RAM-SCB User Manual

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Chapter 1

Introduction

The **R**ing current **A**tmosphere interactions **M**odel with **S**elf **C**onsistent magnetic field (**B**) is a unique code that combines a kinetic model of ring current plasma with a three dimensional force-balanced model of the terrestrial magnetic field. The kinetic portion, RAM, solves the kinetic equation to yield the bounce-averaged distribution function as a function of azimuth, radial distance, energy and pitch angle for three ion species (H^+, He^+, and O^+) and, optionally, electrons. The domain is a circle in the Solar-Magnetic (SM) equatorial plane with a radial span of 2 to 6.5 R_E. It has an energy range of approximately $100\,eV$ to $500\,KeV$. The 3-D force balanced magnetic field model, SCB, balances the $\mathbf{J} \times \mathbf{B}$ force with the divergence of the general pressure tensor to calculate the magnetic field configuration within its domain. The domain ranges from near the Earth's surface, where the field is assumed dipolar, to the shell created by field lines passing through the SM equatorial plane at a radial distance of 6.5 R_E. The two codes work in tandem, with RAM providing anisotropic pressure to SCB and SCB returning the self-consistent magnetic field through which RAM plasma is advected.

RAM-SCB has grown from a research-grade code with limited options and static magnetic field (RAM) to a rich, highly configurable research and operations tool with a multitude of new physics and output products. This manual provides a guide to users who want to learn how to install, configure, and execute RAM-SCB simulations. While the code is designed to make these steps as straight-forward as possible, it is strongly recommended that users review the publications listed in the Bibliography to ensure a thorough understanding of the physics included in the model. Additionally, all users are asked to review the terms of use.

1.1 About This Manual

Users who want to install and begin quickly should start at Chapter 2, which quickly outlines the path from installation to simulation with little detail. The installation process is discussed fully in Chapter 3. Instructions on performing simulations, as well as several example simulations, are given in Chapter 5. An outline of using this code in the Space Weather Modeling Framework is found in Chapter 7. Useful scripts included in the distribution are described, in brief, in Chapter 8. Finally, a complete list of all param file commands is found in Chapter 9.

1.2 TERMS OF USE & DISTRIBUTION POLICY

Use of the RAM-SCB software implies agreement with the terms herein. RAM-SCB is open source software that has been developed at Los Alamos National Laboratory (LANL). The code is based on physics and numerical methods detailed in the following publications and references therein:

1. Jordanova, V. K. et al. (2006), Kinetic simulations of ring current evolution during the Geospace Environment Modeling challenge events, J. Geophys, Res., 111, A11S10, doi:10.1029/2006JA011644.

2. Zaharia, S. et al. (2006), Self-consistent modeling of magnetic fields and plasmas in the inner magnetosphere: Application to the geomagnetic storm, J. Geophys. Res., 111, A11S14, doi:10.1029/2006JA011619.

The references below represent critical development milestones for RAM-SCB. Please consider citing these works to give the developers proper credit.

Citation	Information
Jordanova et al. [1996]	First description of ring current model (RAM) using dipolar magnetic field
Jordanova et al. [2006]	First extension of RAM for non-dipolar magnetic field and coupling with SCB
Zaharia et al. [2006]	Description of SCB model and coupling with RAM
Jordanova et al. [2010]	Full description of RAM extension for non-dipolar magnetic field
Welling et al. [2011]	First full description of one-way coupling with SWMF
Welling et al. [2015]	Description of two-way coupling of RAM-SCB with SWMF

Chapter 2

Quick Start Guide

This section provides a bare-bones approach for getting started with RAM-SCB. The instructions are aimed at users familiar with Unix-like environments. Tutorial-like guides for each step can be found in other chapters.

2.1 Installation

Installation and compilation of RAM-SCB requires a Fortran compiler, Perl interpreter, an MPI library compiled with the preferred Fortran compiler, and several external libraries which have additional requirements (most notably a C compiler, typically standard on Unix-like systems.) The required libraries are:

1. NCAR's Graphics Library
2. NTCC's Princeton Spline (Pspline) library
3. Unidata's NetCDF library

If these libraries are installed in non-standard locations, use environment variables to make them visible to RAM-SCB. The three variables to be set are `NCARGDIR`, `PSPLINEDIR`, and `NETCDFDIR`.

Installation of RAM-SCB is handled via the Config.pl script found in the installation directory. The `-h` option will print help, but the user will almost exclusively use Config.pl as follows:

```
Config.pl -install -compiler=pgf90 -mpi=mpich2
```

Other compiler/MPI combinations are of course available, see Chapter 3.

Finally, use GNU Make to compile and, if desired, test:

```
make
make test
```

2.2 Execution

RAM-SCB is run from *run directories*, create one via `make`:

```
make rundir RUNDIR=~/desired_run_location
```

Run directories are soft-linked to the installation. This means you can move them freely and create multiple run directories for simultaneous simulations. A fresh run directory has everything you need to execute the code immediatetly, simply start the executable.

```
./ram_scb.exe
```

To customize the simulation, edit the PARAM.in file. Chapter 9 has a complete description of all parameters. Additional input files required must be placed in either the `input_ram` or `input_scb` subdirectories; output is located in similarly named locations.

Output data come in two formats: simple ASCII and NetCDF files. While there are a plethora of tools available, we recommend users try the Python-based tools discussed in Chapter 6.

Chapter 3

Installation

Installation and compilation of RAM-SCB requires a Fortran compiler, MPI compiled against the chosen Fortran compiler, a Perl interpreter, and several external libraries which have additional requirements (most notably a C compiler, typically standard on Unix-like systems.) The configuration is done with the Config.pl script; compilation with GNU Make. The Config/Make system follows the tiered makefile standards of the Space Weather Modeling Framework (SWMF) project; users of the SWMF should feel right at home using RAM-SCB.

