-time density-density correlation function.

We can further decompose the r.h.s. into the contribution at $\lambda = 0$ and the other.

$$\begin{aligned} &= \frac{\lambda}{2} \int d^3\mathbf{x} d^3\mathbf{x}' V(\mathbf{x} - \mathbf{x}') \times \left[\imath D^0(\mathbf{x}, t; \mathbf{x}', t) + n^2 - \delta^3(\mathbf{x} - \mathbf{x}') n \right] \\ &\quad + \frac{\lambda}{2} \int d^3\mathbf{x} d^3\mathbf{x}' V(\mathbf{x} - \mathbf{x}') \times \left[\imath D^{\lambda}(\mathbf{x}, t; \mathbf{x}', t) - \imath D^0(\mathbf{x}, t; \mathbf{x}', t) \right] \\ &= \lambda \langle \hat{H}_1 \rangle_{\lambda=0} + \frac{\lambda}{2} \int d^3\mathbf{x} d^3\mathbf{x}' V(\mathbf{x} - \mathbf{x}') \times \left[\imath D^{\lambda}(\mathbf{x}, t; \mathbf{x}', t) - \imath D^0(\mathbf{x}, t; \mathbf{x}', t) \right] \end{aligned}$$

substituting this back into eq.(2.8.1), we obtain¹

$$\begin{aligned}
 E_0(\lambda = 1) &= \langle H_0 \rangle_{\lambda=0} + \int_{\lambda=0}^{\lambda=1} \frac{d\lambda}{\lambda} \left\{ \lambda \langle H_1 \rangle_{\lambda=0} + \frac{\lambda}{2} \int d^3\mathbf{x} d^3\mathbf{x}' V(\mathbf{x} - \mathbf{x}') \right. \\
 &\quad \times \left[\imath D^\lambda(\mathbf{x}, t; \mathbf{x}', t) - \imath D^0(\mathbf{x}, t; \mathbf{x}', t) \right] \left. \right\} \\
 &= \langle H \rangle_{\lambda=0} + \frac{1}{2} \int_0^1 \frac{d\lambda}{\lambda} \int d^3\mathbf{x} d^3\mathbf{x}' \lambda V(\mathbf{x} - \mathbf{x}') \left[\imath D^\lambda(\mathbf{x}, t; \mathbf{x}', t) - \imath D^0(\mathbf{x}, t; \mathbf{x}', t) \right] \\
 &= \underline{\langle H \rangle_{\lambda=0}} + E_{corr}
 \end{aligned} \tag{2.8.3}$$

To calculate the density-density correlation function, let us briefly summarize the Feynman's rule for the two-particle Green's function.

$$G_{\alpha,\beta;\gamma,\delta}(\mathbf{x}_1, t_1, \mathbf{x}_2, t_2; \mathbf{x}'_1, t'_1, \mathbf{x}'_2, t'_2) = (-\imath)^2 \frac{\langle \Psi_0 | \Gamma[\hat{\psi}_\alpha(\mathbf{x}_1, t_1) \hat{\psi}_\beta(\mathbf{x}_2, t_2) \hat{\psi}_\gamma^\dagger(\mathbf{x}'_2, t'_2) \hat{\psi}_\delta^\dagger(\mathbf{x}'_1, t'_1)] | \Psi_0 \rangle}{\langle \Psi_0 | \Psi_0 \rangle} \tag{2.8.4}$$

where $|\Psi_0\rangle$ is the ground state wavefunction of the interacting Fermi system.

So far, these fermion fields are all given in the Heisenberg picture.

Using the same mathematics as we did for eq.(??), we can reexpress this in terms of the interaction picture.

$$\begin{aligned}
 \imath^2 G_{\alpha,\beta;\gamma,\delta}(\mathbf{x}_1, t_1, \mathbf{x}_2, t_2; \mathbf{x}'_1, t'_1, \mathbf{x}'_2, t'_2) &= \sum_{\nu=0}^{+\infty} \left(\frac{-\imath}{\hbar} \right)^\nu \frac{1}{\nu!} \int_{-\infty}^{+\infty} dt_1 \dots \int_{-\infty}^{+\infty} dt_\nu \\
 &\quad \times \frac{\langle \Phi_0 | \Gamma[\hat{H}_1(t_1) \dots \hat{H}_1(t_\nu) \psi_\alpha(x_1) \psi_\beta(x_2) \psi_\gamma^\dagger(x'_2) \psi_\delta^\dagger(x'_1)] | \Phi_0 \rangle}{\langle \Phi_0 | S | \Phi_0 \rangle} \\
 S &= \sum_{\nu=0}^{+\infty} \left(\frac{-\imath}{\hbar} \right)^\nu \frac{1}{\nu!} \int_{-\infty}^{+\infty} dt_1 \dots \int_{-\infty}^{+\infty} dt_\nu \Gamma[\hat{H}_1(t_1) \dots \hat{H}_1(t_\nu)]
 \end{aligned}$$

where $|\Phi_0\rangle$ is the ground state wavefunction in non-interacting system.

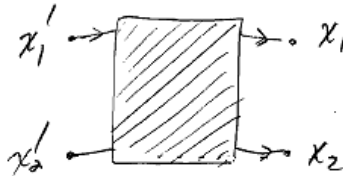
Using the Wick's theorem, we can rewrite this time-ordered product in terms of a sum of many normal-ordered products.

Among them, all the uncontracted normal-ordered products vanish for the ground state average.

So that, both the numerator and the denominator can be expanded in terms of fully-contracted normal-ordered products.

As in the single-particle Green's function, the fully-contracted normal-ordered products are symbolically expressed in terms of Feynman diagram.

Since we have two creation fields and two annihilation fields, which are nothing to do with the interaction potentials, we have two external lines which connects x'_1 and x'_2 into x_1 and x_2 .



As we did in the single-particle Green's function, ν -numbers of interaction Hamiltonians can be divided into n -numbers of those interaction Hamiltonian which are connected to these external lines. And $(\nu - n)$ -numbers of the interaction Hamiltonians which are disconnected from these external lines.

¹The underlined part equals to $\langle \Psi_0(\lambda = 0) | H | \Psi_0(\lambda = 0) \rangle$.

The temporal arguments of the former group of interaction Hamiltonian and those of the latter are nothing to do with each other.

As such, we can rewrite this as

$$\begin{aligned}
& (i)^2 \tilde{G}_{\alpha,\beta;\gamma,\delta}^{(\nu)}(x_1, x_2; x'_1, x'_2) \\
&= \left(\frac{-i}{\hbar}\right)^\nu \frac{1}{\nu!} \int_{-\infty}^{+\infty} dt_1 \dots \int_{-\infty}^{+\infty} dt_\nu \\
&\quad \times \langle \Phi_0 | \Gamma[\hat{H}_1(t_1) \dots \hat{H}_1(t_\nu) \hat{\psi}_\alpha(x_1) \hat{\psi}_\beta(x_2) \hat{\psi}_\gamma^\dagger(x'_2) \hat{\psi}_\delta^\dagger(x'_1)] | \Phi_0 \rangle \\
&= \sum_{n=0}^{\nu} \left(\frac{-i}{\hbar}\right)^\nu \frac{1}{\nu!} \cdot \frac{\nu!}{n!(\nu-n)!} \\
&\quad \times \int_{-\infty}^{+\infty} dt_1 \dots \int_{-\infty}^{+\infty} dt_n \langle \Phi_0 | \Gamma[\hat{H}_1(t_1) \dots \hat{H}_1(t_n) \hat{\psi}_\alpha(x_1) \hat{\psi}_\beta(x_2) \hat{\psi}_\gamma^\dagger(x'_2) \hat{\psi}_\delta^\dagger(x'_1)] | \Phi_0 \rangle \\
&\quad \times \int_{-\infty}^{+\infty} dt_1 \dots \int_{-\infty}^{+\infty} dt_{\nu-n} \langle \Phi_0 | \Gamma[\hat{H}_1(t_1) \dots \hat{H}_1(t_{\nu-n})] | \Phi_0 \rangle
\end{aligned}$$

Here we put this factor, because the numbers of ways of dividing ν -number of interaction Hamiltonians into n -numbers of interaction Hamiltonians and $(\nu - n)$ -numbers of interaction Hamiltonians.

When we take the summation over ν , the denominator can be decomposed into two part:

$$\begin{aligned}
& (i)^2 \sum_{\nu=0}^{+\infty} \tilde{G}_{\alpha,\beta;\gamma,\delta}^{(\nu)}(x_1, x_2; x'_1, x'_2) \\
&= \sum_{n=0}^{+\infty} \left(\frac{-i}{\hbar}\right)^n \frac{1}{n!} \int_{-\infty}^{+\infty} dt_1 \dots \int_{-\infty}^{+\infty} dt_n \\
&\quad \times \langle \Phi_0 | \Gamma[\hat{H}_1(t_1) \dots \hat{H}_1(t_n) \hat{\psi}_\alpha(x_1) \hat{\psi}_\beta(x_2) \hat{\psi}_\gamma^\dagger(x'_2) \hat{\psi}_\delta^\dagger(x'_1)] | \Phi_0 \rangle \\
&\quad \times \sum_{m=0}^{+\infty} \left(\frac{-i}{\hbar}\right)^m \frac{1}{m!} \int_{-\infty}^{+\infty} dt_1 \dots \int_{-\infty}^{+\infty} dt_m \langle \Phi_0 | \Gamma[\hat{H}_1(t_1) \dots \hat{H}_1(t_m)] | \Phi_0 \rangle
\end{aligned}$$

where the latter part can be set off by the denominator.

As a result, the two-particle Green's function can be also given by the sum of all connected Feynman diagrams.

$$\begin{aligned}
& (i)^2 G_{\alpha,\beta;\gamma,\delta}(x_1, x_2; x'_1, x'_2) \\
&= \sum_{\nu=0}^{+\infty} \left(\frac{-i}{\hbar}\right)^\nu \frac{1}{\nu!} \int_{-\infty}^{+\infty} dt_1 \dots \int_{-\infty}^{+\infty} dt_\nu \\
&\quad \times \langle \Phi_0 | \Gamma[\hat{H}_1(t_1) \dots \hat{H}_1(t_\nu) \hat{\psi}_\alpha(x_1) \hat{\psi}_\beta(x_2) \hat{\psi}_\gamma^\dagger(x'_2) \hat{\psi}_\delta^\dagger(x'_1)] | \Phi_0 \rangle
\end{aligned}$$

Feynman's rule for the two particle Green's function can be readily derived as a generalization of the Feynman's rule for the single-particle Green's function.

For the n -th order contribution to the two-particle Green's function $G_{\alpha,\beta;\gamma,\delta}(x_1, x_2; x'_2, x'_1)$:

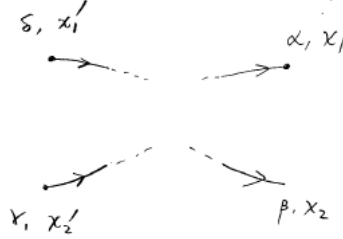
- Draw all topologically distinct connected diagrams with n interaction lines, ν and $(2n + 2)$ Green's function G^0 .
- Label each vertex with a four dimensional space-time coordinate x_i .
- Each solid line represents a Green's function $G_{\alpha,\beta}^0(x, y)$, running from y to x .

- (d) Each wavy line represents an interaction

$$U_{\lambda,\lambda';\mu,\mu'}(x,y) = V_{\lambda,\lambda';\mu,\mu'}(\mathbf{x},\mathbf{y})\delta(t_x - t_y)$$

- (e) Integrate all internal variables over space & time.
 (f) Take the sum over all the spin indices along each fermion line.

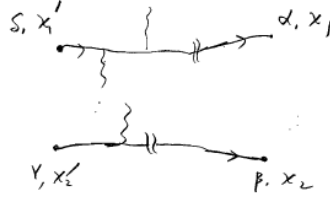
Now that we have four external variables associated with these four fermion field operators.



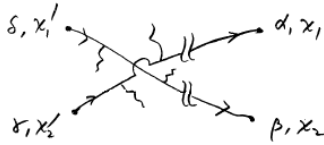
Two fermion lines which start from these two points should be terminated to either α, x_1 or β, x_2 .

We have two way of connecting these two creation fields with these two annihilation fields.

One is to connect (δ, x_1') with (α, x_1) and (γ, x_2') with (β, x_2) .



The other is to connect (δ, x_1') with (β, x_2) and (γ, x_2') with (α, x_1)



- (g) We affix a sign factor $(-1)^F$ to the former type of Feynman diagram where “ F ” is the number of closed fermion loops in the diagram, while we affix a sign factor $(-1)^{F+1}$ to the ladder type of Feynman diagram.

* This additional (-1) factor simply comes from the (-1) factor which we get when exchanging these two creation operators.

The n -th order term has an explicit numerical factor $(\frac{-i}{\hbar})^n$, while the $(2n+2)$ contractions of field operators contribute an additional factor $(i)^{2n+2}$.

On top of this, we have $(-i)^2$ factor, in the definition of the two-particle Green's function.

- (h) Thus, to compute $G_{\alpha,\beta;\gamma,\delta}(x_1, x_2; x_2', x_1')$, we assign a factor $(-i)^2 \cdot (\frac{-i}{\hbar})^n (i)^{2n+2} = (\frac{i}{\hbar})^n$ to every n -th order term.

- (i) Finally, for a Green's function with equal time variables must be interpreted as $G_{\alpha,\beta}^0(\mathbf{x}, t; \mathbf{x}', t+0)$.

These are all about the Feynman's rule for two-particle Green's function in the real space representation.

With this Feynman's rule in mind, let us go back to the density-density correlation function:

$$\iota D(x, x') = \frac{\langle \Psi_0 | \Gamma[\tilde{n}_H(x) \tilde{n}_H(x')] | \Psi_0 \rangle}{\langle \Psi_0 | \Psi_0 \rangle}$$

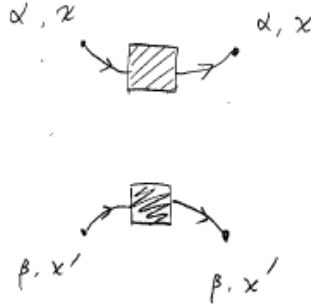
where we omit the fictitious mechanical parameter λ .

The r.h.s. can be calculated as follows:

$$\begin{aligned} \iota D(x, x') &= \frac{1}{\langle \Psi_0 | \Psi_0 \rangle} \left\{ \frac{\langle \Psi_0 | \Gamma[\{\sum_{\alpha} \psi_{H,\alpha}^{\dagger}(x) \psi_{H,\alpha}(x) - \langle n_H(x) \rangle\} \{\sum_{\beta} \psi_{H,\beta}^{\dagger}(x') \psi_{H,\beta}(x') - \langle n_H(x') \rangle\}] | \Psi_0 \rangle}{\langle \Psi_0 | \Psi_0 \rangle} \right\} \\ &= \frac{1}{\langle \Psi_0 | \Psi_0 \rangle} \sum_{\alpha, \beta} \langle \Psi_0 | \Gamma[\psi_{H,\alpha}^{\dagger}(x) \psi_{H,\alpha}(x) \psi_{H,\beta}^{\dagger}(x') \psi_{H,\beta}(x')] | \Psi_0 \rangle \\ &\quad - \langle n_H(x) \rangle \langle n_H(x') \rangle \\ &= - \left(\sum_{\alpha, \beta} G_{\alpha, \beta; \beta, \alpha}(x, x'; x, x') - \sum_{\alpha} G_{\alpha, \alpha}(x, x) \sum_{\beta} G_{\beta, \beta}(x', x') \right) \end{aligned}$$

Now both the two-particle Green's function here and the one-particle Green's function here are given by the sum of connected Feynman diagrams.

Especially, two-particle Green's function contain those Feynman diagrams in which the external line from x, α to x, α and that from x', β to x', β are disconnected.



Connected Feynman diagrams to $G_{\alpha, \beta; \beta, \alpha}(x, x'; x, x')$, in which an external line from (α, x) to (α, x) and that from (β, x') to (β, x') are disconnected.

Since each of these two are described by single particle Green's functions $G_{\alpha, \alpha}(x, x)$ and $G_{\beta, \beta}(x', x')$ respectively, such Feynman diagrams are completely cancelled by this second term.

Such a cancellation becomes possible, because any n -th order term in this class of diagrams are accompanied by $(\frac{i}{\hbar})^n (-1)^F$ where F denotes the number of fermion loop.

Such a factor can be always decomposed into two part:

$$\left(\frac{i}{\hbar}\right)^n (-1)^F = \left(\frac{i}{\hbar}\right)^{n_1} (-1)^{F_1} \cdot \left(\frac{i}{\hbar}\right)^{n_2} (-1)^{F_2}$$

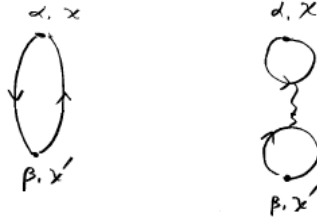
each of which can be ascribed to these two part respectively.

Thus.

$$\begin{aligned} -iD(x, x') &= \sum_{\alpha, \beta} G_{\alpha, \beta; \beta, \alpha}(x, x'; x, x') - \sum_{\alpha} G_{\alpha, \alpha}(x, x) \cdot \sum_{\beta} G_{\beta, \beta}(x', x') \\ &= \sum_{\alpha, \beta} G_{\alpha, \beta; \beta, \alpha}^{(C)}(x, x'; x', x) \end{aligned}$$

where this superscript “(C)” means that the fermion field at (α, x) and the fermion field at (β, x') is always connected in these Feynman diagrams.

Examples of such Feynman diagrams include,

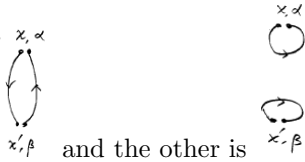


Here, I put together two fermion fields whose arguments are same.

As in eq.(2.8.3), to obtain the ground state energy, we need the density-density correlation function as a function of λ , where λ denotes a mechanical(fictitious) parameter we put in front of the interaction potential.

When this λ is taken to be zero, we don't have any interaction potential. In such a case, we only have two way of contracting the fermion fields [here](#).

$$-iD^{\lambda=0}(x, x') = \sum_{\alpha, \beta} G_{\alpha, \beta; \beta, \alpha}^{\lambda=0}(x, x'; x, x') - \sum_{\alpha} G_{\alpha, \alpha}^{\lambda=0}(x, x) \cdot \sum_{\beta} G_{\beta, \beta}^{\lambda=0}(x', x')$$



One is x, α and the other is x', β .

The latter one is cancelled by the 2nd term($\sum_{\beta} G_{\beta, \beta}^{\lambda=0}(x', x')$), so that the density-density correlation function is given by the former one.

Following the Feynman's rule described so far, the analytic expression for the former one is given as follows:

$$\begin{aligned} -iD^{\lambda=0}(x, x') &= G_{\alpha, \beta; \beta, \alpha}^{\lambda=0, (C)}(x, x'; x, x') \\ &= \underline{(-)} G_{\beta, \alpha}^0(x', x) G_{\alpha, \beta}^0(x, x') \end{aligned}$$

where this underlined $(-)$ sign comes from the interchange between these two fields.

$$D^{\lambda=0}(x, x') = -iG_{\alpha, \beta}^0(x, x') G_{\beta, \alpha}^0(x', x) \equiv \hbar \Pi^0(x, x')$$

where we call this as a bare polarization part.

Similarly, we can define polarization part as the density-density correlation function.

$$D(x, x') \equiv \hbar \Pi(x, x') \quad (2.8.4')$$

(or $D^\lambda(x, x') \equiv \hbar \Pi^\lambda(x, x')$)

Since our system is spatially and temporally homogeneous, the correlation function between two points depends only on their relative distance:

$$D^\lambda(x, x') = D^\lambda(\mathbf{x} - \mathbf{x}', t - t')$$

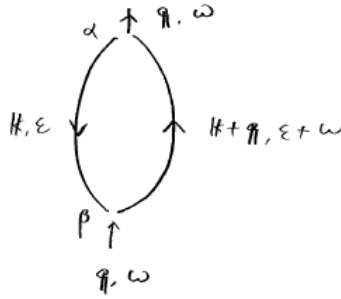
so that we can introduce its Fourier series.

$$= \int \frac{d\mathbf{q}}{(2\pi)^3} \int \frac{d\omega}{2\pi} e^{i\mathbf{q}(\mathbf{x}-\mathbf{x}') - i\omega(t-t')} D^\lambda(\mathbf{q}, \omega)$$

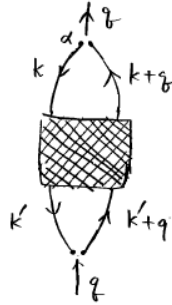
In the momentum representation, the bare polarization part is given as follows.

$$\hbar \Pi^0(\mathbf{q}, \omega) = -i \int \frac{d\mathbf{k}}{(2\pi)^3} \int \frac{d\epsilon}{2\pi} G_{\alpha,\beta}^0(\mathbf{k} + \mathbf{q}, \epsilon + \omega) G_{\beta,\alpha}^0(\mathbf{k}, \epsilon)$$

and: the corresponding Feynman diagram can be described like this



Similarly, Feynman diagrams for the full polarization part can be schematically described as follows:



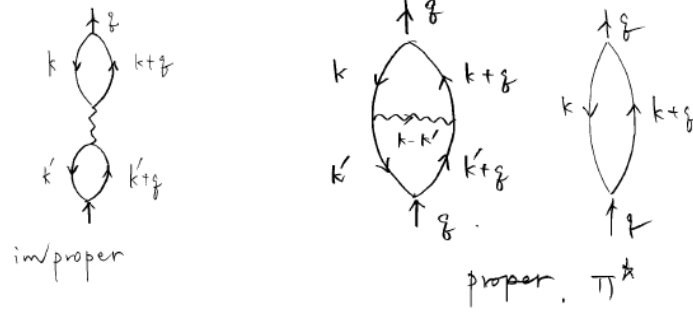
where k, k', q here are all four-dimensional momentum vector.

This shadow region contains all possible Feynman diagrams which connect the upper part & the lower part.

Such diagrams can be generally classified into proper part and improper part, in the same way as we did for the self-energy.

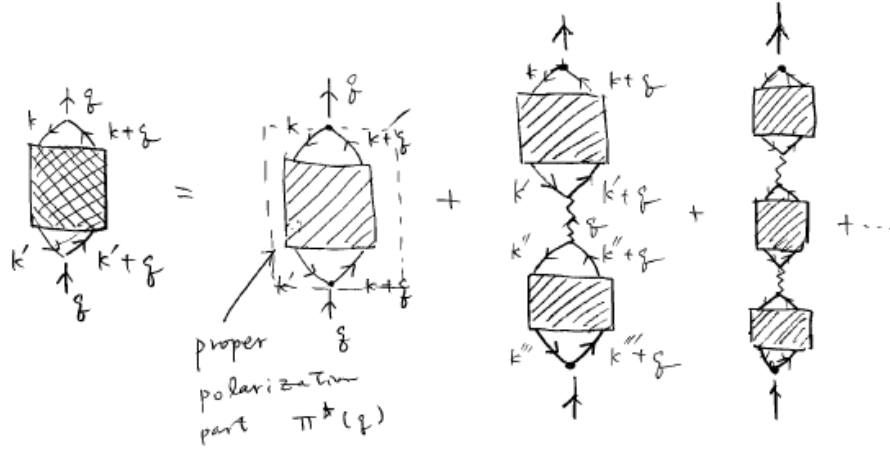
The proper part is a polarization part that **cannot** be separated into two part by cutting a single interaction line. While the improper part is a polarization part that **can** be divided into the upper part and lower part by cutting a single interaction line.

For example, this is a proper polarization part, while this is improper, because we can cut this into two polarization part by removing this interaction line.



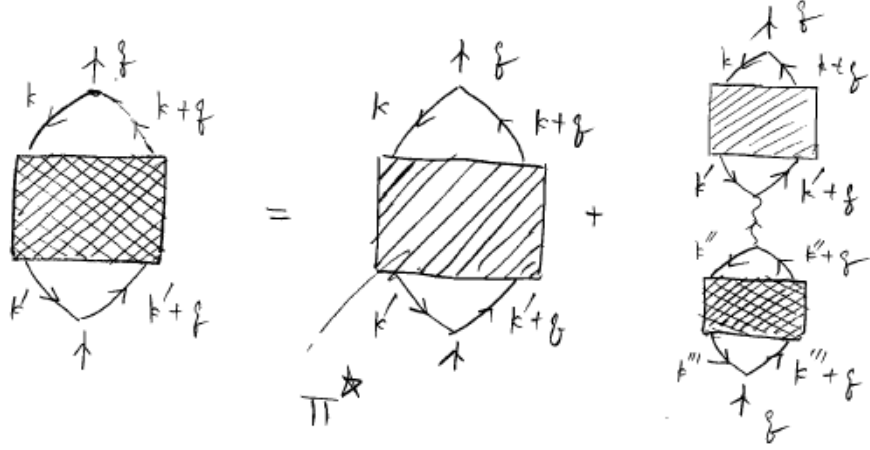
Then, in the same way as we did for the Dyson equation for self-energy, we can also derive the Dyson equation for the polarization part:

- Firstly, the full polarization part consist of a sum of all possible repetitions of the proper polarization part.



where this part denotes the proper polarization part, while this wavy line is for interaction potential.

