

# Quantum Mechanics II

## Assignment II

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### 1 Problem 7.11

Find the lowest bound on the ground state of hydrogen you can get using a gaussian trial wave function

$$\psi(\mathbf{r}) = Ae^{-br^2}, \quad (1.1)$$

where  $A$  is determined by normalisation and  $b$  is an adjustable parameter. *Answer:* -11.5 eV.

#### Solution:

First of all, let's write down the Hamiltonian for the Hydrogen

$$H_{Bohr} = -\frac{\hbar^2}{2m}\nabla^2 - \frac{e^2}{4\pi\epsilon_0}\frac{1}{r} \quad (1.2)$$

Now, we have to follow the next steps to address this problem

1. Normalise our trial wavefunction:

$$\begin{aligned} 1 = \langle \psi | \psi \rangle &= |A|^2 \int e^{-2br^2} r^2 \sin\theta dr d\theta d\phi \\ &= |A|^2 \int_0^\infty e^{-2br^2} r^2 dr \underbrace{\int_0^\pi \sin\theta d\theta \int_0^{2\pi} d\phi}_{4\pi} \\ &= |A|^2 4\pi \int_0^\infty e^{-2br^2} r^2 dr \end{aligned}$$

Using

$$\boxed{\int_0^{+\infty} r^m e^{-ar^2} dr = \frac{\Gamma\left(\frac{m+1}{2}\right)}{2a^{\frac{m+1}{2}}}} \quad (1.3)$$

$$\begin{aligned} 1 &= 4\pi |A|^2 \frac{1}{4} \sqrt{\frac{\pi}{(2b)^3}} = |A|^2 \left(\frac{\pi}{2b}\right)^{3/2} \\ \Rightarrow A &= \left(\frac{2b}{\pi}\right)^{3/4} \end{aligned} \quad (1.4)$$

2. Compute  $\langle \psi | \hat{T} | \psi \rangle$

$$\langle \psi | \hat{T} | \psi \rangle = \int \psi \left( -\frac{\hbar^2}{2m} \nabla^2 \psi \right) dr \sin\theta d\theta d\phi \quad (1.5)$$

Firs we need to compute

$$\begin{aligned} \nabla^2 \psi(\mathbf{r}) &= \frac{1}{r^2} \frac{d}{dr} \left( r^2 \frac{d}{dr} \psi \right), \quad \text{no angular dependence!} \\ &= A \frac{1}{r^2} \frac{d}{dr} \left( -2b r^3 e^{-2br^2} \right) \\ &= -2b A \frac{1}{r^2} (3r^2 - 2r^4) e^{-2br^2} \\ &= \frac{-2b}{r^2} (3r^2 - 2r^4) \psi(\mathbf{r}) \end{aligned} \quad (1.6)$$

Then, Equation 1.5 becomes

$$\langle \psi | \hat{T} | \psi \rangle = -\frac{\hbar^2}{2m} |A|^2 (-2b) 4\pi \int_0^\infty (3r^2 - 2br^4) e^{-2br^2} dr \quad (1.7)$$

We can solve the integral using Equation 1.3, then

$$\begin{aligned} \langle \psi | \hat{T} | \psi \rangle &= \frac{4\hbar^2 \pi b}{m} \left( \frac{2b}{\pi} \right)^{3/2} \left[ 3 \frac{1}{8b} \sqrt{\frac{\pi}{2b}} - 2b \frac{3}{32b^2} \sqrt{\frac{\pi}{2b}} \right] \\ &= \frac{4\hbar^2 \pi b}{m} \left( \frac{2b}{\pi} \right) \left( \frac{3}{8b} - \frac{3}{16b} \right) \\ &= \frac{4\hbar^2 \pi b}{m} \left( \frac{2b}{\pi} \right) \left( \frac{3}{16b} \right) = \boxed{\frac{3\hbar^2 b}{2m} = \langle \hat{T}(b) \rangle} \end{aligned} \quad (1.8)$$