3.1 Installation of Required Libraries

There are three required libraries that must be installed to use RAM-SCB: NCAR Graphics, Unidata's NetCDF library and NTCC's Princeton Spline (Pspline) library. It is possible to find pre-compiled binaries for each library, allowing the user to skip the configure, make, and install steps. However, if a binary cannot be found that matches your system and compilers, you will be forced to install from source. Some systems may already have these libraries installed; be sure to check before going through unneccessary installation work.

3.1.1 Installation From Source

Tarballs of the source code can be obtained through these URLs:

1. http://www.ncl.ucar.edu/Download/ (login required)
2. http://w3.pppl.gov/rib/repositories/NTCC/files/pspline.tar.gz
3. http://www.unidata.ucar.edu/downloads/netcdf/ftp/netcdf-4.0.1.tar.gz

Unpack the tarballs in the usual fashion:

```
tar -xzvf netcdf-4.0.1.tar.gz
```

Because these libraries are not owned or maintained by the authors of RAM-SCB, only minimal installation instructions are provided here. The reader is encouraged to explore the websites and help files associated with each package.

Ncarg Quick Install

Installation of the NCAR Graphics library from source is an involved process. It is strongly suggested that the user first check their system to see if the library is already installed or install via binary.

NetCDF Quick Install

Unpack the NetCDF tarball, `cd` into the unzipped directory, and use the following commands:

```
./configure --prefix=/Users/ram_user/netcdf_lib
make install
```

This will install the NetCDF library in `/Users/ram_user/netcdf_lib` (the home directory for user *ram_user*); change the value of `--prefix` to select the installation destination. For installation, NetCDF requires only a C compiler. `gcc` is strongly suggested.

Pspline Quick Install

Pspline requires NetCDF, so be sure that NetCDF is correctly installed before continuing. After unpacking the tarball, compile the library using

```
make FORTRAN_VARIANT=Portland NETCDF_DIR=~/netcdf_lib
```

It is important to specify the Fortran compiler that coincides with the compiler you will use to make the RAM-SCB executable. Other options accepted by the `FORTRAN_VARIANT` variable are `PathScale`, `GCC`(gfortran and g95), `NagWare`, `Fujitsu`, `Intel`, and `Absoft`. Ensure that the value of `NETCDF_DIR` coincides with the installation directory of NetCDF.

`make` may fail during the creation of the Pspline test utilities. If this happens, check for the creation of a `lib` directory and several `*.a` files contained within. If they exist, disregard the error message and continue to the next step.

Install Pspline as follows (changing the value of `PREFIX` to select installation directory):

```
make install PREFIX=~/psline_portland
```

On machines where you will use different compilers, it is helpful to add the compiler name to the name of the installation directory.

3.1.2 Libraries From Modules

Well maintained clusters often use the `module` interface for loading and unloading software packages. Be sure to explore the available packages on a new system before going through the arduous task of installing from source!

```
module avail
```

will list all available software libraries on the system. You can load a library, unload a library, and list loaded software using the module commands:

```
module load package_name
module unload package_name
module list
```

As an example, here is what you would type on NASA's Pleiades cluster to get started with RAM-SCB:

```
module load comp/intel/10.1.021_64
module load mpi/mpt.1.25
module load netcdf/4.0-i10.1
module load ncarg/4.4.2/intel
module load python/2.6.1
```

Intel Fortran, MPI, NetCDF and NCAR Graphics are now loaded and ready to use. Place such commands your shell configuration script (`~/.cshrc` or `~/.bashrc` depending on your shell) to load modules upon login. Do not forget to point RAM-SCB to these libraries as listed below.

3.2 Pointing RAM-SCB to External Libraries

RAM-SCB must be explicitly pointed to libraries installed in non-standard places. There are several ways for RAM-SCB to find the required external libaries, summarized in Table 3.1. The most permanant and convenient method is to set an environment variable that RAM-SCB will search for upon installation. They can be set with the following commands:

Bash Shell: `export NETCDFDIR=`*installation path*

C Shell: `setenv NETCDFDIR ''`*installation path*`''`

Placing these commands into configuration files (e.g. `~/.bashrc` or `~/.cshrc`) will load the environment variables at login. This is preferable to typing the commands every session. Note that the installation path listed should point to the top level directory for the installation. For example, `/Users/ram_user/ncarg` is sufficient but `/Users/ram_user/ncarg/lib` will cause an error.

Alternatively, it is possible to set the install paths for the external libraries using Config.pl during installation. The use of this method is detailed in the next section. Even if you have set the environment variables, using the Config.pl switches will override them. These switches are *only* recognized in conjunction with the `-install` switch of Config.pl.

Finally, if you set neither the environment variables or the Config.pl switches, the Fortran compiler will look for the libaries in the default directories. This may be useful to test if the libraries are already properly installed by the system administrator.

If you change the library locations by any method, you must reinstall the code.

Library	Environment Variable	Config.pl Switch
NCAR Graphics	`NCARGDIR=[...]`	`-ncarg=[...]`
NetCDF	`NETCDFDIR=[...]`	`-netcdf=[...]`
Pspline	`PSPLINEDIR=[...]`	`-psline=[...]`

Table 3.1: List of required libraries and methods for expressing their location to RAM-SCB. Note that the Config.pl switches override environment variables. If none are given, the Fortran compiler will search in the default library location.

3.3 Installation, Configuration, and Compiling

`Config.pl` handles the installation and configuration of RAM-SCB. To view the installation and configuration status, or to view help, use the following commands:

```
Config.pl
Config.pl -h
```

To install the code, use the `-install` flag:

```
Config.pl -install
```

Although RAM-SCB will try to use reasonable defaults based on your system, there are a number of flags that allow you to customize your installation:

- Use `-compiler` to select the Fortran 90 compiler. Common choices include pgf90, ifort, gfortran, and f95(for NAG).
- The `-mpi` flag allows the user to pick which version of MPI to use. Choose from mpich, mpich2, Altix, and openmpi. Alternatively, the `-nompi` flag may be set with no value. This option is fragile.
- Set `-ncdf` to the path of the NetCDF library installation. If used, this flag overrides the environment variable NETCDFDIR.

