

Shear Mapping in Python (SMPy): Modular, Extensible, and Accessible Dark Matter Mapping

Georgios N. Vassilakis¹, Jacqueline E. McCleary¹, Maya Amit¹, and Sayan Saha¹

¹ Department of Physics, Northeastern University, Boston, MA, USA ¶ Corresponding author

DOI: [10.xxxxxx/draft](https://doi.org/10.xxxxxx/draft)

Software

- [Review](#)
- [Repository](#)
- [Archive](#)

Editor: [Open Journals](#)

Reviewers:

- [@openjournals](#)

Submitted: 01 January 1970

Published: unpublished

License

Authors of papers retain copyright and release the work under a Creative Commons Attribution 4.0 International License ([CC BY 4.0](#)).

Summary

Understanding the universe's large-scale distribution of dark matter is a central objective in the era of precision cosmology. A key technique for the study of dark matter is weak gravitational lensing: a phenomenon where light from distant galaxies is sheared as it passes through the gravitational field of a massive object, like a galaxy cluster. This shear, which manifests as a slight (weak) distortion of shapes over thousands of galaxies, allows astrophysicists to infer the distribution of total matter, including both luminous and dark matter.

Obtaining a mass distribution from a catalog of galaxy shears requires an intermediate step. A common tool for this step is the mapping of convergence (κ), which quantifies how much a gravitational lens converges the light from distant galaxies, resulting in a magnification of their shapes. This value is directly proportional to the projected mass density, enabling easy visualization of the overall mass distribution. For a comprehensive review of weak gravitational lensing refer to (Umetsu, 2020).

The **Shear Mapping in Python (SMPy)** package provides a standardized, well-documented, and open-source solution for creating convergence maps from weak lensing galaxy shear measurements. SMPy was initially developed to support the Superpressure Balloon-borne Imaging Telescope (SuperBIT), a stratospheric, near-UV to near-IR observing platform which completed its 45-night observing run in spring 2023 with over 30 galaxy cluster observations (Gill et al., 2024), (Sirks et al., 2023). SMPy has since evolved into a general-purpose tool suitable for analyzing the weak lensing data from cosmological.

Statement of Need

The weak lensing community is served by publicly available mapping tools like lenspack and jax-lensing (Remy et al., 2022), each with their own strengths. jax-lensing excels at neural network-based approaches and deep learning methods, while lenspack has a well-documented module specifically for traditional mass-mapping implementations. While both tools are powerful, the steep learning curve of jax-lensing and the rigid architecture of lenspack motivated the development of SMPy as an accessible and extensible alternative.

Thus, SMPy addresses an outstanding need for the lensing community: an accessible, well-documented, and extensible tool to construct publication-quality mass maps from galaxy shear data. Built on standard scientific Python packages, it provides an easy entry point for researchers new to mass mapping. It offers specialized and unique features valuable for mass mapping, such as flexible coordinate system support (both celestial and pixel space) and comprehensive signal-to-noise analysis with multiple randomization techniques. Its modular architecture also enables future contributions of new mapping methods. An example convergence map, created from simulated SuperBIT galaxy cluster observations (McCleary et al., 2023), is shown in

41 Figure 1. SMPy is, to our knowledge, the first convergence mapping software to prioritize both
42 accessibility and advanced features.

43 SMPy was built with the following design principles in mind:

- 44 1. **Accessibility:** SMPy is written entirely in Python and deliberately relies only on widely-used
45 scientific Python packages (numpy, scipy, pandas, astropy, matplotlib, and pyyaml).
46 This choice of standard dependencies ensures that users can easily install the packages
47 without complex dependency chains, and that the codebase is maintainable and familiar
48 to the scientific Python community.
- 49 2. **Extensibility:** SMPy is built with a modular architecture that allows for easy implementation
50 of new mass mapping techniques beyond the currently implemented Kaiser-Squires
51 inversion algorithm (Kaiser & Squires, 1993). For example, we are planning to add
52 aperture mass mapping (Leonard et al., 2012) and KS+ (Pires, 2020) algorithms to the
53 codebase.
- 54 3. **Usability:** Creating convergence maps with SMPy requires minimal input—users need to
55 only provide a catalog of galaxies with their associated shear components and coordinates.
56 This straightforward input requirement makes the tool accessible to researchers at all
57 levels. A flexible configuration system is integrated via a single YAML file that defines
58 file paths, convergence map algorithm settings, plotting parameters, and more.
- 59 4. **Robustness:** SMPy is designed to be mathematically and algorithmically accurate, allowing
60 the user to create convergence maps with any galaxy shear data. The coordinate
61 system abstraction handles both celestial coordinates (with proper spherical geometry
62 approximations) or pixel-based coordinates through a unified interface. To quantify the
63 significance of the weak lensing detection, multiple noise realizations can be generated
64 using either spatial shuffling (randomizing galaxy positions while preserving shear values)
65 or orientation shuffling (randomizing shear orientations while preserving positions). These
66 noise realizations are used to create a signal-to-noise map with the observed convergence.

