

# Atomic Physics

→ Rutherford's atomic model

No. of  $\alpha$ -particles scattered is  $N = \frac{Qnt Z^2 e^4}{(8\pi\epsilon_0)^2 r^2 KE \sin^2(\theta/2)}$   
 impact parameter  $b = \frac{1}{4\pi\epsilon_0} \frac{Z_1 Z_2 e^2}{\frac{1}{2}mv^2} \cot(\theta/2)$ , doesn't apply,  $r_0 = \frac{1}{4\pi\epsilon_0} \frac{Z_1 Z_2 e^2}{KE}$

→ Bohr's atomic model:

i)  $L = mvr = \frac{nh}{2\pi}$

ii)  $\Delta E = E_2 - E_1 = h\nu = \frac{hc}{\lambda}$

iii) Radius of  $n^{th}$  orbit ( $r_n$ ) =  $0.53 \frac{n^2}{Z} \text{ \AA} = \frac{n^2 h^2 \epsilon_0}{4\pi m Z e^2}$

iv) velocity of revolving  $e^-$  in  $n^{th}$  orbit  $v_n = \frac{c}{137} \frac{Z}{n} = \frac{e^2}{2\epsilon_0 h} \left(\frac{Z}{n}\right)$   
 $v_n = 2.18 \times 10^6 \frac{Z}{n} \text{ (m/s)}$

v) Angular momentum ( $w_n$ ) =  $41.5376 \times 10^{15} \times \frac{Z^2}{n^3} \frac{\text{rad}}{\text{sec}}$   
 $w = \frac{\pi m e^4}{2\epsilon_0^2 h^3} \frac{Z^2}{n^3}$

vi) frequency ( $f_n$ ) or ( $\nu_n$ ) =  $6.62 \times 10^{15} \times \frac{Z^2}{n^3} \text{ Hz}$

vii) time period ( $T_n$ ) =  $1.51 \times 10^{-16} \frac{n^3}{Z^2} \text{ sec} = \frac{4\epsilon_0^2 h^3}{m e^4} \left(\frac{n^3}{Z^2}\right)$

viii) acceleration =  $a_c = \frac{v^2}{r} = \frac{\left[\frac{c}{137} \left(\frac{Z}{n}\right)\right]^2}{0.53 \frac{n^2}{Z}}$  ( $\therefore$  centripetal acceleration)

Recoil velocity =  $p = mv$   
 $\sqrt{2} p/m$   $a_c \propto \frac{Z^3}{n^4}$

ix) electric current ( $I$ ) =  $1.06 \left(\frac{Z^2}{n^3}\right)$  milliamperes,  $I \propto m^2 e^5$

x) Magnetic induction ( $B_n$ ) =  $\frac{12.58 \times Z^3}{n^5} \text{ Tesla (T)}, B_n \propto m^2 e^4$

xi) Magnetic moment ( $M$ ) =  $9.26 \times 10^{-24} n$   $\text{Am}^2$   
 ( $\therefore n$  = no. of orbits)  $M \propto \frac{e}{m}$

xii)  $KE = \frac{1}{8\pi\epsilon_0} \left(\frac{Ze^2}{r}\right)$ ,  $PE = \frac{-Ze^2}{4\pi\epsilon_0 r}$

$T.E = PE + K.E$

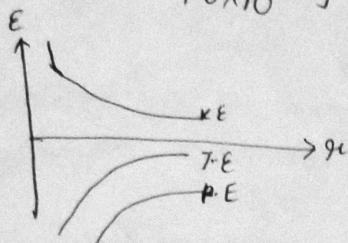
$T.E = \frac{-Ze^2}{8\pi\epsilon_0 r} = \frac{-me^4}{8\epsilon_0^2 h^3 c} (ch) \frac{Z^2}{n^2}$

$$KE: PE: T.E = 1:2:1 \quad T.E = -13.6 \left( \frac{Z^2}{n^2} \right) eV$$

$$E = h\nu \quad h = 6.624 \times 10^{-34} Js = 4.14 \times 10^{-15} eVs \quad T.E = \frac{-me^4 (ch)^2 Z^2}{8\epsilon_0^2 h^2 c^3 n^2}$$

$$1eV = 1.6 \times 10^{-19} J, \quad 1J = \frac{1}{1.6 \times 10^{-19}} eV$$

xiii)



$$E \propto m \propto Z^2 \propto \frac{1}{n^2}$$

xiv) Rydberg constant ( $R$ ) =  $1.097 \times 10^7 m^{-1}$

xv) Rydberg energy =  $Rch = 2.17 \times 10^{-18} J = 13.6 eV$

xvi) wave number  $\bar{\nu} = \frac{1}{\lambda} = RZ^2 \left[ \frac{1}{n_1^2} - \frac{1}{n_2^2} \right]$

xvii) conservative force ( $F$ ) =  $-\left(\frac{dU}{dx}\right), P.E = -\int \vec{F} \cdot d\vec{r}$   
 $PE = -\int F(r) dr$

$\rightarrow n^{th}$  excited state of  $e^- = (n+1)$  orbit of  $e^-$

$\rightarrow$  no of spectral lines =  $\frac{n(n-1)}{2}$

if  $n_1 \neq 1, n_2 \neq 1$  then spectral lines =  $\frac{(n_2 - n_1)(n_2 - n_1 + 1)}{2}$

$\rightarrow$  De-Broglie wave eqn:  $\lambda = \frac{h}{mv} = \frac{h}{\sqrt{2mKE}} = \frac{h}{\sqrt{2mV_0}}$

$\lambda_{e^-} = \frac{12.27}{\sqrt{V}} \text{ \AA}, \quad \lambda_{p^+} = \frac{0.286}{\sqrt{V}} \text{ \AA}, \quad \lambda_n = \frac{0.286}{\sqrt{KE}} \text{ \AA}, \quad \lambda_{\alpha} = \frac{0.202}{\sqrt{KE}} \text{ \AA}$

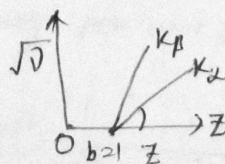
$\lambda_{\alpha} = \frac{0.101}{\sqrt{V}} \text{ \AA},$

Wave nature of  $e^-$  is used in  $e^-$  microscope of  $10^5$  magnification.

$\rightarrow$  Bragg's eqn:  $2d \sin \theta = n\lambda$

$\rightarrow$  Moseley's law:  $\sqrt{\nu} = a(Z-b)$   $\sqrt{\nu} \propto Z$

$E_{K\alpha} > E_{K\beta} > E_{K\gamma}$



$b =$  screening const,  $b=1$  for  $K$   
 $b=7.4$  for  $L$

$\lambda$  for  $X$ -rays  $\frac{1}{\lambda} = R(Z-b)^2 \left[ \frac{1}{n_1^2} - \frac{1}{n_2^2} \right]$

$\frac{\lambda_{K\alpha}}{\lambda_{K\beta}} = \frac{32}{27}$