- Set `-pspline` to the path of the Pspline library installation. If used, this flag overrides the environment variable PSPLINEDIR.

To exemplify a typical installation, imagine a machine with several different Fortran compilers available. For each compiler available, the user has a corresponding installation of Pspline. The user will use this command to properly install RAM-SCB:

```
Config.pl -install -compiler=pgf90 -mpi=mpich2 -pspline=~/libs/pspline_portland/
```

There are other Config.pl options that set up real precision, debug flags, and optimization level. Use `Config.pl -h` to learn about the available options.

After the code has been properly configured, compiliation is simple:

```
make
```

Compilation is most likely to fail for two reasons. The first is RAM-SCB not finding MPI or another key library. The second is MPI or an external library that is installed using a mix of different Fortran compilers. Be vigilant when installing each library!

To remove object files before a fresh compilation, use

```
make clean
```

To uninstall RAM-SCB, simply use

```
Config.pl -uninstall
```

3.4 Testing the Installation

Running the RAM-SCB tests is an excellent way to evaluate the success and stability of your installation. To run the tests, simply type

```
make test
```

This will compile RAM-SCB, create a run directory, perform a short simulation and compare the results to a reference solution. If there is a significant difference between the test and reference solution, the test will fail. Details can be found in `*.test` files in the RAM-SCB directory. For a full description of running and interpreting test results, see Chapter 4.

3.5 Building Documentation

To generate a PDF of the latest User Manual, type

```
make PDF
```

The document will be located in the `doc/` directory.

Chapter 4

Testing RAM-SCB

RAM-SCB comes packaged with a test suite to determine if the current installation operates correctly and produces the expected results. Different tests evaluate different code capabilities and options. Testing is an especially powerful development tool for evaluating the impact, intentional or not, of changes to the source code. A summary of the tests and associated commands can be generated by using `make test_help`.

4.1 Using and Interpreting Test Results

Tests are called through GNU `make`. To run all available tests, simply use the command

```
make test
```

in the RAM-SCB installation directory. To run a specific test (listed below), call that test by name:

```
make test1
```

A test performs the following actions:

1. Compile the code with any options required by the particular test.
2. Create a run directory entitled `run_test`. If on exists, it will be deleted, so use caution when using this name for any other purpose besides automated testing.
3. Copy the required input and parameter files into the new test run directory.
4. Run the code. Test simulations are only long enough to perform the features being tested.
5. Perform a specialized version of `diff`, included in the distribution, to compare the results produced by the test simulation against reference solutions stored in the output directory of the installation location. If the two files have values that differ by a certain amount (default is an absolute difference of 10^{-30}), the test will fail.

Each of these steps can be called individually by using `make` *nametest_stepname*, where the step names are `compile`, `rundir`, `run`, and `check`.

If, at any point, the test fails, the details will be recorded in *nametest*.diff in the installation directory. If the test is successful, this file will be created but be empty. Tests that fail at the comparitive stage will list all of the differences. When this happens, additional tools to examine the file differences are recommended. An excellent visualization of file differences can be rendered with `tkdiff`, an open source tool that is already installed on many Linux machines.

4.2 Description of Available Tests

4.2.1 Test 1

Command: `make test1`

This is the most basic test for RAM-SCB. It runs RAM-SCB for 300 seconds using Volland-Stern electric field, constant dipole magnetic field (no SCB calculation), and LANL fluxes. **SHIELDS-RC** users should rely on this test.

4.2.2 Test 2

Command: `make test2`

Test 2 repeats the previous test, but stops half way through the simulation, writes restart files, and restarts the run. This excercises the code's ability to seamlessly restart a simulation at any arbitrary point.

4.2.3 Test 3

Command: `make test3`

Test 3 activates SCB, uses the more complicated Tsyganenko 89 boundary conditions and Weimer 2000 empirical electric field. This is the base test for SCB functionality and ionosphere-to-equator electric potential mapping.

Chapter 5

Performing Simulations

Running RAM-SCB occurs not in the installation directory, but through *run directories* – special directories that keep your simulations separate from the installation and other simulations. To create a run directory, simply use the `make` interface:

```
make rundir
```

`make` will unpack some default inputs and organize a new directory aptly named `run`. Note that if a run directory exists that is called "run", RAM-SCB will not over write it and `make` will complain.

The first thing you should do is rename your run directory and move it to a place that makes sense given the constraints of your system. For example, let's pretend we're using NASA's Pleiades system and doing a set of validation simulations. On this system, you typically install code on your home directory, but must run from designated directories called "nobackupXX" (where XX is the number of the directory and nobackup tells you why you didn't install your code there.) From the install directory, make and move your first run directory:

```
make rundir
mv run /nobackup10/username/run_valid
```

Run directories soft link to the installation directory, so as long as you don't move, uninstall, or otherwise molest your installion, you can put the run directory where ever you want. Furthermore, you are not limited to a single run directory. Adding more allows you to run more simulations simultaneously from a single installation. Using the `PARAM` interface, described below, you can perform many very different simulations at once without re-compiling the code. Let's add another run directory, but using `make` variable syntax to make, name, and move the directory in one line:

```
make rundir RUNDIR=/nobackup10/username/run_valid2
```

The variable `RUNDIR` sets the location without any extra hassle. Each of the two run directories created in this example use the same installation of RAM-SCB but have different inputs and outputs. They can be run simultaneously without interfering with each other.

Let's look inside a typical run directory:

Figure 5.1 shows the default run directory organization. The IM directory is where most input and output occurs; subdirectories shaded green are soft-linked to the top run level for convenience. These directories are self explanatory outside of `output_swmf`, which is for runs that use SWMF output in stand-alone mode. Enough inputs are given in the input directories for a sample simulation right away. The red box in Figure 5.1 denotes the linked executable. The final files, denoted as yellow boxes, are the final input files. `w2k.dat` is the Weimer 2000 empirical model data file, `omni.txt` is solar wind input file from the OMNI database, and `RamIndices.txt` is a list of geomagnetic indices. The necessity of these three files is discussed below.