3. Compute  $\langle \psi | \hat{V} | \psi \rangle$

$$\begin{aligned} \langle \psi | \hat{V} | \psi \rangle &= -\frac{e^2}{4\pi\epsilon_0} |A|^2 4\pi \int_0^\infty e^{-2br^2} \frac{1}{r} r^2 dr \\ &= -\frac{e^2}{4\pi\epsilon_0} \left( \frac{2b}{\pi} \right)^{3/2} \left( \frac{4\pi}{4b} \right) \\ &= \boxed{-\frac{e^2}{4\pi\epsilon_0} 2\sqrt{\frac{2b}{\pi}} = \langle \hat{V}(b) \rangle} \end{aligned} \quad (1.9)$$

4. Compute  $\langle \hat{H} \rangle = \langle \hat{T}(b) \rangle + \langle \hat{V}(b) \rangle = E_{gs}(b)$

Using the results from Equation 1.8 and Equation 1.9 write

$$\langle \hat{H} \rangle = \frac{3\hbar^2 b}{2m} - \frac{e^2}{4\pi\epsilon_0} 2\sqrt{\frac{2b}{\pi}} = E_{gs}(b) \quad (1.10)$$

5. Now, we need to minimise  $E_{gs}(b)$

$$\frac{\partial}{\partial b} E_{gs}(b) = \frac{3\hbar^2}{2m} - \frac{e^2}{4\pi\epsilon_0} \sqrt{\frac{2}{\pi}} \frac{1}{\sqrt{b}} = 0 \quad (1.11)$$

solving for  $b$

$$\sqrt{b} = \frac{e^2}{4\pi\epsilon_0} \sqrt{\frac{2}{\pi}} \left( \frac{2m}{3\hbar^2} \right) \Rightarrow b_{min} = \left( \frac{e^2}{4\pi\epsilon_0} \right)^2 \frac{2}{\pi} \left( \frac{2m}{3\hbar^2} \right)^2 \quad (1.12)$$

Replacing the result from Equation 1.12 into Equation 1.10 we get

$$\begin{aligned} E_{gs}(b_{min}) &= \frac{3\hbar^2}{2m} \left( \frac{e^2}{4\pi\epsilon_0} \right)^2 \frac{2}{\pi} \frac{4m^2}{9\hbar^4} - \frac{e^2}{4\pi\epsilon_0} 2 \sqrt{\frac{2}{\pi}} \frac{e^2}{4\pi\epsilon_0} \sqrt{\frac{2}{\pi}} \left( \frac{2m}{3\hbar^2} \right) \\ &= \left( \frac{e^2}{4\pi\epsilon_0} \right)^2 \frac{m}{\hbar^2} \left( \frac{4}{3\pi} - \frac{8}{3\pi} \right) \\ &= \left( \frac{e^2}{4\pi\epsilon_0} \right)^2 \frac{m}{\hbar^2} \left( -\frac{4}{3\pi} \right) \\ &= -\frac{m}{2\hbar^2} \left( \frac{e^2}{4\pi\epsilon_0} \right)^2 \left( \frac{8}{3\pi} \right), \quad E_1 = -\frac{m}{2\hbar^2} \left( \frac{e^2}{4\pi\epsilon_0} \right)^2 \\ &= \boxed{\frac{8}{3\pi} E_1 = \frac{8}{3\pi} (-13.6 \text{ eV}) = -11.5 \text{ eV}} \end{aligned} \quad (1.13)$$

## 2 Problem 7.12

If the photon had nonzero mass ( $m_\gamma \neq 0$ ), the Coulomb potential would be replaced by a **Yukawa potential**, of the form

$$V(\mathbf{r}) = -\frac{e^2}{4\pi\epsilon_0} \frac{e^{-\mu r}}{r} \quad (2.1)$$

where  $\mu = \frac{m_\gamma c}{\hbar}$ . With a trial wave function of your own devising, estimate the binding energy of a "hydrogen" atom with this potential. Assume  $\mu a \ll 1$ , and give your answer correct to order  $(\mu a)^2$ .

