

Special
Relativity

Theory of special relativity:

Inertial frames and the principle of relativity

→ frames are one of the most essential concept in the theory of relativity, we are free to mark any steady frame S and a moving frame w.r.t S as S'.

There exist a class of preferred reference systems called inertial frames in which a free particle is either at rest or moving with a constant velocity.



This specifically means that it is not changing direction

∴ In cartesian coordinates it means $\left\{ \frac{d^2x}{dt^2} = \frac{d^2y}{dt^2} = \frac{d^2z}{dt^2} = 0 \right\}$

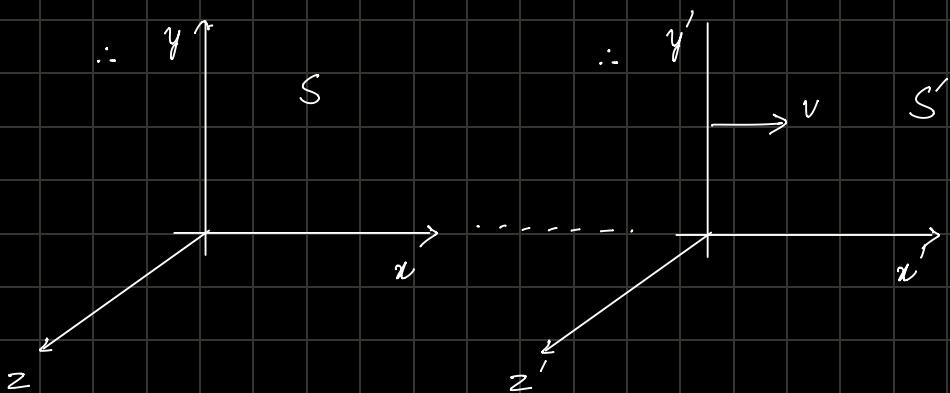
∴ In absence of gravity, if we have two inertial frames then S' can differ from S only by

i) Translation

ii) Rotation

iii) A motion of one frame w.r.t the other frame at a constant velocity.

These inertial frames are the fundamental concept of the special relativity. In relativity a point in a 4-dimensional space is described by (ct, x, y, z) .



• Transformations for S & S' are

$$\left. \begin{array}{l} t' = At + Bx \\ x' = Ct + Dx \\ y' = y \\ z' = z \end{array} \right\} \text{Don't worry, this will be proven later.}$$

$$x' = 0 \Rightarrow x = vt \quad \therefore x = 0 \Rightarrow x' = -vt'$$

$$\therefore t' = At \quad \text{at } x=0 \quad \therefore x' = 0 \quad Ct + vD t = 0 \\ x = 0 \Rightarrow x' = Ct = \frac{Ct'}{A} \Rightarrow \left[\frac{C}{A} = -v \right]$$

$$\therefore C + vD = 0 \Rightarrow -Av + vD = 0 \Rightarrow \boxed{A = D}$$

\therefore Finally the transformations looks like,

$$\begin{aligned} t' &= At + Bx \\ x' &= A(x - vt) \\ y' &= y \\ z' &= z \end{aligned}$$

I am assuming how does classical relativity works (i.e $v \ll c$) !!

So, let's move to the topic of geometry of space-time.

In SR, Einstein removed the concept of absolute time and replaced it by saying that the speed of light c is the same in all inertial frames.

\therefore By applying this postulate together with the principle of relativity we will be able of connect coordinates of an event P in two different cartesian inertial frames, S and S' .

• Consider a photon being emitted at $t=0$ & $t'=0$ and travelling in arbitrary dx^n .