- This description becomes possible, especially because any n -th order Feynman diagram is accompanied by $\left(\frac{i}{\hbar}\right)^n \cdot (-1)^{F'}$, where F' stands for the number of fermion loop.
- As is clear from this repetitive structure, this can be rewritten into a self-consistent equation:



where this singly shaded one is a collection of all possible proper polarization part.

while this doubly shaded one denote a collection of all possible polarization part.

This equation is called as Dyson's equation for polarization part and its analytic expression is given like this:

$$(-i\hbar)\Pi(q) = (-i\hbar)\Pi^\star(q) + (-i\hbar)\Pi^\star(q) \cdot \frac{i}{\hbar} V(q) (-i\hbar)\Pi(q)$$

Here $V(q)$ stands for the Fourier transform of the interaction potential, which corresponds to this wavy line.

As I mentioned already, we assumed spin-independent interaction potential.

$$V_{\lambda,\lambda';\mu,\mu'}(\mathbf{x}, \mathbf{x}') = V(\mathbf{x} - \mathbf{x}') \delta_{\lambda\lambda'} \delta_{\mu\mu'}$$

with $V(\mathbf{q}) = \int d^3\mathbf{x} e^{-i\mathbf{q}\cdot\mathbf{x}} V(\mathbf{x})$.

so that the product between $\Pi_0(q)$ and $V(\mathbf{q})$ and $V(\mathbf{q})$ and $\Pi(q)$ reduces to a single scalar product.

This interaction potential is accompanied by $(\frac{i}{\hbar})$, which is this. Remember also that both Π_0 and Π are always accompanied by factor $(-i\hbar)$ here. This factor $(-i\hbar)$ comes from our definition of polarization part:¹

$$-i\hbar\Pi(x, x') \equiv -iD(x, x') \equiv (-i)^2 \frac{\langle \Psi_0 | \Gamma[\tilde{n}_H(x) \tilde{n}_H(x')] | \Psi_0 \rangle}{\langle \Psi_0 | \Psi_0 \rangle}$$

Namely, I put this factor in front of the polarization part, in order to make this corresponding r.h.s. to be consistent with definition of two-particle Green's function, for which we can apply Feynman's rule.

With these factors in mind, one can see that this equation can be simplified into this form:

$$\Pi(q) = \Pi^\star(q) + \Pi^\star(q) V(q) \Pi(q) \quad (2.8.5)$$

which can be solved exactly as

$$\Pi(q) = \frac{\Pi^\star(q)}{1 - \Pi^\star(q) V(q)} \quad (2.8.5')$$

Then, the density-density correlation function is given as follows:

$$D^\lambda(x, x') = \hbar\Pi^\lambda(x, x') = \frac{\hbar\Pi^{\star,\lambda}(q)}{1 - \lambda\Pi^{\star,\lambda}(q)V(q)} = \frac{\hbar\Pi^{\star,\lambda}(q)}{K^\lambda(q)} \quad (2.8.6)$$

¹ see eq. (2.8.4)

Here, I put λ in order to make explicit the λ -dependence of the correlation function. substituting this back into eq.(2.8.3), we finally obtain

$$\begin{aligned} E_{corr} &= \frac{i}{2} V \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4 q}{(2\pi)^4} \lambda V(q) [D^\lambda(q) - D^0(q)] \\ &= \frac{i}{2} \hbar V \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4 q}{(2\pi)^4} \left[\frac{\lambda V(q) \Pi^{\star, \lambda}(q)}{1 - \lambda V(q) \Pi^{\star, \lambda}(q)} - \lambda V(q) \Pi^0(q) \right] \end{aligned} \quad (2.8.7)$$

$$\begin{aligned} &= \frac{i}{2} \hbar V \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4 q}{(2\pi)^4} \left[-1 + \frac{1}{1 - \lambda V(q) \Pi^{\star, \lambda}(q)} - \lambda V(q) \Pi^0(q) \right] \\ &= \frac{i}{2} \hbar V \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4 q}{(2\pi)^4} \left[[K^\lambda(q)]^{-1} - 1 - \lambda V(q) \Pi^0(q) \right] \end{aligned} \quad (2.8.8)$$

where V here denote the total volume of the system.

Eq.(2.8.8) can be applied to any homogeneous system with spin-independent interaction.

In the following, however, we specialize to a degenerate electron gas with long-range Coulomb interaction, whose Hamiltonian is already defined in eq.(??).

As I have already explained there, the uniform positive background precisely cancels the $\mathbf{q} = 0$ term in the interaction potential $V(\mathbf{q})$, so that $V(\mathbf{q} = 0)$ vanishes.

To understand the structure of eq.(2.8.8), let us expand [...] in λ (or interaction potential): ¹

$$[\dots] = \frac{\lambda V(q) \Pi^{\star, \lambda}(q)}{1 - \lambda V(q) \Pi^{\star, \lambda}(q)} - \lambda V(q) \Pi^0(q)$$

Generally, the proper polarization part depends on λ in a non-trivial manner.

But, we can expand $\Pi^{\star, \lambda}$ in the powers of λ :

$$\Pi^{\star, \lambda}(q) = \Pi_0^\star(q) + \lambda \Pi_1^\star(q) + \lambda^2 \Pi_2^\star(q) + \dots$$

where


1

↳ because they cannot be transformed to each other, by any kind of deformation of wavy lines



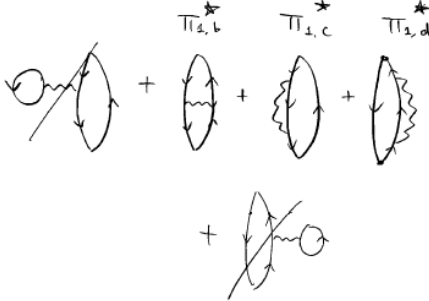
these two should be regarded as topologically distinct. ↖

•



$$\Pi_0^\star(q) = \quad (\equiv \Pi_0(q))$$

•



$$\Pi_1^\star(q) =$$

The 1st term and last term vanish because $V(\mathbf{q} = 0) = 0$.

•

$$\Pi_2^\star(q) = \dots$$

With this expansion in mind, we can expand this in powers of λ as:

$$[\dots] = \lambda V(q) \Pi_0^\star(q) + \lambda^2 V(q) \Pi_1^\star(q) + \lambda^2 V(q) \Pi_0^\star(q) V(q) \Pi_0^\star(q) + \dots - \lambda V(q) \Pi^0(q)$$

Now $\Pi_0^\star = \Pi^0$, so that this starts from 2nd order in λ .

$$[\dots] = \lambda^2 V(q) \Pi_1^\star(q) + \lambda^2 (V(q) \Pi^0(q))^2$$

Correspondingly, we have the expansion of ground state energy as

$$E_{corr} = E_2^r + E_2^b + E_2^c + E_2^d + \dots$$

where

$$E_2^r = \frac{1}{2} i V \hbar \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4 q}{(2\pi)^4} [\lambda V_0(q) \Pi^0(q)]^2$$

$$E_2^{b,c,d} = \frac{1}{2} i V \hbar \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4 q}{(2\pi)^4} [\lambda V_0(q) \Pi_{1,b,c,d}^\star(q)]$$

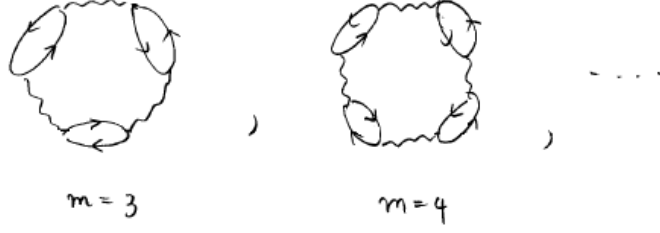
where we named these 3 as $\Pi_{1,b}$, $\Pi_{1,c}$, $\Pi_{1,d}$ respectively.

One can explicitly show that the correlation energy due to these proper polarization parts are finite.

On the one hand, the correlation energy due to this improper polarization part contain the logarithmic divergence of the q -integral, which we will see later.

This divergence comes from the singular behaviour of the Coulomb potential $V(q) \equiv \frac{4\pi e^2}{\mathbf{q}^2}$ for small \mathbf{q} region. Namely, E_2^r contains two $V(q)$, which gives its integrand with \mathbf{q}^{-4} behaviour.

A similar divergence also occurs in all orders in λ , because there is always a single m -th order term with the integrand $[V(q) \Pi^0(q)]^m$.



An idea of “ring diagram approximation” is to sum up all these singular terms up to all order in the perturbation theory by introducing an effective interaction potential as

$$U_r(q) = V(q) + V(q)\Pi^0(q)V(q) + [V(q)\Pi^0(q)]^2V(q) + [V(q)\Pi^0(q)]^3V(q) + \dots$$

$$= \frac{V(q)}{1 - V(q)\Pi^0(q)}$$

Namely, with this effective interaction potential, these singular terms can be summed up in all order in λ as

$$E_r = \frac{1}{2}iV\hbar \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4q}{(2\pi)^4} \sum_{n=2}^{\infty} [\lambda V_0(q)\Pi^0(q)]^n$$

$$= \frac{1}{2}iV\hbar \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4q}{(2\pi)^4} \frac{[\lambda V_0(q)\Pi^0(q)]^2}{1 - \lambda V_0(q)\Pi^0(q)} \quad (2.8.9)$$

$$= \frac{1}{2}iV\hbar \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4q}{(2\pi)^4} \lambda V_0(q)\Pi^0(q) \underline{U_r^\lambda(q)} \Pi^0(q)$$

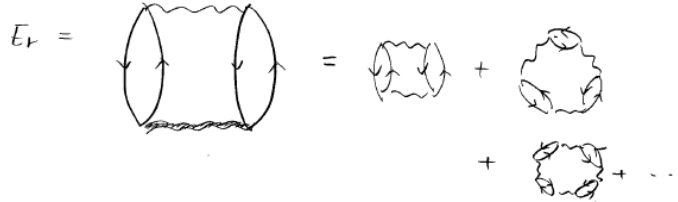
where the underlined term is the effective interaction potential.

To summarize, the correlation energy of a degenerate electron gas can be given as

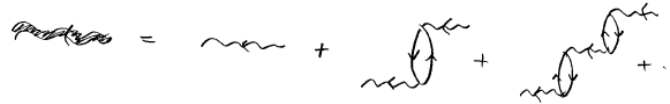
$$E_{corr} = E_r + E_2^b + E_2^c + E_2^d + \dots$$

where E_r is the collection of leading contributions, while $E_2^{b,c,d}$ are at most finite and less leading contribution.

E_r can be schematically given by the following diagrams:



where the bold wavy line represents the effective interaction potential $U_r(q)$;



Taking into account only these singular diagrams is called a “ring-diagram approximation”.

Ring-diagram approximation is equivalent to employing Dyson equation with its proper polarization part being replaced by the zero-th order contribution.

Namely, when we replace $\Pi^{\star,\lambda}(q)$ in eq.(2.8.7) by its zero-th order contribution $\Pi^0(q)$, we safely have E_r :

$$\begin{aligned} E_{corr} &= \frac{i}{2} \hbar V \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4 q}{(2\pi)^4} \left[\frac{\lambda V(q) \Pi^{\star,\lambda}(q)}{1 - \lambda V(q) \Pi^{\star,\lambda}(q)} - \lambda V(q) \Pi^0(q) \right] \\ &\approx \frac{i}{2} \hbar V \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4 q}{(2\pi)^4} \left[\frac{\lambda V(q) \Pi^0(q)}{1 - \lambda V(q) \Pi^0(q)} - \lambda V(q) \Pi^0(q) \right] \\ &= \frac{i}{2} \hbar V \int_0^1 d\lambda \cdot \lambda^{-1} \int \frac{d^4 q}{(2\pi)^4} \frac{[\lambda V(q) \Pi^0(q)]^2}{1 - \lambda V(q) \Pi^0(q)} \\ &= E_r \end{aligned}$$

Correspondingly, the effective interaction potential $U_r(q)$ can be rewritten in terms of a dielectric constant.

$$U_r(q) = \frac{V(q)}{K_r(q)} \quad (2.8.10)$$

$$K_r(q) = 1 - V(q) \Pi^0(q) \quad (2.8.11)$$

Where $K_r(q)$ can be regarded as the ring-diagram approximation to the exact dielectric constant. We next evaluate this lowest-order proper polarization part. $\Pi_0(q)$

$$\Pi_0(q) = -\frac{2i}{\hbar} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \int \frac{d\epsilon}{2\pi} G^0(\mathbf{k} + \mathbf{q}, \epsilon + \omega) G^0(\mathbf{k}, \epsilon) \quad (2.8.12)$$

where this factor 2 comes from the summation over the spin index, using this

$$G^0(\mathbf{k}, \epsilon) = \frac{\theta(|\mathbf{k}| - k_F)}{\omega - \omega_{\mathbf{k}} + i\eta} + \frac{\theta(k_F - |\mathbf{k}|)}{\omega - \omega_{\mathbf{k}} - i\eta} \quad (2.8.13)$$

we have

$$\begin{aligned} \Pi_0(q) &= -\frac{2i}{\hbar} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \int_{-\infty}^{+\infty} \frac{d\epsilon}{2\pi} \times \left\{ \frac{\theta(|\mathbf{k} + \mathbf{q}| - k_F)}{\epsilon + \omega - \omega_{\mathbf{k}+\mathbf{q}} + i\eta} + \frac{\theta(k_F - |\mathbf{k} + \mathbf{q}|)}{\epsilon + \omega - \omega_{\mathbf{k}+\mathbf{q}} - i\eta} \right\} \\ &\quad \times \left\{ \frac{\theta(|\mathbf{k}| - k_F)}{\epsilon - \omega_{\mathbf{k}} + i\eta} + \frac{\theta(k_F - |\mathbf{k}|)}{\epsilon - \omega_{\mathbf{k}} - i\eta} \right\} \end{aligned}$$

Out of this integrand, we have four terms, 2 of which vanish under the integration over ϵ .

Namely, product between these two and that between these two have their poles only in the lower half plane and only in the upper half plane respectively.

Such terms reduce to zero, when we close the contour in the opposite side of the half plane.

Accordingly, we end up a product between these two and that between these two.

$$\begin{aligned} \Pi_0(q) &= -\frac{2i}{\hbar} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \int_{-\infty}^{+\infty} \frac{d\epsilon}{2\pi} \times \left\{ \frac{\theta(|\mathbf{k} + \mathbf{q}| - k_F)}{\epsilon + \omega - \omega_{\mathbf{k}+\mathbf{q}} + i\eta} \cdot \frac{\theta(k_F - |\mathbf{k}|)}{\epsilon - \omega_{\mathbf{k}} - i\eta} \right. \\ &\quad \left. + \frac{\theta(k_F - |\mathbf{k} + \mathbf{q}|)}{\epsilon + \omega - \omega_{\mathbf{k}+\mathbf{q}} - i\eta} \cdot \frac{\theta(|\mathbf{k}| - k_F)}{\epsilon - \omega_{\mathbf{k}} + i\eta} \right\} \\ &= \frac{2}{\hbar} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \left\{ \frac{\theta(|\mathbf{k} + \mathbf{q}| - k_F) \theta(k_F - |\mathbf{k}|)}{\omega + \omega_{\mathbf{k}} - \omega_{\mathbf{k}+\mathbf{q}} + 2i\eta} - \frac{\theta(k_F - |\mathbf{k} + \mathbf{q}|) \theta(|\mathbf{k}| - k_F)}{\omega + \omega_{\mathbf{k}} - \omega_{\mathbf{k}+\mathbf{q}} - 2i\eta} \right\} \end{aligned}$$

As for the 2nd term, we can define $\mathbf{k}' = -\mathbf{k} - \mathbf{q}$, so that, \mathbf{k} and $\mathbf{k} + \mathbf{q}$ are replaced by $-\mathbf{k}' - \mathbf{q}$ and $-\mathbf{k}'$ respectively.

Now $\omega_{\mathbf{k}}$ are an even function in momentum. these two can be replaced by $\mathbf{k}' + \mathbf{q}$ and \mathbf{k}' respectively. \mathbf{k}' as well as \mathbf{k} are dummy index, we can replace \mathbf{k}' by \mathbf{k} .

$$= \frac{2}{\hbar} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \theta(|\mathbf{k} + \mathbf{q}| - k_F) \theta(k_F - |\mathbf{k}|) \times \left\{ \frac{1}{\omega + \omega_{\mathbf{k}} - \omega_{\mathbf{k}+\mathbf{q}} + i\eta} - \frac{1}{\omega - \omega_{\mathbf{k}} + \omega_{\mathbf{k}+\mathbf{q}} - i\eta} \right\}$$

This clearly suggests that $\Pi^0(\mathbf{q}, \omega)$ is an even function of ω .

$$\Pi^0(\mathbf{q}, \omega) = \Pi^0(\mathbf{q}, -\omega) \quad (2.8.14)$$

so that we have only to study $\Pi^0(\mathbf{q}, \omega)$ for positive ω

Let us begin with the real part of $\Pi^0(\mathbf{q}, \omega)$:¹

$$\begin{aligned} Re\Pi^0(\mathbf{q}, \omega) &= \frac{2}{\hbar} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \theta(|\mathbf{k} + \mathbf{q}| - k_F) \theta(k_F - |\mathbf{k}|) \\ &\times \left\{ \frac{\mathcal{P}}{\omega + \omega_{\mathbf{k}} - \omega_{\mathbf{k}+\mathbf{q}}} - \frac{\mathcal{P}}{\omega - \omega_{\mathbf{k}} + \omega_{\mathbf{k}+\mathbf{q}}} \right\} \end{aligned} \quad (2.8.15)$$

$$\begin{aligned} Im\Pi^0(\mathbf{q}, \omega) &= -\frac{1}{\hbar} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \theta(|\mathbf{k} + \mathbf{q}| - k_F) \theta(k_F - |\mathbf{k}|) \\ &\times \{ \delta(\omega + \omega_{\mathbf{k}} - \omega_{\mathbf{k}+\mathbf{q}}) + \delta(\omega - \omega_{\mathbf{k}} + \omega_{\mathbf{k}+\mathbf{q}}) \} \end{aligned} \quad (2.8.16)$$

By replacing $\theta(|\mathbf{k} + \mathbf{q}| - k_F)$ by $1 - \theta(k_F - |\mathbf{k} + \mathbf{q}|)$, we have

$$Re\Pi^0(\mathbf{q}, \omega) = \frac{2}{\hbar} \mathcal{P} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} (1 - \theta(k_F - |\mathbf{k} + \mathbf{q}|)) \theta(k_F - |\mathbf{k}|) \times \left\{ \frac{1}{\omega + \omega_{\mathbf{k}} - \omega_{\mathbf{k}+\mathbf{q}}} - \frac{1}{\omega - \omega_{\mathbf{k}} + \omega_{\mathbf{k}+\mathbf{q}}} \right\}$$

This part vanishes under the \mathbf{k} -integral.

One can see this by introducing again $\mathbf{k}' = -\mathbf{k} - \mathbf{q}$.

$$Re\Pi^0(\mathbf{q}, \omega) = \frac{2}{\hbar} \mathcal{P} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \theta(k_F - |\mathbf{k}|) \times \left\{ \frac{1}{\omega + \omega_{\mathbf{k}} - \omega_{\mathbf{k}+\mathbf{q}}} - \frac{1}{\omega - \omega_{\mathbf{k}} + \omega_{\mathbf{k}+\mathbf{q}}} \right\}$$

$\omega_{\mathbf{k}}$ denotes the kinetic energy of electron gas;

$$\omega_{\mathbf{k}} = \frac{\epsilon_{\mathbf{k}}}{\hbar} = \frac{\hbar \mathbf{k}^2}{2m}$$

so that we have

$$\begin{aligned} \omega_{\mathbf{k}+\mathbf{q}} - \omega_{\mathbf{k}} &= \frac{\hbar}{2m} ((\mathbf{k} + \mathbf{q})^2 - \mathbf{k}^2) \\ &= \frac{\hbar}{2m} (2\mathbf{k} \cdot \mathbf{q} + \mathbf{q}^2) \\ &= \frac{\hbar}{m} \left(\mathbf{k} \cdot \mathbf{q} + \frac{\mathbf{q}^2}{2} \right) \end{aligned} \quad (2.8.17)$$

using this, the real part is given by

$$Re\Pi^0(\mathbf{q}, \omega) = \frac{2}{\hbar} \mathcal{P} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \theta(k_F - |\mathbf{k}|) \times \left\{ \frac{1}{\omega - \frac{\hbar}{m} (\mathbf{q} \cdot \mathbf{k} + \frac{\mathbf{q}^2}{2})} - \frac{1}{\omega + \frac{\hbar}{m} (\mathbf{q} \cdot \mathbf{k} + \frac{\mathbf{q}^2}{2})} \right\}$$

¹This “ \mathcal{P} ” stands for the principal integral(value).

By making ω to be dimensionless variable and measuring all wave vectors in the unit of k_F ;

$$\nu = \frac{m}{\hbar k_F^2} \omega \quad \begin{pmatrix} \mathbf{q}_{old} = k_F \mathbf{q}_{new} \\ \mathbf{k}_{old} = k_F \mathbf{k}_{new} \end{pmatrix} \quad (2.8.18)$$

we have

$$\begin{aligned} Re\Pi_{old}^0(k_F \mathbf{q}, \frac{\hbar k_F^2}{m} \nu) &\equiv Re\Pi_{new}^0(\mathbf{q}, \nu) \\ &= \frac{2mk_F}{\hbar^2} \mathcal{P} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \theta(1 - |\mathbf{k}|) \times \left\{ \frac{1}{\nu - (\mathbf{q} \cdot \mathbf{k} + \frac{q^2}{2})} - \frac{1}{\nu - (-\mathbf{q} \cdot \mathbf{k} - \frac{q^2}{2})} \right\} \\ Re\Pi^0(\mathbf{q}, \nu) &= \frac{2mk_F}{\hbar^2} \int_0^1 k^2 \cdot dk \int_{-1}^1 dt \times \left\{ \frac{\mathcal{P}}{\nu - (qkt + \frac{q^2}{2})} - \frac{\mathcal{P}}{\nu - (-qkt - \frac{q^2}{2})} \right\} \end{aligned} \quad (2.8.19)$$

Since these integrals are Cauchy principal integral, we need to pay attention to those possibilities of these denominators being zero.

To clarify these possibilities, notice that this quantity(the underlined part) is bounded like this:

$$-q + \frac{q^2}{2} \leq qkt + \frac{q^2}{2} \leq q + \frac{q^2}{2}$$

while this quantity is bounded by this form

$$-q - \frac{q^2}{2} \leq -qkt - \frac{q^2}{2} \leq q - \frac{q^2}{2}$$

As such, depending on ν and q , we have following five cases:

- (i) $\nu > \frac{q^2}{2} + q > \frac{q^2}{2} - q, -\frac{q^2}{2} + q > -\frac{q^2}{2} - q$
- (ii) $\frac{q^2}{2} + q > \nu > \frac{q^2}{2} - q, -\frac{q^2}{2} + q > -\frac{q^2}{2} - q$
- (iii) $\frac{q^2}{2} + q > \max(\frac{q^2}{2} - q, -\frac{q^2}{2} + q) > \nu > \min(\frac{q^2}{2} - q, -\frac{q^2}{2} + q) > -\frac{q^2}{2} - q$
- (iv) $\frac{q^2}{2} + q > \frac{q^2}{2} - q, -\frac{q^2}{2} + q > \nu > -\frac{q^2}{2} - q$
- (v) $\frac{q^2}{2} + q > \frac{q^2}{2} - q, -\frac{q^2}{2} + q > -\frac{q^2}{2} - q > \nu$

In the case of (i), these are no possibilities of these denominators being zero, so that we can freely integrate this:

$$Re\Pi^0(\mathbf{q}, \nu) = -\frac{2mk_F}{\hbar^2} \frac{1}{q} \int_0^1 k \cdot dk \times \left\{ \left[\log(\nu - qkt - \frac{q^2}{2}) \right]_{t=-1}^{t=+1} - \left[\log(\nu + qkt + \frac{q^2}{2}) \right]_{t=-1}^{t=+1} \right\}$$

we obtain ¹

¹Using

$$\int x \log(\beta - \alpha x) dx = \frac{x^2}{2} \log(\beta - \alpha x) - \frac{\beta^2}{2\alpha^2} \log(\beta - \alpha x) - \frac{x^2}{4} - \frac{\beta}{2\alpha} x$$

$$Re\Pi^0(\mathbf{q}, \nu) = \frac{2mk_F}{\hbar^2} \frac{1}{4\pi^2} \times \left\{ -1 + \frac{1}{2q} \left(1 - \left(\frac{\nu}{q} - \frac{q}{2} \right)^2 \right) \log \left[\frac{\nu - \frac{q^2}{2} + q}{\nu - \frac{q^2}{2} - q} \right] \right. \\ \left. - \frac{1}{2q} \left(1 - \left(\frac{\nu}{q} + \frac{q}{2} \right)^2 \right) \log \left[\frac{\nu + \frac{q^2}{2} + q}{\nu + \frac{q^2}{2} - q} \right] \right\}$$

As for other cases, we can do a similar integral, from which we obtain.

$$Re\Pi^0(\mathbf{q}, \nu) = \frac{2mk_F}{\hbar^2} \frac{1}{4\pi^2} \times \left\{ -1 + \frac{1}{2q} \left(1 - \left(\frac{\nu}{q} - \frac{q}{2} \right)^2 \right) \log \left[\frac{|\nu - \frac{q^2}{2} + q|}{|\nu - \frac{q^2}{2} - q|} \right] \right. \\ \left. - \frac{1}{2q} \left(1 - \left(\frac{\nu}{q} + \frac{q}{2} \right)^2 \right) \log \left[\frac{|\nu + \frac{q^2}{2} + q|}{|\nu + \frac{q^2}{2} - q|} \right] \right\} \quad (2.8.20)$$

for any cases.

