

## Wave Optics

$$\rightarrow \Delta\phi = \frac{2\pi}{\lambda} \Delta x = \frac{2\pi}{\lambda} (\mu \Delta x) \text{ if med. other than air } \mu \neq 1$$

$$\rightarrow \vec{y}_{\text{res}} = \vec{y}_1 + \vec{y}_2 + \dots + \vec{y}_n, I_{\text{res}} = A \cos(\omega t + \phi) = R \cos(\Delta t + \phi)$$

$$\rightarrow \text{for coherent source } A = \sqrt{A_1^2 + A_2^2 + 2A_1 A_2 \cos\phi} \quad I \propto A^2 \propto (\text{intensity})^2$$

$$I = I_1 + I_2 + 2\sqrt{I_1 I_2} \cos\phi, \quad \text{for incoherent source } I = I_1 + I_2 + \dots + I_n$$

$\rightarrow$  When  $\Delta\phi = 0^\circ \text{ or } 360^\circ \rightarrow$  constructive interference

$$A_{\text{max}} = A_1 + A_2, I_{\text{max}} = [I_1 + I_2]^2$$

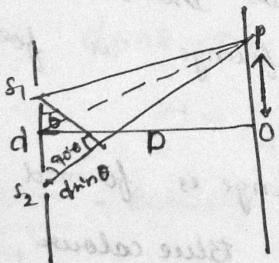
$\rightarrow$  When  $\Delta\phi = 180^\circ \rightarrow$  destructive interference

$$A_{\text{min}} = A_1 - A_2, I_{\text{min}} = [\sqrt{I_1} - \sqrt{I_2}]^2$$

$$\frac{A_{\text{max}}}{A_{\text{min}}} = \frac{A_1 + A_2}{A_1 - A_2} \Rightarrow \frac{I_{\text{max}}}{I_{\text{min}}} = \frac{[\sqrt{I_1} + \sqrt{I_2}]^2}{[\sqrt{I_1} - \sqrt{I_2}]^2} \frac{A_{\text{max}}^2}{A_{\text{min}}^2} = \frac{(A_1 + A_2)^2}{(A_1 - A_2)^2}$$

$$\rightarrow I_{\text{avg}} = I_1 + I_2 + 2\sqrt{I_1 I_2} \cos\phi \quad \text{but for long time } \langle \cos\phi \rangle \approx 0$$

YDSE



$$\Delta x = d \sin \theta \approx d \tan \theta = \frac{yD}{D} = \frac{yD}{d} = \frac{D^2}{d} - D$$

$$\Delta x = (\mu_r - 1)t$$

$$\Delta\phi = \frac{2\pi}{\lambda} (\Delta x)$$

$$R_{\text{eff}} = 2A \cos(\phi/2), \quad I_{\text{res}} = 4I_0 \cos^2(\phi/2)$$

$\rightarrow$  Bright fringe (max),  $\Delta\phi = 2n\pi, \Delta x = n\lambda, I_{\text{max}} = 4I_0$

$$y_n = \frac{n\lambda D}{d}$$

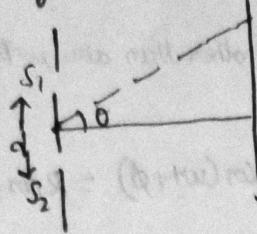
$\rightarrow$  Dark fringe (min),  $\Delta\phi = (2n-1)\pi, \Delta\phi = (2n-1)\pi, I_{\text{min}} = 0$

$$y_n = \frac{(2n-1)\lambda D}{2d}$$

$\rightarrow$  Band width ( $\Delta\phi$ ) fringe width  $\beta = \frac{\lambda D}{d}$

if YDSE conducted in other medium  $B_{\text{med}} = \frac{B_{\text{free}}}{\mu_r}$

Angular fringe width ( $\theta$ )



$$\theta = \frac{y}{d} = \frac{\lambda}{D} = \frac{P}{D}$$

$$\theta_{\text{med}} = \frac{\theta_{\text{free}}}{n}$$

$b$  = size of object  
 $D$  = distance of object  
 $\lambda$  = wavelength of light

- If  $\frac{b^2}{\lambda D} \gg 1$  then geometrical optics is satisfied.
- Validity of Ray optics  $Z_f = \frac{a^2}{\lambda}$
- If  $\frac{b^2}{\lambda D} < 1$  or  $\frac{b^2}{\lambda D} \approx 1$  then wave optics is satisfied.
- If  $n_1$  fringes are observed with light of wavelength  $\lambda_1$  &  $n_2$  fringes are observed with light of wavelength  $\lambda_2$  (bright bands)

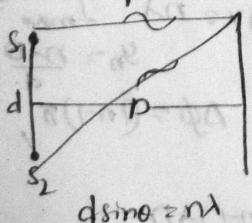
$$n_1 \lambda_1 = n_2 \lambda_2 \quad n_1 \beta_1 = n_2 \beta_2$$

$$\text{min. } n_1 \lambda_1 = (2n_2 - 1) \lambda_2$$

→ If white light rays is used in YDSE then all colours travels equally towards the centre of screen & hence white coloured fringe is formed at centre of screen.

→  $\lambda_R > \lambda_B$ . Red colour fringe is formed far away from centre of screen. Blue colour fringe is visible nearer to centre of screen.

→ Screen placed  $11\ell$  to plane of slits

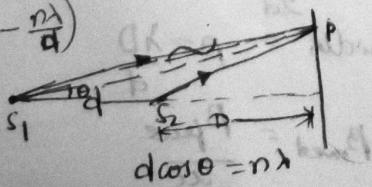


If point sized sources are used hyperbolic fringe is formed

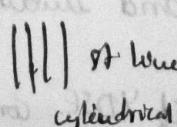


If cylindrical size

$$= D \sqrt{2(1 - \frac{n}{a})}$$



point-size - concentric



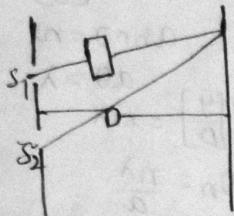
st line cylindrical

$$\rightarrow \text{Visibility (V)} = \frac{I_{\max} - I_{\min}}{I_{\max} + I_{\min}} = \frac{2\sqrt{I_1 I_2}}{I_1 + I_2}, 0 \leq V \leq 1$$

$I_{\max} = I_{\min} \Rightarrow V=0$  we can't observe dark & bright fringes clearly.

$I_{\min} = 0 \Rightarrow V=1$  we can observe dark & bright fringe clearly.

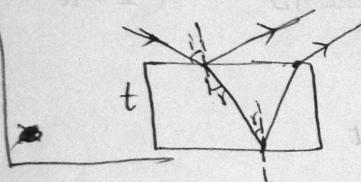
$\rightarrow$  Lateral displacement of fringes:



$$y = \frac{(u-1)tD}{d} = \frac{(u-1)t\beta}{\lambda}, \beta = \frac{(u-1)tD}{d}$$

$$\text{no. of fringes shifted } n = \frac{(u-1)t}{\lambda}$$

Inference in thin films:



$$2ut \cos r = \cancel{[n+1/2]\lambda} = [2m-1/2]\lambda$$

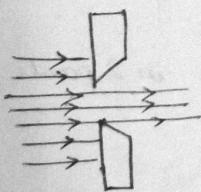
$$2ut \cos r = m\lambda \quad \begin{matrix} \text{max} \\ \text{min} \end{matrix} \Rightarrow m=0, 1, 2, \dots$$

$\rightarrow$  Finding no. of maxima in YDSE

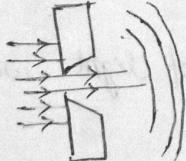
$$ds \sin \theta = n\lambda \Rightarrow n \sin \theta = \frac{n\lambda}{d}$$

$$-1 \leq \frac{n\lambda}{d} \leq 1 \Rightarrow -\frac{d}{\lambda} \leq n \leq \frac{d}{\lambda} \quad \begin{matrix} \text{minima is} \\ \text{formed b/w 2} \\ \text{maxima.} \end{matrix}$$

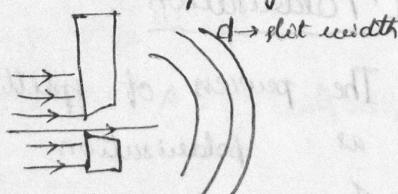
II. Diffraction - bending of light at sharp edges.



No diffraction  
 $\lambda \ll d$

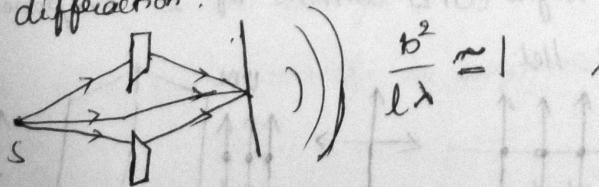


diffraction is  
noticeable  
 $\lambda \ll d \text{ or } \lambda \approx d$



diffraction is  
more  
 $\lambda > d$

$\rightarrow$  Fresnel diffraction:

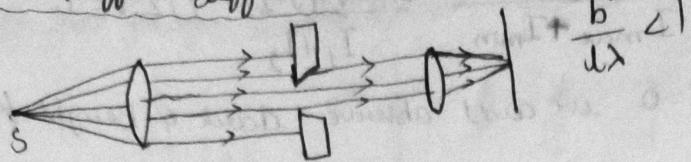


$$\frac{b^2}{l\lambda} \approx 1$$

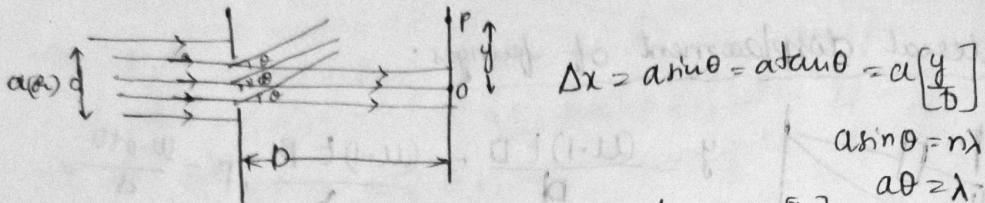
source & screen are  
very close

Mathematical approach is difficult.

Stein-Hoffer defecations



## Single slit diffraction



$$\Delta x = a \sin \theta = a \tan \theta = a \left[ \frac{y}{x} \right]$$

$$a \sin \theta = n\lambda$$

$$\sin \theta = \tan \theta = \alpha \left[ \frac{y}{P} \right] = n \lambda$$

$$\text{for } n^{\text{th}} \text{ order minima } y_n = \frac{n\lambda D}{a} \quad \theta_n = \frac{n\lambda}{a}$$

$$\text{for } n^{\text{th}} \text{ order maxima} \quad \alpha_{\text{max}} = \left[ \frac{2n+1}{2} \right] \lambda$$

$$y_n = \frac{(2n+1)}{2a} \lambda D \quad \theta_n = \left( \frac{2n+1}{2} \right) \frac{\lambda}{a}$$

$$\text{Resolving power} = \frac{1}{\text{Resolving limit}}$$

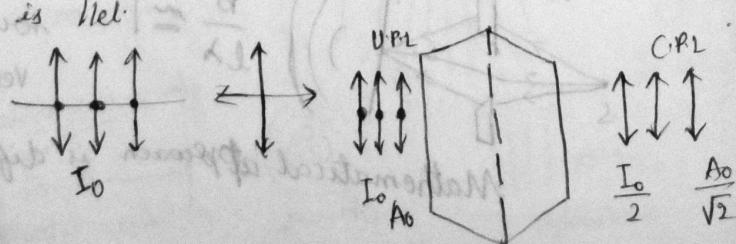
Human eye  $\theta = 1^\circ \Rightarrow$  limit

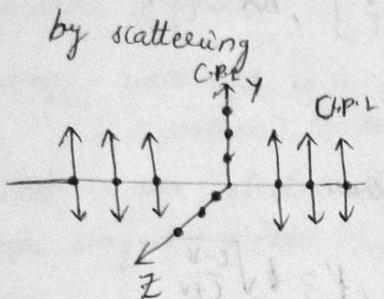
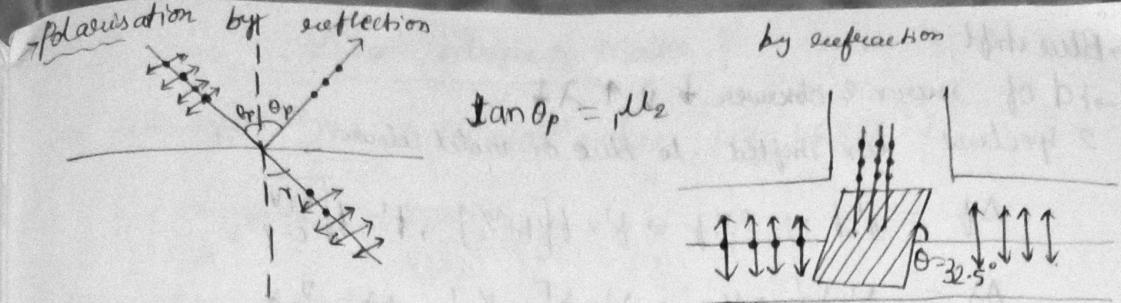
$$\text{Telescope} \quad \theta = \frac{1.22\lambda}{d}$$

$$\text{Microscope } \theta = \frac{0.6 k\lambda}{\text{mm dia}} \quad \text{if } \frac{600}{B} \geq 1$$