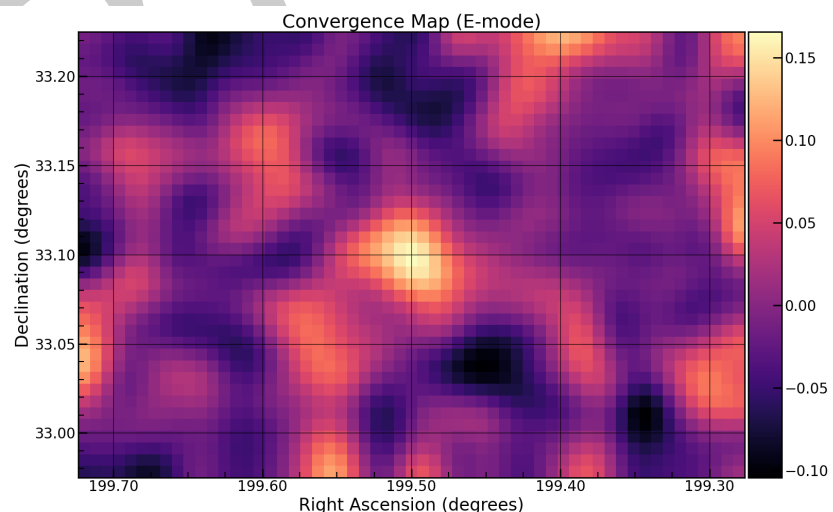


Figure 1: Example convergence map created with SMPy showing the mass distribution of a simulated galaxy cluster. The map was generated using the Kaiser-Squires inversion method on simulated weak lensing data from SuperBIT. The color scale represents the dimensionless surface mass density (convergence), with brighter regions indicating higher mass concentrations.

67 Software References

68 SMPy is written in Python 3.8+ and uses the following packages:

- 69 ▪ NumPy ([Harris et al., 2020](#))
- 70 ▪ SciPy ([Virtanen et al., 2020](#))
- 71 ▪ Pandas ([team, 2024](#))
- 72 ▪ Astropy ([Astropy Collaboration et al., 2022](#)) ([Astropy Collaboration et al., 2018](#)) ([Astropy](#)
- 73 [Collaboration et al., 2013](#))
- 74 ▪ Matplotlib ([Hunter, 2007](#))
- 75 ▪ PyYAML ([Simonov, 2024](#))

Acknowledgements

This material is based upon work supported by a Northeastern University Undergraduate Research and Fellowships PEAK Summit Award.

References

- Astropy Collaboration, Price-Whelan, A. M., Lim, P. L., Earl, N., Starkman, N., Bradley, L., Shupe, D. L., Patil, A. A., Corrales, L., Brasseur, C. E., Nöthe, M., Donath, A., Tollerud, E., Morris, B. M., Ginsburg, A., Vaher, E., Weaver, B. A., Tocknell, J., Jamieson, W., ... Astropy Project Contributors. (2022). The Astropy Project: Sustaining and Growing a Community-oriented Open-source Project and the Latest Major Release (v5.0) of the Core Package. *935*(2), 167. <https://doi.org/10.3847/1538-4357/ac7c74>
- Astropy Collaboration, Price-Whelan, A. M., Sipőcz, B. M., Günther, H. M., Lim, P. L., Crawford, S. M., Conseil, S., Shupe, D. L., Craig, M. W., Dencheva, N., Ginsburg, A., VanderPlas, J. T., Bradley, L. D., Pérez-Suárez, D., de Val-Borro, M., Aldcroft, T. L., Cruz, K. L., Robitaille, T. P., Tollerud, E. J., ... Astropy Contributors. (2018). The Astropy Project: Building an Open-science Project and Status of the v2.0 Core Package. *156*(3), 123. <https://doi.org/10.3847/1538-3881/aabc4f>
- Astropy Collaboration, Robitaille, T. P., Tollerud, E. J., Greenfield, P., Droettboom, M., Bray, E., Aldcroft, T., Davis, M., Ginsburg, A., Price-Whelan, A. M., Kerzendorf, W. E., Conley, A., Crighton, N., Barbary, K., Muna, D., Ferguson, H., Grollier, F., Parikh, M. M., Nair, P. H., ... Streicher, O. (2013). Astropy: A community Python package for astronomy. *558*.
- Gill, A. S., Benton, S. J., Damaren, C. J., Everett, S. W., Fraisse, A. A., Hartley, J. W., Harvey, D., Holder, B., Huff, E. M., Jauzac, M., Jones, W. C., Lagattuta, D., Leung, J. S.-Y., Li, L., Luu, T. V. T., Massey, R., McCleary, J. E., Nagy, J. M., Netterfield, C. B., ... Vitorelli, A. Z. (2024). SuperBIT superpressure flight instrument overview and performance: Near-diffraction-limited astronomical imaging from the stratosphere. *The Astronomical Journal*, *168*(2), 85. <https://doi.org/10.3847/1538-3881/ad5840>
- Harris, C. R., Millman, K. J., Walt, S. J. van der, Gommers, R., Virtanen, P., Cournapeau, D., Wieser, E., Taylor, J., Berg, S., Smith, N. J., Kern, R., Picus, M., Hoyer, S., Kerkwijk, M. H. van, Brett, M., Haldane, A., Río, J. F. del, Wiebe, M., Peterson, P., ... Oliphant, T. E. (2020). Array programming with NumPy. *Nature*, *585*(7825), 357–362. <https://doi.org/10.1038/s41586-020-2649-2>
- Hunter, J. D. (2007). Matplotlib: A 2D graphics environment. *Computing in Science & Engineering*, *9*(3), 90–95. <https://doi.org/10.1109/MCSE.2007.55>
- Kaiser, N., & Squires, G. (1993). Mapping the dark matter with weak gravitational lensing. *404*, 441–450. <https://doi.org/10.1086/172297>
- Leonard, A., Pires, S., & Starck, J.-L. (2012). Fast calculation of the weak lensing aperture mass statistic. *Monthly Notices of the Royal Astronomical Society*, *423*(4), 3405–3412. <https://doi.org/10.1111/j.1365-2966.2012.21133.x>

