If we only take the lowest order effects of the motion of mass sources into account and neglect stresses, then we can define two vector fields, \vec{E} and \vec{B} , which are similar to the electric and magnetic field. We can define $A_{\mu} = -\frac{1}{4}c_0\bar{h}_{0\mu} = (\varphi/c_0, \vec{A})$, then A_{μ} , which are similar to the electromagnetic 4-potential, will satisfy

 $\partial^{\nu}\partial_{\nu}A_{\mu} = -\frac{4\pi G_0}{c_0^2}J_{\mu}, \quad \partial^{\mu}A_{\mu} = 0, \tag{1}$

where $J_{\mu}=-c_0T_{0\mu}/c_0^2=(c_0\rho,\vec{j})=\rho(c_0,\vec{v})$ is 4-momentum density. If we define $\vec{E}=-\vec{\nabla}\varphi-\frac{\partial}{\partial t}\vec{A},\,\vec{B}=\vec{\nabla}\times\vec{A},\,$ and $\varepsilon_{\rm G0}=\frac{1}{4\pi G_0},\,\mu_{\rm G0}=\frac{4\pi G_0}{c_0^2},\,$ then there will be

$$\begin{cases} \vec{\nabla} \cdot (\varepsilon_{G0}\vec{E}) = \rho, \\ \vec{\nabla} \cdot \vec{B} = 0, \\ \vec{\nabla} \times \vec{E} = -\frac{\partial}{\partial t}\vec{B}, \\ \vec{\nabla} \times (\mu_{G0}^{-1}\vec{B}) = \vec{j} + \frac{\partial}{\partial t}(\varepsilon_{G0}\vec{E}), \end{cases}$$
(2)

and the acceleration of mass particle \vec{a} will satisfy

$$\vec{a} = -\vec{E} - 4\vec{v} \times \vec{B}.\tag{3}$$

Since (2) are identical to Maxwell's equations, and (3) is identical to the Lorentz force equation except for an overall minus sign and a factor of 4 in the "magnetic force" term, we assume that the constants $\varepsilon_{\rm G0}$ and $\mu_{\rm G0}$ can simply vary as constants $\varepsilon_{\rm 0}$ and $\mu_{\rm 0}$ in electromagnetic theories, which means, (3) is still valid and (2) becomes

$$\begin{cases} \vec{\nabla} \cdot \vec{D} = \rho, \\ \vec{\nabla} \cdot \vec{B} = 0, \\ \vec{\nabla} \times \vec{E} = -\frac{\partial}{\partial t} \vec{B}, \\ \vec{\nabla} \times \vec{H} = \vec{j} + \frac{\partial}{\partial t} \vec{D}, \end{cases}$$
(4)

where $\vec{D} = \varepsilon_{\rm G} \vec{E}$, $\vec{B} = \mu_{\rm G} \vec{H}$. Also, we define $c = \frac{1}{\sqrt{\varepsilon_{\rm G} \mu_{\rm G}}}$, $G = \frac{1}{4\pi\varepsilon_{\rm G}}$, together with $\vec{E} = -\vec{\nabla}\varphi - \frac{\partial}{\partial t}\vec{A}$, $\vec{B} = \vec{\nabla}\times\vec{A}$, and $\bar{h}_{0\mu} = -4A_{\mu}/c = -4(\varphi/c,\vec{A})/c$. We assume that the source of GW is located in a spherical coordinate origin

We assume that the source of GW is located in a spherical coordinate origin and $c=c(r),\ G=G(r)$, We suppose that $\bar{h}_{0\mu}(\rho,\vec{j};c_0,G_0)$ is the solution of $\bar{h}_{0\mu}$ when ρ and \vec{j} are assumed and $c\equiv c_0,\ G\equiv G_0$ everywhere, then if $c\equiv c_s$, $G\equiv G_s$, where c_s and G_s are two constants, within a region $r\leq R$, then $\bar{h}_{0\mu}(\rho,\vec{j};c_s,G_s)$ can be a solution of $\bar{h}_{0\mu}$ within the region $r\leq R$, and now we can solve the equations

$$\begin{cases} \vec{\nabla} \cdot \vec{D} = 0, \\ \vec{\nabla} \cdot \vec{B} = 0, \\ \vec{\nabla} \times \vec{E} = -\frac{\partial}{\partial t} \vec{B}, \\ \vec{\nabla} \times \vec{H} = +\frac{\partial}{\partial t} \vec{D}, \end{cases}$$
 (5)

with the boundary conditions $\vec{E}|_{r=R} = \vec{E}(\rho, \vec{j}; c_{\rm s}, G_{\rm s}), \ \vec{B}|_{r=R} = \vec{B}(\rho, \vec{j}; c_{\rm s}, G_{\rm s}),$ where $\vec{E}(\rho, \vec{j}; c_{\rm s}, G_{\rm s})$ and $\vec{B}(\rho, \vec{j}; c_{\rm s}, G_{\rm s})$ can be derived from $\bar{h}_{0\mu}(\rho, \vec{j}; c_{\rm s}, G_{\rm s})$.