### III. Polarisation

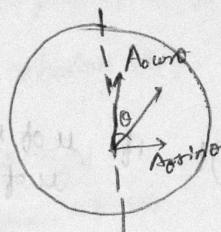
- The process of splitting of light wave in 2 parts is known as polarisation.
  - In electromagnetic wave the components of electrical vectors are mutually Iar.
  - Unpolarised light (UPL) consists of 2 components I & Lar





degree of polarisation depends on the angle of scattering

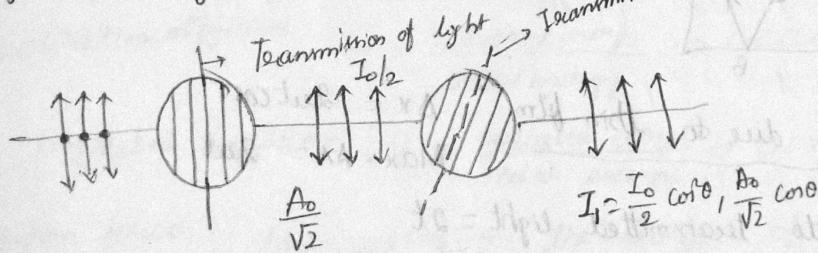
→ Effect of polariser on polarised light: only 1 polariser



$$I = \frac{I_0}{2} \quad A = \frac{A_0}{\sqrt{2}}$$



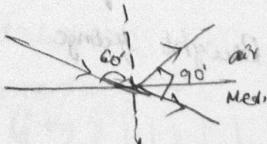
→ Effect of analyser on polarised light Transmission axis of analyser



$$I_1 = \frac{I_0}{2} \cos^2 \theta_1 \cos^2 \theta_2 \quad A = \frac{A_0}{\sqrt{2}} \cos \theta_1 \cos \theta_2$$

$$I_2 = \frac{I_0}{2} \cos^2(\pi - \theta)$$

Brewster's Law



Doppler effect

1. If d changes  $\nu, \lambda$  also changes.

2. Doppler effect in sound is asymmetric. In light is symmetric.

3. Only blue & red shifts are possible.

→ Blue shift

→ i) d of source & observer  $\rightarrow v \uparrow \lambda \downarrow$

2. spectral line shifted to blue & violet colour

$$\Delta f = f' - f = +\frac{v}{c} f \Rightarrow f' = f \left[ 1 + \frac{v}{c} \right], t' = t \sqrt{\frac{c+v}{c-v}}$$

$$\Delta \lambda = \lambda' - \lambda = +\frac{v}{c} \lambda \Rightarrow \lambda' = \lambda \left[ 1 + \frac{v}{c} \right], \Delta \lambda = \frac{v}{c} \lambda$$

→ Red shift

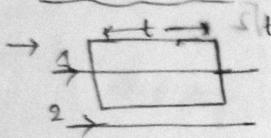
i) d of source & observer  $\uparrow v \downarrow \text{ & } \lambda \uparrow$

2. Spectral line shifted to red colour

$$① \Delta f = f' - f = -\frac{v}{c} f \Rightarrow f' = f \left[ 1 - \frac{v}{c} \right], t' = t \sqrt{\frac{c-v}{c+v}}$$

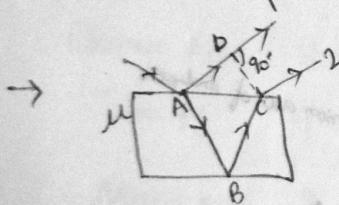
$$② \Delta \lambda = \lambda' - \lambda = -\frac{v}{c} \lambda \Rightarrow \lambda' = \lambda \left[ 1 - \frac{v}{c} \right]$$

→ Interference



$$\Delta x = (n_1 - 1)t$$

$$\Delta x = \left( \frac{n_2}{n_1} - 1 \right) t \quad \text{if } n_1 \text{ of slab} = n_2 \\ n_2 \text{ of med.} = n_1$$



$$t = n\lambda$$

⇒ YDSE due to thin film  $\Delta x = 2nt \cos r$   
Max.  $\Delta x = 2nt$

① Due to transmitted light  $= 2t$

i) Bright fringe  $2nt \cos r = n\lambda, n=1, 2, 3, \dots$

ii) dark fringe  $2nt \cos r = (2n-1)\lambda_b, n=1, 2, 3, \dots$

② Due to reflection

i) Bright fringe  $2nt \cos r + \lambda_b = n\lambda, n=1, 2, 3, \dots$

$2nt \cos r = (2n-1)\lambda_b, n=1, 2, 3, \dots$

ii) dark fringe  $2nt \cos r + \lambda_b = (2n+1)\lambda_b, n=0, 1, 2, \dots$

$2nt \cos r = n\lambda, n=1, 2, 3, \dots$

→ for scattering wavelength, and size of obstacle should be smaller than wavelength, and here