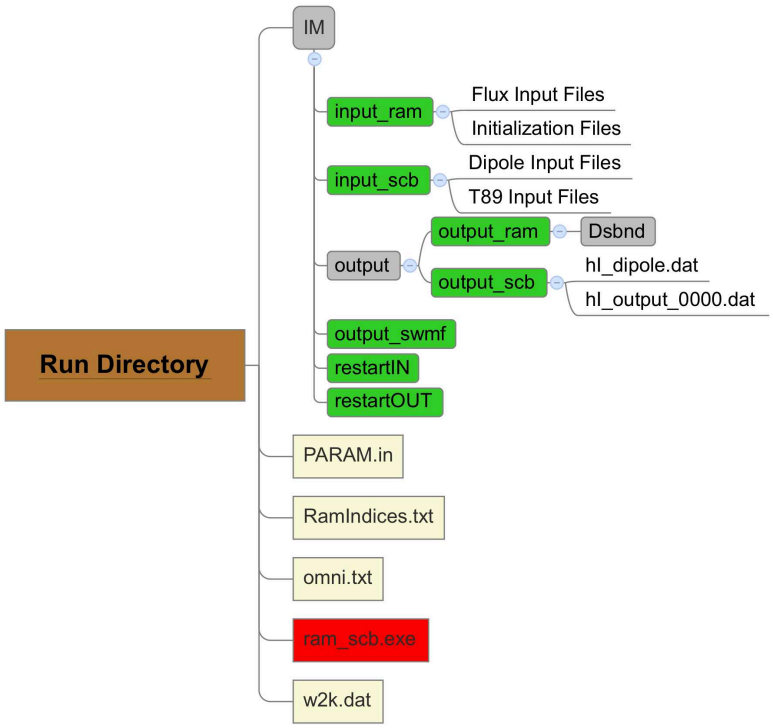


Figure 5.1: Run directory layout. Gray and green rounded boxes indicate subdirectories; green boxes are directories that are soft linked one level up for convenience. Yellow boxes are key input files, the red box is the linked RAM-SCB executable, and other data input files are listed without boxes.

Finally, there is `PARAM.in`, the run-time configuration file that is used to tell RAM-SCB what and how it should simulate. The `PARAM` interface, identical to the one used by the SWMF, is simple. Any line of text in the file not preceeded by a # symbol is ignored, so comment your file to your heart's content. Any text preceeded by #, however, is a interpreted as a `PARAM` command. A generic command takes the form,

```
#COMMANDNAME
Parameter1       ParameterName
Parameter2       ParameterName
...
ParameterN       ParameterName
```

The exact values of the parameters are command specific; the parameter names are simply tab-delimited from the parameters and are not needed (keeping them is recommended to keep your file clear and easily understood!) Here are a few commands in action:

```
#STARTTIME
2006 iYear
7 iMonth
19 iDay
0 iHour
```

```
0 iMinute
0 iSecond

#OUTERBOUNDARY
LANL NameBoundPlasma
T89C NameBoundMag

#VARIABLEDT
T       DoVariableDt

#STOP
-1      MaxIteration
300     tSimulationMax
```

First, take note that the parameter values can take many forms: rational numbers, integers, logical values, or strings of text. Remember that the trailing descriptions are not read by RAM-SCB; they are there to help you remember what each value does. In order, this list of `PARAM` commands sets the start date and time of the simulation, sets the plasma and magnetic field outer boundary conditions, turns on variable timestepping, and tells the code to run for unlimited iterations but only 300 seconds. All commands are described in detail in Chapter 9.

Once your `PARAM` file is customize, execute RAM-SCB.

```
./ram_scb.exe
```

RAM-SCB will begin; a lot of information will be written to screen that updates the user of its progress. Experienced Unix users know all of the tricks to capture that output; the following copies the output to file while preserving the output to screen:

```
./ram_scb.exe | tee runlog.txt
```

5.1 Required Inputs

Ring current models have three basic requirements: plasma fluxes at the outer boundary, magnetic field, and electric field throughout the equatorial plane. RAM-SCB fulfills these requirements slightly differently than other ring current models thanks to its sophisticated self-consistent magnetic field. Electric field may be specified either in the equatorial plane directly or as an ionospheric potential that is mapped along SCB field lines. Magnetic field is supplied as a shell surrounding the SCB domain, which consists of the body containing all field lines passing through the RAM equatorial domain. Plasma fluxes remain a simple specification about the outer boundary of RAM. In RAM-SCB, there are many ways to specify these; refer to Chapter 9 for commands that select inputs.

Plasma fluxes can be provided one of two ways: the first is coupling with the SWMF, covered in Chapter 7. The second is to use "GeoMlt" files, a product of LANL geosynchronous observations. These files can be obtained on request from the data caretakers; sample files are provided with the distribution. These files come in pairs per day; an ion and electron file exists for each day for which there is data coverage. Because the source observations do not readily differentiate between different ion species, the $K_P/F_{10.7}$ dependent empirical relationship of *Young et al.,* [1982] is used internally to divide up the Hydrogen flux in the input files. The K_P and $F_{10.7}$ indices are packaged in the RamIndices.txt input file; the user need not worry about them for historical simulations.

Convection electric potential can be specified in a plethora of ways. The simplest is to use the K_P dependent Volland-Stern electric field. This requires no additional input files beyond RamIndices.txt. The Weimer 2000 empirical model can also be used; like Volland-Stern, it is an internal calculation. Because

this model depends on solar wind conditions, users must supply an input file that contains this data from the OMNI database. The Weimer 2000 potential is mapped along SCB field lines to the equatorial plane automatically. Finally, SWMF electric fields can be used; this is covered later. Though it is possible to supply Weimer 2000 and SWMF potentials in the equatorial plane without SCB mapping, this approach requires extensive work by the user as is not recommended. See the entry on the command `#EFIELD` for additional details on file formats and obtaining these inputs.