Again, since (2) are identical to Maxwell's equations, we can use the Liénard–Wiechert potentials formula to calculate the \vec{E} and \vec{B} field produced by a moving particle, whose position is \vec{r}' , in vacuum, and the results are

$$\vec{E} = \frac{m}{4\pi\varepsilon_{G0}} \left[\frac{(\vec{n}' - \vec{\beta})}{(1 - \vec{\beta} \cdot \vec{n}')^3 \gamma^2 r'^2} + \frac{\vec{n}' \times \{(\vec{n}' - \vec{\beta}) \times \dot{\vec{\beta}}\}\}}{(1 - \vec{\beta} \cdot \vec{n}')^3 c_0 r'} \right]_{\text{ret}}, \vec{B} = \left[\frac{\vec{n}' \times \vec{E}}{c_0} \right]_{\text{ret}},$$
(6)

where $r'=|\vec{r}'|, \ \vec{n}'=\vec{r}'/r', \ \vec{\beta}=\dot{\vec{r}}'/c_0, \ \gamma=1/\sqrt{1-|\vec{\beta}|^2}, \ \text{and "ret" denotes}$ that the fields are "retarded fields". Both \vec{E} and \vec{B} in (6) can be divided into two parts: one, called "inherent field part", is proportional to $1/r'^2$ and the other one, called "radiation part", is proportional to 1/r'. We only take the lowest order effects of the motion of mass sources into account, therefore we can assume that $|\vec{\beta}|\ll 1$. The "inherent field part" of \vec{E} will be parallel to \vec{r}' and the "inherent field part" of \vec{B} will be nearly zero vector. The "radiation part" of \vec{E} and \vec{B} will be perpendicular to \vec{r}' . Since (5) are linear equations, we can discuss two parts of \vec{E} and \vec{B} separately for a binary GW source. If we assume that the distance between two components of binary is much less than R mentioned above, then the "inherent field part" of \vec{E} will have only r component, and the "radiation part" of \vec{E} and \vec{B} will have no r component.

For the "inherent field part", (5) will become $\frac{1}{r^2} \frac{\partial}{\partial r} (r^2 D_r) = 0$, then $D_r \propto 1/r^2$, and $E_r \propto \varepsilon_{\rm G}^{-1}(1/r^2)$. As mentioned below, for the "radiation part", $E \propto c^{-1}\varepsilon_{\rm G}^{-1/2}(1/r)$, therefore the "inherent field part" will vanish in the region far away from source.

For the "radiation part", from the third and fourth equation of (5), we can derive

$$\begin{cases}
[\vec{e}_r \cdot (\vec{\nabla} \times \vec{E})] \vec{e}_r - \frac{1}{r} \frac{\partial}{\partial r} (r E_{\phi}) \vec{e}_{\theta} + \frac{1}{r} \frac{\partial}{\partial r} (r E_{\theta}) \vec{e}_{\phi} = -\mu_{G} \frac{\partial}{\partial t} (H_{\theta} \vec{e}_{\theta} + H_{\phi} \vec{e}_{\phi}), \\
[\vec{e}_r \cdot (\vec{\nabla} \times \vec{H})] \vec{e}_r - \frac{1}{r} \frac{\partial}{\partial r} (r H_{\phi}) \vec{e}_{\theta} + \frac{1}{r} \frac{\partial}{\partial r} (r H_{\theta}) \vec{e}_{\phi} = +\varepsilon_{G} \frac{\partial}{\partial t} (E_{\theta} \vec{e}_{\theta} + E_{\phi} \vec{e}_{\phi}),
\end{cases}$$
(7)

and then

$$\begin{cases}
\frac{1}{r}\frac{\partial}{\partial r}(rE_{\theta}) = -\mu_{G}\frac{\partial}{\partial t}H_{\phi}, \\
\frac{1}{r}\frac{\partial}{\partial r}(rE_{\phi}) = +\mu_{G}\frac{\partial}{\partial t}H_{\theta}, \\
\frac{1}{r}\frac{\partial}{\partial r}(rH_{\theta}) = +\varepsilon_{G}\frac{\partial}{\partial t}E_{\phi}, \\
\frac{1}{r}\frac{\partial}{\partial r}(rH_{\phi}) = -\varepsilon_{G}\frac{\partial}{\partial t}E_{\theta},
\end{cases} (8)$$

$$\begin{cases} \frac{\partial}{\partial r}(rE_{\theta}) = -\mu_{G} \frac{\partial}{\partial t}(rH_{\phi}), \\ \frac{\partial}{\partial r}(rE_{\phi}) = +\mu_{G} \frac{\partial}{\partial t}(rH_{\theta}), \\ \frac{\partial}{\partial r}(rH_{\theta}) = +\varepsilon_{G} \frac{\partial}{\partial t}(rE_{\phi}), \\ \frac{\partial}{\partial r}(rH_{\phi}) = -\varepsilon_{G} \frac{\partial}{\partial t}(rE_{\theta}), \end{cases}$$
(9)