Let us next calculate the imaginary part of $\Pi^0(\mathbf{q}, \omega)$.

$$Im\Pi^0(\mathbf{q}, \omega) = -\frac{1}{\hbar} \int \frac{d^3\mathbf{k}}{(2\pi)^3} \theta(|\mathbf{k} + \mathbf{q}| - k_F) \theta(k_F - |\mathbf{k}|) \\ \times \{ \delta(\omega + \omega_{\mathbf{k}} - \omega_{\mathbf{k}+\mathbf{q}}) + \delta(\omega - \omega_{\mathbf{k}} + \omega_{\mathbf{k}+\mathbf{q}}) \} \quad (2.8.16)$$

Since $\Pi^0(\mathbf{q}, \omega)$ is an even function of ω , we consider positive ω case.

For $\omega > 0$, we have

$$Im\Pi^0(\mathbf{q}, \omega) = -\frac{1}{\hbar} \int \frac{d^3\mathbf{k}}{(2\pi)^3} \theta(|\mathbf{k} + \mathbf{q}| - k_F) \theta(k_F - |\mathbf{k}|) \delta(\omega - (\omega_{\mathbf{k}+\mathbf{q}} - \omega_{\mathbf{k}}))$$

because, under this condition, $\omega_{\mathbf{k}+\mathbf{q}}$ is always greater than $\omega_{\mathbf{k}}$.¹

$$\delta(\omega - (\omega_{\mathbf{k}+\mathbf{q}} - \omega_{\mathbf{k}})) \\ = \delta\left(\frac{\hbar k_F^2}{m} \nu - \frac{\hbar k_F^2}{m} \left(\frac{\mathbf{k}}{k_F} \cdot \frac{\mathbf{q}}{k_F} + \frac{1}{2} \left(\frac{q}{k_F} \right)^2 \right)\right) \\ = \frac{m}{\hbar k_F^2} \delta\left(\nu - \left(\mathbf{k} \cdot \mathbf{q} + \frac{q^2}{2} \right)\right)$$

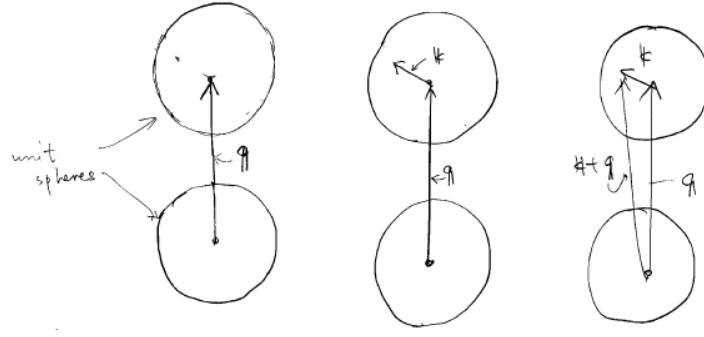
substituting this expression into this, we have

$$Im\Pi^0(k_F \mathbf{q}, \frac{\hbar k_F^2}{m} \nu) = Im\Pi_{new}^0(\mathbf{q}, \nu) \\ = -mk_F \frac{1}{(2\pi\hbar)^2} \int d^3\mathbf{k} \theta(|\mathbf{q} + \mathbf{k}| - 1) \theta(1 - |\mathbf{k}|) \times \delta\left(\nu - \left(\mathbf{k} \cdot \mathbf{q} + \frac{q^2}{2} \right)\right)$$

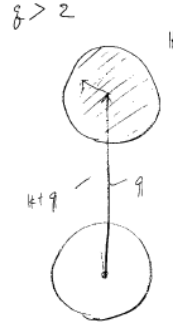
To evaluate this integral, let us introduce two unit spheres whose centers are separated by a vector \mathbf{q} .

¹

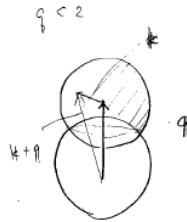
$$k_F \mathbf{k}_{new} = \mathbf{k}_{old} \\ k_F \mathbf{q}_{new} = \mathbf{q}_{old}$$



- We consider the vector \mathbf{k} from the center of this unit sphere.
- Then a vector, which connects this point and the other center is nothing but $\mathbf{k} + \mathbf{q}$.
- These two factors require that the allowed integral region for \mathbf{k} is within this unit sphere, while \mathbf{k} should be outside the other unit sphere.
- In this picture, these two conditions are always compatible, so that \mathbf{k} can be freely integrated within this sphere.



- This situation holds true, whenever the amplitude of \mathbf{q} is greater than 2.



- Otherwise, these two unit spheres have a finite overlap, where the integral region for \mathbf{k} is restricted within this shadow region.

Let us first consider $q > 2$ case.

In such a case, this part is bounded by this form:¹

$$-q + \frac{q^2}{2} \leq \mathbf{k} \cdot \mathbf{q} + \frac{q^2}{2} \leq q + \frac{q^2}{2}$$

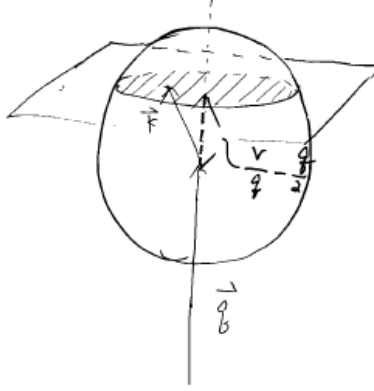
¹equality holds true when $|\mathbf{k}| = 1, \mathbf{k} \cdot \mathbf{q} < 0, \mathbf{k} \parallel \mathbf{q}$ or $|\mathbf{k}| = 1, \mathbf{k} \cdot \mathbf{q} > 0, \mathbf{k} \parallel \mathbf{q}$

As such, the right hand side remains finite, only when ν is within this regions.

$$\underline{-q + \frac{q^2}{2} < \nu < q + \frac{q^2}{2}}$$

Otherwise, $Im\Pi^0(\mathbf{q}, \nu)$ always reduces to zero.

If this (underlined) is the case, the integrand remains finite only when k is within the plane which satisfies $\frac{q^2}{2} + \mathbf{q} \cdot \mathbf{k} = \nu$,



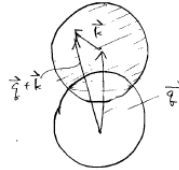
where the distance between this plane and the center of this unit sphere is given by

$$k_d = \left| \frac{\nu}{q} - \frac{q}{2} \right|$$

With this in mind, we have

$$\begin{aligned} Im\Pi^0(\mathbf{q}, \nu) &= -\frac{mk_F}{4\pi^2\hbar^2} 2\pi \int_{|\frac{\nu}{q} - \frac{q}{2}|}^1 k^2 \cdot dk \int_{-1}^{+1} dt \delta(\nu - qkt - \frac{q^2}{2}) \\ &= -\frac{mk_F}{4\pi^2\hbar^2} 2\pi \int_{|\frac{\nu}{q} - \frac{q}{2}|}^1 k^2 \cdot dk \int_{-1}^{+1} dt \frac{1}{qk} \delta(t - \frac{\nu}{qk} + \frac{q}{2k}) \\ &= -\frac{mk_F}{4\pi^2\hbar^2} \frac{2\pi}{q} \left[\frac{k^2}{2} \right]_{k=|\frac{\nu}{q} - \frac{q}{2}|}^{k=1} \\ &= -\frac{mk_F}{\hbar^2} \cdot \frac{1}{4\pi} \left[1 - \left(\frac{\nu}{q} - \frac{q}{2} \right)^2 \right] \end{aligned} \tag{2.8.21}$$

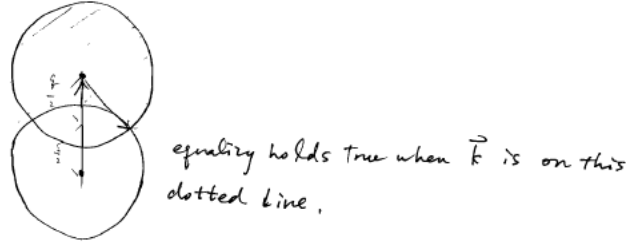
Let us next consider $q < 2$ case.



In this case, the integration over k is restricted within this shaded region.

In such a case, this part $(\frac{q^2}{2} + \mathbf{q} \cdot \mathbf{k})$ is bounded by this form

$$0 \leq \frac{q^2}{2} + \mathbf{q} \cdot \mathbf{k} \leq \frac{q^2}{2} + q$$



Therefore, the integral remains finite only when ν is bounded by this region

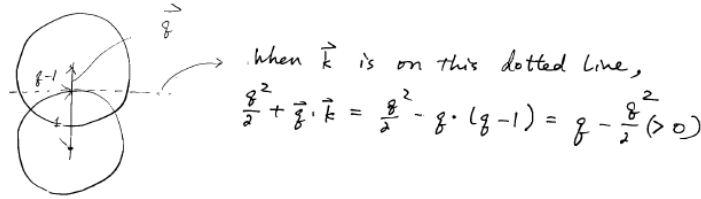
$$0 \leq \nu \leq \frac{q^2}{2} + q$$

To perform this integral explicitly, it is technically useful to divide this region into these two regions.

$$(i) \quad q - \frac{q^2}{2} \leq \nu (= \frac{q^2}{2} + \mathbf{q} \cdot \mathbf{k}) \leq \frac{q^2}{2} + q$$

$$(ii) \quad 0 \leq \nu (= \frac{q^2}{2} + \mathbf{q} \cdot \mathbf{k}) \leq q - \frac{q^2}{2}$$

When \mathbf{k} is along the plane which touches with this lower unit sphere, this term becomes equal to this quantity.

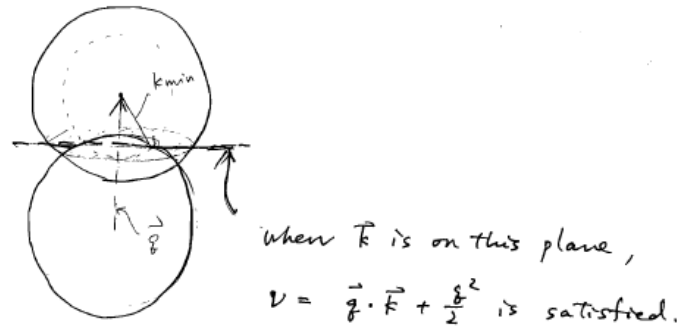


In the former case, the integral can be performed exactly in the same way as we did for $q > 2$ case.

$$\begin{aligned} Im\Pi^0(q, \nu) &= -\frac{mk_F}{4\pi^2\hbar^2} 2\pi \int_{|\frac{\nu}{q} - \frac{q}{2}|}^1 k^2 \cdot dk \int_{-1}^1 dt \delta(\nu - qkt - \frac{q^2}{2}) \\ &= -\frac{mk_F}{\hbar^2} \frac{1}{4\pi} \left[1 - \left(\frac{\nu}{q} - \frac{q}{2} \right)^2 \right] \end{aligned} \quad (2.8.22)$$

In the latter case, the integration becomes a little involved, because the plane also enters inside the lower unit sphere, which is the forbidden region.

$$Im\Pi^0(\mathbf{q}, \nu) = -\frac{mk_F}{(2\pi\hbar)^2} \int d^3k \cdot \theta(1 - |k|) \theta(|\mathbf{q} + \mathbf{k}| - 1) \delta(\nu - \mathbf{q} \cdot \mathbf{k} - \frac{q^2}{2})$$



When \mathbf{k} is within this shaded annulus, the integrand remains finite.

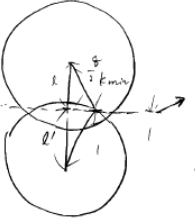
Thus we define k_{min} as a distance between the center of the upper sphere and the inner boundary of the annulus.

In terms of k_{min} thus defined, the integral is given by

$$Im\Pi^0(\mathbf{q}, \nu) = -\frac{mk_F}{\hbar^2} \frac{1}{2\pi} \int_{k_{min}}^1 k^2 \cdot dk \int_{-1}^1 dt \cdot \delta(\nu - qkt - \frac{q^2}{2})$$

By a simple geometry, k_{min} is given by ν as follows:

$$k_{min} = (1 - 2\nu)^{\frac{1}{2}}$$



$$\begin{aligned} \nu &= \left(\frac{q}{2} \cdot k\right) + \frac{q^2}{2} \\ k_{min}^2 &= 1 - (l')^2 + l^2 \\ l &= \frac{q}{2} - \frac{\nu}{q} \\ l' &= \frac{q}{2} - l = \frac{q}{2} + \frac{\nu}{q} \\ k_{min}^2 &= 1 - \left(\frac{q}{2} + \frac{\nu}{q}\right)^2 + \left(\frac{q}{2} - \frac{\nu}{q}\right)^2 \\ &= 1 - \left(\frac{q^2}{4} + \nu + \frac{\nu^2}{q^2}\right) + \left(\frac{q^2}{4} - \nu + \frac{\nu^2}{q^2}\right) = 1 - 2\nu. \end{aligned}$$

Thus

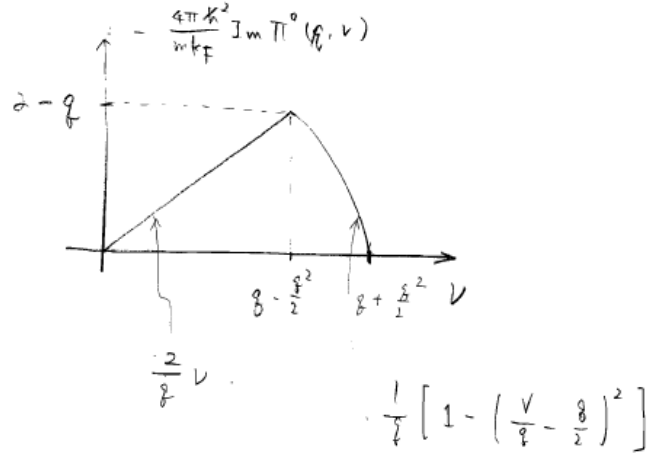
$$\begin{aligned} Im\Pi^0(\mathbf{q}, \nu) &= -\frac{mk_F}{\hbar^2} \frac{1}{2\pi q} \int_{(1-2\nu)^{\frac{1}{2}}}^1 k \cdot dk \int_{-1}^1 dt \cdot \delta\left(\frac{\nu}{qk} - \frac{q}{2k} - t\right) \\ &= -\frac{mk_F}{\hbar^2} \frac{1}{2\pi q} \left[\frac{k^2}{2} \right]_{k=(1-2\nu)^{\frac{1}{2}}}^{k=1} \\ &= -\frac{mk_F}{\hbar^2} \frac{1}{4\pi q} [1 - (1 - 2\nu)] \\ &= -\frac{mk_F}{\hbar^2} \frac{1}{4\pi q} 2\nu \end{aligned} \tag{2.8.23}$$

To summarize, so far, the imaginary part can be given as follows

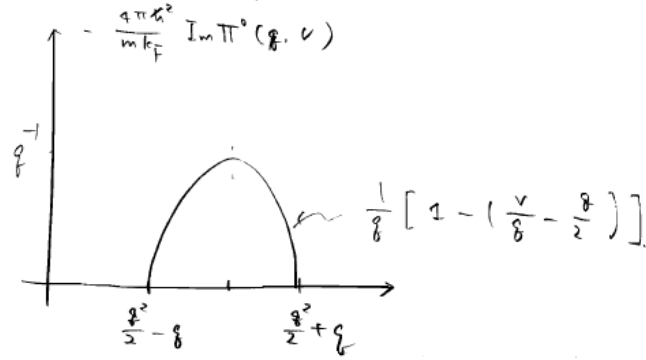
$$Im\Pi^0(q, \nu) = \begin{cases} -\frac{mk_F}{\hbar^2} \frac{1}{4\pi q} \left[1 - \left(\frac{\nu}{q} - \frac{q}{2} \right)^2 \right] & \begin{pmatrix} (i) q > 2 & \& \frac{q^2}{2} - q \leq \nu \leq \frac{q^2}{2} + q \\ or \\ (ii) q < 2 & \& -\frac{q^2}{2} + q \leq \nu \leq \frac{q^2}{2} + q \end{pmatrix} \\ -\frac{mk_F}{\hbar^2} \frac{1}{4\pi q} 2\nu & \left((iii) q < 2 \quad \& \quad 0 \leq \nu \leq -\frac{q^2}{2} + q \right) \\ 0 & (otherwise) \end{cases} \tag{2.8.24}$$

When viewed as a function of ν , the Imaginary part can be drawn like this:

- For $(0 <) q < 2$ case:



- For $2 < q$ case:



For many applications, it is useful to enumerate limiting forms of $\Pi^0(\mathbf{q}, \nu)$:

- One is called as a static limit, in which we take ν to be zero, while q is fixed:

Fixed q , while $\nu \rightarrow 0$.

In these two graphs, this limit corresponds to approaching the origin in this way or approaching the origin in this way.

Therefore, $\text{Im}\Pi^0(\mathbf{q}, \nu)$ reduces to zero in this limit.

$$\text{Im}\Pi^0(\mathbf{q}, \nu) = 0 \quad (2.8.25.a)$$

The real part of $\Pi^0(\mathbf{q}, \nu)$ can be also easily calculated from eq.(F)

$$\text{Re}\Pi^0(\mathbf{q}, \nu) = \frac{mk_F}{\hbar^2} \cdot \frac{1}{2\pi^2} \times \left\{ -1 + \frac{1}{q} \left(1 - \frac{q^2}{4} \right) \log \left(\frac{1 - \frac{q}{2}}{1 + \frac{q}{2}} \right) \right\} \quad (2.8.25.b)$$

- Another physically important limit is the dynamic limit, in which we take q to be zero, with fixed ν :

Fixed ν , while $q \rightarrow 0$

When q goes to zero, these upper bounds and lower bounds go to zero.

Since we fix ν to be finite, this limit ends up in outside of these regions, where the imaginary part reduces to zero.

$$Im\Pi^0(\mathbf{q}, \nu) = 0 \quad (2.8.26.a)$$

To calculate the real part in the dynamical limit, we need to Taylor-expand this logarithmic part in small q :

$$\log\left(\frac{\nu + q \mp \frac{q^2}{2}}{\nu - q \mp \frac{q^2}{2}}\right) = \frac{2q}{\nu} \pm \frac{q^3}{\nu^2} + \frac{2}{3} \frac{q^3}{\nu^3} + \frac{q^5}{2\nu^3} \pm \frac{q^5}{\nu^4} + \frac{2}{5} \frac{q^5}{\nu^5} + o(q^7)$$

In terms of these expansions, eq.(2.8.20) can be calculated as

$$\begin{aligned} Re\Pi^0(\mathbf{q}, \nu) &= \frac{2mk_F}{\hbar^2} \cdot \frac{1}{4\pi^2} \times \\ &\quad \left\{ -1 + \frac{1}{2q} \left(1 - \left(\frac{\nu}{q}\right)^2 + \nu - \frac{q^2}{4} \right) \times \left(\frac{2q}{\nu} + \frac{q^3}{\nu^2} + \frac{2}{3} \frac{q^3}{\nu^3} + \frac{q^5}{2\nu^3} + \frac{q^5}{\nu^4} + \frac{2}{5} \frac{q^5}{\nu^5} \right) \right. \\ &\quad \left. - \frac{1}{2q} \left(1 - \left(\frac{\nu}{q}\right)^2 - \nu - \frac{q^2}{4} \right) \times \left(\frac{2q}{\nu} - \frac{q^3}{\nu^2} + \frac{2}{3} \frac{q^3}{\nu^3} + \frac{q^5}{2\nu^3} - \frac{q^5}{\nu^4} + \frac{2}{5} \frac{q^5}{\nu^5} \right) \right\} \\ &= \frac{2mk_F}{\hbar^2} \cdot \frac{1}{4\pi^2} \times \\ &\quad \left\{ -1 + \frac{1}{2q} \left(1 - \left(\frac{\nu}{q}\right)^2 - \frac{q^2}{4} \right) \times \left(\frac{2q^3}{\nu^2} + \frac{2q^5}{\nu^4} \right) + \frac{\nu}{2q} \left(\frac{4q}{\nu} + \frac{4q^3}{3\nu^3} + o(q^5) \right) \right\} \\ &= \frac{2mk_F}{\hbar^2} \cdot \frac{1}{4\pi^2} \times \left\{ -1 - 1 + 2 + \frac{2}{3} \frac{q^2}{\nu^2} + o(q^4) \right\} \\ Re\Pi^0(\mathbf{q}, \nu) &= \frac{2mk_F}{\hbar^2} \cdot \frac{1}{4\pi^2} \cdot \frac{2}{3} \frac{q^2}{\nu^2} + o(q^4) \quad (2.8.26.b) \end{aligned}$$

- The third limit is an intermediate limit between the static limit & the dynamic limit.

More specifically, we take q & ν to be zero simultaneously, while the ratio between q and ν to be fixed:

Fixed $\frac{\nu}{q} \equiv x$, while $q \rightarrow 0$

This limit corresponds to this case, so that the imaginary part is given by

$$Im\Pi^0(q, \nu = qx) = \begin{cases} -\frac{mk_F}{\hbar^2} \cdot \frac{1}{4\pi q} 2\nu & (0 \leq \nu \leq q) \\ 0 & (q \leq \nu) \end{cases}$$

$$Im\Pi^0(q, \nu = qx) = -\frac{mk_F}{\hbar^2} \frac{x}{2\pi} \theta(1-x) \quad (2.8.27.a)$$

To calculate the real part, we Taylor-expand these logarithmic parts in small q , while $\frac{\nu}{q}$ being regarded as a constant:

$$\log\left|\frac{1 + \frac{\nu}{q} - \frac{q}{2}}{1 - \frac{\nu}{q} + \frac{q}{2}}\right| = \log\left|\frac{1 + \frac{\nu}{q}}{1 - \frac{\nu}{q}}\right| - \frac{\frac{q}{2} \cdot 2}{1 - \left(\frac{\nu}{q}\right)^2} + o(q^2)$$

$$\log \left| \frac{1 + \frac{\nu}{q} + \frac{q}{2}}{1 - \frac{\nu}{q} - \frac{q}{2}} \right| = \log \left| \frac{1 + \frac{\nu}{q}}{1 - \frac{\nu}{q}} \right| + \frac{\frac{q}{2} \cdot 2}{1 - (\frac{\nu}{q})^2} + o(q^2)$$

Using these expansions, we obtain the leading order expression for the real part as follows:

$$\begin{aligned} \text{Re}\Pi^0(\mathbf{q}, \nu = qx) &= \frac{2mk_F}{\hbar^2} \frac{1}{4\pi^2} \times \\ &\quad \left\{ -1 + \frac{1}{2q} \left(1 - \left(\frac{\nu}{q} \right)^2 + \nu - \frac{q^2}{4} \right) \times \left(\log \left| \frac{1 + \frac{\nu}{q}}{1 - \frac{\nu}{q}} \right| - \frac{q}{1 - (\frac{\nu}{q})^2} + o(q^2) \right) \right. \\ &\quad \left. - \frac{1}{2q} \left(1 - \left(\frac{\nu}{q} \right)^2 - \nu - \frac{q^2}{4} \right) \times \left(\log \left| \frac{1 + \frac{\nu}{q}}{1 - \frac{\nu}{q}} \right| + \frac{q}{1 - (\frac{\nu}{q})^2} + o(q^2) \right) \right\} \\ &= \frac{mk_F}{\hbar^2} \frac{1}{2\pi^2} \left\{ -1 + \frac{\nu}{2q} \log \left| \frac{1 + \frac{\nu}{q}}{1 - \frac{\nu}{q}} \right| \times 2 - 1 + o(q) \right\} \\ \text{Re}\Pi^0(\mathbf{q}, \nu = qx) &= \frac{mk_F}{\hbar^2} \frac{1}{2\pi^2} \left\{ -2 + x \cdot \log \left| \frac{1+x}{1-x} \right| + o(q) \right\} \end{aligned} \quad (2.8.27.b)$$

Using these expressions for the zero-th order proper polarization part, let us next calculate the dielectric constant given within the ring-diagram approximation.

