高能天体物理笔记

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第一章 高能辐射机制简述

1.1 辐射传播相关物理量的定义

 $I_{\nu}(\theta,\phi)$: specific intensity (brightness, 亮度), 单位面积, 单位时间, 单位频率, 法线方向单位立体角穿过的能量,

$$I_{\nu}(\theta,\phi) = \frac{dE}{dA\,dt\,d\nu\,d\Omega}.\tag{1.1}$$

若法线方向沿 z 轴正方向, 且不要求法线方向穿过,

$$I_{\nu}(\theta,\phi) = \frac{dE}{(dA\cos\theta)\,dt\,d\nu\,d\Omega}.$$
(1.2)

 J_{ν} : 平均强度,

$$J_{\nu} = \frac{1}{4\pi} \int I_{\nu}(\theta, \phi) d\Omega. \tag{1.3}$$

在自由空间,沿视线方向,辐射强度不变.

 F_{ν} : monochromatic energy flux (单色能流量): 单位面积, 单位时间, 在单位频率间隔穿过的能量,

$$F_{\nu} = \int I_{\nu}(\theta, \phi) \cos \theta \, d\Omega. \tag{1.4}$$

 \mathcal{F} : energy fluence (能流): 单位面积穿过的能量,

$$\mathcal{F} = \int F_{\nu} \, dt \, d\nu. \tag{1.5}$$

 $u_{\nu}(\theta,\phi)$: radiative energy density 辐射能量密度,

$$u_{\nu}(\theta,\phi) = \frac{dE}{(dA\,c\,dt)\,d\nu\,d\Omega},\tag{1.6}$$

$$u_{\nu}(\theta,\phi) = \frac{1}{c} I_{\nu}(\theta,\phi). \tag{1.7}$$

$$u_{\nu} = \frac{4\pi}{c} J_{\nu}.\tag{1.8}$$

整理: specific intensity / brightness

$$dE_{\nu}(\Omega) = I_{\nu}(\Omega) dA_{\perp}(\Omega) dt d\Omega d\nu, \qquad (1.9)$$

net flux

$$dE_{\nu} = F_{\nu} \, dA \, dt \, d\nu \tag{1.10}$$

$$= \int_{\Omega} I_{\nu}(\Omega) dA_{\perp}(\Omega) dt d\Omega d\nu \qquad (1.11)$$

$$= \int_{\Omega} I_{\nu}(\Omega) dA \cos \theta dt d\Omega d\nu \qquad (1.12)$$

$$= \left(\int_{\Omega} I_{\nu}(\Omega) \cos \theta \, d\Omega \right) \, dA \, dt \, d\nu, \tag{1.13}$$

momentum flux

$$dP_{\perp\nu}(\Omega) = dp_{\perp\nu}(\Omega) \, dA \, dt \, d\nu, \tag{1.14}$$

$$dP_{\perp\nu}(\Omega) = dP_{\nu}(\Omega)\cos\theta \tag{1.15}$$

$$= (dE_{\nu}(\Omega)/c)\cos\theta \tag{1.16}$$

$$= (I_{\nu}(\Omega)/c)\cos^2\theta \,d\Omega \,dA \,dt \,d\nu, \tag{1.17}$$

energy density

$$dE_{\nu}(\Omega) = u_{\nu}(\Omega) \, dV_{\perp}(\Omega) \, d\Omega \, d\nu \tag{1.18}$$

$$= u_{\nu}(\Omega) dA_{\perp}(\Omega) cdt d\Omega d\nu, \qquad (1.19)$$

$$dE_{\nu}(\Omega) = I_{\nu}(\Omega) dA_{\perp}(\Omega) dt d\Omega d\nu, \qquad (1.20)$$