Magnetic field boundary conditions can be supplied via three sources: a simple dipole model, from one of the Tsyganenko empirical models, or from the SWMF. When using a dipole, it is possible to use a dipole with or without the SCB calculation. Tsyganenko 89 (T89) input files are provided with the distribution and require no additional input from the user. Input from more complex Tsyganenko models require substantial work to acquire, they are not recommended for most users. See the entry for the `#OUTERBOUNDARY` command in Chapter 9 for more details.

All files listed above must be placed in their respective input files in the run directory. Other input files required by advanced options (e.g. virtual satellites) should be placed in the top level of the run directory. Each run directory requires its own set of inputs.

5.2 Restarting Simulations

It is often desirable to split a simulation up into several parts. For example, you may want to split a very long simulation up into separate executions, or change parameters at certain parts of a storm, or perhaps a simulation was interrupted undesireably. Restart files allow you to continue a simulation seamlessly.

Throughout a simulation, restart files are being written to the `restartOUT` folder of the run directory. The files are written at a set frequency and at the end of a successful run (both of these options are configuarble through the `PARAM` interface.) Only the most recent restart is saved, users can periodically pull these files out of `restartOUT` if they prefer a back log of restarts. These files contain the full distribution functions as well as the start time, current time, and current iteration at the point that the restart was written.

To restart a run, first move restart files from `restartOUT` to `restartIN`. Then, either create a new `PARAM.in` file or edit the existing one such that no `#STARTTIME` command is present (remember, the restart file knows this information already), the run time exceeds the current time listed in the restart file, and the command `#RESTART` is included. This tells RAM-SCB to restart the simulation rather than start anew. Restarts are carefully implemented to preserve files in the output directories and continue the simulation as if there was no interruption.

For an example of restarting a simulation, run Test 2 and inspect the `PARAM.in` files as well as the contents of `restartIN` and `restartOUT`.

5.3 Output

RAM-SCB, by default, has a rich output set. This can be expanded by activating other output file types via the `PARAM` interface, be sure to review the commands listed in Chapter 9. Table 5.1 summarizes the output that can be generated by RAM-SCB. Visualization of these files is covered in Chapter 6.

Type	Extension	Format	Default?	Contents
Log file	`*.log`	ASCII	Yes	Dst, integrated values
Pressure files	`*.in`	ASCII	Yes	$\perp$ and $\parallel$ partial pressures
Trapped files	`*.t`	ASCII	Yes	Averaged trapped fluxes
E-Field files	`*.in`	ASCII	Yes	Equatorial electric potential
Boundary files	`*.dat`	ASCII	Yes	Fluxes at the outer boundary
Full flux files	`*.nc`	NetCDF	No	Full equatorial flux information
3D flux files	`*.nc`	NetCDF	No	Fluxes in the full 3D domain
Virtual Satellites	`*.nc`	NetCDF	No	Satellite specific values
Field integral files	`*.dat`	ASCII	Yes	h, I geometric integrals
Magnetic field file	`*.nc`	NetCDF	No	Full 3D SCB field
Potential file	`*.nc`	NetCDF	No	Ionospheric and equatorial electric potential
3D pressure file	`*.nc`	NetCDF	No	Full 3D anisotropic pressure

Table 5.1: List of available output files.

Chapter 6

Output Visualization

RAM-SCB output can be visualized in a number of different ways depending on the user's tastes and preferences. However, a standard library for opening, manipulating, and plotting exists as a sub-module of Spacepy. Users may find this freely available Python library at spacepy.lanl.gov. Alternatively, development versions can be obtained via the project's Source Forge page. The RAM-SCB module is nested under `spacepy.pybats.ram`.

Stand-alone code snippets for both Python and Matlab can be found in the viz directories. These routines are largely out-of-date and not maintained. We recommend caution when using them.

Chapter 7

RAM-SCB and SWMF

RAM-SCB may be used as a component to the Space Weather Modeling Framework. Under this mode, RAM-SCB receives initial and boundary conditions from the Framework's other components and returns plasma properties to create a two-way coupled system.

7.1 Installing RAM-SCB as a SWMF Component

Obtain a copy of the SWMF and, if necessary, unpack the tarball. Descend into the Inner Magnetosphere (IM) directory of the Framework. This is where all inner magnetosphere-type codes used by the Framework are located. Copy the entire RAM-SCB directory here as `RAM_SCB`. It is important to name the directory correctly. CVS users should note that checking out RAM-SCB directly into the SWMF directories will be tricky because of conflicting CVS files located in each Framework directory. Use CVS carefully!

Next, move or remove the share directory. RAM-SCB will be using the version obtained through the SWMF. Additionally, making RAM-SCB's share directory unavailable signals RAM-SCB to go into component mode.

Installing the SWMF and RAM-SCB happens concurrently using the Config.pl script from the top-level directory of the SWMF.

Chapter 8

Included Scripts

RAM-SCB comes packaged with many helpful scripts that aid the user. These scripts are written either in Perl or Python with the standard libraries only. This allows them to be run on many systems without needing to install additional software as both languages are ubiquitous in Unix-like environments. The subdirectory `Scripts` contains several helpful scripts that are unique to RAM-SCB while the subdirectory `Share/Scripts` contains scripts designed for SWMF-like applications but often useful for RAM-SCB as well.

All scripts are designed to be called from the command line and have a common help feature:

```
ScriptName -h
```

will print out the script's help text. This info is typically more complete and up-to-date than the info listed here.

While all RAM-SCB scripts are listed here, only a handful of SWMF scripts are described. Be sure to explore the `Share/Scripts` directory before you constructing your own solutions to common problems (e.g. endian conversions, PARAM-checking, etc.)

8.1 Config.pl (SWMF Script)

`Config.pl` is part of the SWMF *Config* system for installing and pre-configuring RAM-SCB in both stand-alone and component modes. Use of `Config.pl` is covered extensively in Chapter 3.3.

8.2 DiffNum.pl (SWMF Script)

`DiffNum.pl` is a powerful, quantitative re-write of the popular `diff` utility. It compares two files, finds and quantifies differences in any numerical entries, and, if any are found, lists the differences and raises an exception. The main purpose of `DiffNum.pl` is to find and quantify failures in RAM-SCB tests.

Usage:

```
DiffNum.pl [options] File1 File2
```

Common options include `-a=VALUE` and `-r=VALUE`, which allow the user to ignore absolute and relative differences less than `VALUE` and `-t` which turns off the comparison of text.

8.3 CatLog.py (RAM-SCB Script)

A common problem in both RAM-SCB and many SWMF modules is many fractured, separate log files from a single simulation that required several restarts. Often, these log files overlap in time because a simulation

did not complete and restarting results in re-simulating a small portion of the run. Manually concatenating these log files together into a single seamless, monotonic file can be time consuming.

`CatLog.py` concatenates many log files into a single file. It assumes that the first column of the log file is either run iteration or run time and uses this info to check for and remove overlapping entries.

Usage:

```
CatLog.py [options] log1 log2 [log3] [log4]...[logN]
```

Files 2 through N will be appended to the first file. Unix wild-card characters can be used to get file-globbing effects. If the headers of any of the trailing log files does not match the leading file, it is discarded. If the leading file includes a wild-card character, the files are arranged and appended in alpha-numeric order. Available options include `-debug` (print debug information), `-rm` (remove all but first log file), and `-nocheck` (deactivate checking for overlapping entries.) See `CatLog.py -h` for examples.

Chapter 9

Complete List of Input Commands

The content of this chapter is generated from the PARAM.XML file. The XML file can be read with an editor and can be used for creating PARAM.in files by copying small parts from them.

The transformation of the XML format into LaTex is done with the share/Scripts/XmlToTex.pl script. This script generates index terms for all commands, which are used to create an alphabetical index at the end of this chapter.

9.1 Input Commands for the RAM_SCB: IM Component

List of IM commands used in the PARAM.in file

9.1.1 Operational mode

#SHIELDSRC command

```
#SHIELDSRC
```

This command activates **SHIELDS-RC** mode, an operational version of RAM-SCB and the Ring Current portion of the SHIELDS project. This mode is best suited for long term simulations with operational implications (e.g. monitoring D_{st}, satellite-specific charging environments, etc.)

In this mode,

- SCB is deactivated and a simple dipole magnetic field is used.
- K_P-based Volland-Stern electric field is used.
- Only LANL Geosynchronous flux files can be used for the plasma outer boundary conditions.
- No full-domain flux files are saved.
- Log file write frequencies are lowered to five minutes.
- Variable time steps are used.
- Output files use date/time naming scheme to prevent overwriting.

These options can NOT be altered in SHIELDS-RC mode.

9.1.2 Stand alone

These commands should only be used in Stand-Alone mode. Many, if not all, have analogs in the SWMF PARAM system. If that is the case, RAM-SCB will get values from those commands.

#DESCRIPTION command

```
#DESCRIPTION
Simulation of the Sep. 1st, 2005 storm.
```

The StringDescription string can be used to describe the simulation for which the parameter file is written. The #DESCRIPTION command and the StringDescription string are saved into the restart file, which helps in identifying the restart files. It is often added to NetCDF files to help describe the output.

#STARTTIME command

```
#STARTTIME
2006                    iYear
7                       iMonth
19                      iDay
0                       iHour
0                       iMinute
0                       iSecond
0.0                     fracSecond
```

The STARTTIME command sets the integer year, month, day, hour, minute and second at which the simulation begins. This command is only used in standalone mode and only for the first session.

#RESTART command

```
#RESTART
```

If present, RAM-SCB will restart a simulation using the information stored in the restart files that the user must place in the `restartIN` folder in the run directory. Start time and simulation conditions are read from the restart files. This command can not be used with `#STARTTIME`.

See Chapter 5.2 for details on restarting simulations.

#STOP command

```
#STOP
100                     MaxIteration
10                      tSimulationMax [sec]
```

This command is only used in stand alone mode.

The MaxIteration variable contains the maximum number of iterations *since the beginning of the current run* (in case of a restart, the time steps done before the restart do not count). If nIteration reaches this value the session is finished. The tSimulationMax variable contains the maximum simulation time relative to the initial time determined by the #STARTTIME command. If tSimulation reaches this value the session is finished.

Using a negative value for either variables means that the corresponding condition is not checked. The default values are MaxIteration=0 and tSimulationMax = 0, so the #STOP command must be used in every session.

9.1.3 Numerical scheme

Options for configuring the schemes and solvers, boundary conditions, and other numerical options.

#OUTERBOUNDARY command

```
#OUTERBOUNDARY
LANL            NameBoundPlasma
T89C            NameBoundMag
MAXW            NameDistribution
```

Set the outer boundary conditions for RAM-SCB. Based on these settings, different files are expected in the respective input directories to suppy magnetic field and energy flux at the outer boundary of the codes.
Current options for plasma boundary conditions:

1. **LANL**: Use fluxes calculated from LANL Geosynchronous measurements. Composition is determined from the Young et al. empirical relationship based on Kp and F10.7.
2. **SWMF**: Use fluxes calculated from the Space Weather Modeling Framework. If in coupled mode, values are calculated on-the-fly and no input files are required.
3. **TM03**: Use fluxes calculated from the Tsyganenko-Mukai plasma sheet empirical model.

Current options for magnetic field boundary conditions:

1. **DIPS**: Use a simple dipole to constrain the SCB field.
2. **DIPL**: Use a simple dipole throughout; no SCB calculation.
3. **LANL**: Use LANL geosynchronous magnetic field files.
4. **SWMF**: Use the magnetic field from the Space Weather Modeling Framework. If RAM-SCB is in coupled mode, these values are calculated and obtained on-the-fly rather than read from input files.
5. **T89C**: Use the Tsyganenko 89c empirical model. As this field depends only on Kp, input files are provided in the RAM-SCB distribution.
6. **TS04**: Use the Tsyganenko 2004 empirical model.