### Solution:

We know that the ground state of the Hydrogen is given by

$$\psi = \frac{1}{\sqrt{\pi} a_0^3} e^{-r/a_0} \quad (2.2)$$

So, let's suppose the "new" hydrogen atom has different Bohr radius, given the new potential. Then Equation 2.2 becomes

$$\psi = \frac{1}{\sqrt{\pi} b^3} e^{-r/b} \quad (2.3)$$

where  $b$  is the radius parameter we must minimise to find the binding energy of the "hydrogen" atom in the new potential.

Now, let's write our Hamiltonian

$$H = -\frac{\hbar^2}{2m}\nabla^2 - \frac{e^2}{4\pi\epsilon_0} \frac{e^{-\mu r}}{r} \quad (2.4)$$

Since our trial wavefunction (Equation 2.4) is already normalised, we can continue to compute  $\langle T \rangle$  and  $\langle V \rangle$

$$\begin{aligned} \langle T \rangle &= \frac{1}{\pi b^3} \int e^{r/b} \left( \frac{-\hbar^2}{2m} \nabla^2 e^{-r/b} \right) dr^3 \\ &= -\frac{\hbar^2}{2m \pi b^3} \int e^{-r/b} (\nabla^2 e^{-r/b}) dr^3 \quad (\text{use Equation 1.6}) \\ &= \frac{\hbar^2}{2m \pi b^4} \int \frac{1}{r^2} e^{-2r/b} \left( 2r - \frac{r^2}{b} \right) dr^3 \\ &= \frac{\hbar^2}{2m \pi b^4} \int_0^\infty e^{-2r/b} \left( 2r - \frac{r^2}{b} \right) dr \underbrace{\int_0^\pi \sin\theta d\theta \int_0^{2\pi} d\phi}_{4\pi} \\ &= \frac{4\pi \hbar^2}{2m \pi b^4} \left[ 2\frac{b^2}{4} - \frac{b^2}{4} \right] = \boxed{\frac{\hbar^2}{2m b^2} = \langle T \rangle} \end{aligned} \quad (2.5)$$

For the integral in Equation 2.5, I used the following formula

$$\int_0^\infty x^n e^{-ax} dx = \frac{n!}{a^{n+1}} \quad (2.6)$$

$$\begin{aligned} \langle V \rangle &= \left( \frac{1}{\pi b^3} \right) \left( -\frac{e^2}{4\pi\epsilon_0} \right) \int \frac{1}{r} e^{-2r/b} e^{-\mu r} dr^3 \\ &= \left( \frac{1}{\pi b^3} \right) \left( -\frac{e^2}{4\pi\epsilon_0} \right) 4\pi \int_0^\infty e^{-r(2/b+\mu)} r dr \\ &= -\frac{e^2}{4\pi\epsilon_0} \frac{4}{b^3} \frac{1}{(2/b+\mu)^2} \\ &= \boxed{-\frac{e^2}{4\pi\epsilon_0} \frac{1}{b \left( 1 + \frac{\mu b}{2} \right)^2} = \langle V \rangle} \end{aligned} \quad (2.7)$$

Now, use results from Equation 2.5 and Equation 2.7 to define  $\langle H \rangle$

$$\langle H \rangle = \frac{\hbar^2}{2m b^2} - \frac{e^2}{4\pi\epsilon_0} \frac{1}{b \left( 1 + \frac{\mu b}{2} \right)^2} \quad (2.8)$$

