John Clouse IMD HW5 problem 3

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Initialize

```
clearvars -except hw_pub function_list
close all;
```

Initial conditions

```
IC\_set(:,1) = [
   -0.08
-0.03
0.01
3.5
-3.1
-0.1
26];
IC\_set(:,2) = [0.05]
-0.05
0
4.0
2.6
251;
IC\_set(:,3) = [0.8300]
0.114062816271683
0.229389507175582
15];
IC\_set(:,4) = [-0.05]
-0.02
4.09
-5.27
15];
% Constants
```

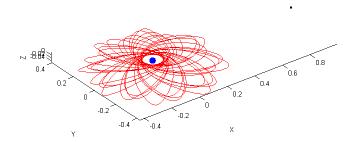
```
mu = 0.012150585609624;

dunit = 384747.962856037;
```

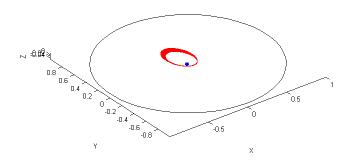
Loop through the conditions

```
for ii = 1:4
   X = IC set(1:end-1,ii);
   T = 6.192169331319632;
    [T_out,X_out] = ode45(@CRTBP, [0,IC_set(end,ii)], X, odeset(),mu);
   figure('Position', hw_pub.figPosn)
   subplot(2,1,1);
   plot3(X_out(:,1), X_out(:,2), X_out(:,3), 'r')
   hold on
   rad_vec = [0:0.1:2*pi, 2*pi];
   my_circ = [cos(rad_vec); zeros(1, length(rad_vec)); sin(rad_vec)]';
   for ang = rad_vec
        for blah = 1:length(my_circ)
            new_circ(blah,:) = (Euler2DCM('3', ang)*my_circ(blah,:)')';
        end
        earth = new circ * 6378.1/dunit;
       moon = (new_circ * 1737/dunit);
       plot3(earth(:,1) - mu, earth(:,2), earth(:,3))
       plot3(moon(:,1) + 1-mu, moon(:,2), moon(:,3), 'k')
    end
   axis equal; xlabel('X'); ylabel('Y'); zlabel('Z');
   title(['IC ' num2str(ii) ', Rotating Frame'])
   % For the inertial plots
   X_inrt = X_out;
   X_{inrt}(:,1) = X_{inrt}(:,1) + mu;
   for jj = 1:length(T_out)
       t = T_out(jj);
        ang = t;
        X_{inrt(jj,1:3)} = (Euler2DCM('3', -ang)*X_{inrt(jj,1:3)')';
   end
   subplot(2,1,2);
   plot3(X_inrt(:,1), X_inrt(:,2), X_inrt(:,3), 'r')
   hold on
   for ang = rad_vec
   for blah = 1:length(my_circ)
       new_circ(blah,:) = (Euler2DCM('3', ang)*my_circ(blah,:)')';
   end
   earth = new_circ * 6378.1/dunit;
   plot3(earth(:,1) - mu, earth(:,2), earth(:,3))
   end
   plot3(my_circ(:,1), my_circ(:,3), my_circ(:,2), 'k')
   axis equal; xlabel('X'); ylabel('Y'); zlabel('Z');
   title(['IC ' num2str(ii) ', Inertial Frame'])
end
```

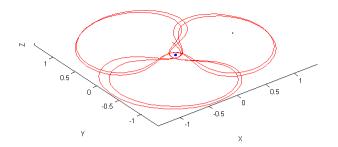
IC 1, Rotating Frame



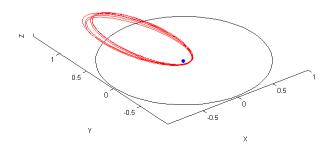
IC 1, Inertial Frame



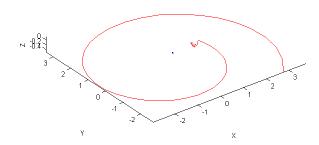
IC 2, Rotating Frame



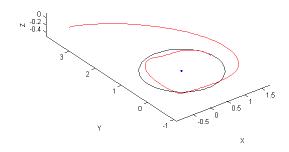
IC 2, Inertial Frame



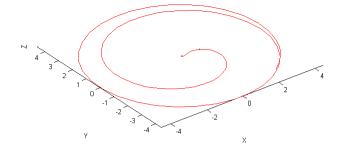
IC 3, Rotating Frame



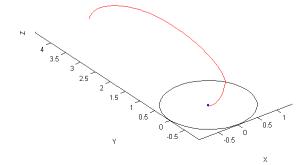
IC 3, Inertial Frame



IC 4, Rotating Frame



IC 4, Inertial Frame



Plot 1 Conclusion

For the first case, a two-body, point-mass propagation would yeild an elliptical (0<e<1) orbit that would follow previous orbit passes exactly, not changing any of the Keplerian orbital elements. The lunar perturbation is evident in the raising of the apogee as time progresses.

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