Namely, using eq.(2.8.10), we have

$$\begin{aligned} K_r(k_F \mathbf{q}, \omega = \frac{\hbar k_F^2}{m} \nu) &= 1 - V(k_F \mathbf{q}) \Pi^0(\mathbf{q}, \nu) \\ &= 1 - \frac{4\pi e^2}{k_F^2} \frac{1}{q^2} (\text{Re}\Pi^0(\mathbf{q}, \nu) + i \text{Im}\Pi^0(\mathbf{q}, \nu)) \end{aligned} \quad (2.8.28)$$

where the real part and imaginary part are given in eq.(2.8.20) & eq.(2.8.24) respectively.

- In the static limit, the dielectric constant thus obtained takes the following asymptotic form:

- static limit ($\nu \rightarrow 0$, while fixing q)

$$K_r(q, 0) = 1 - \frac{4\pi e^2}{k_F^2} \frac{1}{q^2} \cdot \frac{mk_F}{\hbar^2} \frac{1}{2\pi^2} \times \left\{ -1 + \frac{1}{q} \left(1 - \frac{q^2}{4} \right) \log \left| \frac{1 - \frac{q}{2}}{1 + \frac{q}{2}} \right| \right\} \quad (2.8.29)$$

The underlined coefficient part is proportional to a ratio between Fermi wavelength and Bohr radius.

$$\begin{aligned} \frac{4\pi e^2}{k_F^2} \cdot \frac{mk_F}{\hbar^2} \frac{1}{2\pi^2} &= \frac{2}{\pi} \cdot \frac{1}{k_F} \cdot \frac{me^2}{\hbar^2} = \frac{2}{\pi} \frac{1}{k_F} \frac{1}{a_B} \\ &= \frac{1}{\pi^2} \frac{\lambda_F}{a_B} \end{aligned} \quad (2.8.30)$$

Fermi wavelength is proportional to the interparticle spacing in degenerate electron gas:

$$\lambda_F = \frac{2\pi}{k_F} = 2\pi \left(\frac{4}{9\pi} \right)^{\frac{1}{3}} r_0 \equiv 2\pi \alpha r_0 \quad (2.8.31)$$

Thus, this coefficient is a dimensionless quantity which measures the density of a degenerate electron gas.

$$= \frac{2}{\pi} \alpha \left(\frac{r_0}{a_B} \right) \equiv \frac{2}{\pi} \alpha r_S \quad (2.8.32)$$

Especially, the ratio between the interparticle spacing and Bohr radius is often called as r_S , which plays role of the most important coupling constant in a degenerate electron gas.

By its definition, r_S goes to zero in the high-density limit, while r_S goes infinity in the low-density limit.

$$r_S \rightarrow 0 \quad (\text{high-density limit})$$

In terms of r_S , the dielectric constant in the static limit is given by

$$K_r(q, 0) = 1 + \frac{2\alpha r_S}{\pi q^2} \left[1 - \frac{1}{q} \left(1 - \frac{q^2}{4} \right) \ln \left| \frac{1 - \frac{q}{2}}{1 + \frac{q}{2}} \right| \right] \quad (2.8.33)$$

- In the dynamic limit, the dielectric constant is given as follows within the ring-diagram approximation:

- dynamic limit ($q \rightarrow 0$, while fixing ν)

$$\begin{aligned} K_r(0, \nu) &= 1 - \frac{4\pi e^2}{k_F^2} \frac{1}{q^2} \frac{2mk_F}{\hbar^2} \cdot \frac{1}{4\pi^2} \frac{2}{3} \frac{q^2}{\nu^2} \\ &= 1 - \frac{4\alpha r_S}{3\pi \nu^2} \end{aligned} \quad (2.8.34)$$

- In the intermediate limit, the dielectric constant takes the following asymptotic form:

$$\begin{aligned} K_r(q, qx) &= 1 - \frac{4\pi e^2}{k_F^2} \frac{1}{q^2} \times \left\{ -\frac{mk_F}{\hbar^2} \cdot \frac{1}{2\pi^2} \left(2 - x \ln \left| \frac{1+x}{1-x} \right| \right) - \frac{mk_F}{\hbar^2} \cdot \frac{x}{2\pi} \theta(1-x) \right\} \\ &= 1 + \frac{4\alpha r_S}{\pi q^2} \left(1 - \frac{x}{2} \ln \left| \frac{1+x}{1-x} \right| \right) + \frac{2\alpha r_S}{q} x \theta(1-x) \end{aligned} \quad (2.8.35)$$

These three expressions for the dielectric constant turns out to be very useful to understand physical properties of high-density degenerate electron gas.

In the latter sections, we will use these expressions, in order to understand many important properties of the degenerate electron gas.

In the remain part of this section, we will go back to the evaluation of correlation energy within the ring-diagram approximation.

For its comparison, we will also evaluate the second-order term.

$$E_r = \frac{1}{2} i V \hbar \int \frac{d^4 q}{(2\pi)^4} \int_0^1 d\lambda \cdot \lambda^{-1} \frac{[\lambda V_0(q) \Pi^0(q)]^2}{1 - \lambda V_0(q) \Pi^0(q)} \quad (2.8.9)$$

$$E_2^r = \frac{1}{2} i V \hbar \int \frac{d^4 q}{(2\pi)^4} \int_0^1 d\lambda \cdot \lambda^{-1} [\lambda V_0(q) \Pi^0(q)]^2 \quad (2.8.9.a)$$

Equivalently

$$E_r = \text{bubble with 2 lines} + \text{bubble with 4 lines} + \text{bubble with 6 lines} + \dots$$

$$E_2 = \text{bubble with 2 lines}$$

Let us begin with the 2nd-order term first:

$$E_2^r = \frac{1}{4} i V \hbar \int \frac{d^3 \mathbf{q}}{(2\pi)^3} \int_{-\infty}^{+\infty} \frac{d\omega}{2\pi} V_0^2(\mathbf{q}) [\Pi^0(\mathbf{q}, \omega)]^2$$

As before, we introduce dimensionless frequency and momentum

$$\nu = \frac{m}{\hbar k_F^2} \omega, \quad \mathbf{q}_{new} = \frac{\mathbf{q}_{old}}{k_F} \quad (2.8.18)$$

In terms of these new variables, E_2^r is given as follows

$$\begin{aligned} E_2^r &= \frac{1}{4} i V \hbar \frac{k_F^3}{16\pi^4} \frac{\hbar k_F^2}{m} \int d^3 \mathbf{q} \int d\nu \cdot [V_0 \Pi^0]^2 \\ &\stackrel{k_F^3 = \frac{3\pi^2 N}{V}}{=} \frac{1}{4} i \frac{\hbar^2 k_F^2}{m e^2} \cdot \frac{3e^2 N}{16\pi^2} \int d^3 \mathbf{q} \int d\nu \cdot [V_0 \Pi^0]^2 \\ &\stackrel{\frac{\hbar^2}{m e^2} = a_0}{=} \frac{N e^2}{2a_0} \left(\frac{3i}{32\pi^2} (k_F a_0)^2 \int d^3 \mathbf{q} \int d\nu [V_0 \Pi^0]^2 \right) \\ E_2^r &= \frac{N e^2}{2a_0} \left(\frac{3i}{16\pi^2 (\alpha r_S)^2} \int d^3 \mathbf{q} \int_0^{+\infty} d\nu [V_0 \Pi^0]^2 \right) \end{aligned} \quad (2.8.9.b)$$

expressions for the polarization part obtained in the dynamic limit and the intermediate limit.

One can also confirm that the expression in the intermediate limit (eq.(2.8.27.a) & eq.(2.8.27.b)), includes that of the dynamic limit (eq.(2.8.26.a) & eq.(2.8.26.b)) as a limiting situation.

Therefore, we may use the expression obtained in the intermediate limit,

$$\begin{cases} \text{Im} \Pi^0(|\mathbf{q}|, \nu = |\mathbf{q}|x) = -\frac{m k_F}{\hbar^2} \frac{x}{2\pi} \theta(1-x) \\ \text{Re} \Pi^0(|\mathbf{q}|, \nu = |\mathbf{q}|x) = \frac{m k_F}{\hbar^2} \frac{1}{2\pi^2} \left\{ -2 + x \ln \left| \frac{1+x}{1-x} \right| \right\} \end{cases}$$

Now that the energy must be always real-valued, we only focus on the real part of the right hand side.

As mentioned before, E_2^r diverges logarithmically due to the singular behaviour of the Coulomb potential: $V_0(q) = \frac{4\pi e^2}{q^2}$ around $q = 0$.

To see this situation, let us separate the momentum integral into two parts:

$$\epsilon_2^r = \frac{3i}{16\pi^2 \alpha^2 r_S^2} \left(\int_{|\mathbf{q}| < q_c} d^3 \mathbf{q} + \int_{|\mathbf{q}| > q_c} d^3 \mathbf{q} \right) \int_0^{+\infty} d\nu \cdot V_0^2(\mathbf{q}) \Pi^{02}(\mathbf{q}, \nu)$$

and let us focus only on the first integral.

$$\epsilon_2^{r,1} = \frac{3i}{16\pi^2 \alpha^2 r_S^2} \int_{|\mathbf{q}| < q_c} d^3 \mathbf{q} \int_0^{+\infty} d\nu \cdot V_0^2(\mathbf{q}) \Pi^{02}(\mathbf{q}, \nu) \quad (2.8.9.c)$$

When q_c is chosen to be much smaller than 1, we may approximately employ the ^{1 2}

$$\begin{aligned}
\epsilon_2^{r,1} &= -\frac{3}{8\pi^2\alpha^2 r_S^2} \int_{|\mathbf{q}| < q_c} d^3\mathbf{q} \int_0^{+\infty} d\nu \cdot V_0^2(\mathbf{q}) \text{Re}\Pi^0(\mathbf{q}, \nu) \text{Im}\Pi^0(\mathbf{q}, \nu) \\
&= -\frac{3}{2\pi\alpha^2 r_S^2} \int_0^{q_c} q^2 \cdot d\mathbf{q} \int_0^{+\infty} d(qx) \cdot V_0^2(\mathbf{q}) \times \text{Re}\Pi^0(\mathbf{q}, qx) \text{Im}\Pi^0(\mathbf{q}, qx) \\
&= -\frac{6}{\pi} \int_0^{q_c} q^3 \cdot d\mathbf{q} \int_0^{+\infty} dx \cdot \frac{1}{q^4} x \frac{2}{\pi} \left\{ 1 - \frac{x}{2} \ln \left| \frac{1+x}{1-x} \right| \right\} \theta(1-x) \\
\epsilon_2^{r,1} &= -\frac{6}{\pi} \int_0^{q_c} \frac{1}{q} \cdot d\mathbf{q} \int_0^1 dx \cdot x f(x)
\end{aligned} \tag{2.8.9.d}$$

From this final expression, we can see that this momentum integral exhibits logarithmic divergence in the infrared regime.

Moreover, when it comes to the 3rd-order term $\left(\text{diagram} \right)$, the integrand contains $V_0^3(q)(\Pi^0(q))^3$, so that this momentum integral in the infrared regime will be replaced by the following form.

$$\int_0^{q_c} q^3 \cdot \frac{1}{q^6} \cdot d\mathbf{q} \int_0^1 dx (axf^2(x) + bx^3) \tag{2.8.9.e}$$

which exhibits more singular divergence at $q = 0$.

An idea of “ring-diagram approximation” is to sum up all these singular terms up to all order in the perturbation theory:

$$\begin{aligned}
E_r &= \text{diagram 1} + \text{diagram 2} + \text{diagram 3} + \dots \\
E_r &= \frac{1}{2} iV\hbar \int \frac{d^4q}{(2\pi)^4} \int_0^1 d\lambda \cdot \lambda \frac{[V_0(q)\Pi^0(q)]^2}{1 - \lambda V_0(q)\Pi^0(q)}
\end{aligned} \tag{2.8.9}$$

Remarkably, when summed over up to all order, these singularities cancel each other, which results in a finite contribution.

To see this situation, let us next calculate E_r :

$$\begin{aligned}
E_r &= \frac{1}{2} iV\hbar \int \frac{d^4q}{(2\pi)^4} \int_0^1 d\lambda \cdot \lambda \frac{V_0^2 \Pi^0{}^2}{1 - \lambda V_0 \Pi^0} \\
&= \frac{1}{2} iV\hbar \int \frac{d^4q}{(2\pi)^4} \int_0^1 d\lambda V_0 \Pi^0 \left(-1 + \frac{1}{1 - \lambda V_0 \Pi^0} \right) \\
&= -\frac{1}{2} iV\hbar \int \frac{d^4q}{(2\pi)^4} \left[(\log[1 - \lambda V_0 \Pi^0])_{\lambda=0}^{\lambda=1} + V_0 \Pi^0 \right] \\
E_r &= -\frac{1}{2} iV\hbar \int \frac{d^3\mathbf{q}}{(2\pi)^3} \int \frac{d\omega}{2\pi} \{ \log[1 - V_0 \Pi^0] + V_0 \Pi^0 \} \\
&= -\frac{1}{2} iV\hbar \int \frac{d^3\mathbf{q}}{(2\pi)^3} \int \frac{d\omega}{2\pi} \{ \log[K_r(\mathbf{q}, \omega)] + 1 - K_r(\mathbf{q}, \omega) \}
\end{aligned} \tag{2.8.9'}$$

¹ $\left(\frac{mk_F}{\hbar^2} \right)^2 \left(\frac{4\pi e^2}{k_F^2} \right)^2 = \left(\frac{4\pi}{k_F a_0} \right)^2 = (4\pi)^2 (\alpha r_S)^2$

² $f(x) = \frac{2}{\pi} \left\{ 1 - \frac{x}{2} \ln \left| \frac{1+x}{1-x} \right| \right\}$

Using the dimensionless frequency and momentum, this can be given by the dimensionless integral.

$$E_r = \frac{Ne^2}{2a_0} \epsilon_r \quad (2.8.36)$$

$$\epsilon_r = -\frac{3i}{16\pi^2\alpha^2 r_S^2} \int d^3\mathbf{q}' \int_{-\infty}^{+\infty} d\nu \left\{ \log K_r(q', \nu) + 1 - K_r(q', \nu) \right\}$$

(with $\mathbf{q}' = \frac{\mathbf{q}}{k_F}$, $\nu = \frac{m}{\hbar k_F^2} \omega$)

Since K_r depends on the norm of \mathbf{q}' and it is an even function of ν , we have

$$\epsilon_r = -\frac{3i}{2\pi\alpha^2 r_S^2} \int_0^{+\infty} q'^2 \cdot dq' \int_0^{+\infty} d\nu \left\{ \log K_r(q', \nu) + 1 - K_r(q', \nu) \right\}$$

The energy must be real-valued, so that we focus only on the imaginary part of this parenthesis.

$$\epsilon_r = \frac{3}{2\pi\alpha^2 r_S^2} \int_0^{+\infty} q^2 \cdot dq \int_0^{+\infty} d\nu \left\{ \tan^{-1} \left(\frac{ImK_r(q, \nu)}{ReK_r(q, \nu)} \right) - ImK_r(q, \nu) \right\} \quad (2.8.37)$$

with

$$\log(ReK + iImK) = \log \sqrt{ReK^2 + ImK^2} + i \tan^{-1} \left(\frac{ImK}{ReK} \right) \quad (2.8.38)$$

In order to see how the singular behaviour of the integrand at $q = 0$ is removed by the summation of the ring-diagrams, we separate this momentum integral into two parts:

$$\epsilon_r = \epsilon_{r,1} + \epsilon_{r,2}$$

$$\epsilon_{r,1} = \frac{3}{2\pi\alpha^2 r_S^2} \int_0^{q_c} q^2 \cdot dq \int_0^{+\infty} d\nu \left\{ \tan^{-1} \left(\frac{ImK}{ReK} \right) - ImK \right\} \quad (2.8.39)$$

$$\epsilon_{r,2} = \frac{3}{2\pi\alpha^2 r_S^2} \int_{q_c}^{+\infty} q^2 \cdot dq \int_0^{+\infty} d\nu \left\{ \tan^{-1} \left(\frac{ImK}{ReK} \right) - ImK \right\} \quad (2.8.40)$$

$\epsilon_{r,2}$ is free from infrared divergence, so that it results in a finite contribution.

To evaluate $\epsilon_{r,1}$, we again choose q_c to be sufficiently smaller than 1, so that the integrand can be evaluated in terms of eqs.

$$\begin{aligned} \epsilon_{r,1} &= \frac{3}{2\pi\alpha^2 r_S^2} \int_0^{q_c} q^3 \cdot dq \int_0^{+\infty} dx \left\{ \tan^{-1} \left[\frac{2\alpha r_S x \theta(1-x)}{q^2 + 2\alpha r_S f(x)} \right] - \frac{2\alpha r_S x \theta(1-x)}{q^2} \right\} \\ &= \frac{3}{2\pi\alpha^2 r_S^2} \int_0^{q_c} q^3 \cdot dq \int_0^1 dx \left\{ \tan^{-1} \left[\frac{2\alpha r_S x}{q^2 + 2\alpha r_S f(x)} \right] - \frac{2\alpha r_S x}{q^2} \right\} \\ &\quad + \frac{3}{2\alpha^2 r_S^2} \int_0^{q_c} q^3 \cdot dq \int_1^{+\infty} dx \theta(-q^2 - 2\alpha r_S f(x)) \end{aligned} \quad (2.8.41)$$

where θ denotes the step function:

$$\theta(x) = \begin{cases} 1 & x > 0 \\ 0 & x < 0 \end{cases}$$

The second term is because $\tan^{-1} \left(\frac{K_2}{K_1} \right)$ takes π , when $K_2 = 0$ and K_1 is negative.

The singular behaviour of the integrand in small q regime is expected from the former part.

In fact, when we expand the former term with respect to $\alpha r_S \propto e^2$, we have

$$\begin{aligned} &= \frac{3}{2\pi\alpha^2 r_S^2} \int_0^{q_c} q^3 dq \int_0^1 dx \left\{ \frac{2\alpha r_S x}{q^2} - \frac{4\alpha^2 r_S^2 x f(x)}{q^4} + \dots - \frac{2\alpha r_S x}{q^2} \right\} \\ &= \frac{3}{2\pi} \int_0^{q_c} dq \int_0^1 dx \left\{ -\frac{4x f(x)}{q} + \dots \right\} \end{aligned} \quad (2.8.41')$$

where the first term is nothing but $\epsilon_2^{r,1}$ (given in eq.(2.8.9.d)) which exhibits logarithmic divergence at $q = 0$.

However, when we sum up all the ring-type diagram, we have this:

$$\epsilon_{r,1} = \frac{3}{2\pi\alpha^2 r_S^2} \int_0^{q_c} q^3 \cdot dq \int_0^1 dx \left\{ \underbrace{\tan^{-1} \left[\frac{2\alpha r_S x}{q^2 + 2\alpha r_S f(x)} \right] - \frac{2\alpha r_S x}{q^2}} + \dots \right.$$

in which the q^{-4} dependence of this underlined part is cut off for $q^2 < 2\alpha r_S f(x)$.

This integral can be evaluated for small r_S as follows:

$$\epsilon_{r,1} = \frac{2}{\pi^2} (1 - \ln 2) \ln r_S + \text{const} \quad (2.8.42)$$

Here ‘constant’ part means either some number which is free from r_S or some function which reduces to zero in small r_S limit.¹

Thus, the leading order contribution in the small r_S limit is given by:

$$E_r = \frac{N e^2}{2a_0} \left(\frac{2}{\pi^2} (1 - \ln 2) \ln r_S + \text{const} \right)$$

Later, many people calculated the subleading order correction to the correlation energy in the high density limit:

$$\frac{E_{corr}}{N} = \frac{e^2}{2a_0} \left(\frac{2}{\pi^2} (1 - \ln 2) \ln r_S - 0.094 + o(r_S \ln r_S) \right) \quad (2.8.43)$$

¹W.Macke, Z.Naturforsch 5a:192(1950)

Chapter 3

Application to electron gas at the zero temperature

3.1 Screening effect in a degenerate electron gas

In the preceding section, we have introduced the Dyson equation for the density-density correlation function, where the proper polarization part is calculated at the zero-th order perturbation theory.

These results are useful to understand many important physical properties of degenerate electron gas, such as screening effect & plasma oscillation.

In this section, we will study these properties within the ring-diagram approximation.

As a preliminary for this, let us first introduce a theory of linear response to a weak external perturbation.

Consider an interacting many-particles system with a time-independent Hamiltonian

$$i\hbar \frac{\partial}{\partial t} |\Psi_S(t)\rangle = H |\Psi_S(t)\rangle$$

$|\Psi_S(t)\rangle$ is a state vector in the Schrödinger picture, which is given by

$$|\Psi_S(t)\rangle = e^{-i\frac{H}{\hbar}t} |\Psi_S(0)\rangle$$

Suppose that the system is perturbed at $t = t_0$, by introducing an additional time-dependent Hamiltonian

$$i\hbar \frac{\partial}{\partial t} |\bar{\Psi}_S(t)\rangle = (H + H^{ex}(t)) |\bar{\Psi}_S(t)\rangle \quad (3.1.1)$$

This can be also solved symbolically:

$$|\bar{\Psi}_S(t)\rangle = e^{-i\frac{H}{\hbar}t} \hat{A}(t) |\Psi_S(t)\rangle \quad (3.1.2)$$

Here, $\hat{A}(t)$ reduces to a unit, when t is less than t_0 .

$$\hat{A}(t) = 1 \quad (t \leq t_0) \quad (3.1.3)$$

because

$$H^{ex}(t) = 0 \quad (t \leq t_0)$$

substituting this solution into this Schrödinger's equation, we obtain the equation of motion for $\hat{A}(t)$:

$$e^{-i\frac{H}{\hbar}t} \hat{H} \hat{A}(t) |\Psi_S(0)\rangle + e^{-i\frac{H}{\hbar}t} (i\hbar \frac{\partial \hat{A}(t)}{\partial t}) |\Psi_S(0)\rangle = (H + H^{ex}(t)) e^{-i\frac{H}{\hbar}t} \hat{A}(t) |\Psi_S(0)\rangle$$

These two are set off by each other. So that we have

$$\begin{aligned} i\hbar \frac{\partial}{\partial t} \hat{A}(t) &= e^{i\frac{H}{\hbar}t} H^{ex}(t) e^{-i\frac{H}{\hbar}t} \hat{A}(t) \\ &\stackrel{def}{=} H_H^{ex}(t) \cdot \hat{A}(t) \end{aligned} \quad (3.1.4)$$

where $H_H^{ex}(t)$ is the operator in the usual Heisenberg picture.

As we did many times in sec., the equation of motion is solved in favor for $\hat{A}(t)$ iteratively:

$$\hat{A}(t) = 1 - \frac{i}{\hbar} \int_{t_0}^t dt' \hat{H}_H^{ex}(t') + \left(\frac{-i}{\hbar}\right)^2 \int_{t_0}^t dt' \int_{t_0}^{t'} dt'' \hat{H}_H^{ex}(t') \hat{H}_H^{ex}(t'') + \dots$$

This clearly observes the boundary condition. Since $H^{ex}(t) = 0$ for $t \leq t_0$, we can also rewrite this as

$$\hat{A}(t) = 1 - \frac{i}{\hbar} \int_{-\infty}^t dt' \hat{H}_H^{ex}(t') + \left(\frac{-i}{\hbar}\right)^2 \int_{-\infty}^t dt' \int_{-\infty}^{t'} dt'' \hat{H}_H^{ex}(t') \hat{H}_H^{ex}(t'') + \dots \quad (3.1.5)$$

Correspondingly, we have

$$\begin{aligned} |\bar{\Psi}_S(t)\rangle &= e^{-i\frac{H}{\hbar}t} |\Psi_S(0)\rangle - \frac{i}{\hbar} e^{-i\frac{H}{\hbar}t} \int_{-\infty}^t dt' \hat{H}_H^{ex}(t') |\Psi_S(0)\rangle \\ &\quad + \left(\frac{-i}{\hbar}\right)^2 e^{-i\frac{H}{\hbar}t} \int_{-\infty}^t dt' \int_{-\infty}^{t'} dt'' \hat{H}_H^{ex}(t') \hat{H}_H^{ex}(t'') |\Psi_S(0)\rangle + \dots \end{aligned} \quad (3.1.5')$$

With this in mind, we can calculate an expectation value of a physical observable at any given time t :

$$\langle O(t) \rangle_{ex} = \frac{\langle \bar{\Psi}_S(t) | O_S | \bar{\Psi}_S(t) \rangle}{\langle \bar{\Psi}_S(t) | \bar{\Psi}_S(t) \rangle} \quad (3.1.6)$$

The numerator is calculated as follows:

$$\begin{aligned} \langle \bar{\Psi}_S(t) | O_S | \bar{\Psi}_S(t) \rangle &= \langle \Psi_S(0) | \left[1 + \frac{i}{\hbar} \int_{-\infty}^t dt' H_H^{ex}(t') + \dots \right] \times e^{i\frac{H}{\hbar}t} \hat{O}_S e^{-i\frac{H}{\hbar}t} \times \\ &\quad \left[1 - \frac{i}{\hbar} \int_{-\infty}^t dt' H_H^{ex}(t') + \dots \right] | \Psi_S(0) \rangle \\ &= \langle \Psi_S(0) | \hat{O}_H(t) | \Psi_S(0) \rangle + \\ &\quad \frac{i}{\hbar} \int_{-\infty}^t dt' \langle \Psi_S(0) | \left[H_H^{ex}(t') \hat{O}_H(t) - \hat{O}_H(t) H_H^{ex}(t') \right] | \Psi_S(0) \rangle \end{aligned} \quad (3.1.7)$$

As far as the first order in external perturbation, we can easily see that the denominator is equal to the norm at $t = 0$.

$$\begin{aligned} \langle \bar{\Psi}_S(t) | \bar{\Psi}_S(t) \rangle &= \langle \Psi_S(0) | \left[1 + \frac{i}{\hbar} \int_{-\infty}^t H_H^{ex}(t') dt' + \dots \right] \left[1 - \frac{i}{\hbar} \int_{-\infty}^t H_H^{ex}(t') dt' + \dots \right] | \Psi_S(0) \rangle \\ &= \langle \Psi_S(0) | \Psi_S(0) \rangle + O((H^{ex})^2) \end{aligned}$$

As such, the expectation value is evaluated within the first order in the external perturbation as follows.

$$\langle O(t) \rangle_{ex} = \frac{\langle \Psi_S(0) | \hat{O}_H(t) | \Psi_S(0) \rangle}{\langle \Psi_S(0) | \Psi_S(0) \rangle} + \frac{i}{\hbar} \int_{-\infty}^t dt' \frac{\langle \Psi_S(0) | \left[H_H^{ex}(t'), \hat{O}_H(t) \right] | \Psi_S(0) \rangle}{\langle \Psi_S(0) | \Psi_S(0) \rangle} + o((H^{ex})^2) \quad (3.1.8)$$

where the first term in the right hand side stands for the expectation value of the physical observable without external perturbation.¹

Thus, linear response of the ground state expectation value of the observable is given by

$$\begin{aligned}\delta\langle\hat{O}(t)\rangle &\equiv \langle O(t)\rangle_{ex} - \langle O(t)\rangle \\ &= \frac{i}{\hbar} \int_{-\infty}^t dt' \frac{\langle\Psi_0| [H_H^{ex}(t'), O_H(t)] |\Psi_0\rangle}{\langle\Psi_0|\Psi_0\rangle}\end{aligned}\quad (3.1.9)$$

where $|\Psi_S(0)\rangle$ was replaced by the ground-state wavefunction of H , which we call $|\Psi_0\rangle$.

