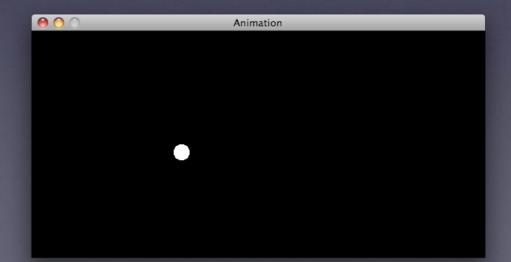
- processing makes animation very easy by using the function draw ()
- If we define the draw() function in our application, it will be continuously executed in a constant loop (in a different thread)
- Initially, draw() repetition rate is fixed to 60 times per second This behaviour can be chnaged with frameRate() function
- frameRate() function establish an objective or target, but
- to achieve that depends on the possibilities of the machine in executing the draw() function at the specified framerate

```
// A classic example:
// A bouncing ball
// If we want animation we have to
// define setup() and draw() functions
// Position
int px, py;
// Velocity
int vx, vy;
// Ball diametre
int diametre = 20;
void setup()
  size(600, 300);
  fill(255);
 noStroke();
  // Ball initial
  // position
  px = width/4;
  py = height/2;
  // Initial velocity
  vx = vy = 1;
```

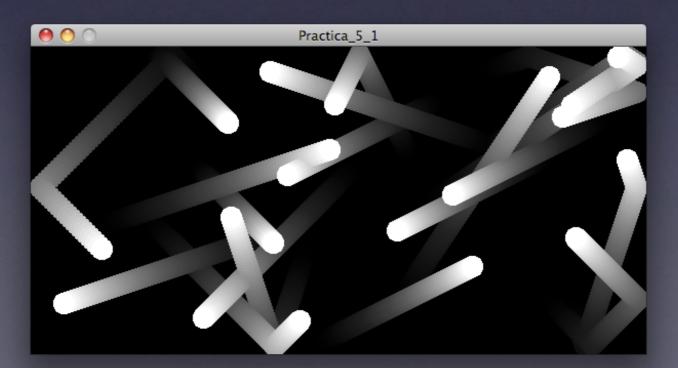
```
// Draw
void draw()
 background(0);
  // Detecting collisions and
  // bouncing
  if (px + diametre/2 > width - 1 ||
     px - diametre/2 < 0)
   vx *= -1;
  if (py + diametre/2 > height - 1 ||
     py - diametre/2 < 0)
   vv *= -1;
  // Updating positions
 px += vx;
 py += vy;
  // Drawing the ball
  ellipse(px, py, diametre, diametre);
```



```
// Fading effect...
// By substituting background(0)
// with the following:
void draw()
{
     // We fill all the window
     // with the RGB color (0, 0, 0) and
     // transparency 20
     fill(0, 20);
     rect(0, 0, width, height);
     // Fill to 255 to draw
     // the ball
                                    Animation
     fill(255);
```

Practice 5-1

- Modify the previous application to visualise n boucing balls.
 Use arrays to manage the velocity and position values of each ball
- Initial positions should be random taking into account the diameter of the balls and the window size
- Velocities should also be random with values from -4 to 4 (zero not included)



Considering gravity force:



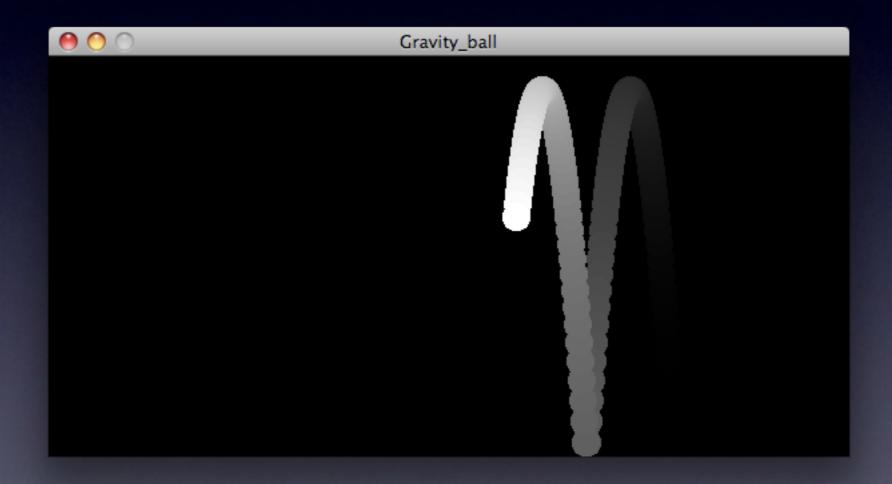
Euler integration:

$$a = g$$

 $v(t+1) = v(t) + a e(t + 1) = e(t) + v(t+1)$

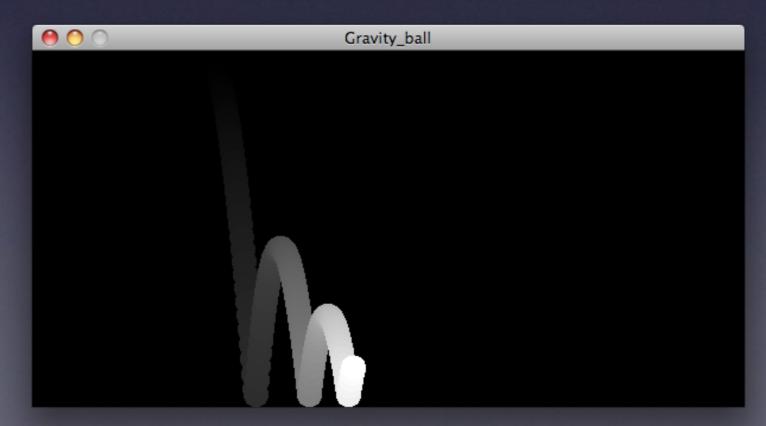
```
// We consider the
// effect.
// Position
float px, py;
// Velocity
float vx, vy;
// Ball diameter
int diametre = 20;
// gravity
float gravity = 0.5;
void setup()
  size(600, 300);
  fill(255);
  noStroke();
  // Ball initial position
  px = width/4;
  py = diametre/2;
  // Initial velocity
  vx = vy = 1.0;
  // Initial background
  background(0);
```

```
void draw()
  fill(0, 20);
  rect(0, 0, width, height);
  fill(255);
  // Detecting collisions and
  // bouncing
 if (px + diametre/2 > width - 1)
   vx *= -1;
   px = width - 1 - diametre/2;
  if (px - diametre/2 < 0)
   vx *= -1;
   px = diametre/2;
  if (py + diametre/2 > height - 1)
   vy *= -1;
   py = height - 1 - diametre/2;
  if (py - diametre/2 < 0)
   vy *= -1;
   py = diametre/2;
  // Updating positions
  vy += gravity;
  px += vx;
  py += vy;
  // Drawing
  ellipse(px, py, diametre, diametre);
```

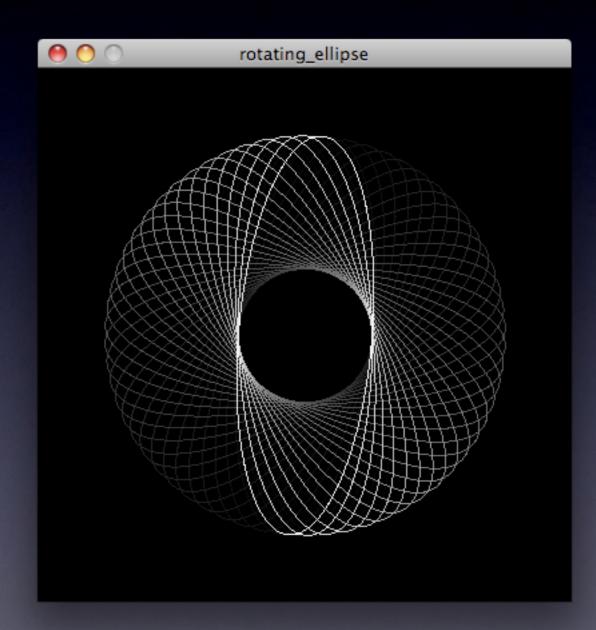


- To achieve a similar behaviour to the real one, the damping effect due to the pass of the ball through a fluid (atmosphere) has to be considered
- This effect is achieved just by damping the velocity down, as follows:

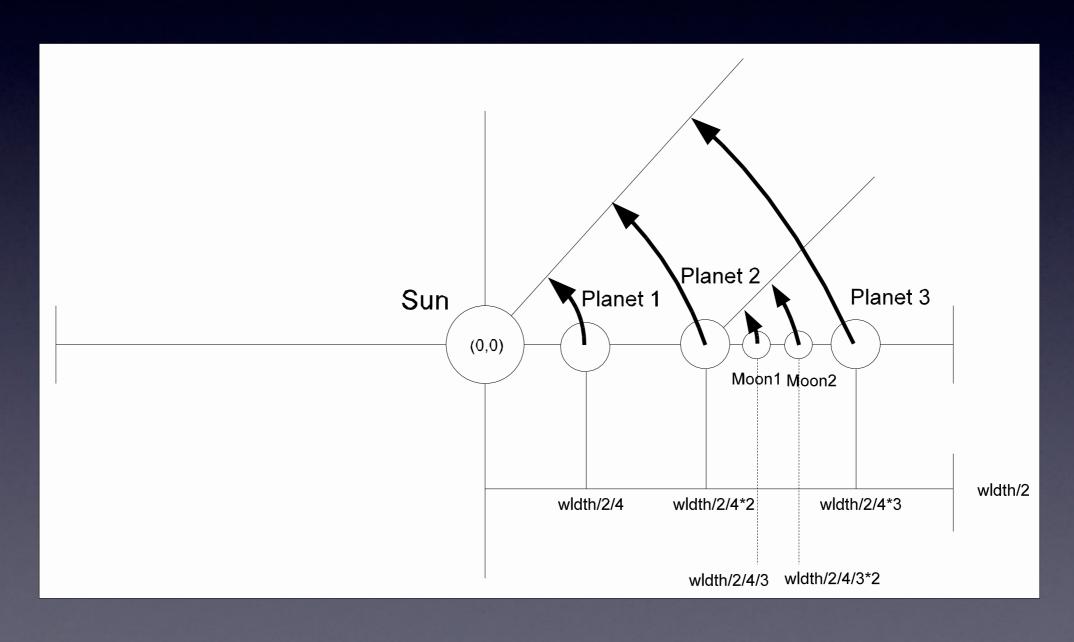
```
// We update positions
vy += gravedad;
vy *= 0.98; // This is new !
px += vx;
py += vy;
```



```
// An ellipse rotating continuously
float angle = 0.0;
void setup()
 size(400, 400);
// Look out !
// draw() initiates model/view
// matrix to the identity
void draw()
 fill(0, 20);
  noStroke();
  rect(0, 0, width, height);
 noFill();
  stroke (255);
  // We rotate around the middle
  // of the window, so we have to
  // translate to (0,0).
  // We must take into account
  // the order of the operations
  // (from bottom to top)
  translate(width/2, height/2);
  rotate(angle+=0.1);
  translate(-width/2, -height/2);
  ellipse(width/2, height/2, 100, 300);
```