Now, we need to minimise it

$$\begin{aligned}
\frac{\partial \langle H \rangle}{\partial b} &= -\frac{\hbar^2}{m b^3} + \frac{e^2}{4\pi\epsilon_0} \left[ \frac{1}{b^2(1 + \mu b/2)^2} + \frac{\mu}{b(1 + \mu b/2)^3} \right] = 0 \\
0 &= -\frac{\hbar^2}{m b^3} + \frac{e^2}{4\pi\epsilon_0} \left[ \frac{b(1 + \mu b/2)^3 + \mu b^2(1 + \mu b/2)^2}{b^3(1 + \mu b/2)^5} \right] \\
&= -\frac{\hbar^2}{m b^3} + \frac{e^2}{4\pi\epsilon_0} \left[ \frac{(1 + 3\mu b/2)}{b^2(1 + \mu b/2)^3} \right] \\
\frac{\hbar^2 b}{m} \frac{4\pi\epsilon_0}{e^2} &= \frac{(1 + 3\mu b/2)}{(1 + \mu b/2)^3} \\
\frac{\hbar^2}{m} \left( \frac{4\pi\epsilon_0}{e^2} \right) &= b \frac{(1 + 3\mu b/2)}{(1 + \mu b/2)^3} \Rightarrow a_0 = b \frac{(1 + 3\mu b/2)}{(1 + \mu b/2)^3}
\end{aligned} \tag{2.9}$$

So, we got a relationship between the Bohr radius, and our new radius. But, there is a cubic relation, then we need to simplify even more using the fact that  $\mu a_0 \ll 1$  so we can say that  $\mu b \ll 1$  as well.

Using binomial theorem

$$(1 + x)^{-n} = 1 - nx + \frac{n(n+1)}{2!}x^2 + \dots \tag{2.10}$$

Therefore

$$\left( \frac{\mu b}{2} + 1 \right)^{-3} = 1 - \frac{3\mu b}{2} + \frac{3(\mu b)^2}{2} \tag{2.11}$$

Now, rewrite Equation 2.9

$$\begin{aligned}
a_0 &\approx b \left[ \left( 1 + \frac{3\mu b}{2} \right) \left( 1 - \frac{3\mu b}{2} + \frac{3(\mu b)^2}{2} \right) \right] \\
&\approx b \left( 1 + \frac{3\mu b}{2} - \frac{3\mu b}{2} - \frac{9(\mu b)^2}{4} + \frac{3(\mu b)^2}{2} + \frac{9(\mu b)^3}{4} \right) \\
&= b \left( 1 - \frac{3(\mu b)^2}{4} \right) \quad (\text{neglect third order}) \\
b &= \frac{a_0}{\left( 1 - \frac{3(\mu b)^2}{4} \right)}
\end{aligned} \tag{2.12}$$

Expanding in the denominator in the same way, we get

$$b \approx a_0 \left( 1 + \frac{3(\mu b)^2}{4} \right) \tag{2.13}$$

Since  $\frac{3(\mu b)^2}{4}$  is a really small number, we can write

$$b \cong a_0 \left( 1 + \frac{3(\mu a_0)^2}{4} \right) \tag{2.14}$$

Now, replacing this result into Equation 2.8

$$\langle H_{min} \rangle = \frac{\hbar^2}{2m} \frac{1}{a_0^2 \left(1 + \frac{3(\mu a_0)^2}{4}\right)^2} - \frac{e^2}{4\pi\epsilon_0} \frac{1}{a_0 \left(1 + \frac{3(\mu a_0)^2}{4}\right) \left(1 + \frac{\mu a_0}{2}\right)^2} \quad (2.15)$$

Expanding the denominators

$$\begin{aligned} \langle H_{min} \rangle &\approx \frac{\hbar^2}{2m a_0^2} \left[1 - 2\frac{3(\mu a_0)^2}{4}\right] - \frac{e^2}{4\pi\epsilon_0 a_0} \left[1 - \frac{3(\mu a_0)^2}{4}\right] \left[1 - 2\frac{\mu a_0}{2} + 3\left(\frac{\mu a_0}{2}\right)^2\right] \\ &= \frac{\hbar^2}{2m a_0^2} \left[1 - \frac{3(\mu a_0)^2}{2}\right] - \frac{e^2}{4\pi\epsilon_0 a_0} \left[1 - \mu a_0 + \frac{3(\mu a_0)^2}{4} - \frac{3(\mu a_0)^2}{4}\right] \\ &= -E_1 \left[1 - \frac{3(\mu a_0)^2}{2}\right] + 2E_1[1 - \mu a_0] \\ &= \boxed{E_1 \left[1 - 2\mu a_0 + \frac{3}{2}(\mu a_0)^2\right]} \end{aligned} \quad (2.16)$$

### 3 Problem 7.13

Suppose you are given a quantum system whose Hamiltonian  $H_0$  admits just two eigenstates,  $\psi_a$  (with energy  $E_a$ ), and  $\psi_b$  (with energy  $E_b$ ). They are orthogonal, normalised, and nondegenerate (assume  $E_a$  is the smaller of the two). Now, we turn in a perturbation  $H'$ , with the following matrix elements:

$$\langle \psi_a | H' | \psi_a \rangle = \langle \psi_b | H' | \psi_b \rangle = 0; \quad \langle \psi_a | H' | \psi_b \rangle = \langle \psi_b | H' | \psi_a \rangle = h \quad (3.1)$$

- Find the exact eigenvalues of the perturbed Hamiltonian.
- Estimate the energies of the perturbed system using second-order perturbation theory.
- Estimate the ground state energy of the perturbed system using variational principle, with a trial function of the form

$$\psi = (\cos\phi)\psi_a + (\sin\phi)\psi_b \quad (3.2)$$

where  $\phi$  is an adjustable parameter. (Note that writing the linear combination in this way guarantees that  $\psi$  is normalised.)

- Compare your answers to (a), (b), and (c). Why is the variational principle so accurate in this case?

### Solution:

- To find the eigenvalues, we use the following relation

$$\det |H - \lambda I| = 0 \quad (3.3)$$

But first we need to define  $H$

$$H = H_0 + H' \quad (3.4)$$

Since the eigenstates of the system are orthogonal, and nondegenerate, we can write

$$H_0 = \begin{pmatrix} E_a & 0 \\ 0 & E_b \end{pmatrix} \quad (3.5)$$

and  $H'$  is given by the problem as

$$H' = \begin{pmatrix} 0 & h \\ h & 0 \end{pmatrix} \quad (3.6)$$

then,  $H$  has the form

$$H = \begin{pmatrix} E_a & h \\ h & E_b \end{pmatrix} \quad (3.7)$$

Now, let's find the eigenvalues. Using Equation 3.3

$$\begin{aligned} \det \begin{vmatrix} E_a - \lambda & h \\ h & E_b - \lambda \end{vmatrix} &= 0 = (E_a - \lambda)(E_b - \lambda) - h^2 \\ &= E_a E_b - \lambda(E_a + E_b) + \lambda^2 - h^2 \\ \lambda &= \frac{E_a + E_b}{2} \pm \frac{\sqrt{(E_a + E_b)^2 - 4(E_a E_b - h^2)}}{2} \end{aligned} \quad (3.8)$$

Finally

$$E_{\pm} = \frac{E_a + E_b}{2} \pm \frac{\sqrt{(E_a - E_b)^2 + 4h^2}}{2} \quad (3.9)$$

b) Second order perturbation theory is given by

$$E_n = E_n^0 + \langle \psi_n | H' | \psi_n \rangle + \sum_{m \neq n} \frac{|\langle \psi_m | H' | \psi_n \rangle|^2}{E_n^0 - E_m^0} \quad (3.10)$$

Then, for the given eigenenergies, and using the matrix elements from Equation 3.1

$$\begin{aligned} \tilde{E}_a &= E_a + \frac{|\langle \psi_b | H' | \psi_a \rangle|^2}{E_a - E_b} \\ E_- &= E_a - \frac{h^2}{E_b - E_a} \end{aligned} \quad (3.11)$$

$$\begin{aligned} \tilde{E}_b &= E_b + \frac{|\langle \psi_a | H' | \psi_b \rangle|^2}{E_b - E_a} \\ E_+ &= E_b + \frac{h^2}{E_b - E_a} \end{aligned} \quad (3.12)$$