$$\begin{cases}
\mu_{G} \frac{\partial}{\partial r} [\mu_{G}^{-1} \frac{\partial}{\partial r} (rE_{\theta})] = -\mu_{G} \frac{\partial}{\partial t} \frac{\partial}{\partial r} (rH_{\phi}) = \varepsilon_{G} \mu_{G} \frac{\partial}{\partial t} \frac{\partial}{\partial t} (rE_{\theta}), \\
\mu_{G} \frac{\partial}{\partial r} [\mu_{G}^{-1} \frac{\partial}{\partial r} (rE_{\phi})] = +\mu_{G} \frac{\partial}{\partial t} \frac{\partial}{\partial r} (rH_{\theta}) = \varepsilon_{G} \mu_{G} \frac{\partial}{\partial t} \frac{\partial}{\partial t} (rE_{\phi}), \\
\varepsilon_{G} \frac{\partial}{\partial r} [\varepsilon_{G}^{-1} \frac{\partial}{\partial r} (rH_{\theta})] = +\mu_{G} \frac{\partial}{\partial t} \frac{\partial}{\partial r} (rE_{\phi}) = \varepsilon_{G} \mu_{G} \frac{\partial}{\partial t} \frac{\partial}{\partial t} (rH_{\theta}), \\
\varepsilon_{G} \frac{\partial}{\partial r} [\varepsilon_{G}^{-1} \frac{\partial}{\partial r} (rH_{\phi})] = -\mu_{G} \frac{\partial}{\partial t} \frac{\partial}{\partial r} (rE_{\theta}) = \varepsilon_{G} \mu_{G} \frac{\partial}{\partial t} \frac{\partial}{\partial t} (rH_{\phi}),
\end{cases} \tag{10}$$

therefore

$$\begin{cases}
\frac{\partial}{\partial r}\frac{\partial}{\partial r}(rE_{\theta}) - \frac{\partial}{\partial r}(\ln\mu_{G})\frac{\partial}{\partial r}(rE_{\theta}) - \frac{\partial}{\partial(ct)}\frac{\partial}{\partial(ct)}(rE_{\theta}) = 0, \\
\frac{\partial}{\partial r}\frac{\partial}{\partial r}(rE_{\phi}) - \frac{\partial}{\partial r}(\ln\mu_{G})\frac{\partial}{\partial r}(rE_{\phi}) - \frac{\partial}{\partial(ct)}\frac{\partial}{\partial(ct)}(rE_{\phi}) = 0, \\
\frac{\partial}{\partial r}\frac{\partial}{\partial r}(rH_{\theta}) - \frac{\partial}{\partial r}(\ln\varepsilon_{G})\frac{\partial}{\partial r}(rH_{\theta}) - \frac{\partial}{\partial(ct)}\frac{\partial}{\partial(ct)}(rH_{\theta}) = 0, \\
\frac{\partial}{\partial r}\frac{\partial}{\partial r}(rH_{\phi}) - \frac{\partial}{\partial r}(\ln\varepsilon_{G})\frac{\partial}{\partial r}(rH_{\phi}) - \frac{\partial}{\partial(ct)}\frac{\partial}{\partial(ct)}(rH_{\phi}) = 0.
\end{cases} (11)$$

All equations in (11) have the form like

$$\frac{\partial^2}{\partial r^2} f(r,t) - p(r) \frac{\partial}{\partial r} f(r,t) - \frac{\partial^2}{\partial (ct)^2} f(r,t) = 0.$$
 (12)

If we solve (12) by the method of separation of variables, we can suppose that $f(r,t)=f(r)e^{-ikct}$, and then

$$\frac{d^{2}}{dr^{2}}f(r) - p(r)\frac{d}{dr}f(r) + k^{2}f(r) = 0.$$
(13)

If p is constant, then

$$f(r) = e^{(p/2)r} \left[C_{+} e^{i\sqrt{k - (p/2)^{2}}r} + C_{-} e^{-i\sqrt{k - (p/2)^{2}}r} \right], \tag{14}$$

where C_{+} and C_{-} are two constants, and

$$f(r,t) = e^{(p/2)r} \left[C_{+} e^{i(+\sqrt{(\omega/c) - (p/2)^2}r - \omega t)} + C_{-} e^{i(-\sqrt{(\omega/c) - (p/2)^2}r - \omega t)} \right], \quad (15)$$

where $\omega = kc$. If $\mathrm{d}p/\mathrm{d}r$ is small, we can divide the region r > R into many spherical shells, and in each of them p is nearly constant. To make f(r,t) continuous anytime, we suppose that the "angular frequency" ω in each shell is same, and then there will be a approximate solution

$$f(r,t) = e^{\int (p/2)dr} \left[C_{+} e^{i(+\int \sqrt{(\omega/c) - (p/2)^{2}} dr - \omega t)} + C_{-} e^{i(-\int \sqrt{(\omega/c) - (p/2)^{2}} dr - \omega t)} \right].$$
(16)