For example, consider that a system with charge e per particle is subjected under an external scalar potential $\varphi^{ex}(\mathbf{x}, t)$:

$$\hat{H}_H^{ex}(t) = \int d^3\mathbf{x} \hat{n}_H(\mathbf{x}, t) e\varphi^{ex}(\mathbf{x}, t) \quad (3.1.10)$$

where $\hat{n}_H(\mathbf{x}, t) \equiv e^{i\frac{H}{\hbar}t} \hat{n}(\mathbf{x}) e^{-i\frac{H}{\hbar}t}$ and $\hat{n}(\mathbf{x})$ denotes particle density operator at \mathbf{x} .

We are interested in how the particle density will be perturbed by the external scalar potential.

Within the linear response theory described so far, it is given by

$$\delta\langle\hat{n}(\mathbf{x}, t)\rangle = \frac{i}{\hbar} \int_{-\infty}^t dt' \int d^3\mathbf{x}' e\varphi^{ex}(\mathbf{x}', t') \frac{\langle\Psi_0| [\hat{n}_H(\mathbf{x}', t'), \hat{n}_H(\mathbf{x}, t)] |\Psi_0\rangle}{\langle\Psi_0|\Psi_0\rangle}$$

Using the deviation operator,

$$\begin{aligned}\tilde{n}_H(\mathbf{x}, t) &\stackrel{def}{=} \hat{n}_H(\mathbf{x}, t) - \langle\hat{n}_H(\mathbf{x}, t)\rangle \\ &= \hat{n}_H(\mathbf{x}, t) - n\end{aligned}\quad (3.1.11)$$

we can rewrite this into the followings

$$\begin{aligned}\delta\langle\hat{n}(\mathbf{x}, t)\rangle &= \frac{i}{\hbar} \int_{-\infty}^t dt' \int d^3\mathbf{x}' e\varphi^{ex}(\mathbf{x}', t') \frac{\langle\Psi_0| [\tilde{n}_H(\mathbf{x}', t'), \tilde{n}_H(\mathbf{x}, t)] |\Psi_0\rangle}{\langle\Psi_0|\Psi_0\rangle} \\ &= \frac{1}{\hbar} \int_{-\infty}^{+\infty} dt' \int d^3\mathbf{x}' D^R(\mathbf{x}, t; \mathbf{x}', t') e\varphi^{ex}(\mathbf{x}', t')\end{aligned}\quad (3.1.12)$$

where we define the retarded density-density correlation function in analogy with the time-ordered density-density correlation function defined in eq.(2.8.2):

- Retarded density-density correlation function

$$iD^R(\mathbf{x}, t; \mathbf{x}', t') = \theta(t - t') \times \frac{\langle\Psi_0| [\tilde{n}_H(\mathbf{x}, t), \tilde{n}_H(\mathbf{x}', t')] |\Psi_0\rangle}{\langle\Psi_0|\Psi_0\rangle} \quad (3.1.12')$$

- Time-ordered density-density correlation function

$$iD(\mathbf{x}, t; \mathbf{x}', t') = \frac{\langle\Psi_0| \Gamma[\tilde{n}_H(\mathbf{x}, t), \tilde{n}_H(\mathbf{x}', t')] |\Psi_0\rangle}{\langle\Psi_0|\Psi_0\rangle} \quad (2.8.2')$$

¹ $\langle O(t)\rangle = \frac{\langle\Psi_S(0)|\hat{O}_H(t)|\Psi_S(0)\rangle}{\langle\Psi_S(0)|\Psi_S(0)\rangle}$

If the system without external perturbation is spatially and temporally homogeneous, we can assume the translational symmetry in the density-density correlation function:

$$D^R(\mathbf{x}, t; \mathbf{x}', t') = D^R(\mathbf{x} - \mathbf{x}', t - t') \quad (3.1.13)$$

So that we can take the Fourier transform of this(...) with

$$\begin{cases} \varphi^{ex}(\mathbf{k}, \omega) \equiv \int d^3\mathbf{x} \int dt \cdot e^{-i\mathbf{k}\cdot\mathbf{x}} e^{i\omega t} \varphi^{ex}(\mathbf{x}, t) \\ \delta\langle \hat{n}(\mathbf{k}, \omega) \rangle = \int d^3\mathbf{x} \int dt \cdot e^{-i\mathbf{k}\cdot\mathbf{x}} e^{i\omega t} \delta\langle \hat{n}(\mathbf{x}, t) \rangle \\ D^R(\mathbf{k}, \omega) = \int d^3\mathbf{x} \int dt \cdot e^{-i\mathbf{k}\cdot\mathbf{x}} e^{i\omega t} D^R(\mathbf{x}, t) \end{cases} \quad (3.1.14)$$

With this, we have

$$\delta\langle \hat{n}(\mathbf{k}, \omega) \rangle = \frac{1}{\hbar} D^R(\mathbf{k}, \omega) e\varphi^{ex}(\mathbf{k}, \omega) \quad (3.1.15)$$

This indicates that the system responds at the same wave vector and same frequency as the external perturbation.

From this, we can define a generalized susceptibility

$$\chi_{nn}(\mathbf{k}, \omega) \equiv \frac{\delta\langle \hat{n}(\mathbf{k}, \omega) \rangle}{e\varphi^{ex}(\mathbf{k}, \omega)} = \frac{1}{\hbar} D^R(\mathbf{k}, \omega) \quad (3.1.15')$$

Contrary to the time-order density correlation function, retarded function cannot be calculated directly with Feynman-Dyson perturbation series, because Wick's theorem can be applicable to time-ordered products of operators.

Thus, we next relate the retarded function with the time-ordered density correlation function.

In a similar way as we did for single-particle Green's function, we can do this by using the Lehmann representation.

Lehmann representation of the time-ordered density-density correlation function is obtained in the following way.

$$\begin{aligned} iD(\mathbf{x}, t; \mathbf{x}', t') &= \theta(t - t') \langle \Psi_0 | \tilde{n}_H(\mathbf{x}, t) \tilde{n}_H(\mathbf{x}', t') | \Psi_0 \rangle \\ &\quad + \theta(t' - t) \langle \Psi_0 | \tilde{n}_H(\mathbf{x}', t') \tilde{n}_H(\mathbf{x}, t) | \Psi_0 \rangle \\ &= \sum_n \left\{ \theta(t - t') e^{-i(E_n - E_0)(t - t')/\hbar} \langle \Psi_0 | \tilde{n}_H(\mathbf{x}) | \Psi_n \rangle \langle \Psi_n | \tilde{n}_H(\mathbf{x}') | \Psi_0 \rangle \right. \\ &\quad \left. - \theta(t' - t) e^{i(E_n - E_0)(t - t')/\hbar} \langle \Psi_0 | \tilde{n}_H(\mathbf{x}') | \Psi_n \rangle \langle \Psi_n | \tilde{n}_H(\mathbf{x}) | \Psi_0 \rangle \right\} \end{aligned}$$

where $|\Psi_n\rangle$ is an eigenstate of H , which belongs an eigenvalue E_n .

Using

$$\tilde{n}_H(\mathbf{x}) = e^{-i\hat{\mathbf{P}}\cdot\hat{\mathbf{x}}/\hbar} \tilde{n}_H(0) e^{i\hat{\mathbf{P}}\cdot\hat{\mathbf{x}}/\hbar}$$

where $\hat{\mathbf{P}}$ denote the total momentum operator, we have

$$\begin{aligned} &= \sum_n \left\{ \theta(t - t') e^{-i(E_n - E_0)(t - t')/\hbar} e^{i\mathbf{P}_n\cdot(\mathbf{x} - \mathbf{x}')/\hbar} \langle \Psi_0 | \tilde{n}_H(0) | \Psi_n \rangle \langle \Psi_n | \tilde{n}_H(0) | \Psi_0 \rangle \right. \\ &\quad \left. - \theta(t' - t) e^{i(E_n - E_0)(t - t')/\hbar} e^{-i\mathbf{P}_n\cdot(\mathbf{x} - \mathbf{x}')/\hbar} \langle \Psi_0 | \tilde{n}_H(0) | \Psi_n \rangle \langle \Psi_n | \tilde{n}_H(0) | \Psi_0 \rangle \right\} \end{aligned}$$

Here, we used a fact that $|\Psi_n\rangle$ is also an eigenstate of the total momentum operator $\hat{\mathbf{P}}$, because $\hat{\mathbf{P}}$ and $\hat{\mathbf{H}}$ commute with each other.

\mathbf{P}_n is an eigenvalue for $|\Psi_n\rangle$

By taking the Fourier transformation, we obtain

$$\begin{aligned} \imath D(\mathbf{k}, \omega) &= \int d^3(\mathbf{x} - \mathbf{x}') \int dt(t - t') e^{-\imath \mathbf{k} \cdot (\mathbf{x} - \mathbf{x}')} e^{\imath \omega(t - t')} \imath D(\mathbf{x}, t; \mathbf{x}', t') \\ &= \imath V \sum_n \delta_{\mathbf{k}, \frac{\mathbf{P}_n}{\hbar}} \frac{\langle \Psi_0 | \tilde{n}_H(0) | \Psi_n \rangle \langle \Psi_n | \tilde{n}_H(0) | \Psi_0 \rangle}{\omega - \hbar^{-1}(E_n - E_0) + \imath \eta} \\ &\quad - \imath V \sum_n \delta_{\mathbf{k}, -\frac{\mathbf{P}_n}{\hbar}} \frac{\langle \Psi_0 | \tilde{n}_H(0) | \Psi_n \rangle \langle \Psi_n | \tilde{n}_H(0) | \Psi_0 \rangle}{\omega + \hbar^{-1}(E_n - E_0) - \imath \eta} \end{aligned} \quad (3.1.16)$$

In a similar way, the Fourier transform of the retarded density correlation function is obtained as follows.

$$\begin{aligned} \imath D^R(\mathbf{k}, \omega) &= \imath V \sum_n \delta_{\mathbf{k}, \frac{\mathbf{P}_n}{\hbar}} \frac{\langle \Psi_0 | \tilde{n}_H(0) | \Psi_n \rangle \langle \Psi_n | \tilde{n}_H(0) | \Psi_0 \rangle}{\omega - \hbar^{-1}(E_n - E_0) + \imath \eta} \\ &\quad - \imath V \sum_n \delta_{\mathbf{k}, -\frac{\mathbf{P}_n}{\hbar}} \frac{\langle \Psi_0 | \tilde{n}_H(0) | \Psi_n \rangle \langle \Psi_n | \tilde{n}_H(0) | \Psi_0 \rangle}{\omega + \hbar^{-1}(E_n - E_0) - \imath \eta} \end{aligned} \quad (3.1.17)$$

Since the summation over $|\Psi_n\rangle$ is taken over an excited state of H , $E_n - E_0$ is always positive.

Therefore, these two expressions leads to the following general relation between $D^R(\mathbf{k}, \omega)$ and $D(\mathbf{k}, \omega)$.¹

$$\begin{cases} Re D^R(\mathbf{k}, \omega) = Re D(\mathbf{k}, \omega) \\ Im D^R(\mathbf{k}, \omega) = sgn(\omega) \cdot Im D(\mathbf{k}, \omega) \end{cases} \quad (3.1.18)$$

Using these relation, we can calculate the linear response function out of the time-ordered density-density correlation function.

Based on this general framework, let us first discuss about the screening effect in a degenerate electron gas.

Here, we are interested in how electron density will response, when a static impurity with positive charge Ze is introduced into an electron gas. Namely, we consider a static scalar potential in eq.(3.1.10)

$$\varphi^{ex}(\mathbf{x}, t) = \frac{Ze}{|\mathbf{x}|} \quad (3.1.19)$$

or equivalently

$$\varphi^{ex}(\mathbf{q}, \omega) = \frac{8\pi^2 Ze}{\mathbf{q}^2} \delta(\omega) \quad (3.1.19')$$

According to the linear response theory, the induced charge density is given by

$$\begin{aligned} \delta \langle n(\mathbf{x}) \rangle &= - \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \int \frac{d\omega}{2\pi} e^{\imath \mathbf{k} \cdot \mathbf{x}} e^{-\imath \omega t} \delta \langle n(\mathbf{k}, \omega) \rangle \\ &= - \frac{1}{\hbar} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \int \frac{d\omega}{2\pi} e^{\imath \mathbf{k} \cdot \mathbf{x}} e^{-\imath \omega t} D^R(\mathbf{k}, \omega) e \varphi^{ex}(\mathbf{k}, \omega) \\ &= - \int \frac{d^3 \mathbf{k}}{(2\pi)^3} e^{\imath \mathbf{k} \cdot \mathbf{x}} \frac{4\pi Ze^2}{\hbar \mathbf{k}^2} D^R(\mathbf{k}, 0) \end{aligned} \quad (3.1.20)$$

According to eq.(2.8.4') & eq.(2.8.5'), the time-ordered density-density correlation function is given by the proper polarization function in the following way.

$$D(\mathbf{k}, \omega) = \hbar \Pi(\mathbf{k}, \omega) = \frac{\hbar \Pi^\star(\mathbf{k}, \omega)}{1 - \Pi^\star(\mathbf{k}, \omega) V(\mathbf{k})}$$

¹ $sgn \omega \stackrel{def}{=} \frac{\omega}{|\omega|}$

Here $V(\mathbf{k})$ denotes the bare Coulombic potential.

$$V(\mathbf{k}) = \frac{4\pi e^2}{\mathbf{k}^2}$$

so that, we have

$$\begin{aligned} \frac{4\pi e^2}{\hbar k^2} D(\mathbf{k}, \omega) &= \frac{\Pi^\star(\mathbf{k}, \omega) V(\mathbf{k})}{1 - \Pi^\star(\mathbf{k}, \omega) V(\mathbf{k})} \\ &= \frac{1}{K(\mathbf{k}, \omega)} - 1 = \frac{ReK - iImK}{ReK^2 + ImK^2} - 1 \end{aligned} \quad (3.1.21)$$

Then, using the relation between the retarded density correlation function and time-ordered correlation function, we obtain.

$$\frac{4\pi e^2}{\hbar k^2} D^R(\mathbf{k}, \omega) = \frac{ReK(\mathbf{k}, \omega) - i \operatorname{sgn}(\omega) ImK(\mathbf{k}, \omega)}{ReK^2(\mathbf{k}, \omega) + ImK^2(\mathbf{k}, \omega)} - 1 \quad (3.1.22)$$

Within the ring-diagram approximation, this dielectric constant $K(\mathbf{k}, \omega)$ was evaluated explicitly, such as in eq.(2.8.33) & eq.(2.8.34) & eq.(2.8.35).

Especially, in the static limit where $\omega = 0$, the imaginary part of the dielectric constant reduces to zero, and we obtain

$$K_r(\mathbf{k}, 0) = 1 + 4\alpha r_S k_F^2 \frac{1}{\pi |\mathbf{k}|^2} g\left(\frac{|\mathbf{k}|}{k_F}\right) \quad (3.1.23)$$

where a real-valued function g is given by this.

$$g(x) = \frac{1}{2} - \frac{1}{2x} \left(1 - \frac{x^2}{4}\right) \ln \left| \frac{1 - \frac{x}{2}}{1 + \frac{x}{2}} \right| \quad (3.1.24)$$

By substituting eq.(3.1.23) into eq.(3.1.22) & eq.(3.1.20), we obtain the induced charge density as

$$\begin{aligned} \delta\langle\rho(\mathbf{x})\rangle_{ring} &= -e\delta\langle n(\mathbf{x})\rangle_{ring} \\ &= Ze \int \frac{d^3\mathbf{k}}{(2\pi)^3} e^{i\mathbf{k}\cdot\mathbf{x}} \frac{1 - K_{ring}(\mathbf{k}, 0)}{K_{ring}(\mathbf{k}, 0)} \\ &= -Ze \int \frac{d^3\mathbf{k}}{(2\pi)^3} e^{i\mathbf{k}\cdot\mathbf{x}} \frac{\frac{4\alpha r_S}{\pi} \left(\frac{k_F^2}{k^2}\right) g\left(\frac{k}{k_F}\right)}{1 + \frac{4\alpha r_S}{\pi} \left(\frac{k_F^2}{k^2}\right) g\left(\frac{k}{k_F}\right)} \\ &= -Ze \int \frac{d^3\mathbf{k}}{(2\pi)^3} e^{i\mathbf{k}\cdot\mathbf{x}} \frac{\frac{4\alpha r_S}{\pi} g\left(\frac{k}{k_F}\right)}{\left(\frac{k}{k_F}\right)^2 + \frac{4\alpha r_S}{\pi} g\left(\frac{k}{k_F}\right)} \end{aligned}$$

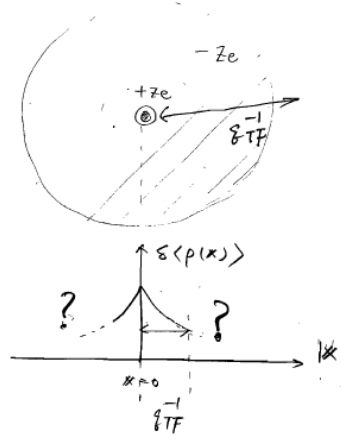
This expression has several important features.

- First of all, the total induced charge is equal to $-Ze$.

$$\begin{aligned} \delta Q_r &= \int d^3\mathbf{x} \delta\langle\rho(\mathbf{x})\rangle_{ring} \\ &= -Ze \int d^3\mathbf{k} \delta^3(\mathbf{k}) \frac{\frac{4\alpha r_S}{\pi} g\left(\frac{k}{k_F}\right)}{\left(\frac{k}{k_F}\right)^2 + \frac{4\alpha r_S}{\pi} g\left(\frac{k}{k_F}\right)} \\ &= -Ze \end{aligned} \quad (3.1.25)$$

which means that induced charge density screens the impurity with $+Ze$ charge completely.

- The second feature is that the induced charge density is finite everywhere and it has maximum at $\mathbf{x} = 0$.



$$\delta\langle\rho(\mathbf{x})\rangle_{ring} \leq \delta\langle\rho(\mathbf{x})\rangle_{ring} < +\infty \quad (3.1.26)$$

- Thirdly, we are interested in how the induced charge density will decay when \mathbf{x} goes away from the origin.

To calculate, let us first define Thomas-Fermi wavenumber as

$$q_{TF} \stackrel{def}{=} \sqrt{\frac{4\alpha r_S}{\pi}} \cdot k_F \quad (3.1.27)$$

In terms of this, the induced charge density is given by

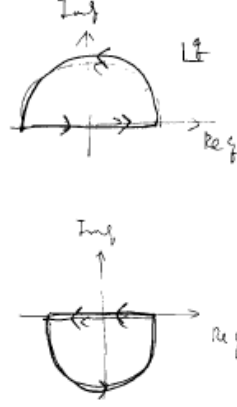
$$\begin{aligned} \delta\langle\rho(\mathbf{x})\rangle &= -Ze \int \frac{d^3\mathbf{q}}{(2\pi)^3} e^{i\mathbf{q}\cdot\mathbf{x}} \frac{q_{TF}^2 g(q/k_F)}{q^2 + q_{TF}^2 g(q/k_F)} \\ &= -Ze \int \frac{dq}{(2\pi)^2} \int_{-1}^1 dt q^2 e^{iqxt} \frac{q_{TF}^2 g(q/k_F)}{q^2 + q_{TF}^2 g(q/k_F)} \\ &= -\frac{Ze}{2\pi^2 x} \int dq \cdot q \sin(qx) \frac{q_{TF}^2 g(q/k_F)}{q^2 + q_{TF}^2 g(q/k_F)} \\ \delta\langle\rho(\mathbf{x})\rangle &= -\frac{Ze}{4\pi^2 x} \int_{-\infty}^{+\infty} dq \cdot q \sin(qx) \frac{q_{TF}^2 g(q/k_F)}{q^2 + q_{TF}^2 g(q/k_F)} \end{aligned} \quad (3.1.28)$$

Since $g(x)$ is a function of order unit, the second term in the denominator becomes important only when $q \ll q_{TF}$.

In the high-density limit where $r_S \rightarrow 0$, we can approximate $g(\frac{q}{k_F})$ by $g(0)$ for those q which is much smaller than the Thomas-Fermi wavenumber.

$g(0) = 1$, so that we can approximately evaluate this integral as

$$r_S \rightarrow 0 \approx -\frac{Ze}{4\pi^2 x} \int_{-\infty}^{+\infty} dq \cdot q \left(\frac{e^{iqx} - e^{-iqx}}{2i} \right) \frac{q_{TF}^2}{q^2 + q_{TF}^2}$$



$$\begin{aligned}
&= -\frac{Ze}{4\pi^2 x} \oint_{\Gamma_{upperhalf}} dq \frac{1}{2} \frac{e^{iqx}}{2i} \left\{ \frac{1}{q - iq_{TF}} + \frac{1}{q + iq_{TF}} \right\} \\
&\quad - \frac{Ze}{4\pi^2 x} \oint_{\Gamma_{lowerhalf}} dq \frac{1}{2} \frac{e^{-iqx}}{2i} \left\{ \frac{1}{q - iq_{TF}} + \frac{1}{q + iq_{TF}} \right\} \\
&= -\frac{Ze}{8\pi x} q_{TF}^2 e^{-q_{TF}x} \times 2 \\
&= -\frac{Ze}{4\pi x} q_{TF}^2 e^{-q_{TF}x} \tag{3.1.29}
\end{aligned}$$

$\delta\langle\hat{\rho}(\mathbf{x})\rangle_{ring}$ thus calculated is called Thomas-Fermi result, which signifies that induced charge density is roughly extended within $|\mathbf{x}| \leq q_{TF}^{-1}$.

The Thomas-Fermi result is not quite correct, especially when x become sufficiently large.

To obtain accurate asymptotic behaviour of the induced charge density, notice first that $g(x)$ has singularity at $x = \pm 2$, because of this logarithmic function:

$$g(x) = \frac{1}{2} - \frac{1}{2x} \left(1 - \frac{x^2}{4} \right) \ln \left| \frac{1 - \frac{x}{2}}{1 + \frac{x}{2}} \right|$$

To put the conclusion first, this singularity gives rise to an algebraic asymptotic dependence on x of $\delta\langle\rho(\mathbf{x})\rangle_{ring}$.