c) So, we need to compute  $\langle H \rangle$

$$\begin{aligned}
\langle H \rangle &= \langle \psi | H_0 | \psi \rangle + \langle \psi | H' | \psi \rangle \\
&= \langle \cos \phi \psi_a + \sin \phi \psi_b | H_0 | \cos \phi \psi_a + \sin \phi \psi_b \rangle \\
&+ \langle \cos \phi \psi_a + \sin \phi \psi_b | H' | \cos \phi \psi_a + \sin \phi \psi_b \rangle \\
&= \cos^2 \phi \langle \psi_a | H_0 | \psi_a \rangle + \sin^2 \phi \langle \psi_b | H_0 | \psi_b \rangle \\
&+ \sin \phi \cos \phi \langle \psi_b | H' | \psi_a \rangle + \sin \phi \cos \phi \langle \psi_a | H' | \psi_b \rangle \\
&= E_a \cos^2 \phi + E_b \sin^2 \phi + 2h \sin \phi \cos \phi
\end{aligned} \tag{3.13}$$

Now, we have to minimise  $\langle H \rangle$

$$\begin{aligned}
\frac{\partial \langle H \rangle}{\partial \phi} &= -E_a 2 \cos \phi \sin \phi + E_b 2 \sin \phi \cos \phi + 2h(\cos^2 \phi - \sin^2 \phi) \\
0 &= (E_b - E_a) \sin(2\phi) + 2h \cos(2\phi) \\
\tan(2\phi) &= -\frac{2h}{(E_b - E_a)} = -\epsilon, \quad \text{where} \quad \epsilon = \frac{2h}{E_b - E_a}
\end{aligned}$$

we can write  $\epsilon$  as follows

$$\begin{aligned}
-\epsilon &= \frac{\sin(2\phi)}{\sqrt{1 - \sin^2(2\phi)}} \Rightarrow \sin^2(2\phi) = \epsilon^2(1 - \sin^2(2\phi)) \\
&\Rightarrow \sin^2(2\phi)[1 + \epsilon^2] = \epsilon^2
\end{aligned}$$

Solving for  $\sin^2(2\phi)$

$$\sin^2(2\phi) = \frac{\epsilon^2}{1 + \epsilon^2} \Rightarrow \sin(2\phi) = \frac{\pm \epsilon}{\sqrt{1 + \epsilon^2}} \tag{3.14}$$

We can also write

$$\cos^2(2\phi) = 1 - \sin^2(2\phi) = 1 - \frac{\epsilon^2}{1 + \epsilon^2} = \frac{1}{1 + \epsilon^2} \Rightarrow \cos(2\phi) = \frac{\mp 1}{\sqrt{1 + \epsilon^2}} \tag{3.15}$$

the sign is given by  $\tan(\phi)$ .

But we need to find an expression for  $\sin^2 \phi$  and  $\cos^2 \phi$  so we can plug them in our Hamiltonian, therefore i will use the following trig. identities

$$\cos^2 \phi = \frac{1}{2} (1 + \cos(2\phi)) = \frac{1}{2} \left( 1 \mp \frac{1}{\sqrt{1 + \epsilon^2}} \right) \tag{3.16}$$

$$\sin^2 \phi = \frac{1}{2} (1 - \cos(2\phi)) = \frac{1}{2} \left( 1 \pm \frac{1}{\sqrt{1 + \epsilon^2}} \right) \tag{3.17}$$

taking into account that  $2 \sin \phi \cos \phi = \sin(2\phi)$ , and using the results from 3.14, 3.16, let's rewrite our Hamiltonian

$$\begin{aligned}
\langle H \rangle_{min} &= \frac{1}{2} E_a \left( 1 \mp \frac{1}{\sqrt{1 + \epsilon^2}} \right) + \frac{1}{2} E_b \left( 1 \pm \frac{1}{\sqrt{1 + \epsilon^2}} \right) \pm h \frac{\epsilon}{\sqrt{1 + \epsilon^2}} \\
\langle H \rangle_{min} &= \frac{1}{2} \left[ E_a + E_b \pm \frac{(E_b - E_a + 2h\epsilon)}{\sqrt{1 + \epsilon^2}} \right]
\end{aligned} \tag{3.18}$$