Current options for plasma distribution shape:

1. **MAXW**: Assume a Maxwellian distribution.
2. **KAPA**: Assume a Kappa distribution.

#MULTISPECIESBCS command

```
#MULTISPECIESBCS
F       DoMultiSpeciesBcs
F       DoElectrons
```

In standalone mode, it may be desireable to still take advantage of multispecies boundary conditions from the Space Weather Modeling Framework through the use of several .swf files (one for each species desired.) Use this command to force this behavior. Note that it disables the application of Young et al. composition specification.

In coupled mode, or when the outer boundary plasma conditions are not set to SWMF, this command has no effect.

The parameter DoElectrons sets the calculation of the electron distribution. Note that the electron version of RAM is still under development.

#EFIELD command

```
#EFIELD
IESC                    NameEfield
F                       UseEfInd
```

Set the source for the convective electric field in RAM-SCB. The choice made will set the type of input file required at runtime.

Current options for the electric field:

1. **IE89**: SWMF IE-component electric field mapped to the equatorial plane via T89c field lines.

2. **IESC**: SWMF IE-component electric field mapped to the equatorial plane via RAM-SCB field lines.

3. **VOLS**: K_P-based Volland-Stern empirical electric field (internal VS calculation).

4. **WE01**: Weimer 2001 empirical electric field specified at the equatorial plane by user (read from file).

5. **WESC**: Weimer 2001 empirical electric field mapped to the equatorial plane via RAM-SCB field lines (internal W2K calculation).

The parameter UseEfInd turns the use of induced electric field on or off. Default is no induced electric field.

The **IE89** and **WE01** options require an additional input file for every five minutes of simulation time and at the start time. These files are three-column ascii files that give the potential at each RAM grid point (Local Time - L-Shell combination).

The **VOLS** and **WESC** are *internal* calculations and do not require these additional files, but carry the requirements of their respective underlying models. The Volland-Stern model requires the K_P index, which is provided for historical simulations. The Weimer 2000 empirical model requires upstream solar wind conditions, which can be obtained from the OMNI database. This data must be placed into the run directory in a file named *omni.txt*.

Note that these input files can be obtained via this web form.

#USESCB command

```
#USESCB
T       DoUseScb
F       DoWriteB
F       DoWriteCur
```

This command toggles the use of the 3D Equilibrium Magnetic field code. Activation of the 3DEQ creates the Self-Consistent Magnetic Field (SCB) version of RAM-SCB. Default is to use 3DEQ. If switched off, be sure to have the hI files available to RAM (from any source you choose) in the IM/output_scb directory.

#USEPLANE_SCB command

```
#USEPLANE_SCB
F       DoUsePlane_SCB
```

Turn on plasmaspheric density calculation based on an updated *Rasmussen et al.* [1993] model. Further sub-options for the plasmasphere model are specified in ModRamPl_Ne.f90.

#USEWPI command

```
#USEWPI
F       DoUseWPI
F       DoUseBASdiff
F       DoUseKpdiff
```

Flag to turn on pitch angle diffusion. The default setting for the Kp-based version uses Kp=0 coefficients. Turning this flag on interpolates Kp-dependent diffusion coefficents for Kp in the options are available if selecting DoUseBASDiff.

#VARIABLEDT command

```
#VARIABLEDT
T       DoVariableDt
```

Select between a variable timestep that is governed by the CFL number or a semi-static timestep that only decreases during periods of high geomagnetic activity as determined by the Kp index. The former will allow the timestep to grow as large as the CFL number will permit, while the latter will use a timestep of 5 seconds during low (less than 5) Kp or 1 second during higher Kp.

#RAMLIMITER command

```
#RAMLIMITER
1.5     BetaLim
```

Set the beta factor for the MC limiter implemented in RAM. A value of 1 is equivalent to using the Min-Mod limiter; a value of 2 is equivalent to the superbee limiter. 1.5 is the default; many codes that implement this limiter use a value of 1.5.

#MAXTIMESTEP command

```
#MAXTIMESTEP
5.0     MaxHalfStep
```

Set the maximum half time step that can be taken by RAM. Even if the CFL limit allows for a larger time step, the code will not surpass the limit set here. Note that this is the *half* time step taken in the time-splitting scheme; a full timestep is two times the half time step.

Default vaule is 5 seconds.

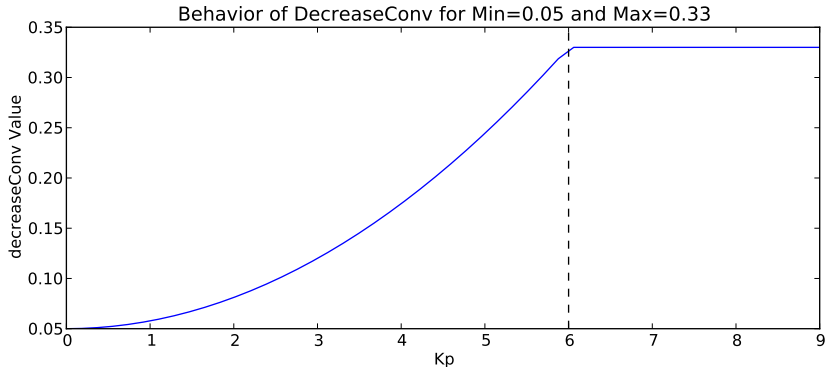
#SCBSCHEME command

```
#SCBSCHEME
0.05    DecreaseConvAlphaMin
0.33    DecreaseConvAlphaMax
0.05    DecreaseConvPsiMin
0.33    DecreaseConvPsiMax
0.20    BlendAlpha
0.20    BlendPsi
```

Set options for the 3D-Equillibrium solver. Default values are 0.05 for all DecreaseConv-type variables, 0.20 for all Blend variables. Alpha and Psi refer to the different Euler potential surfaces. DecreaseConv variables set the maximum change in the solution between subsequent iterations before it is assumed that convergence

has been reached. Blending factors control the amount of blending of the previous solution with the current one.