- 114 McCleary, J. E., Everett, S. W., Shaaban, M. M., Gill, A. S., Vassilakis, G. N., Huff, E.
115 M., Massey, R. J., Benton, S. J., Brown, A. M., Clark, P., Holder, B., Fraisse, A.
116 A., Jauzac, M., Jones, W. C., Lagattuta, D., Leung, J. S.-Y., Li, L., Luu, T. V. T.,
117 Nagy, J. M., ... Tam, S. I. (2023). Lensing in the blue. II. Estimating the sensitivity of
118 stratospheric balloons to weak gravitational lensing. *The Astronomical Journal*, 166(3),
119 134. <https://doi.org/10.3847/1538-3881/ace7ca>
- 120 Pires, S. (2020). Euclid: Reconstruction of weak-lensing mass maps for non-gaussianity studies.
121 *Astronomy & Astrophysics*, 638, A141. <https://doi.org/10.1051/0004-6361/201936865>
- 122 Remy, B., Lanusse, F., Jeffrey, N., Liu, J., Starck, J.-L., Osato, K., & Schrabback, T. (2022).
123 Probabilistic mass-mapping with neural score estimation. *Astronomy & Astrophysics*, 672.
124 <https://doi.org/10.1051/0004-6361/202243054>
- 125 Simonov, K. (2024). PyYAML. <https://pyyaml.org/>
- 126 Sirks, E. L., Massey, R., Gill, A. S., Anderson, J., Benton, S. J., Brown, A. M., Clark, P.,
127 English, J., Everett, S. W., Fraisse, A. A., Franco, H., Hartley, J. W., Harvey, D., Holder,
128 B., Hunter, A., Huff, E. M., Hynous, A., Jauzac, M., Jones, W. C., ... Vassilakis, G. N.
129 (2023). Data downloaded via parachute from a NASA super-pressure balloon. *Aerospace*,
130 10(11). <https://doi.org/10.3390/aerospace10110960>
- 131 team, T. pandas development. (2024). *Pandas-dev/pandas: pandas* (Version v2.2.2). Zenodo.
132 <https://doi.org/10.5281/zenodo.10957263>
- 133 Umetsu, K. (2020). Cluster–galaxy weak lensing. *The Astronomy and Astrophysics Review*,
134 28(1), 106. <https://doi.org/10.1007/s00159-020-00129-w>
- 135 Virtanen, P., Gommers, R., Oliphant, T. E., Haberland, M., Reddy, T., Cournapeau, D.,
136 Burovski, E., Peterson, P., Weckesser, W., Bright, J., van der Walt, S. J., Brett, M., Wilson,
137 J., Millman, K. J., Mayorov, N., Nelson, A. R. J., Jones, E., Kern, R., Larson, E., ... SciPy
138 1.0 Contributors. (2020). SciPy 1.0: Fundamental Algorithms for Scientific Computing in
139 Python. *Nature Methods*, 17, 261–272. <https://doi.org/10.1038/s41592-019-0686-2>