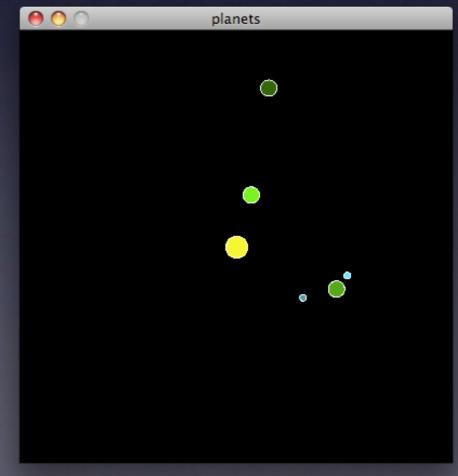


A solar system:



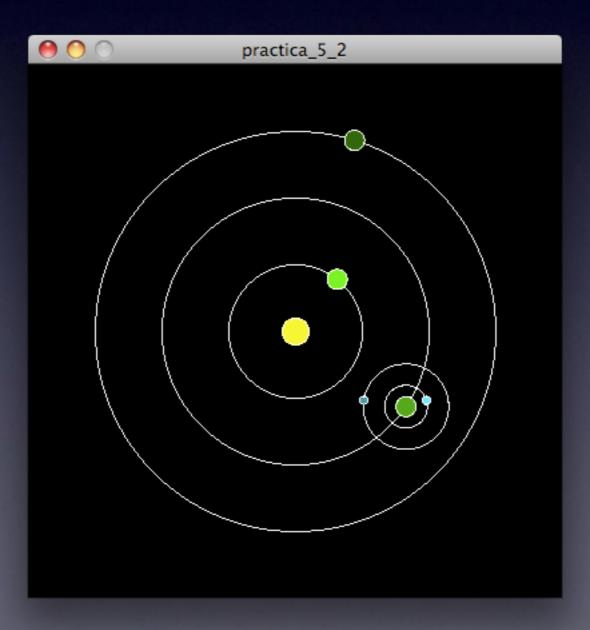
```
// A planetary system
// in 2D
// A sun, 3 planets and two moons
float angPlanet1 = 0.0,
      angPlanet2 = PI/3.0,
      angPlanet3 = 2.0*PI/3.0,
      angMoon1 = 0.0,
      angMoon2 = PI;
void setup()
  size(400, 400);
  stroke (255);
  frameRate(30);
void draw()
  background(0);
  // We will draw everything centered
  // at (0,0) and will solve
  // their final positions through
  // 2D transformations
  // The sun in the center of
  // our universe
  translate(width/2, height/2);
  // Sun
  fill(#F1FA03); // Hex. using color selector
  ellipse(0, 0, 20, 20);
  pushMatrix();
  // Planet 1
  rotate(angPlanet1 += 0.1);
  translate(width/2/4, 0);
  fill(#05FA03);
  ellipse(0, 0, 15, 15);
```

```
// Planet 2
popMatrix();
pushMatrix();
rotate(angPlanet2 += 0.05);
translate(width/2/4*2, 0);
fill(#0BA00A);
ellipse(0, 0, 15, 15);
// Moon 1
pushMatrix();
rotate(angMoon1 += 0.1);
translate(width/2/4/3, 0);
fill(#08E4FF);
ellipse(0, 0, 6, 6);
// Moon 2
popMatrix();
rotate(angMoon2 += 0.05);
translate (width/2/4/3*2, 0);
fill(#118998);
ellipse(0, 0, 6, 6);
// Planet 3
popMatrix();
rotate(angPlanet3 += 0.025);
translate(width/2/4*3, 0);
fill(#075806);
ellipse(0, 0, 15, 15);
```



Practice

 Modify the planetary system in order to draw the orbits of the planets and moons:



Practice

 Modify the previous application to allow the sun to orbit as well:

