

- Shear Mapping in Python (SMPy): Modular,
- Extensible, and Accessible Dark Matter Mapping
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#### Software

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# Summary

Understanding the universe's large-scale distribution of dark matter is a central objective in the era of precision cosmology. A key tool for the study of dark matter is weak gravitational lensing; a phenomenon where light from distant galaxies is sheared as it passes though the gravitational field of a massive object, like a galaxy cluster. Measuring this slight (weak) distortion of shapes over thousands of galaxies allows astrophysicists to infer the distribution of total matter, which includes both luminous and dark matter.

A common tool to infer the matter distribution is the mapping of convergence ( $\kappa$ ), which quantifies how much a gravitational lens converges the light from distant galaxies, resulting in a magnification of their shapes. [Placeholder Sentence for shear].

The Shear Mapping in Python (SMPy) package provides a standardized, well-documented, and open-source solution for creating convergence maps from weak lensing galaxy shear measurements. SMPy was initially developed to support the Superpressure Balloon-borne Imaging Telescope (SuperBIT), a stratospheric, near-UV to near-IR observing platform which completed its 45-night observing run in spring 2023 with over 30 galaxy cluster observations (Gill et al., 2024). SMPy has since evolved into a general-purpose tool suitable for analyzing the weak lensing data from any source of galaxies.

# 23 Background

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24 For a comprehensive review of weak gravitational lensing, please refer to (Umetsu, 2020).

## <sub>25</sub> Statement of Need

- Mass maps are a critical deliverable of many cosmological analyses (Darwish et al., 2020; Jeffrey et al., 2021; Oguri et al., 2017). SMPy addresses an outstanding need for the lensing community: A robust, well-documented, and open-source tool to construct publication quality mass maps from galaxy shear data. SMPy was built with multiple design directions in mind:
  - Accessibility: SMPy is written entirely in Python and deliberately relies only on widely-used scientific Python packages (numpy, scipy, pandas, astropy, matplotlib, and pyyaml). This choice of standard dependencies ensures that users can easily install the packages without complex dependency chains, and that the codebase is maintainable and familiar to the scientific Python community.
  - 2. Extensibility: SMPy is built with a modular architecture that allows for easy implementation of new mass mapping techniques beyond the currently implemented Kaiser-Squires inversion algorithm (Kaiser & Squires, 1993). An example convergence map is shown in



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Figure 1, created from simulated galaxy cluster observations from SuperBIT (McCleary et al., 2023). Aperture mass mapping (Leonard et al., 2012) and KS+ (Pires, 2020) algorithms are currently planned to be added to the codebase.

- 3. **Usability:** Creating convergence maps with SMPy requires minimal input users need only provide a catalog of galaxies with their associated shears (g1 & g2) and coordinates. This straightforward input requirement makes the tool accessible to researchers at all levels. A flexible configuration system is integrated via a single YAML file that defines file paths, convergence map algorithm settings, plotting parameters, and more.
- 4. Robustness: Designed to be mathematically and algorithmically accurate, allowing the user to create convergence maps with any galaxy shear data. The coordinate system abstraction handles both RA/Dec celestial coordinates (with proper spherical geometry approximations) or pixel-based coordinates through a unified interface. Signal-to-noise maps can be generated using either spatial shuffling (randomizing galaxy positions while preserving shear values) or orientation shuffling (randomizing shear orientations while preserving positions) to distinguish real signals from noise.

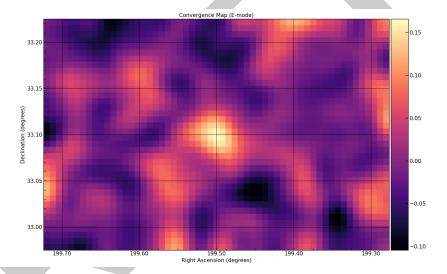


Figure 1: Example convergence map created with SMPy showing the mass distribution of a simulated galaxy cluster. The map was generated using the Kaiser-Squires inversion method on simulated weak lensing data from SuperBIT. The color scale represents the dimensionless surface mass density (convergence), with brighter regions indicating higher mass concentrations.

To our knowledge, SMPy is the first open-source, well-documented convergence mapping software that can compute convergence in either astrometric or pixel space. [NEEDS EDIT/EX-PANSION]

## Software References

57 SMPy requires and uses the following packages:

- NumPy
- SciPy
- Pandas

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- ₁ Astropy
- 62 Matplotlib
  - PyYAML



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