It will satisfy the following eqⁿ

$$\left\{ c^2 t^2 - x^2 - y^2 - z^2 = c^2 t'^2 - x'^2 - y'^2 - z'^2 = 0 \right\}$$

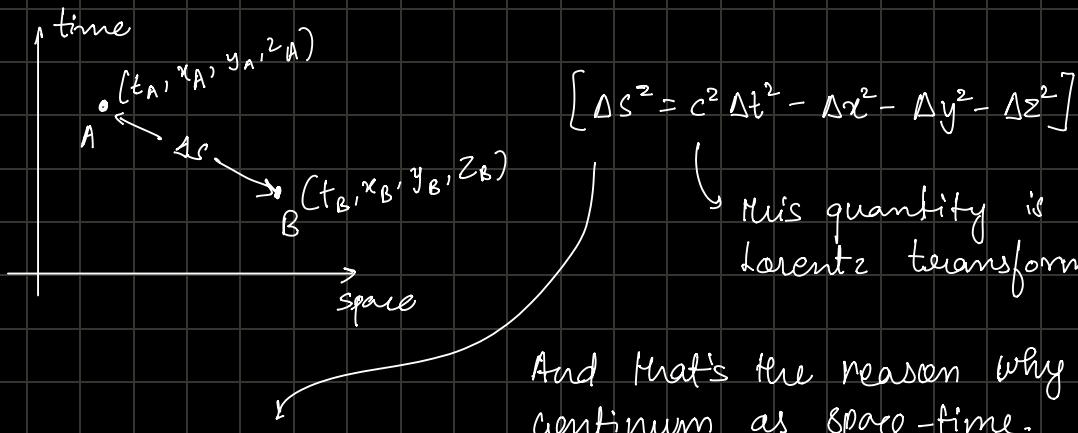
Plugging the above transformation here will give us A and B.

$$\begin{aligned}ct' &= \gamma(ct - \beta x) \\x' &= \gamma(x - \beta ct) \\y' &= y \\z' &= z\end{aligned}$$

where $\left\{\beta = \frac{v}{c}\right\}$ and $\left\{\gamma = \frac{1}{\sqrt{1 - v^2/c^2}}\right\}$

This is called Lorentz transformation and it is particularly boost in x-direction.

for $\beta \ll 1 \Rightarrow$ It becomes Galilean transformation.

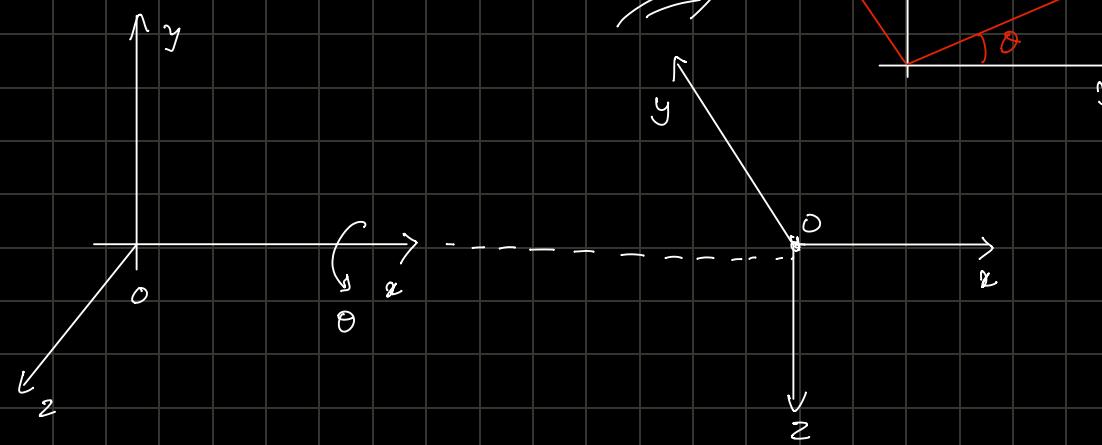


And that's the reason why we call this 4D continuum as space-time.

This is often called as 'Minkowski Geometry'.

* For fixed t, it becomes Euclidean.