If DecreaseConv* max and min values are not the same, the values are scaled as a function of Kp from the minimum to the maximum. This increases code stability as the magnetic field becomes more complex with increased activity. Scaling is parabolic up to Kp=6.0; see figure below. Because default Max and Min are identical, no activity-based scaling occurs until the defaults are changed.



9.1.4 Input/output

#NAMEFORMAT command

```
#NAMEFORMAT
F        UseNewFormat
```

Historically, RAM-SCB output files have the format *prefix_IIII.suffix* where *IIII* is a four digit integer that gives the number of intervals passed. The interval is file specific, often set to five minutes or hours. This format is simple but useful for marking time passed since the beginning of the run when examining a single event. When performing long simulations, however, this format is not only unwieldy but unfeasible as it ultimately limits the number of files that can be written.

This command toggles the output file name format from the legacy format described above to *prefix_dYYYYMMDD_tHHMNSS.suffix*, where the simulation date and time are used to organize and differentiate individual files of the same *prefix* and *suffix*. The default is to use the legacy system.

Developer's Note: Module `ModRamIO.f90` contains tools to build file names of both types with minimal effort.

#SAVERESTART command

```
#SAVERESTART
3600.0           DtSaveRestart
T                DoSaveFinalRestart
```

Configure when restart files are saved. The first parameter sets the frequency, in seconds, that restart files are saved during a simulation. The second parameter toggles saving restarts at the successful completion of a simulation. Defaults are shown.

#LOGFILE command

```
#LOGFILE
60.0     DtWrite
```

Specify the write frequency of the RAM-SCB log file. Default value is shown. Note that frequent writes for any output file may limit the maximum timestep taken by RAM-SCB.

#SAVEFLUX command

```
#SAVEFLUX
F       DoSaveFlux
```

Toggle writing of equatorial flux values for all pitch angles and energies to NetCDF file. These files are powerful research tools but large and may be undesirable for long simulations. Default is to not write flux files.

#DUMP3DFLUX command

```
#DUMP3DFLUX
F       DoDump3dFlux
```

Toggle the writing of the full 3D (spatial), pitch angle, energy, and species dependent flux from RAM-SCB. Default is to not write these values.

Activating this capability comes with heavy caveats due to the immense size of these output files. First, ensure that you have sufficient disk space as each file is on the order of tens of gigabytes. Second, be sure to compile on a 64-bit machine in a 64-bit compatable mode. The arrays being written are so massive that the NetCDF library will choke if it is not in 64-bit mode. Finally, be ready for unexpected behavior due to the large amount of data being processed to file. At current time, this capability is best reserved for debugging purposes. Satellite tracing provides a much more compact way to sample the distribution function at discrete points.

#OMNIFILE command

```
#OMNIFILE
omni.txt        NameOmniFile
```

The WESC (Weimer electric field traced along SCB field lines) electric field selection calculates Weimer's empirical electric field on-the-fly. To do this, solar wind inputs are required from the Omni database. The ascii file that contains these inputs should either be called "omni.txt" and be located in the run directory (default behavior) or this command should be used to point the code in the correct location.

#SATELLITE command

```
#SATELLITE
60.0            DtOutput
2               nSatellite
satellite1.dat  NameTrajectoryFile
satellite2.dat  NameTrajectoryFile
F               DoUseVAPini
```

The solution of RAM-SCB can be extracted along a satellite's path as the simulation progresses by using the #SATELLITE command. Simply set the number of virtual satellites to be included in the simulation, the time frequency to write the output in seconds (minimum/default is 10), and the location of each corresponding Trajectory File. For each Trajectory File given, RAM-SCB will produce a NetCDF file containing the solution along the satellite's trajectory at a time frequency of sixty seconds. Note that reducing DtOutput to small values can reduce the maximum RAM-SCB time step as a result of ensuring that the file is written every DtOutput.

Satellite Trajectory Files contain the trajectory of the satellite. They should have the following format:

```
#START
 2004  6  24   0   0  58   0  2.9  -3.1 - 3.7
 2004  6  24   0   1  58   0  2.8  -3.2 - 3.6
```

The file containing the satellite trajectory should include data in the following order:

```
yr mn dy hr min sec msec x y z
```

with the position variables in units of the body radii or the length scale normalization. Note that this is the same format as BATS-R-US trajectory files with one important feature: the position is assumed to be in SM coordinates. If the input coordinates are in another coordinate system, use the #COOR command to specify what system is used in the file. RAM-SCB will convert from that system to SM.

The maximum number of satellite files allowed is 100. The maximum number of lines in a given satellite file (past the #START command) is 100,000. Both of these can be changed by editing the value of MaxRamSat and MaxRamSatLines (respectively) in ModRamSats.f90

Each satellite listed under the #SATELLITE command will produce a NetCDF file in the output_ram folder. Simulation meta data, such as run parameters, start times, etc., are contained within this file as well as energy and pitch angle grid information. Magnetic field, flux, spacecraft position, and other values are saved for every DtOutput for which the satellite is within the 3D Equilibrium code's domain. Each record written takes about ten kilobytes of hard drive space.

9.1.5 Testing

#TEST command

```
#TEST
CON_session::do_session         StringTest
```

A space separated list of subroutine and function names to be tested. Only subroutines containing the 'call CON_set_do_test(...)' statement can be tested. The first argument is the name of the subroutine, usually defined as the string parameter 'NameSub'. Default is an empty string.

This feature and its conventions has been adopted from the Space Weather Modeling Framework.

9.1.6 Other

#INCLUDE command

```
#INCLUDE
Commands.in     NameIncludeFile
```

The NameIncludeFile parameter contains the name of the file to be included. This file can contain any additional PARAM commands and will be included as if the text of NameIncludeFile was located within PARAM.in place of the #INCLUDE command which points to it. This is useful for organizing large PARAM.in files into separate, smaller chunks. The file name may be followed with a trailing comment if it is separated with at least 3 spaces or one TAB character. The #INCLUDE command can be used anywhere in the parameter file, even in the sections which contain the component specific parameters.

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Chapter 10

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