To see this, let us rewrite this logarithmic function in this way

$$\log \left| \frac{q - 2k_F}{q + 2k_F} \right| = \lim_{\eta \rightarrow 0} \frac{1}{2} \log \left(\frac{(q - 2k_F)^2 + \eta^2}{(q + 2k_F)^2 + \eta^2} \right) \tag{3.1.30}$$

substituting this expression into eq.(3.1.28), we have

$$\begin{aligned}
 \delta\langle\rho(\mathbf{x})\rangle &= -\frac{Ze}{4\pi^2 x i} \lim_{\eta \rightarrow 0} \int_{-\infty}^{+\infty} dq \cdot q e^{iqx} \\
 &\quad \left(\frac{q^2}{q^2 + q_{TF}^2 \left\{ \frac{1}{2} - \frac{k_F}{2q} \left(1 - \frac{1}{4} \left(\frac{q}{k_F} \right)^2 \right) \frac{1}{2} \log \left(\frac{(q-2k_F)^2 + \eta^2}{(q+2k_F)^2 + \eta^2} \right) \right\}} - 1 \right) \\
 &= -\frac{Ze}{4\pi^2 x i} \lim_{\eta \rightarrow 0} \int_{-\infty}^{+\infty} dq e^{iqx} \\
 &\quad \times \frac{q^3}{q^2 + q_{TF}^2 \left\{ \frac{1}{2} - \frac{k_F}{2q} \left(1 - \frac{1}{4} \left(\frac{q}{k_F} \right)^2 \right) \frac{1}{2} \log \left(\frac{(q-2k_F)^2 + \eta^2}{(q+2k_F)^2 + \eta^2} \right) \right\}}
 \end{aligned}$$

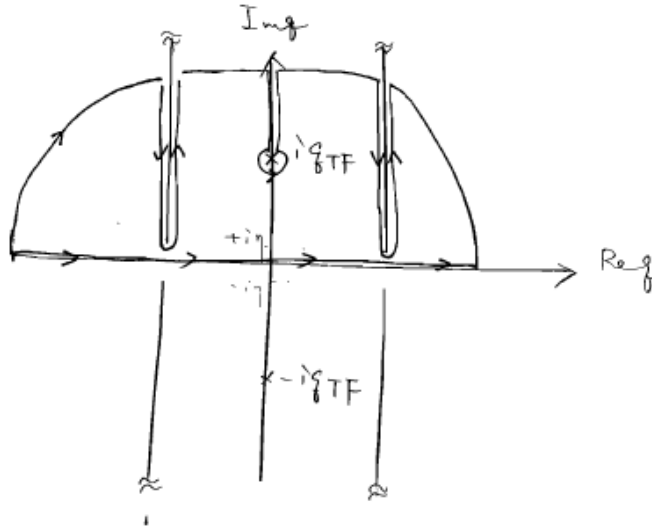
The integrand has a pole at $q = \pm i q_{TF}$, which gives rises to a Thomas-Fermi result, as described above.

On top of these poles, the integrand requires us to introduce a branch cut associated with these logarithmic functions.

Namely, when $q = \pm 2k_F + iV$, the arguments of these logarithmic function become negative when absolute value of V is greater than η , e.g.

$$\begin{aligned}
 \log \left[\frac{(q-2k_F)^2 + \eta^2}{(q+2k_F)^2 + \eta^2} \right]_{q=2k_F+iV} &= \log [(q-2k_F)^2 + \eta^2]_{q=2k_F+iV} - \log [(q+2k_F)^2 + \eta^2]_{q=2k_F+iV} \\
 &= \log [-V^2 + \eta^2] - \log [\dots]
 \end{aligned} \tag{3.1.31}$$

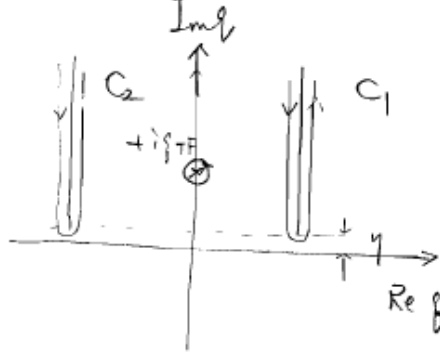
As such, we introduce four branch cuts in the complex- q plane.



Correspondingly, we will deform the integral path along the real axis into this way. We choose the upper half plane, because of this e^{iqx} where x is positive definite.

When the contour takes to be infinitely far from the origin, we have

$$\delta\langle\rho(\mathbf{x})\rangle = -\frac{Ze}{4\pi^2 x} \lim_{\eta \rightarrow 0} \left(\oint_{|q-iq_{TF}|=\epsilon} + \oint_{C_1} + \oint_{C_2} \right) \frac{q^3}{q^2 + q_{TF}^2 \left\{ \frac{1}{2} - \frac{k_F}{2q} \left(1 - \frac{1}{4} \left(\frac{q}{k_F} \right)^2 \right) \frac{1}{2} \log \left(\frac{(q-2k_F)^2 + \eta^2}{(q+2k_F)^2 + \eta^2} \right) \right\}} \quad (3.1.32)$$

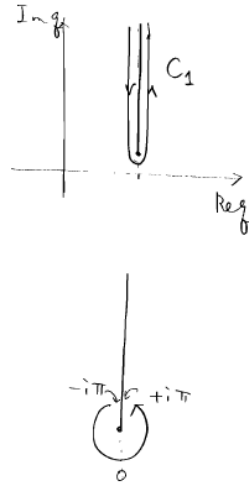


Where this corresponds to the Thomas-Fermi pole contribution while these two contour-integrals are taken along these two branch cuts.

As we have already see above, the Thomas-Fermi pole contribution vanishes exponentially for larger x .

On the one hand, these two contour integrals give rise to “algebraic decay for larger x ” which is much more dominant than the exponential decay.

To see this algebraic decay, let us first calculate this integral.



According to this branch cut, the imaginary part of this logarithmic function takes $-i\pi$ on the left side of C_1 , while takes $+i\pi$ on the right side of C_1 .

With this in mind, we can calculate this integral as

$$\begin{aligned}
& \oint_{C_1} dq e^{iqx} \frac{q^3}{q^2 + q_{TF}^2 \left\{ \frac{1}{2} - \frac{k_F}{2q} \left(1 - \frac{1}{4} \left(\frac{q}{k_F} \right)^2 \right) \frac{1}{2} \log \left(\frac{(q-2k_F)^2 + \eta^2}{(q+2k_F)^2 + \eta^2} \right) \right\}} \\
&= i \int_{+\infty}^{\eta} dV e^{i(2k_F + iV)x} \frac{(2k_F + iV)^3}{(2k_F + iV)^2 + q_{TF}^2 g \left(\frac{2k_F - \epsilon + iV}{k_F} \right)} \\
&\quad + i \int_{\eta}^{+\infty} dV e^{i(2k_F + iV)x} \frac{(2k_F + iV)^3}{(2k_F + iV)^2 + q_{TF}^2 g \left(\frac{2k_F + \epsilon + iV}{k_F} \right)} \\
&= i \int_{\eta}^{+\infty} dV e^{i(2k_F + iV)x} \frac{(2k_F + iV)^3 (q_{TF}^2 g_- - q_{TF}^2 g_+)}{[(2k_F + iV)^2 + q_{TF}^2 g_+][(2k_F + iV)^2 + q_{TF}^2 g_-]}
\end{aligned} \tag{3.1.33}$$

where we take q to be $2k_F + iV - \epsilon$ in the integral along the left side of C_1 , while take q to be $2k_F + iV + \epsilon$ in the integral along the right side of C_1 .

Here ϵ is a infinitesimally small real and positive quantity.

g_{\pm} stands for these two function respectively

$$g_{\pm} \equiv g \left(\frac{2k_F + iV \pm \epsilon}{k_F} \right)$$

with

$$g(x) = \frac{1}{2} - \frac{1}{2x} \left(1 - \frac{1}{4} x^2 \right) \frac{1}{2} \ln \left[\frac{(x-2)^2 + \eta^2}{(x+2)^2 + \eta^2} \right]$$

As mentioned before, g_+ and g_- differ from each other along the branch cut, because of this logarithmic part:

$$\begin{aligned}
g_+ - g_- &= \frac{1}{2} \frac{k_F}{2k_F + iV} \left(1 - \frac{1}{4} \left(\frac{2k_F + iV}{k_F} \right)^2 \right) \times \frac{1}{2} \left[\ln \left[\frac{(x-2)^2 + \eta^2}{(x+2)^2 + \eta^2} \right] \right]_{x=\frac{2k_F + iV + \epsilon}{k_F}}^{x=\frac{2k_F + iV - \epsilon}{k_F}} \\
&= \frac{\pi i}{2} \frac{k_F}{2k_F + iV} \left(1 - \frac{1}{4} \left(\frac{2k_F + iV}{k_F} \right)^2 \right)
\end{aligned}$$

substituting this into eq.(3.1.33), we have

$$\begin{aligned}
\text{eq. (3.1.33)} &= i \int_{\eta}^{+\infty} dV e^{i2k_F x} e^{-Vx} \frac{(2k_F + iV)^3 q_{TF}^2}{\left[(2k_F + iV)^2 + q_{TF}^2 g \left(\frac{2k_F + iV}{k_F} \right) \right]^2} \times \\
&\quad \left(-\frac{\pi i}{2} \right) \frac{k_F}{2k_F + iV} \left(1 - \frac{1}{4} \left(\frac{2k_F + iV}{k_F} \right)^2 \right)
\end{aligned}$$

Since this integral is a bit involved, we will expand the integrand in power of V . We will retain only the lowest power of V .

As will be clear soon, the power of V corresponds to the power of $\frac{1}{x}$ in the final expression for the induced charge density.

As such, we have only to retain the lowest power of V , in order to obtain the leading asymptotic form for the larger x .

$$\begin{aligned}
\text{eq. (3.1.33)} &= e^{i2k_F x} \int_{\eta}^{+\infty} dV e^{-Vx} \frac{\pi}{2} q_{TF}^2 \times \\
&\quad \left\{ \frac{4k_F^3}{\left(4k_F^2 + \frac{q_{TF}^2}{2}\right)^2} \cdot \left(1 - \frac{1}{4} \left(4 + 4i \frac{V}{k_F}\right)\right) + o(V^2) \right\} \\
&= e^{i2k_F x} \int_{\eta}^{+\infty} dV e^{-Vx} \left(-i \frac{V}{k_F} + o(V^2)\right) \times \frac{\pi q_{TF}^2}{2} \frac{4k_F^3}{\left(4k_F^2 + \frac{q_{TF}^2}{2}\right)^2} \\
&= -ie^{i2k_F x} \frac{2\pi q_{TF}^2 k_F^2}{\left(4k_F^2 + \frac{q_{TF}^2}{2}\right)^2} \times \int_{\eta}^{+\infty} dV e^{-Vx} (V + o(V^2)) \\
&= -ie^{i2k_F x} \frac{2\pi q_{TF}^2 k_F^2}{\left(4k_F^2 + \frac{q_{TF}^2}{2}\right)^2} \times \left(\left(-\frac{\partial}{\partial x}\right) \left(\int_{\eta}^{+\infty} dV e^{-Vx}\right) + a \frac{\partial^2}{\partial x^2} \left(\int_{\eta}^{+\infty} dV e^{-Vx}\right) + \dots \right) \\
\text{eq. (3.1.33)} &= -ie^{i2k_F x} \frac{2\pi q_{TF}^2 k_F^2}{\left(4k_F^2 + \frac{q_{TF}^2}{2}\right)^2} \times \left(\frac{1}{x^2} + o\left(\frac{1}{x^3}\right) \right) \tag{3.1.34}
\end{aligned}$$

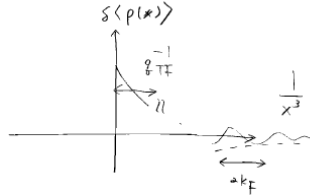
Similarly, we can also calculate the other contour integral.

$$\begin{aligned}
&\oint_{C_2} dq e^{iqx} \frac{q^3}{q^2 + q_{TF}^2 \left\{ \frac{1}{2} - \frac{k_F}{2q} \left(1 - \frac{1}{4} \left(\frac{q}{k_F}\right)^2\right) \frac{1}{2} \log \left(\frac{(q-2k_F)^2 + \eta^2}{(q+2k_F)^2 + \eta^2} \right) \right\}} \\
&= -ie^{-i2k_F x} \frac{2\pi q_{TF}^2 k_F^2}{\left(4k_F^2 + \frac{q_{TF}^2}{2}\right)^2} \times \left(\frac{1}{x^2} + o\left(\frac{1}{x^3}\right) \right) \tag{3.1.35}
\end{aligned}$$

Then, substituting these two into eq. (3.1.32), we finally obtain the leading-order asymptotic form of the induced charge density.

$$\begin{aligned}
\delta\langle\rho(\mathbf{x})\rangle &= -\frac{Ze}{4\pi^2 x i} \times \frac{2\pi q_{TF}^2 k_F^2}{\left(4k_F^2 + \frac{q_{TF}^2}{2}\right)^2} \frac{1}{x^2} \left\{ -ie^{2ik_F x} - ie^{-2ik_F x} \right\} + o\left(\frac{1}{x^4}\right) \\
&= \frac{Ze}{\pi} \frac{2\xi}{[4 + \xi]^2} \frac{\cos(2k_F x)}{x^3} + o\left(\frac{1}{x^4}\right) \quad (\text{for } x \rightarrow +\infty) \tag{3.1.36}
\end{aligned}$$

where $\xi \equiv \frac{q_{TF}^2}{2k_F^2}$.



This result was first obtained by Langer & Vosko.¹

This expression is clearly different from Thomas-Fermi results and exhibits long-range oscillation with a radial wavelength $\frac{\pi}{k_F}$ and an envelop proportional to x^{-3} .

From physical point of view, this long-range oscillation in the screening charge density profile comes from a Fermi surface.

This oscillation behaviour in correlation functions are also called as Friedel oscillation and have been as a broadening of nuclear magnetic resonance in dilute alloys.

Similar effect is also known as the RKKY interaction in dilute magnetic alloys, where conduction electrons induce an indirect magnetic interaction between magnetic impurities of the form of $x_{ij}^{-3} \cos(2k_F x_{ij})$, where x_{ij} denotes a spatial distance between the two magnetic impurities.

So far, we have discussed about how a test charge introduced in an electron gas will be screened by the induced charge density.

Let us next discuss about collective modes in degenerate electron gas and that in imperfect Fermi gas.

3.2 Collective Modes in degenerate electron gas and imperfect Fermi gas – plasma oscillation & zero sound mode

In section, we studied the Lehmann representation of single-particle Green's function, which tells us about a single-particle excitations, called also as a quasi-particle excitations.

A single-particle excitation is accompanied by an addition of one electron outside the Fermi sea or by a removal of one electron inside the Fermi sea. Namely, a single-particle excitation changes the total number of particles.

On the one hand, most physical systems also have another type of excited states which do not change the total number of particles. These excitations have a boson character and often called as collective modes.

Famous example of collective modes is “phonon”, which describes how lattice vibrations are propagating in space and time, and which we studied in sec. 2.

Another example of collective modes is “spin wave” or “magnon”, which describes how magnetic excitations are propagating in magnets.

Degenerate electron gas and imperfect Fermi gas also have collective modes as their excited states, which are called as plasma oscillation mode and zero sound mode respectively.

Physically, both of these modes describe a collective propagation of density fluctuation, so that these modes are connected to the ground state through the density operator.

As is clear from the Lehmann representation of the density-density correlation function given in eq.(3.1.16), the polarization part $\Pi(\mathbf{q}, \omega) \equiv \frac{1}{\hbar} D(\mathbf{q}, \omega)$ has poles at the exact excitation energy of these collective modes.

$$\begin{aligned} \Pi(\mathbf{q}, \omega) &= \hbar^{-1} D(\mathbf{q}, \omega) \\ &= \frac{V}{i\hbar} \sum_n \delta_{\mathbf{q}, \frac{\mathbf{p}_n}{\hbar}} \frac{\langle \Psi_0 | \tilde{n}_H(0) | \Psi_n \rangle \langle \Psi_n | \tilde{n}_H(0) | \Psi_0 \rangle}{\omega - \hbar^{-1}(E_n - E_0) + i\eta} \\ &\quad - \frac{V}{i\hbar} \sum_n \delta_{\mathbf{q}, -\frac{\mathbf{p}_n}{\hbar}} \frac{\langle \Psi_0 | \tilde{n}_H(0) | \Psi_n \rangle \langle \Psi_n | \tilde{n}_H(0) | \Psi_0 \rangle}{\omega + \hbar^{-1}(E_n - E_0) - i\eta} \end{aligned} \quad (3.1.16')$$

¹ Ref: J.S.Langer & S.H.Vosko J.Phys.chem.Solids 12 196(1960)

Poles of $\Pi(\mathbf{q}, \omega)$ can be detected by the zeros of the dielectric constant $K(\mathbf{q}, \omega)$

$$\frac{1}{K(\mathbf{q}, \omega)} - 1 = U_0(\mathbf{q})\Pi(\mathbf{q}) \quad (3.2.1)$$

According to eq.(2.8.34), the dielectric constant evaluated within the ring-diagram approximation has one obvious zero in the long-wavelength limit $|\mathbf{q}| \rightarrow 0$.

Namely, in the dynamic limit, the dielectric constant is given by

$$K_r(\mathbf{q}, \omega) = \frac{\hbar k_F^2}{m} \nu = 1 - \frac{4\alpha r_S}{3\pi \nu^2} \quad (2.8.34)$$

which reduces to zero at

$$\nu^2 = \frac{4\alpha r_S}{3\pi}$$

or equivalently

$$\omega^2 = \frac{4\alpha r_S}{3\pi} \left(\frac{\hbar k_F^2}{m} \right)^2 = \frac{4\pi n e^2}{m} \quad (3.2.2)$$

where n denote the density of electron ($n = \frac{N}{V}$).

This means that a degenerate electron gas has a collective mode at $\omega = \sqrt{\frac{4\pi n e^2}{m}}$.

To study this collective mode in detail, let us consider that a degenerate electron gas is subjected under an impulsive perturbation.

$$\varphi^{ex}(\mathbf{x}, t) = e^{i\mathbf{q} \cdot \mathbf{x}} \varphi_0 \delta(t) \quad (3.2.3)$$

or equivalently

$$\varphi^{ex}(\mathbf{k}, \omega) = \varphi_0 (2\pi)^3 \delta^3(\mathbf{q} - \mathbf{k}) \quad (3.2.3')$$

According to the linear response theory, the induced charge density is given by

$$\begin{aligned} \delta \langle \hat{n}(\mathbf{x}, t) \rangle &= -\frac{1}{\hbar} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \int \frac{d\omega}{2\pi} e^{i\mathbf{k} \cdot \mathbf{x}} e^{-i\omega t} D^R(\mathbf{k}, \omega) e \varphi_0 (2\pi)^3 \delta^3(\mathbf{q} - \mathbf{k}) \\ &= -e \varphi_0 e^{i\mathbf{q} \cdot \mathbf{x}} \int \frac{d\omega}{2\pi} e^{-i\omega t} \Pi^R(\mathbf{q}, \omega) \end{aligned} \quad (3.2.4)$$

As is clear from the Lehmann representation of the retarded density correlation function, $\Pi^R(\mathbf{q}, \omega)$ has a pole only in the lower half plane in the complex- ω plane.

$$\Pi^R(\mathbf{q}, \omega) = \frac{iA_{\mathbf{q}}}{\omega - (\Omega_{\mathbf{q}} - i\gamma_{\mathbf{q}})} + \dots \quad (3.2.5)$$

substituting this into here, we have

$$\begin{aligned} &= -e \varphi_0 e^{i\mathbf{q} \cdot \mathbf{x}} \int_{-\infty}^{+\infty} \frac{d\omega}{2\pi} e^{-i\omega t} \frac{iA_{\mathbf{q}}}{\omega - (\Omega_{\mathbf{q}} - i\gamma_{\mathbf{q}})} \\ &= -e \varphi_0 A_{\mathbf{q}} e^{i\mathbf{q} \cdot \mathbf{x}} e^{-i\Omega_{\mathbf{q}} t - \gamma_{\mathbf{q}} t} \end{aligned} \quad (3.2.6)$$

This suggests that the real part of the pole(\dots) determines the dispersion relation of the collective mode, while the imaginary part(\dots) is nothing but the inverse of the life time

To obtain an explicit form of $\Omega_{\mathbf{q}}$ and $\gamma_{\mathbf{q}}$, we need to calculate the polarization function.

According to eq.(3.1.18), we have

$$\Pi^R(\mathbf{q}, \omega) = \text{Re} \Pi(\mathbf{q}, \omega) + i \text{sgn}(\omega) \text{Im} \Pi(\mathbf{q}, \omega)$$

Within the ring-diagram approximation, time-ordered polarization part is given by this:

$$\Pi(\mathbf{q}, \omega) = \frac{\Pi^0(\mathbf{q}, \omega)}{1 - \Pi^0(\mathbf{q}, \omega)V(\mathbf{q})} \quad (2.8.5'')$$

Since $V(\mathbf{q})$ is real, we have

$$\Pi^R(\mathbf{q}, \omega) \equiv \frac{\Pi^{0,R}(\mathbf{q}, \omega)}{1 - \Pi^{0,R}(\mathbf{q}, \omega)V(\mathbf{q})} \quad (3.2.7)$$

where

$$\Pi^{0,R}(\mathbf{q}, \omega) = \text{Re}\Pi^0(\mathbf{q}, \omega) + i \text{sgn}(\omega) \text{Im}\Pi^0(\mathbf{q}, \omega)$$

where $\Pi^0(\mathbf{q}, \omega)$ was already calculated in the previous section.

The pole of $\Pi^R(\mathbf{q}, \omega)$ can be determined by the zero of the denominator of the right hand side, so that pole satisfies the following equation:

$$1 - \Pi^{0,R}(\mathbf{q}, \Omega_{\mathbf{q}} - i\gamma_{\mathbf{q}})V(\mathbf{q}) = 0 \quad (3.2.8)$$

When the damping (inverse of the life time) is sufficiently smaller than the resonance frequency $\Omega_{\mathbf{q}}$, which turns out to be the case a priori, we can expand this equation is small $\gamma_{\mathbf{q}}$:

$$1 - \left(\text{Re}\Pi^{0,R}(\mathbf{q}, \Omega_{\mathbf{q}}) + i \text{Im}\Pi^{0,R}(\mathbf{q}, \Omega_{\mathbf{q}}) - i\gamma_{\mathbf{q}} \left\{ \left[\frac{\partial \text{Re}\Pi^{0,R}(\mathbf{q}, \omega)}{\partial \omega} \right]_{\omega=\Omega_{\mathbf{q}}} + i \left[\frac{\partial \text{Im}\Pi^{0,R}(\mathbf{q}, \omega)}{\partial \omega} \right]_{\omega=\Omega_{\mathbf{q}}} \right\} + o(\gamma_{\mathbf{q}}^2) \right) V(\mathbf{q}) = 0 \quad (3.2.9)$$

The real part of this equality is given by

$$1 - \text{Re}\Pi^{0,R}(\mathbf{q}, \Omega_{\mathbf{q}})V(\mathbf{q}) + o(\gamma_{\mathbf{q}}) = 0 \quad (3.2.10)$$

while the imaginary part is given by

$$\left(\text{Im}\Pi^{0,R}(\mathbf{q}, \Omega_{\mathbf{q}}) - \gamma_{\mathbf{q}} \left[\frac{\partial \text{Re}\Pi^{0,R}(\mathbf{q}, \omega)}{\partial \omega} \right]_{\omega=\Omega_{\mathbf{q}}} + o(\gamma_{\mathbf{q}}^2) \right) = 0 \quad (3.2.11)$$

Regarding that $\text{Im}\Pi^{0,R}$ & $\gamma_{\mathbf{q}}$ are same order of small quantity, we can solved these equalities in the leading order in small $\gamma_{\mathbf{q}}$

$$\begin{cases} 1 - \text{Re}\Pi^{0,R}(\mathbf{q}, \Omega_{\mathbf{q}})V(\mathbf{q}) = 0 \\ \gamma_{\mathbf{q}} = \text{Im}\Pi^{0,R}(\mathbf{q}, \Omega_{\mathbf{q}}) \cdot \left[\frac{\partial \text{Re}\Pi^{0,R}(\mathbf{q}, \omega)}{\partial \omega} \right]_{\omega=\Omega_{\mathbf{q}}}^{-1} \end{cases} \quad (3.2.12)$$

The former equalities determines the dispersion relation of the collective mode, while the latter gives the inverse of the life time of the collective mode.

Now that retarded polarization function and time-ordered one is related by eq.(3.2.8), these two can be rewritten as follows

$$\begin{cases} 1 - \text{Re}\Pi^0(\mathbf{q}, \omega)V(\mathbf{q}) = 0 \\ \gamma_{\mathbf{q}} = \text{sgn}(\omega) \text{Im}\Pi^0(\mathbf{q}, \omega) \cdot \left[\frac{\partial \text{Re}\Pi^0(\mathbf{q}, \omega)}{\partial \omega} \right]_{\omega=\Omega_{\mathbf{q}}}^{-1} \end{cases} \quad (3.2.12')$$

For the long-wavelength limit ($|\mathbf{q}| \rightarrow 0$), we have already evaluate an explicit form of Π^0 (zero-th order polarization part)

According to eq.(2.8.34), the imaginary part vanishes, while the real part is given by

$$Re\Pi^0(\mathbf{q}, \omega) = \frac{k_F^3}{3\pi^2} \frac{q^2}{m\omega^2} + o(q^4)$$

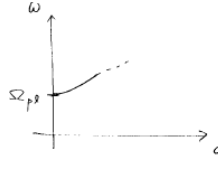
If one further study the high order connection, we have

$$Re\Pi^0(\mathbf{q}, \omega) = \frac{k_F^3}{3\pi^2} \frac{q^2}{m\omega^2} \left[1 + \frac{3}{5} \left(\frac{\hbar k_F q}{m\omega} \right)^2 + \dots \right]$$

Using this, the dispersion relation is obtained in the following way. ¹

$$\Omega_{\mathbf{q}} = \left(\frac{4\pi n e^2}{m} \right)^{\frac{1}{2}} \left[1 + \frac{9}{10} \left(\frac{q}{q_{TF}} \right)^2 + \dots \right] \quad (3.2.13)$$

where the second term is determined by this next leading order term of $Re\Pi^0$.