"Rotations"



$$ct' = ct$$

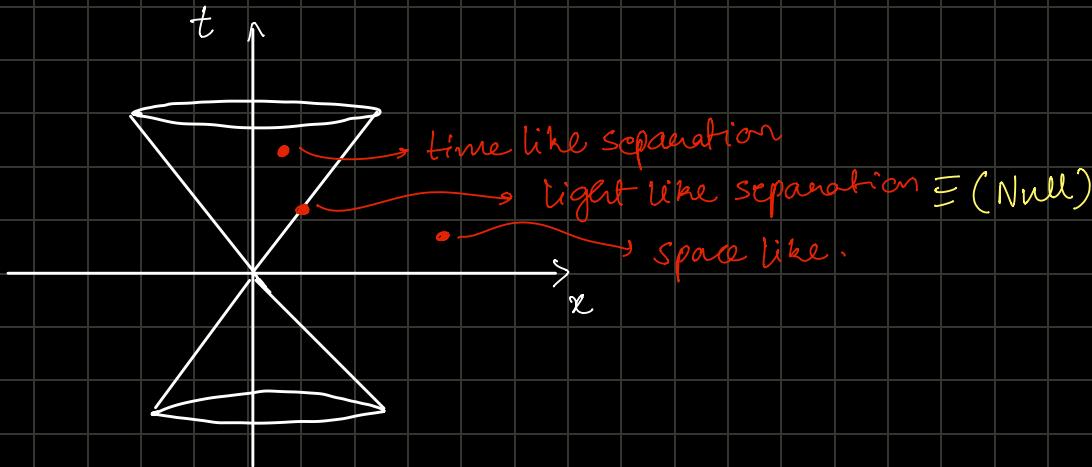
$$x' = x$$

$$y' = y \cos \theta - z \sin \theta$$

$$z' = y \sin \theta + z \cos \theta$$

"Light cone"

$$\Delta s^2 = c \Delta t^2 - \Delta x'^2 - \Delta y'^2 - \Delta z'^2 \Rightarrow \text{For this equation } \Delta s = \begin{cases} > 0 & \text{Space-like} \\ = 0 & \text{Light-like} \\ < 0 & \text{Time-like} \end{cases}$$



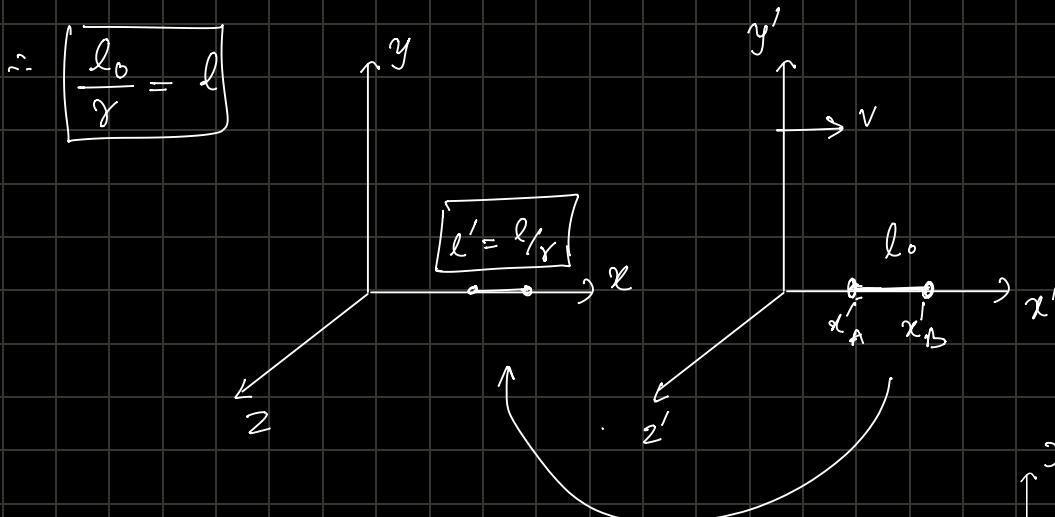
Length contraction and time dilation:

Length contraction:

Assume a rod of length l_0 , where $l_0 = x'_B - x'_A$ s' is at rest

$$x'_A = \gamma(x_A - vt_A) \quad | \quad l'_B = \gamma(l_B) \text{ since } t_A = t_B = t$$

