Table 1.

Summary of AMI-LA Observations of GRB 171010A. Errors are statistical and calibration (5 per cent) combined in quadrature. ΔT is the number of days between T_0 and the observation mid-point. The upper limit at $\Delta T = 86.03$ d is three times the rms noise in the image. All observations were taken at a central frequency of 15.5 GHz.

Date (ΔT)	Obs. Length (h)	Flux Density (mJy)	Error (mJy)
1.35	4	2.26	0.13
2.37	3.5	2.52	0.14
3.31	4	2.48	0.14
4.31	4	2.08	0.13
5.31	4	2.07	0.14
8.33[a]	3	1.11	0.13
12.27	4	0.69	0.07
13.27[a]	4	0.52	0.12
14.27	4	0.70	0.09
19.37	3	0.37	0.07
20.30	4	0.36	0.08
22.36	3	0.43	0.10
26.25	4	0.21	0.05
27.80[b]	8	0.37	0.07
29.24	4	0.25	0.07
30.66[b]	8	0.23	0.08
33.42[b]	8	0.23	0.05
38.24	4	0.19	0.05
46.21	4	0.17	0.04

60.66[a, b]	11	0.13	0.06
86.15	4	<0.09	–

[a]The source in these observations was not well fit by the clean beam due to residual noise. Thus the peak flux density was used.
[b]These observations combined data from multiple days in order to lower the noise level. In this case the time is the centroid of all observations used.