$$J_{\nu} = \frac{\int I_{\nu}(\Omega) d\Omega}{\int d\Omega} \tag{1.21}$$

$$= \int (I_{\nu}(\Omega)/4\pi) d\Omega. \tag{1.22}$$

$$I_{\nu_1}(\Omega) dA_{\perp_1}(\Omega) dt d\Omega_1 d\nu = I_{\nu_2}(\Omega) dA_{\perp_2}(\Omega) dt d\Omega_2 d\nu, \qquad (1.23)$$

$$d\Omega_1 = dA_{\perp 2}(\Omega)/r^2, \quad d\Omega_2 = dA_{\perp 1}(\Omega)/r^2,$$
 (1.24)

$$I_{\nu 1}(\Omega) = I_{\nu 2}(\Omega). \tag{1.25}$$

1.2 辐射转移

$$ddE_{\nu}(\Omega) = \epsilon_{\nu}(\Omega) \, dV_{\perp}(\Omega) \, dt \, d\Omega \, d\nu \tag{1.26}$$

$$= \epsilon_{\nu}(\Omega) dA_{\perp}(\Omega) ds dt d\Omega d\nu \qquad (1.27)$$

$$= dI_{\nu}(\Omega) dA_{\perp}(\Omega) dt d\Omega d\nu \qquad (1.28)$$

$$dI_{\nu}(\Omega) = \epsilon_{\nu}(\Omega) \, ds,\tag{1.29}$$

$$dS_{\perp}(\Omega) = \sigma_{\nu} \, dN_{\perp}(\Omega) \tag{1.30}$$

$$= \sigma_{\nu} n \, dV_{\perp}(\Omega) \tag{1.31}$$

$$= \sigma_{\nu} n \, dA_{\perp}(\Omega) \, ds, \tag{1.32}$$

$$ddE_{\nu}(\Omega) = -I_{\nu}(\Omega) dS_{\perp}(\Omega) dt d\Omega d\nu \qquad (1.33)$$

$$= -I_{\nu}(\Omega)\sigma_{\nu}n \, dA_{\perp}(\Omega) \, ds \, dt \, d\Omega \, d\nu \tag{1.34}$$

$$= -I_{\nu}(\Omega)\alpha_{\nu} dA_{\perp}(\Omega) ds dt d\Omega d\nu \qquad (1.35)$$

$$= dI_{\nu}(\Omega) dA_{\perp}(\Omega) dt d\Omega d\nu \qquad (1.36)$$

$$dI_{\nu}(\Omega) = -I_{\nu}(\Omega)\alpha_{\nu} ds, \qquad (1.37)$$

1.3 辐射机制

1.3.1 黑体辐射

热平衡: 宏观无热量交换. 热辐射: 辐射源热平衡. 黑体辐射: 辐射源热平衡 + 辐射全被辐射源吸收 \Rightarrow 辐射热平衡.

 $\hbar\omega/kT \ll 1, \propto \omega^2 \ \hbar\omega/kT \gg 1, \propto e^{-\hbar\omega/kT}. \ \hbar\omega_W/kT \approx 2.82.$

1.3.2 加速带电粒子的电磁辐射

$$\vec{A}(t,\vec{r}) = \frac{\mu_0}{4\pi} \int \frac{\vec{j}(t - |\vec{r} - \vec{r}'|, \vec{r}')}{|\vec{r} - \vec{r}'|} dV'$$
(1.38)

$$\vec{r}_{e}(t), \, \vec{j}(t, \vec{r}') = q\delta(\vec{r}' - \vec{r}_{e})\vec{v}_{e}, \, K = 1 - \frac{\vec{r} - \vec{r}_{e}}{|\vec{r} - \vec{r}_{e}|}\vec{v}_{e},$$

$$\vec{A}(t, \vec{r}) = \left[\frac{q\vec{v}_{e}}{K|\vec{r} - \vec{r}_{e}|}\right]\Big|_{t=t-|\vec{r} - \vec{r}_{e}|}$$
(1.39)

电场中单电子的辐射

$$I_{\parallel}/I_{\perp} \approx 1/\gamma^2, \, \omega \approx \gamma v/b.$$

热韧致辐射 $au\gg 1,$ $\propto\omega^2,$ $\tau\ll 1\wedge\hbar\omega/kT\ll 1,$ \approx const, $\hbar\omega/kT\gg 1,$ $\propto e^{-\hbar\omega/kT}.$

非热韧致辐射

1.3.3 磁场中单电子的辐射

$$evB = \gamma m\omega^2 r = \gamma m\omega v, \, \omega_{\rm g} = eB/\gamma m = \omega_{\rm cycle}/\gamma.$$

回旋辐射

同步辐射 周期 $2\pi/\omega_{\rm g}$, 宽度 $\approx 1/\gamma^3\omega_{\rm g}$. $\omega/\omega_{\rm c} \ll 1, \propto \omega^{1/3}, \, \omega/\omega_{\rm c} \gg 1, \, \propto e^{-\omega/\omega_{\rm c}}. \, \omega_{\rm c} = 3\gamma^2 eB/2m = (3/2)\gamma^3\omega_{\rm g}.$ $\tau \gg 1, \, I(\omega) \propto \omega^{5/2}, \, \tau \ll 1, \, N(E) \propto E^{-p}, \, I(\omega) \propto \omega^{-(p-1)/2}.$

1.3.4 光子散射

Thomson 散射

$$m_e c^2 = \frac{e^2}{4\pi\epsilon_0 r_e}$$

$$\frac{d\sigma_{\rm T}}{d\Omega} := \frac{\frac{dE_{\rm out}}{dtd\Omega}}{S_{\rm in}} = \frac{\frac{dE_{\rm out}}{dtd\Omega}}{\frac{dE_{\rm in}}{dtdA}}, \quad S_{\rm in}\sigma_{\rm T} = \frac{dE_{\rm out}}{dt}$$

$$(1.40)$$

Compton 散射

$$\hbar\omega/m_ec^2\ll 1,\,\sigma\approx\sigma_{\rm T},\,\hbar\omega/m_ec^2\gg 1,\,\sigma\propto\omega^{-1}.$$

逆 Compton 散射

$$\omega \leq 4\gamma^2 \omega_0$$
. $\omega \ll 4\gamma^2 \omega_0$, $\propto \omega$

第二章 银河系内

2.1 星际介质

$$\begin{split} ?/xD &= \theta, ?/(1-x)D = \alpha, \, \alpha = [x/(1-x)]\theta, \, \theta_{\rm sca} = \theta + \alpha = [?/x + ?/(1-x)]/D = ?/x(1-x)D, \, \theta = (1-x)\theta_{\rm sca}. \\ \sqrt{1+\delta^2} &\approx 1 + \frac{1}{2}\delta^2, \, xD(1+\frac{1}{2}\theta^2) + (1-x)D(1+\frac{1}{2}\alpha^2), \, \frac{1}{2}xD\theta^2 + \frac{1}{2}(1-x)D\alpha^2 = \frac{1}{2}xD\theta^2 + \frac{1}{2}(1-x)D[x^2/(1-x)^2]\theta^2 = \frac{1}{2}xD\theta^2 + \frac{1}{2}D[x^2/(1-x)]\theta^2, \\ \frac{1}{2}x + \frac{1}{2}[x^2/(1-x)] = \frac{1}{2}x/(1-x) \end{split}$$

2.2 XB

第三章 银河系外

3.1 一般星系

²⁶Al 衰变, 1.089 MeV 谱线.

色色图: 较软能段硬度比 $\frac{M-S}{H+M+S}$ -较硬能段硬度比 $\frac{H-M}{H+M+S}$.

3.2 AG

多波段,强度高.

- 3.2.1 AG 分类
- 3.2.2 AGN 统一模型
- 3.2.3 射电宁静 AGN

区域: AGN 核心区域, 靠近 BH, 吸积盘内侧.

- 3.2.4 宽线, 窄线, 反响映射
- 3.2.5 射电噪 AGN

视超光速

Doppler 增亮

3.3 GRB