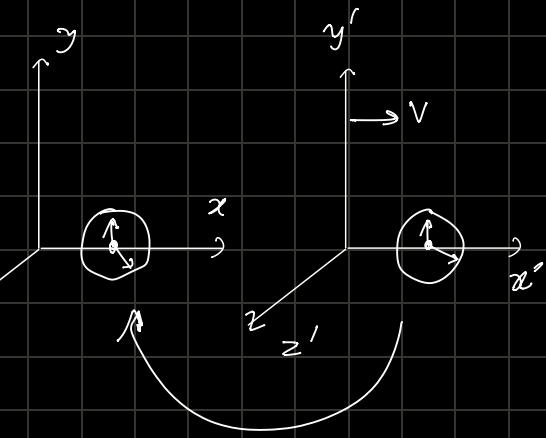
$$x'_B = \gamma(x_B - vt_B)$$



Time-dilation:

$$t_A = \gamma(t'_A + \frac{vx'_A}{c^2}) \quad | \quad t_B - t_A = \gamma(t'_A - t'_B)$$

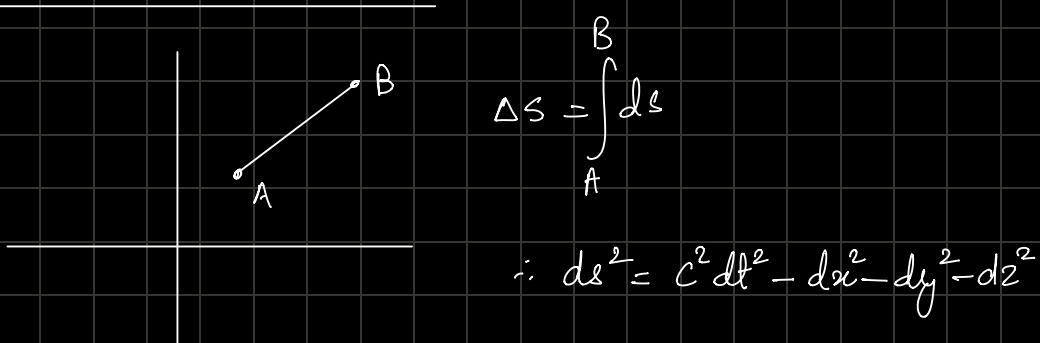
$$t_B = \gamma(t'_B + \frac{vx'_B}{c^2}) \quad | \quad x'_A = x'_B \text{ (clocks are fixed)}$$



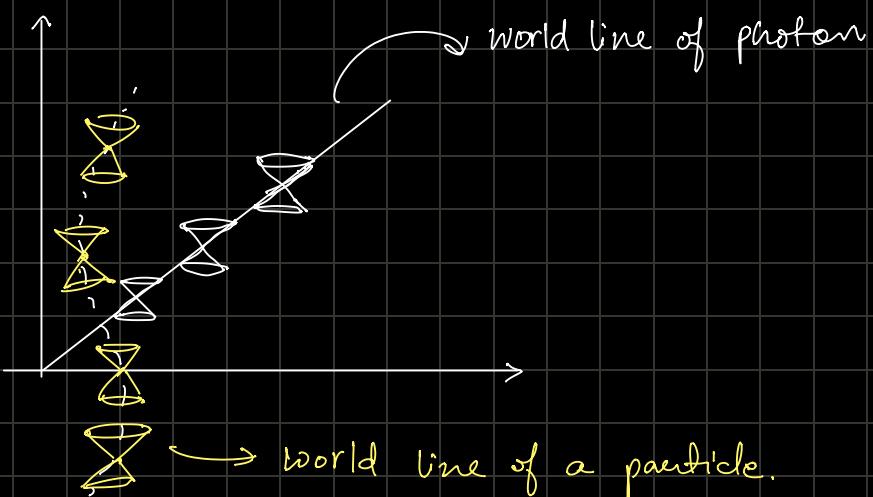
$$\therefore \boxed{\Delta t = \gamma T_0}$$

The faster you move, slower your time click measured by a person at rest.

"Minkowski line-element".