Since $Im\Pi^{0,R}(\mathbf{q}, \Omega_{\mathbf{q}})$ vanishes in the dynamic limit, these gapped collective mode has an infinitely long life time, within the ring-diagram approximation.

As is clear from this derivation, a finite gap of the plasma oscillation mode comes from the long range Coulomb interaction. Namely, the bare Coulomb interaction has $\frac{1}{q^2}$ -divergence in small q region, which is cancelled by q^2 appearing in this part of $Re\Pi^0$. Due to this cancellation, this equation is satisfied by a finite ω in the small q -limit.

On the one hand, the interaction potential in imperfect Fermi gas is short-ranged, where $V(\mathbf{q})$ remains finite in the small q -limit.

In such a case, this equation reduces to

$$1 - \frac{k_F^3}{3\pi^2} \frac{q^2}{m\omega^2} V(0) = 0 \quad (3.2.14)$$

where q^2 in the numerator is not cancelled, because $V(0)$ is finite.

As a result, the resonance frequency of the collective mode in the imperfect Fermi gas has a linear function in q .

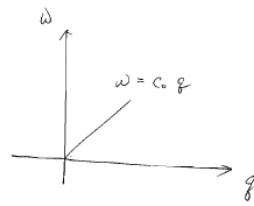
$$\omega = c_0 q \quad (3.2.15)$$

where $c_0 \approx \sqrt{\frac{k_F^3 V(0)}{3\pi^2 m}}$ ("strong coupling" expression)

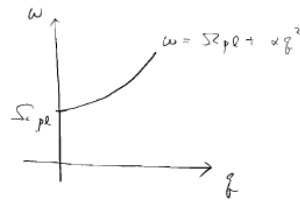
This gapless collective mode in imperfect Fermi gas is called zero sound mode.

¹ $1 = \frac{4\pi n e^2}{m\Omega_{\mathbf{q}}^2} \left[1 + \frac{3}{5} \left(\frac{\hbar k_F q}{m\Omega_{\mathbf{q}}} \right)^2 + \dots \right]$

3.2 Collective Modes in degenerate electron gas and imperfect Fermi gas – plasma oscillation & zero sound mode 87



imperfect Fermi gas
(short-ranged int)



degenerate electron gas
(long-ranged Coulomb int)

Chapter 4

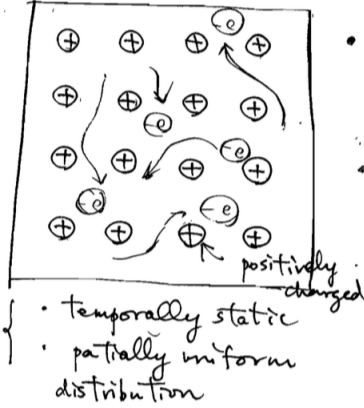
Field theory at finite temperature

- 4.1 Temperature (Matsubara) Green's function
- 4.2 Temperature Green's function in the interaction picture
- 4.3 Wick's theorem for the temperature Green's function
- 4.4 Feynman's rule for the temperature Green's function
- 4.5 Dyson's equation & Hartree-Fock approximation
- 4.6 Specific heat of an imperfect Fermi gas at low-temperature
- 4.7 Real-time Green's function and generalized Lehmann representation
- 4.8 Linear response at finite temperature
- 4.9 Screening effect & plasma oscillation in an electron gas at finite (or high) temperature

Chapter 5

Application to electron-phonon systems

5.1 Non-interacting phonon system and Debye theory of specific heat



In a usual metal, such as alkali metals of column I, outer shell valence electrons are completely separated from their ion cores and form a nearly free electron gas.

In the section 1-8, we have treated the positively charged ion cores as temporally-static and uniformly-distributed positively-charged background.

Thereby, the only play role of ensuring the electrical neutrality of the whole system.

In this section, we will study the dynamics of the positively charged backgrounds, and their effect on the electron gas.

In real metals, the ionic background form a lattice structure, so that it is not exactly the uniformly distributed positive charges.

But, when it comes to the acoustic mode, in which all the ion cores are moving in the same 'phase', the discrete lattice structure is irrelevant and we can treat them as a positively

charged continuous elastic media.

To describe the acoustic wave, let us first define a three-dimensional vector field \mathbf{d} as a function of the spatial coordinate \mathbf{x} , which we call displacement vector:

$\mathbf{d}(\mathbf{x})$: displacement vector

Physically, $\mathbf{d}(\mathbf{x})$ denotes how the ionic core at position \mathbf{x} is displaced:

$$\mathbf{u}(\mathbf{x}) - \langle \mathbf{u}(\mathbf{x}) \rangle \equiv \mathbf{d}(\mathbf{x})$$

Compared to the displacement field φ introduced in section 1-4, such a quantity is characterized by three component real-valued vector in a 3-dimensional elastic media

$$\mathbf{d}(\mathbf{x}) = (d_1(\mathbf{x}), d_2(\mathbf{x}), d_3(\mathbf{x}))$$

According to the Helmholtz-Hodge's decomposition theorem, any vector field in 3-dimensional space can be decomposed into divergence-free vector and a rotation-free vector

$$\mathbf{d}(\mathbf{x}) = \nabla\phi + \nabla \times \vec{A}$$

Although we always have ambiguity when decomposing into the two, let us temporally denote the rotation-free part as $\mathbf{d}'(\mathbf{x})$ and the divergence-free parts as $\mathbf{d}''(\mathbf{x})$ so that

$$\begin{cases} \nabla \times \mathbf{d}' = 0 \\ \nabla \cdot \mathbf{d}'' = 0 \end{cases}$$

Corresponding to this decomposition, the acoustic waves are composed of two different kinds.

One is called the longitudinal acoustic wave, which is described only by the rotation-free vector field \mathbf{d}' (To be precise, $\nabla \cdot \mathbf{d}' = \nabla \cdot \mathbf{d}$): longitudinal(compressional) acoustic wave.

The other is called the transverse acoustic waves, which are described only by the divergence-free part \mathbf{d}'' (to be precise, $\nabla \times \mathbf{d}'' = \nabla \times \mathbf{d}$): transverse(shear) acoustic waves.

$\nabla \cdot \vec{d}$ is nothing but the charge density modulation in a positively charged background

$$\nabla \cdot \vec{d} = -\frac{\delta\rho_{ionic}}{\rho_0}$$

One can easily see this, by drawing a configuration of the displacement vector, which has a finite divergence.

Namely, when these ionic charges are displaced like in this picture, we have effectively a negative charge in the middle.

On the one hand, $\nabla \times \mathbf{d}$ is accompanied by the shear deformation in the positively charged ionic background, e.g., which is not accompanied by any charge density modulation.

Since electrons and positively charged background are coupled with each other by electrostatic Coulomb interaction, we usually regard that the longitudinal acoustic waves are strongly coupled with electrons, while the transverse waves are not.

For this reasoning, we will focus only on the longitudinal (compressional) acoustic wave, where we enforce the displacement vector is rotation-free

$$\nabla \times \vec{d}(\mathbf{x}, t) = 0, \forall \mathbf{x}, t$$

In order to introduce the equation of motion for the charge density modulation

$$\nabla \cdot \vec{d}(\mathbf{x}, t) = \delta\rho_{ion}(\mathbf{x}, t)$$

Let us begin with the Newton's equation of motion for a fluid element

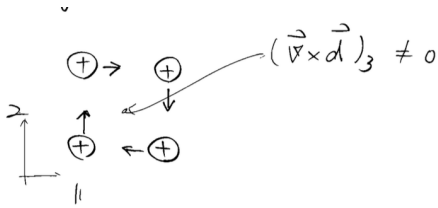
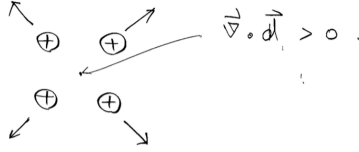
$$\frac{D\vec{p}}{Dt} = \vec{f}$$

Here $\mathbf{p}(\mathbf{x}, t)$ is the momentum density field at position \mathbf{x} at a given time t which is given by a product between mass density field and velocity field

$$\mathbf{p}(\mathbf{x}, t) = \rho(\mathbf{x}, t)\mathbf{v}(\mathbf{x}, t)$$

The temporal derivative in the right hand side is a Legendre derivative often used in the course of fluid mechanics.

$$\frac{D\mathbf{p}}{Dt} \stackrel{def}{=} \lim_{\Delta t \rightarrow 0} \frac{1}{\Delta t} \{\mathbf{p}(\mathbf{x} + \mathbf{v}\Delta t, t + \Delta t) - \mathbf{p}(\mathbf{x}, t)\} = \frac{\partial \mathbf{p}}{\partial t} + (\mathbf{v} \cdot \nabla)\mathbf{p}$$



The force acting on the fluid element is a restoring force which comes from the spatial gradient of the pressure:

$$\frac{\partial}{\partial t}(\rho \mathbf{v}) + (\mathbf{v} \cdot \nabla)(\rho \mathbf{v}) = -\nabla P$$

In the equilibrium, we don't have any velocity field and the mass density is uniformly distributed and static

$$\left. \begin{aligned} \rho(\mathbf{x}, t) &= \rho_0 \\ \mathbf{v}(\mathbf{x}, t) &= 0 \end{aligned} \right\} \text{equilibrium}$$

Out of the equilibrium, we have finite velocity field and mass density is spatially and temporally dependent:

$$\left. \begin{aligned} \rho(\mathbf{x}, t) &= \rho_0 + \delta\rho(\mathbf{x}, t) \\ \mathbf{v}(\mathbf{x}, t) &\neq 0 \end{aligned} \right\} \text{out of equilibrium}$$

When linearizing this equation of motion for a small derivation from the equilibrium, we can regard that the 2nd term in the right hand side is the 2nd order in $\delta\rho$ and thus negligible:

$$\frac{\partial}{\partial t}((\rho_0 + \delta\rho)\mathbf{v}) + (\mathbf{v} \cdot \nabla)((\rho_0 + \delta\rho)\mathbf{v}) = -\nabla P$$

$$\rho_0 \frac{\partial \mathbf{v}}{\partial t} + \textcolor{red}{O((\delta\rho)^2)} = -\nabla p$$

¹
In addition to this Newton's equation of motion, the mass density field and velocity field must satisfy the continuity equation:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

When linearized with respect to a small $\delta\rho$, we have

$$\frac{\partial \delta\rho}{\partial t} = -\rho_0(\nabla \cdot \mathbf{v})$$

Combining these two linearized equation, we have

$$\frac{\partial^2 \delta\rho}{\partial t^2} = -\rho_0(\nabla \cdot \frac{\partial \vec{v}}{\partial t}) = \nabla^2 P$$

A pressure of a given fluid element is determined by the mass density of the element, so that it is quite natural to assume that the pressure depends on the spatial coordinate \mathbf{x} only through ρ .

$$\nabla^2 P = \nabla^2 P(\rho_0 + \delta\rho(\mathbf{x}, t), \dots) \approx \left(\frac{\partial P}{\partial \rho} \right)_{\rho_0} \nabla^2 \delta\rho$$

This leads to the following closed equation of motion for the density fluctuation:

$$\frac{\partial^2 \delta\rho}{\partial t^2} = \left(\frac{\partial P}{\partial \rho} \right)_{\rho_0} \nabla^2 \delta\rho$$

For many elastic continuous media, we can define (adiabatic) bulk modulus.

¹The red part is omit-able.

$$\begin{aligned}
B &= -B \left(\frac{\partial O}{\partial V} \right)_{equilibrium} \\
\rho &= \frac{M}{V}, \quad \frac{d\rho}{dV} = M \left(-\frac{1}{V^2} \right) \\
&= -V \left(\frac{\partial P}{\partial \rho} \right) M \left(-\frac{1}{V^2} \right) \\
&= \frac{M}{V} \left(\frac{\partial P}{\partial \rho} \right)_{eq} \stackrel{def}{=} \rho_m \left(\frac{\partial P}{\partial \rho} \right)_{eq}
\end{aligned}$$

In terms of the bulk modulus, the coefficient of the right hand side is given by

$$\frac{\partial^2 \delta \rho}{\partial t^2} = \frac{B}{\rho_m} \cdot \nabla^2 \delta \rho \equiv c^2 \nabla^2 \delta \rho$$

Here c is nothing but a velocity, with which the density fluctuation propagates in space as time passes by, and which is often called as the “sound velocity” of the longitudinal acoustic wave.

As explained already, the density fluctuation is given by the divergence of the displacement field, the equation of motion for the displacement field is given by

$$\begin{aligned}
\frac{\partial}{\partial t} (\nabla \cdot \vec{d}) &= c^2 \nabla^2 (\nabla \cdot \vec{d}) \\
\nabla \cdot \left(\frac{\partial \vec{d}}{\partial t} - \nabla^2 \vec{d} \right) &= 0
\end{aligned}$$

Remember again that we are interested only in the longitudinal acoustic wave, so that the displacement field is rotation free at any space and time.

$$\nabla \times \vec{d} \equiv \vec{0}$$

The latter condition ensures that

$$\nabla \times \left(\frac{\partial \vec{d}}{\partial t} - \nabla^2 \vec{d} \right) = 0$$

because $\nabla \times$ and ∂_t or ∇^2 commute with each other.

From these two, we have a pair of equation which is satisfies by the displacement field \vec{d} .

$$\begin{cases} \frac{\partial \vec{d}}{\partial t} - c^2 \nabla^2 \vec{d} = 0 \\ \nabla \times \vec{d} = 0 \end{cases}$$

So far, we just describing a classical equation of motion for the displacement field; in order to second-quantize this classical field, let us first introduce the Lagrangian which leads to this equation of motion:

$$L_0 = \frac{1}{2} \sum_{i,j=1,2,3} \int d^3 \mathbf{x} \int dt \left[\rho_m \frac{\partial d_i}{\partial t} \cdot \frac{\partial d_i}{\partial t} - B \frac{\partial d_i}{\partial x_j} \frac{\partial d_i}{\partial x_j} \right]$$

Notice that

$$\int \left[\rho_m \frac{\partial d_i}{\partial t} \cdot \frac{\partial d_i}{\partial t} - B \frac{\partial d_i}{\partial x_j} \frac{\partial d_i}{\partial x_j} \right] = \delta L_0$$

$$\rho_m \delta d_i \left[\frac{\partial^2 d_i}{\partial t^2} - c^2 \sum_{j=1,2,3} \frac{\partial^2 d_i}{\partial x_j^2} \right] = \delta L_0$$

Under the 2nd condition of rotation-free, the Lagrangian can be rewritten into the following:

$$L_0 = \frac{1}{2} \sum_{i,j} \int d^3 \mathbf{x} \int dt \left[\rho_m \frac{\partial d_i}{\partial t} \cdot \frac{\partial d_i}{\partial t} - B \frac{\partial d_i}{\partial x_i} \frac{\partial d_j}{\partial x_j} \right] (= \int dt \tilde{L}_0)$$

To see this, let us begin with the following identity:

5.2 Electron-phonon interaction

Chapter 6

Application to superconductivity

In this chapter, we will study superconductivity which is one of the most remarkable physical phenomenon.

We first review the fundamental physical properties of superconductors.

We then introduce the BCS theory, which is one of the most successful application of quantum many-body techniques.

In terms of the BCS theory, we will see how all the fundamental superconducting properties can be understood from the microscopic point of views.

6.1 Superconducting properties and thermodynamics relations

Resistivity in metals decreases when the temperature goes down (See Fig6.1).

The temperature dependence always exhibits T^2 behavior while the resistivity at the zero temperature remains finite, which we call residual resistivity ρ_{res} .

This residual resistivity stems from the scattering of conduction electron due to randomly distributed impurity and lattice defects.

On the other hand, T^2 -dependence of the resistivity comes from the scattering of conduction electrons due to electron-phonon interaction and electron-electron interactions.

But, in some cases, metals exhibits an electronic phase transition at lower temperatures and resistivity becomes zero.

The low-temperature phase with zero resistivity is called superconductor, or “superconducting phase”.

Simple-substance metals that exhibits superconductivity at lower temperature or under high pressure are shown in the following periodic table, where “black-shaded” simple substance metals show superconductivity (at lower temperature, ■, while at lower temperature and under higher pressure, ▲)

You might as well regard that in a sufficiently clean metal, the resistivity goes to zero at the zero-temperature, which we call “perfect conductor” instead of “superconductor”, the electric conductivity diverges, so that the electric field \mathcal{E} vanish inside the materials

$$(\mathbf{j} = \sigma \mathbf{E})$$

Then, the Maxwell equation tells us that the magnetic field subsequently applied after the system reach the zero-temperature will remains zero inside the material.

$$\frac{\partial \mathbf{B}}{\partial t} = -c \operatorname{rot} \mathbf{E}$$

because, during the application of the external fields, the electric field \mathbf{E} remains zero inside the materials.

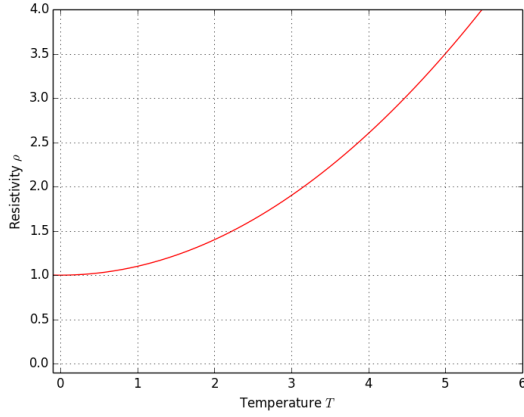


Figure 6.1: Resistivity ρ vs. Temperature T by $\rho = \rho_{res} + aT^2$ where ‘res’ denotes residual resistivity.

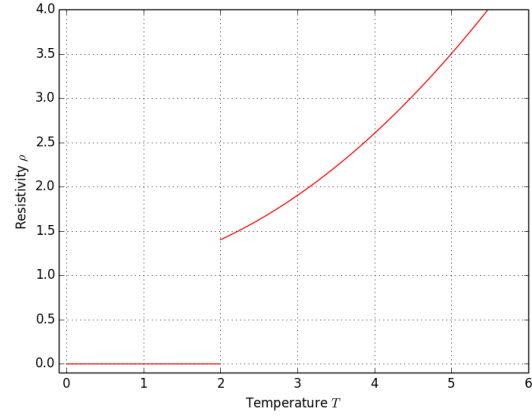


Figure 6.2: Resistivity of superconductor’s ρ vs. Temperature T by $\rho = \rho_{res} + aT^2$ where ‘res’ denotes residual resistivity.

Accordingly, both the perfect conductor at the zero-temperature and the superconductor “expels” the external magnetic field, when the fields are applied subsequently after the system reaches the zero-temperature.

Remember that this is the case when the field is applied, after the temperature reaches zero-temperature.

But, when we reduce the temperature under a finite external magnetic field, this is no longer the case for the “perfect conductor”, while, this is still the case for the “superconductor”.

Namely:

This exclusive of the magnetic field in the field-cooled superconductor is called as “Meissner effect”, which is more essential “defining property” of superconductors than “the zero-resistivity”.

The “Meissner effect” can be understood from a phenomenological equation called London equation:

$$\mathbf{j} = -\frac{ne^2}{mc} \mathbf{A}$$

where \mathbf{j} is the current density and \mathbf{A} is the vector potential.

This equation can be roughly understood by the following argument.

In the superconductor, the electric current flow without any resistivity, so that we might as well begin with the Newton’s equation without dissipation:

$$m\dot{\mathbf{r}} = -e\mathbf{E}$$

Now that the current density is proportional to the velocity of electron

$$\mathbf{j} = -ne\dot{\mathbf{r}}$$

we might as well naively relate the time-derivative of the current with the electric field:

$$\frac{d\mathbf{j}}{dt} = -\frac{ne^2}{m} \mathbf{E} = -\frac{ne^2}{mc} \frac{d\mathbf{A}}{dt}$$

Now that the electric field is given by the time-derivative of the vector potential, we might as well regard that the current is proportional to the vector potential

$$\mathbf{j} = -\frac{ne^2}{mc}\mathbf{A}$$

This London equation is, however, more or less a phenomenological equation and we will later see more microscopic derivation of this equation based on the BCS theory.

Thereby, we will see that this equation holds true for a so-called Cooper pair, which is composed of two electrons, so that, the electron number density “ n ” is replaced by superelectron number density “ n_s ”.

$$\mathbf{j} = -\rho \cdot 2e \cdot \dot{\mathbf{r}} = -n_s e \dot{\mathbf{r}} \quad \text{where } \rho \text{ is the number of cooper pairs.}$$

In any case, this single phenomenological equation captures many important electromagnetic properties including Meissner effect.

For example, the Meissner effect can be derived from the London equation as follows:

From the Maxwell equation:

$$\text{rot}\mathbf{B} = \frac{1}{c} \frac{\partial \mathbf{E}}{\partial t} + \frac{4\pi}{c} \mathbf{j}$$

we consider the temporally static situation:

$$\text{rot}\mathbf{B} = \frac{4\pi}{c} \mathbf{j} = -\frac{4\pi ne^2}{mc^2} \mathbf{A}^t$$

Now $\mathbf{B} = \text{rot}\mathbf{A}$, we have

$$\begin{aligned} \text{rot rot}\mathbf{B} &= -\frac{4\pi ne^2}{mc^2} \text{rot}\mathbf{A}^t \\ &= -\left(\nabla^2 \vec{\mathbf{B}} - \nabla(\nabla \cdot \vec{\mathbf{B}})\right) = -\frac{4\pi ne^2}{mc^2} \vec{\mathbf{b}} \\ \Rightarrow \nabla^2 \mathbf{B} &= \frac{4\pi ne^2}{mc^2} \mathbf{B} = \frac{1}{\lambda^2} \mathbf{B}, \quad \lambda \equiv \sqrt{\frac{mc^2}{4\pi ne^2}} \end{aligned} \quad (\text{A})$$

In order to solve this equation, let us consider a specific geometry of the superconducting material, so that we can define the boundary condition:

Consider that the system is the slab geometry which has two parallel planes boundaries at $x = -L/2$ and at $x = L/2$. For simplicity, we assume that the system is infinitely large along y, z -direction. We suppose L to be very large.

We apply the external field along z direction:

$$\begin{aligned} \mathbf{B} \left(x < -\frac{L}{2} \right) &= B \hat{e}_z \\ \mathbf{B} \left(x > \frac{L}{2} \right) &= B \hat{e}_z \end{aligned}$$

Now that the tangential component of \mathbf{B} should be continuous at the two boundaries (at $x = \pm L/2$)

6.2 BCS theory, Cooper pairing and Gap equation

6.2.1 BCS(Bardeen Copper Schrieffer) Hamiltonian

Based on the experimental observation of the isotope effect, many people in 1950s had worked on the effect of electron-phonon interaction.

As we have seen in the Section 5.2, the electron-phonon coupled field theory is equivalent to the electron system with the attractive electron-electron interaction, so that we begin with the following simplified Hamiltonian called as the BCS Hamiltonian:

$$\hat{K} = \hat{K}_0 + \hat{V} = \int d^3\mathbf{x} \psi_\alpha^\dagger(\mathbf{x}) \left[\frac{1}{2m} \left(-i\hbar\nabla + \frac{e\mathbf{A}(\mathbf{x})}{c} \right)^2 - \mu \right] \psi_\alpha(\mathbf{x}) - \frac{g}{2} \int d^3\mathbf{x} \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) \psi_\beta(\mathbf{x}) \psi_\alpha(\mathbf{x})$$

where $\mathbf{A}(\mathbf{x})$ is a vector potential.

Prior to the BCS theory, Cooper had considered two-electrons problem, which are introduced into a Fermi sea and which are interacting with each other via the attractive interaction.

Without the attractive interaction, the energy cost we need to pay for these creations is

$$\Delta E = \frac{\hbar k_1^2}{2m} + \frac{\hbar k_2^2}{2m} - 2\mu > 0$$

which is positive (so that the metal is stable).

Cooper showed that, once the small attractive interaction between these two are introduced, the energy “cost” for the creating this pair of fermions near the Fermi surface becomes “negative”, due to the binding energy.

between these two particles,

$$\Delta E = \frac{\hbar k_1^2}{2m} + \frac{\hbar k_2^2}{2m} - 2\mu < 0$$

This means that, once the attractive interaction are introduced near the Fermi surface, which is indeed the case with the “effective” BCS Hamiltonian, the Fermi sea (the ground state wave function for normal metal) is stable toward the formations of many pairs of fermions which are bound together by the attractive interaction.

However, Cooper also showed that a finite binding energy between the two fermions comes from the Pauli exclusion due to the background Fermi-sea state.

Therefore, the “true” ground state which we would obtain from the Fermi-sea state is NOT just a pure state of many bound pairs.

Instead, the true ground state is a linear superposition of those states having different number of the bound pairs.

$$|\Psi\rangle = |N - \text{particle}\rangle + |(N + 2) - \text{particle}\rangle + |(N + 4) - \text{particle}\rangle + \cdots + |(N + 2m) - \text{particle}\rangle + \cdots$$

(A)

Contrary to the ground state wavefunction for the normal phase, such a wave function gives a finite expectation value of the square of the creation fields:

$$\begin{aligned}
\langle \Psi | \psi^\dagger \psi^\dagger | \Psi \rangle = & \langle (N+2) - \text{particle} | \psi^\dagger \psi^\dagger | N - \text{particle} \rangle \\
& + \langle (N+4) - \text{particle} | \psi^\dagger \psi^\dagger | (N+2) - \text{particle} \rangle \\
& + \dots \neq 0
\end{aligned}$$

In fact, we will see that the BCS ground state wavefunction is given by this type linear superposition where each of these multiparticle states have a phase coherence

$$|\Psi\rangle = e^{i\phi_N} |N - \text{particle}\rangle + e^{i\phi_{N+2}} |(N+2) - \text{particle}\rangle + e^{i\phi_{N+4}} |(N+4) - \text{particle}\rangle + \dots$$

where ϕ_M are not randomly distributed in M .

Based on this observation, BCS regards $\langle \psi^\dagger \psi^\dagger \rangle$ (or $\langle \psi \psi \rangle$) as an order-parameter for the superconductivity, and construct the Mean-field theory for $\langle \psi^\dagger \psi^\dagger \rangle$.

The BCS Hamiltonian reads

$$\begin{aligned}
\hat{K} = & \int d^3\mathbf{x} \psi_\alpha^\dagger(\mathbf{x}) \left(-\frac{\hbar^2 \nabla^2}{2m} - \mu \right) \psi_\alpha(\mathbf{x}) \\
& - \frac{1}{2} \int d^3\mathbf{x} \hat{\psi}_\alpha^\dagger(\mathbf{x}) \hat{\psi}_\beta^\dagger(\mathbf{x}) \hat{\psi}_\beta(\mathbf{x}) \hat{\psi}_\alpha(\mathbf{x})
\end{aligned}$$

where the interaction part can be rewritten into

$$\begin{aligned}
& \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) \psi_\beta(\mathbf{x}) \psi_\alpha(\mathbf{x}) \\
& = \left(\langle \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) \rangle + \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) - \langle \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) \rangle \right) \\
& \quad \times \left(\langle \psi_\beta(\mathbf{x}) \psi_\alpha(\mathbf{x}) \rangle + \psi_\beta(\mathbf{x}) \psi_\alpha(\mathbf{x}) - \langle \psi_\beta(\mathbf{x}) \psi_\alpha(\mathbf{x}) \rangle \right)
\end{aligned}$$

Suppose that the fluctuation from the mean-field $(\psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) - \langle \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) \rangle)$ is small; we ignore the quadratic in the fluctuation:

$$\begin{aligned}
& \left(\psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) - \langle \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) \rangle \right) \\
& \times (\psi_\alpha(\mathbf{x}) \psi_\beta(\mathbf{x}) - \langle \psi_\alpha(\mathbf{x}) \psi_\beta(\mathbf{x}) \rangle) \rightarrow 0
\end{aligned}$$

The other terms are summarized as

$$\begin{aligned}
& = \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) \langle \psi_\alpha(\mathbf{x}) \psi_\beta(\mathbf{x}) \rangle \\
& \quad + \langle \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) \rangle \psi_\alpha(\mathbf{x}) \psi_\beta(\mathbf{x}) - \langle \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) \rangle \langle \psi_\alpha(\mathbf{x}) \psi_\beta(\mathbf{x}) \rangle
\end{aligned}$$

from which we obtain a mean-field Hamiltonian

$$\begin{aligned}
\hat{K}_{eff} = & \hat{K}_0 - \frac{g}{2} \int d^3\mathbf{x} \left\{ \langle \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) \rangle \psi_\beta(\mathbf{x}) \psi_\alpha(\mathbf{x}) + \psi_\alpha^\dagger(\mathbf{x}) \psi_\beta^\dagger(\mathbf{x}) \langle \psi_\beta(\mathbf{x}) \psi_\alpha(\mathbf{x}) \rangle \right\} + \text{const} \\
= & \hat{K}_0 - \frac{g}{2} \int d^3\mathbf{x} \left\{ \langle \psi_\uparrow^\dagger(\mathbf{x}) \psi_\downarrow^\dagger(\mathbf{x}) \rangle \psi_\downarrow(\mathbf{x}) \psi_\uparrow(\mathbf{x}) + \langle \psi_\downarrow(\mathbf{x}) \psi_\uparrow(\mathbf{x}) \rangle \psi_\uparrow^\dagger(\mathbf{x}) \psi_\downarrow^\dagger(\mathbf{x}) \right\} + \text{const} \\
= & \hat{K}_0 - g \int d^3\mathbf{x} \left\{ \langle \psi_\downarrow(\mathbf{x}) \psi_\uparrow(\mathbf{x}) \rangle \psi_\uparrow^\dagger(\mathbf{x}) \psi_\downarrow^\dagger(\mathbf{x}) + \psi_\downarrow(\mathbf{x}) \psi_\uparrow(\mathbf{x}) \langle \psi_\uparrow^\dagger(\mathbf{x}) \psi_\downarrow^\dagger(\mathbf{x}) \rangle \right\} \\
& + g \int d^3\mathbf{x} \langle \psi_\downarrow^\dagger(\mathbf{x}) \psi_\uparrow^\dagger(\mathbf{x}) \rangle \langle \psi_\uparrow(\mathbf{x}) \psi_\downarrow(\mathbf{x}) \rangle
\end{aligned}$$

In order to make the theory self-consistent, we assume that the mean-field is given by the expectation value of the product of the creation fields, calculated with respect to the ground state wavefunction for mean-field Hamiltonian:

$$\langle \psi_{\downarrow}^{\dagger}(\mathbf{x}) \psi_{\uparrow}^{\dagger}(\mathbf{x}) \rangle \equiv \langle \Phi_0 | \psi_{\uparrow}^{\dagger}(\mathbf{x}) \psi_{\downarrow}^{\dagger}(\mathbf{x}) | \Phi_0 \rangle$$

where $|\Phi_0\rangle$ is the ground state w.f for \hat{K}_{eff} .

$$\begin{aligned} \hat{K}_{eff} = & \int d^3\mathbf{x} \psi_{\alpha}^{\dagger}(\mathbf{x}) \left(-\frac{\hbar^2 \nabla_{\mathbf{x}}^2}{2m} - \mu \right) \psi_{\alpha}(\mathbf{x}) \\ & - g \int d^3\mathbf{x} \{ \langle \psi_{\downarrow}^{\dagger}(\mathbf{x}) \psi_{\uparrow}^{\dagger}(\mathbf{x}) \rangle \psi_{\uparrow}(\mathbf{x}) \psi_{\downarrow}(\mathbf{x}) \\ & + \psi_{\downarrow}^{\dagger}(\mathbf{x}) \psi_{\uparrow}^{\dagger}(\mathbf{x}) \langle \psi_{\uparrow}^{\dagger}(\mathbf{x}) \psi_{\downarrow}^{\dagger}(\mathbf{x}) \rangle \} \\ & + g \int d^3\mathbf{x} \langle \psi_{\downarrow}^{\dagger}(\mathbf{x}) \psi_{\uparrow}^{\dagger}(\mathbf{x}) \rangle \underbrace{\langle \psi_{\uparrow}(\mathbf{x}) \psi_{\downarrow}(\mathbf{x}) \rangle}_{-\Delta(\mathbf{x})} \end{aligned}$$

Provided that the ground state wavefunction does not break the translational symmetry spontaneously, we can regard that the mean-field does not depend on the space coordinate.

$$g \langle \psi_{\downarrow}^{\dagger}(\mathbf{x}) \psi_{\uparrow}^{\dagger}(\mathbf{x}) \rangle \equiv -\Delta^* \equiv |\Delta| e^{-i\theta}$$

With this in mind, the mean-field Hamiltonian is readily diagonalized:

$$\begin{aligned} \hat{K}_{eff} = & \sum_{\mathbf{k}, \alpha} \left(\frac{\xi_{\mathbf{k}}}{2m} - \mu \right) c_{\mathbf{k}, \alpha}^{\dagger} c_{\mathbf{k}, \alpha} \\ & + \Delta^* \sum_{\mathbf{k}} c_{\mathbf{k}\uparrow}^{\dagger} c_{-\mathbf{k}\downarrow} + \Delta \sum_{\mathbf{k}} c_{-\mathbf{k}\downarrow}^{\dagger} c_{\mathbf{k}\uparrow}^{\dagger} - \Delta \sum_{\mathbf{k}} \langle c_{-\mathbf{k}\downarrow}^{\dagger} c_{\mathbf{k}\uparrow}^{\dagger} \rangle \\ = & \sum_{\mathbf{k}} \left\{ (c_{\mathbf{k}\uparrow}^{\dagger} \ c_{-\mathbf{k}\downarrow}) \begin{pmatrix} \xi_{\mathbf{k}} & -e^{i\theta} |\Delta| \\ -e^{-i\theta} |\Delta| & -\xi_{\mathbf{k}} \end{pmatrix} \begin{pmatrix} c_{\mathbf{k}\uparrow}^{\dagger} \\ c_{-\mathbf{k}\downarrow}^{\dagger} \end{pmatrix} + \xi_{\mathbf{k}} \right\} - \Delta \sum_{\mathbf{k}} \langle c_{-\mathbf{k}\downarrow}^{\dagger} c_{\mathbf{k}\uparrow}^{\dagger} \rangle \end{aligned}$$

A unitary transformation gives

$$\begin{pmatrix} c_{\mathbf{k}\uparrow}^{\dagger} \\ c_{-\mathbf{k}\downarrow}^{\dagger} \end{pmatrix} = \begin{pmatrix} u_{\mathbf{k}} & e^{i\theta} v_{\mathbf{k}} \\ -e^{-i\theta} v_{\mathbf{k}} & u_{\mathbf{k}} \end{pmatrix} \begin{pmatrix} \alpha_{\mathbf{k}} \\ \beta_{-\mathbf{k}}^{\dagger} \end{pmatrix}$$

with

$$\begin{cases} 2u_{\mathbf{k}}v_{\mathbf{k}} &= \frac{|\Delta|}{\sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2}} \\ u_{\mathbf{k}}^2 - v_{\mathbf{k}}^2 &= \frac{\xi_{\mathbf{k}}}{\sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2}} \end{cases}$$

Or equivalently

$$\begin{cases} u_{\mathbf{k}} &= \frac{1}{\sqrt{2}} \left(1 + \frac{\xi_{\mathbf{k}}}{\sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2}} \right)^{1/2} \\ v_{\mathbf{k}} &= \frac{1}{\sqrt{2}} \left(1 - \frac{\xi_{\mathbf{k}}}{\sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2}} \right)^{1/2} \end{cases}$$

In terms of this, we have

$$\begin{aligned} \hat{K}_{eff} &= \sum_{\mathbf{k}} \left\{ \sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2} (\alpha_{\mathbf{k}}^\dagger (\alpha_{\mathbf{k}} - \beta_{-\mathbf{k}} \beta_{-\mathbf{k}}^\dagger) + \xi_{\mathbf{k}}) \right\} - \Delta \sum_{\mathbf{k}} \langle c_{-\mathbf{k}\downarrow}^\dagger c_{\mathbf{k}\uparrow}^\dagger \rangle \\ &= V + \sum_{\mathbf{k}} \sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2} (\alpha_{\mathbf{k}}^\dagger \alpha_{\mathbf{k}} + \beta_{\mathbf{k}}^\dagger \beta_{\mathbf{k}}) \end{aligned}$$

where

$$V \equiv \sum_{\mathbf{k}} \{ \xi_{\mathbf{k}} - \sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2} - \Delta \langle c_{-\mathbf{k}\downarrow}^\dagger c_{\mathbf{k}\uparrow}^\dagger \rangle \}$$

Now that \hat{K}_{eff} is given by a quadrotic form in the creation fields $(\alpha_{\mathbf{k}}^\dagger, \beta_{\mathbf{k}}^\dagger)$ and annihilation fields $(\alpha_{\mathbf{k}}, \beta_{\mathbf{k}})$, the eigen state of \hat{K}_{eff} is given by a multi particle state of these new quastic particle.

Noting that these fields as well as the original fields $(c_{\mathbf{k},\sigma}, c_{\mathbf{k},\sigma}^\dagger)$ are fermion fields, the eigen state is given by

$$|(\mathbf{k}_1, \alpha_1), \dots, (\mathbf{k}_n, \beta_n)\rangle = \alpha_{\mathbf{k}_1}^\dagger \cdots \beta_{\mathbf{k}_n}^\dagger |\Phi_0\rangle$$

where $|\Phi_0\rangle$ denotes the vacuum of α -particle and β -particle (called as Bogoliubov particles).

$$\begin{cases} \alpha_{\mathbf{k}} |\Phi_0\rangle &= 0 \\ \beta_{\mathbf{k}} |\Phi_0\rangle &= 0 \end{cases}$$

Since creating the Bogoliubov particles is always accompanied by a positive energy $(\sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2} > 0)$, the vacuum is nothing but the ground state wavefunction of \hat{K}_{eff} , whose energy is V .

$$\hat{K}_{eff} |\Phi_0\rangle = \sum_{\mathbf{k}} \xi_{\mathbf{k}} - \sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2} - \Delta \langle c_{-\mathbf{k}\downarrow}^\dagger c_{\mathbf{k}\uparrow}^\dagger \rangle V |\Phi_0\rangle$$

A group of the lowest excited states is given by “one-particle” state of the Bogoliubov particle

$$\begin{aligned} \alpha_{\mathbf{k}}^\dagger |\Phi_0\rangle &= |(\mathbf{k}, \alpha)\rangle \\ \beta_{\mathbf{k}}^\dagger |\Phi_0\rangle &= |(\mathbf{k}, \beta)\rangle \end{aligned}$$

$$\hat{K}_{eff} \alpha_{\mathbf{k}}^\dagger |\Phi_0\rangle = (V + \sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2}) \alpha_{\mathbf{k}}^\dagger |\Phi_0\rangle$$

which is a gapped excitation (See Fig. 6.3) because of $\sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2} \geq |\Delta|$.

This gapped behaviour of the one-particle excitation of the Bogoliubov particle is the origin of the thermal activation form of the low-temperature specific heat in superconductors. (see later)

To see the mean-field ground state wavefunction that introduced indeed take the form of eq. A, let us construct the $|\Phi_0\rangle$ in terms of the original (electron) creation and annihilation operators:

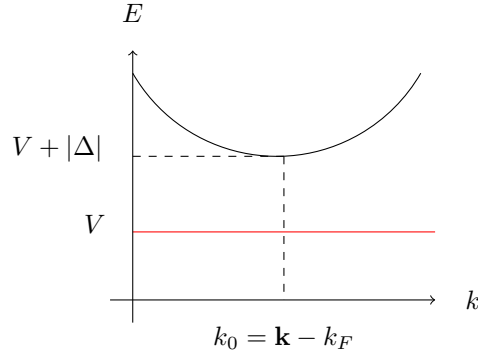


Figure 6.3: Excited spectrum; red line for ground BCS state, and solid line for single excited state; k_0 stands for $\xi_{k_0} = 0$.

Now that $\alpha_{\mathbf{k}}$ and $\beta_{\mathbf{k}}$ are given by a linear combination of $c_{\mathbf{k}\uparrow}$ and $c_{-\mathbf{k}\downarrow}^\dagger$

$$\begin{pmatrix} \alpha_{\mathbf{k}} \\ \beta_{-\mathbf{k}}^\dagger \end{pmatrix} = \begin{pmatrix} u_{\mathbf{k}} & -e^{i\theta} v_{\mathbf{k}} \\ e^{-i\theta} v_{\mathbf{k}} & u_{\mathbf{k}} \end{pmatrix} \begin{pmatrix} c_{\mathbf{k}\uparrow} \\ c_{-\mathbf{k}\downarrow}^\dagger \end{pmatrix}$$

or

$$\alpha_{\mathbf{k}} = u_{\mathbf{k}} c_{\mathbf{k}\uparrow} - e^{i\theta} v_{\mathbf{k}} c_{-\mathbf{k}\downarrow}^\dagger$$

$$\beta_{-\mathbf{k}} = e^{i\theta} v_{\mathbf{k}} c_{\mathbf{k}\uparrow}^\dagger + u_{\mathbf{k}} c_{-\mathbf{k}\downarrow}$$

$$(\beta_{\mathbf{k}} = e^{i\theta} v_{\mathbf{k}} c_{-\mathbf{k}\uparrow}^\dagger + u_{\mathbf{k}} c_{\mathbf{k}\downarrow})$$

This vacuum of such Bogoliubov particles given by: ¹

$$|\Phi_0\rangle = \prod_{\mathbf{k}} (u_{\mathbf{k}} + e^{i\theta} v_{\mathbf{k}} c_{\mathbf{k}\uparrow}^\dagger c_{-\mathbf{k}\downarrow}^\dagger) |vac\rangle$$

where $|vac\rangle$ denotes the vacuum of the original electrons ($c_{\mathbf{k}\sigma} |vac\rangle = 0$ for $\forall \mathbf{k}, \sigma$) and the product over \mathbf{k} is taken over all momentum.

In fact, one can see that

$$\begin{aligned} \alpha_{\mathbf{k}'} |\Phi_0\rangle &= \prod_{\mathbf{k} \neq \mathbf{k}'} (u_{\mathbf{k}} + e^{i\theta} v_{\mathbf{k}} c_{\mathbf{k}\uparrow}^\dagger c_{-\mathbf{k}\downarrow}^\dagger) \times (u_{\mathbf{k}'} c_{\mathbf{k}'\uparrow} - e^{i\theta} v_{\mathbf{k}'} c_{-\mathbf{k}'\downarrow}^\dagger) (u_{\mathbf{k}'} + e^{i\theta} v_{\mathbf{k}'} c_{\mathbf{k}'\uparrow}^\dagger c_{-\mathbf{k}'\downarrow}^\dagger) |vac\rangle \\ &= \prod_{\mathbf{k} \neq \mathbf{k}'} (u_{\mathbf{k}} + e^{i\theta} v_{\mathbf{k}} c_{\mathbf{k}\uparrow}^\dagger c_{-\mathbf{k}\downarrow}^\dagger) \\ &\quad \times (u_{\mathbf{k}'}^2 c_{\mathbf{k}'\uparrow} + e^{i\theta} u_{\mathbf{k}'} v_{\mathbf{k}'} (1 - c_{\mathbf{k}'\uparrow}^\dagger c_{\mathbf{k}'\uparrow}) c_{-\mathbf{k}'\downarrow}^\dagger \\ &\quad - e^{i\theta} u_{\mathbf{k}'} v_{\mathbf{k}'} c_{-\mathbf{k}'\downarrow}^\dagger + e^{i\theta} v_{\mathbf{k}'}^2 c_{\mathbf{k}'\uparrow}^\dagger (c_{-\mathbf{k}'\downarrow}^\dagger)^2) |vac\rangle \\ &= 0 \end{aligned}$$

Similar for applying $\beta_{\mathbf{k}'}$ on this and obviously derive a 0 as well.

Then, the many body wavefunction is given by “a linear superposition of those states having different but even numbers of electrons”.

¹This solution is derived based on guess, not rigorous calculation. The following shows some check for this solution.

$$|\Phi_0\rangle = |vac\rangle + |2\text{-particle}\rangle + |4\text{-particle}\rangle + \dots$$

so that the expectation value of the product of two creation fields can be finite.

To calculate $\langle\psi_\downarrow(\mathbf{x})^\dagger\psi_\downarrow(\mathbf{x})^\dagger\rangle$, notice first that $|\Phi_0\rangle$ is already normalized

$$\begin{aligned}\langle\Phi_0|\Phi_0\rangle &= \langle vac|\prod_{\mathbf{k}}\{(u_{\mathbf{k}} + e^{-i\theta}v_{\mathbf{k}}c_{-\mathbf{k}\downarrow}c_{\mathbf{k}\uparrow})(u_{\mathbf{k}} + e^{i\theta}v_{\mathbf{k}}c_{\mathbf{k}\uparrow}^\dagger c_{-\mathbf{k}\downarrow}^\dagger)\}|vac\rangle \\ &= \langle vac|\prod_{\mathbf{k}}\{u_{\mathbf{k}}^2 + v_{\mathbf{k}}^2 c_{\mathbf{k}\uparrow}c_{-\mathbf{k}\downarrow}c_{\mathbf{k}\uparrow}^\dagger c_{-\mathbf{k}\downarrow}^\dagger\}|vac\rangle \\ &\quad = 1 - c_{\mathbf{k}\uparrow}^\dagger c_{\mathbf{k}\uparrow} \\ &= \langle vac|vac\rangle = 1\end{aligned}$$

Thus the mean-field is readily calculated as follows:

$$\begin{aligned}\langle\psi_\downarrow(\mathbf{x})^\dagger\psi_\uparrow(\mathbf{x})^\dagger\rangle &= \frac{1}{V}\int d^3\mathbf{x}\langle\psi_\downarrow(\mathbf{x})^\dagger\psi_\uparrow(\mathbf{x})^\dagger\rangle = \frac{1}{V}\sum_{\mathbf{k}}\langle\Phi_0|c_{-\mathbf{k}\downarrow}^\dagger c_{\mathbf{k}\uparrow}^\dagger|\Phi_0\rangle \\ &\quad \text{left hand side is independent of } \mathbf{x} \\ &= \frac{1}{V}\sum_{\mathbf{k}}\langle vac|(u_{\mathbf{k}} + e^{-i\theta}v_{\mathbf{k}}c_{-\mathbf{k}\downarrow}c_{\mathbf{k}\uparrow})c_{-\mathbf{k}\downarrow}^\dagger c_{\mathbf{k}\uparrow}^\dagger(u_{\mathbf{k}} + e^{i\theta}v_{\mathbf{k}}c_{\mathbf{k}\uparrow}^\dagger c_{-\mathbf{k}\downarrow}^\dagger)|vac\rangle \\ &= -\frac{1}{V}e^{-i\theta}\sum_{\mathbf{k}}v_{\mathbf{k}}u_{\mathbf{k}}\end{aligned}$$

Multiply g and substitute the expressions for $u_{\mathbf{k}}, v_{\mathbf{k}}$ in terms of Δ , we obtain

$$|\Delta| = \frac{g}{2}\int \frac{d^3\mathbf{k}}{(2\pi)^3} \frac{|\Delta|}{\sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2}}$$

which is called a “gap equation”, determining $|\Delta|$.

Remember that the attraction interaction in the BCS Hamiltonian applies for those two electrons which are lying within an energy shell of the thickness $\hbar\omega_D$ from the Fermi surface (see previous argument).

This restriction tells us that \hat{K}_{eff} with finite $|\Delta|$ holds true only for those momentum \mathbf{k} , which satisfy

$$|\xi_{\mathbf{k}}| \equiv \left|\frac{\hbar^2\mathbf{k}^2}{2m} - \mu\right| < \hbar\omega_D$$

For the other \mathbf{k} , the off-diagonal term connecting $c_{\mathbf{k}\uparrow}$ and $c_{-\mathbf{k}\downarrow}$ should be set to zero:

$$\hat{K}_{eff} = \sum_{\mathbf{k},\alpha}\xi_{\mathbf{k}}c_{\mathbf{k}\alpha}^\dagger c_{\mathbf{k}\alpha} + \Delta^* \sum_{\mathbf{k},|\xi_{\mathbf{k}}|<\hbar\omega_D} c_{\mathbf{k}\uparrow}c_{-\mathbf{k}\downarrow} + \Delta \sum_{\mathbf{k},|\xi_{\mathbf{k}}|<\hbar\omega_D} c_{-\mathbf{k}\downarrow}^\dagger c_{\mathbf{k}\uparrow}^\dagger$$

which leads to

$$\begin{cases} (u_{\mathbf{k}}, v_{\mathbf{k}}) = (1, 0) & \text{for } \xi_{\mathbf{k}} > \hbar\omega_D \\ (u_{\mathbf{k}}, v_{\mathbf{k}}) = (0, 1) & \text{for } \xi_{\mathbf{k}} < -\hbar\omega_D \end{cases}$$

Thus, the right hand side acquires a cut off:

$$|\Delta| = \frac{g}{2}\int_{|\xi_{\mathbf{k}}|\leq\hbar\omega_D} \frac{d^3\mathbf{k}}{(2\pi)^3} \frac{|\Delta|}{\sqrt{\xi_{\mathbf{k}}^2 + |\Delta|^2}}$$

6.3 Electromagnetic property of superconductors**6.4 Microscopic derivation of the Ginzburg Landau equations**

Work Distribution

SiZhen-Wang works for chapter 2 section 6 to 8 and chapter 3. ZeYang-Li works for chapter 1, chapter 5 section 1 and chapter 6.