Simplifying

$$\begin{aligned}
\frac{(E_b - E_a + 2h\epsilon)}{\sqrt{1 + \epsilon^2}} &= \frac{(E_b - E_a) + 2h \frac{2h}{(E_b - E_a)}}{\sqrt{1 + \frac{4h^2}{(E_b - E_a)^2}}} \\
&= \frac{(E_b - E_a)^2 + 4h^2}{\sqrt{(E_b - E_a)^2 + 4h^2}} \\
&= \sqrt{(E_b - E_a)^2 + 4h^2}
\end{aligned} \tag{3.19}$$

So,

$$\langle H \rangle_{min} = \frac{1}{2} \left[ E_a + E_b \pm \sqrt{(E_b - E_a)^2 + 4h^2} \right] \tag{3.20}$$

$$\boxed{E_{\pm} = \frac{1}{2} \left[ E_a + E_b \pm \sqrt{(E_b - E_a)^2 + 4h^2} \right]} \tag{3.21}$$

We can see that this result is similar to that obtained in part a (Equation 3.9). Considering that  $h$  is very small, we can perform binomial expansion (Equation 2.10)

$$\begin{aligned}
E_{\pm} &= \frac{1}{2} \left[ E_a + E_b \pm (E_b - E_a) \sqrt{1 + \frac{4h^2}{(E_b - E_a)^2}} \right] \\
E_{\pm} &\approx \frac{1}{2} \left[ E_a + E_b \pm (E_b - E_a) \left( 1 + \frac{2h^2}{(E_b - E_a)^2} \right) \right] \\
E_{\pm} &= \frac{1}{2} \left[ E_a + E_b \pm (E_b - E_a) \pm \frac{2h^2}{(E_b - E_a)} \right]
\end{aligned}$$

Therefore

$$E_- = E_a - \frac{h^2}{(E_b - E_a)} \tag{3.22}$$

$$E_+ = E_b + \frac{h^2}{(E_b - E_a)} \tag{3.23}$$

- d) As we can see, results from (c), ( (a) follows the same expansion), confirm the perturbation theory results in (b). The variational principle (c) gets the ground state ( $E_-$ ) exactly right, and this is because the trial wavefunction (Equation 3.2) is almost the most general state.

## 4 Problem 7.14

As an explicit example of the method developed in Problem 7.13, consider an electron at rest in a uniform magnetic field  $\mathbf{B} = B_z \hat{k}$ , for which the Hamiltonian is (Equation 4.158)

$$H_0 = \frac{e B_z}{m} S_z \tag{4.1}$$

The eigenspinors,  $\chi_a$  and  $\chi_b$ , and the corresponding energies,  $E_a$  and  $E_b$ , are given in Equation 4.161. Now we turn on a perturbation, in the form of a uniform field in the  $x$  direction:

$$H' = \frac{e B_x}{m} S_x \quad (4.2)$$

- a) Find the matrix elements of  $H'$ , and confirm that they have the structure of Equation 7.55. What is  $\hbar$ ?
- b) Using your result in Problem 7.13 (b), find the new ground-state energy, in second-order perturbation theory.
- c) Using your result in Problem 7.13 (c), find the variational principle bound on the ground-state energy.

### Solution:

For the electron:

$$\gamma = \frac{-e}{m} \quad (4.3)$$

and, from Equation 4.161, we have that

$$E_{\pm} = \pm \frac{e B_z \hbar}{2m} \quad (4.4)$$

Following the conditions of the Problem 7.13, let  $E_a < E_b$ , so

$$\chi_a = \chi_- = \begin{pmatrix} 0 \\ 1 \end{pmatrix} \quad (4.5)$$

$$\chi_b = \chi_+ = \begin{pmatrix} 1 \\ 0 \end{pmatrix} \quad (4.6)$$

$$E_a = E_- = -\frac{e B_z \hbar}{2m} \quad (4.7)$$

$$E_b = E_+ = \frac{e B_z \hbar}{2m} \quad (4.8)$$

$$S_x = \frac{\hbar}{2} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \quad (4.9)$$

a)