- : $ds^2 > 0$ Time like
- $ds^2 = 0$ Light like
- $ds^2 < 0$ space-like. \Rightarrow we cannot have $v > c$.



A four dimensional way of describing a world line is to give the four coordinates (t, x, y, z) as functions of parameter λ , that varies monotonically along the world line.

For a massive particle, such parameter is called its proper time.

$$c^2 d\tau^2 = ds^2 = c^2 dt^2 - dx^2 - dy^2 - dz^2$$

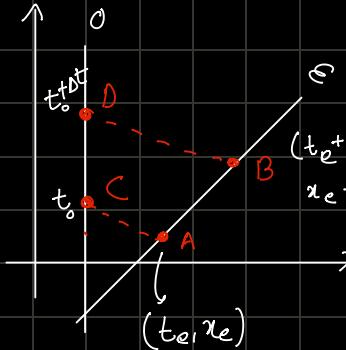
\hookrightarrow we define τ as for infinitesimally separated events.

$$\therefore d\tau^2 = \left(1 - \frac{v^2}{c^2}\right) dt^2 \Rightarrow \left\{ d\tau = \left(1 - \frac{v^2}{c^2}\right)^{1/2} dt = \frac{dt}{\gamma} \right\}$$

$$\left\{ \Delta\tau = \int_a^b d\tau = \int_a^b \left(1 - \frac{v^2}{c^2}\right)^{1/2} dt \right\}$$

proper time τ is just the coordinate time t measured by clocks at rest in S.

"Doppler effect"



$$\Delta\tau_{AB} = \left(1 - \frac{v^2}{c^2}\right)^{1/2} \Delta t_e$$

$$\Delta\tau_{CD} = \Delta t_0$$

worldline representing photons : $ds^2 = 0 \Rightarrow c^2 dt^2 - dx^2 - dy^2 - dz^2 = 0$

$$\therefore \text{World line connecting } A \& C \Rightarrow \int_{t_e}^{t_0} c dt = \int_{x_e}^{x_0} dx$$

because photon is travelling from $A \rightarrow C$ (-ive X direction)

$$\therefore c(t_0 - t_e) = -(x_0 - x_e)$$

similarly along the world line B, D

$$\int_{t_e + \Delta t_e}^{t_0 + \Delta t_0} c dt = - \int_{x_e + \Delta x_e}^{x_0} dx \Rightarrow c \Delta t_0 - c \Delta t_e = \Delta x_e$$

(add both these integrals)

$$\therefore \Delta t_0 = \left(1 + \frac{1}{c} \frac{\Delta x_e}{\Delta t_e}\right) \Delta t_e = \left(1 + \frac{v}{c}\right) \Delta t_e$$

$$\therefore \Delta t_{CD} = (1 + \beta) \Delta t_e ; \quad \Delta\tau_{AB} = \left(1 - \beta^2\right)^{1/2} \Delta t_e$$

$$\therefore \frac{\Delta t_{CD}}{\Delta\tau_{AB}} = \left(\frac{1 + \beta}{1 - \beta}\right)^{0.5} \Rightarrow \frac{v_{AB}}{v_{CD}} = \frac{v_e}{v_0}$$

"Addition of velocities in SR"

$$u_x = \frac{dx}{dt} \quad v_y = \frac{dy}{dt} \quad u_z = \frac{dz}{dt}$$

$$dt' = \gamma_v (dt - \frac{v dx}{c^2}) ; \quad dx' = \gamma_v (dx - v dt) ; \quad dy' = dy ; \quad dz' = dz$$

$$\therefore \left[\begin{array}{l} u'_x = \frac{u_x - v}{\left(1 - \frac{u_x v}{c^2}\right)} \\ \vdots \\ i=y,z \end{array} \right] \quad \left[\begin{array}{l} u'_i = \frac{u_i}{\gamma_v \left(1 - \frac{u_x v}{c^2}\right)} \\ \vdots \\ i=y,z \end{array} \right]$$

And similarly we can find acceleration. $\left[\begin{array}{l} a_i = \frac{du'_i}{dt} \\ \vdots \\ i=x,y,z \end{array} \right]$