

Week 4 - Azra Ozyurt

15 Aralık 2019 Pazar 13:38

TASK 1

Complete the modified example of simplified wall calculations that you went through in the assignment of week 3 and find the total heat transfer through wall.

Week 3 results:

$$R'_{withWood} = 0.03 + 0.14 + 0.11 + 0.63 + 0.079 + 0.12 \\ = 1.109 \text{ m}^2 \cdot ^\circ \frac{C}{W}$$

$$R'_{withIns} = 0.03 + 0.14 + 0.11 + 3.528 + 0.079 + 0.12 \\ = 4.007 \text{ m}^2 \cdot ^\circ \frac{C}{W}$$

$$R'_{total} = 5.116 \text{ m}^2 \cdot ^\circ \frac{C}{W}$$

$$U_{tot} = \frac{1}{R} = U$$

$$U_{tot} = 0.9 + 0.249 = 1.149$$

$$Q_{tot} = U_{tot} \times A_{tot} \times \Delta T = 1.149 * 100 * 24 = 1065.3 \text{ W}$$

TASK 2

RADIATION

Radiation can happen in vacuumed area, which explains that it doesn't need material space in order to occur, because of the nature of wavelengths. It can happen with solids, liquids and gasses.

When something changes in a material in molecular level, the difference of energy will create waves. These waves are electromagnetic waves and characterised by wavelengths and frequency.

$$\lambda = \frac{c}{\nu}$$

Type of electromagnetic radiation that is related to the heat transfer is thermal radiation.

Thermal radiation is always emitted by surroundings and the emissivity will increase with the increasing temperature.

Radiative Heat Transfer

One of the ways for the heat to transfer from one body to another is radiative heat transfer. It will occur with waves carrying the energy. So the heat transfer will be defined and occur with wavelengths and frequencies. All matters which have the temperature above the absolute zero always emits thermal radiation. The energy that is emitting from one body will be absorbed by another. The waves does not require material in between space to carry the heat, so radiation can occur in vacuumed areas. Though it will take place in all forms of material such as solids, liquids and gasses.

BLACK BODY

A blackbody is the ultimate condition theory of radiation. A black body surface will emit and absorb the maximum radiation amount possible. It is the perfect emitter and the perfect absorber surface.

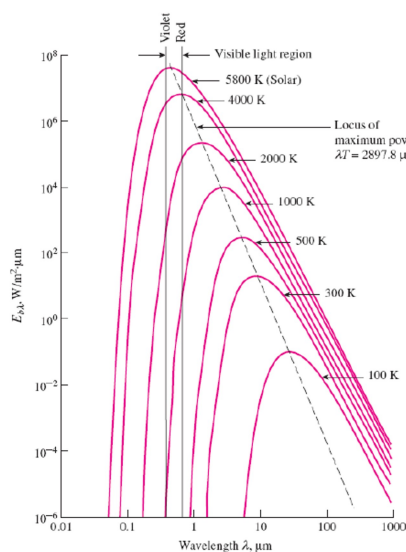
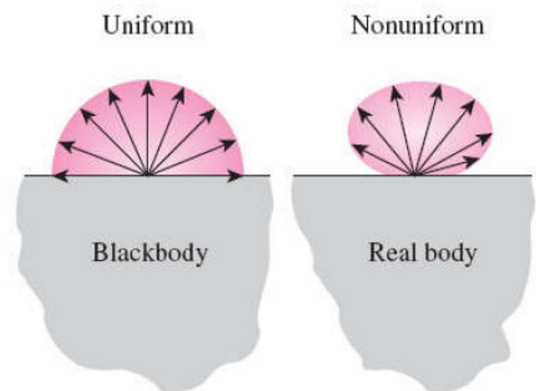
The radiation energy emitted by a blackbody:

$$E_b(T) = \sigma T^4 \quad (\text{W/m}^2)$$

Blackbody emissive power

$$\sigma = 5.670 \times 10^{-8} \text{ W/m}^2 \cdot \text{K}^4$$

Stefan-Boltzmann constant



The variation of the blackbody emissive power with wavelength for several temperatures.

As mentioned before, the emissivity and absorbance of radiation depends on temperature.