$$\langle \chi_a | H' | \chi_a \rangle = \frac{e B_x \hbar}{m} \frac{1}{2} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \end{pmatrix} = \frac{e B_x \hbar}{2m} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \end{pmatrix} = 0 \quad (4.10)$$

$$\langle \chi_b | H' | \chi_b \rangle = \frac{e B_x \hbar}{m} \frac{1}{2} \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \end{pmatrix} = 0 \quad (4.11)$$

$$\langle \chi_b | H' | \chi_a \rangle = \frac{e B_x \hbar}{m} \frac{1}{2} \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \end{pmatrix} = \frac{e B_x \hbar}{2m} \quad (4.12)$$

$$\langle \chi_a | H' | \chi_b \rangle = \frac{e B_x \hbar}{m} \frac{1}{2} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \end{pmatrix} = \frac{e B_x \hbar}{2m} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \end{pmatrix} = \frac{e B_x \hbar}{2m} \quad (4.13)$$

Therefore

$$\boxed{h = \frac{eB_x\hbar}{2m}} \quad (4.14)$$

and the conditions of Problem 7.13 are met.

b) From Problem 7.13

$$E_{gs} \approx E_a - \frac{h^2}{(E_b - E_a)} = -\frac{eB_z\hbar}{2m} - \frac{(eB_x\hbar/2m)^2}{eB_z\hbar/m} = \boxed{-\frac{e\hbar}{2m} \left( B_z + \frac{B_x^2}{2B_z} \right)} \quad (4.15)$$

c) From Problem 7.13(c),

$$E_{gs} = \frac{1}{2} \left[ E_a + E_b - \sqrt{(E_b - E_a)^2 + 4h^2} \right] \quad (4.16)$$

$$E_{gs} = -\frac{1}{2} \sqrt{\left( \frac{eB_z\hbar}{m} \right)^2 + 4 \left( \frac{eB_x\hbar}{2m} \right)^2} = \boxed{-\frac{e\hbar}{2m} \sqrt{B_z^2 + B_x^2}} \quad (4.17)$$

## 5 Problem 7.17

The fundamental problem in harnessing nuclear fusion is getting the two particles (say, two deuterons) close enough together for the attractive (but short-range) nuclear force to overcome the Coulomb repulsion. The "brute force" method is to heat the particles to fantastic temperatures and allow the random collisions to bring them together. A more exotic proposal is **muon catalysis**, in which we construct a "hydrogen molecule ion", only with deuterons in place of protons, and a (muon) in place of the electron. Predict the equilibrium separation distance between the deuterons in such a structure, and explain why muons are superior to electrons for this purpose.

### Solution:

We have to consider the reduced mass concept. Let reduced mass of the muon be  $m_r$ , and  $m_\mu$  be the mass of the muon

$$m_r = \frac{m_\mu m_d}{m_\mu + m_d} = \frac{m_\mu 2m_p}{m_\mu + 2m_p} = \frac{m_\mu}{1 + m_\mu/2m_p}, \quad m_\mu = 207m_e \quad (5.1)$$

$$1 + \frac{m_\mu}{2m_p} = 1 + \frac{207(9.11 \times 10^{-31})}{2(1.67 \times 10^{-27})} = 1.056$$

Therefore

$$m_r = \frac{207 m_e}{1.056} = 196 m_e \quad (5.2)$$

, we know that the Bohr radius is defined as

$$a_0 = \frac{4\pi\epsilon_0\hbar^2}{e^2 m_e}$$

but in this case,  $m_e \rightarrow m_\mu$ , therefore, the molecule would shrink by a factor of 196!, bringing the deuterons much closer together, as desired. Thus  
For the electron case:

$$R = 2.493 a_0$$

therefore, for muons:

$$R_\mu = 2.493 \frac{a_0}{196}$$

$$R_\mu = 2.493 \frac{0.529 \times 10^{-10}}{196} = 6.73 \times 10^{-13} m$$

We can see that the use of muon, yields to a shorter muonic radius than electronic Bohr radius. This brings the nuclei so close so the strong nuclear force can bind both nuclei together very easily.