




# Shear Mapping in Python (SMPy): Modular, Extensible, and Accessible Dark Matter Mapping

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## Software

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## Summary

Understanding the universe's structure, particularly the nature of dark matter, is a central challenge in modern astrophysics. A key approach to this problem is the study of weak gravitational lensing, where light from distant galaxies is bent as it passes through the gravitational field of a massive object, like a galaxy cluster. Measuring this slight (weak) bending of light over thousands of galaxies allows astrophysicists to infer the distribution of matter, including dark matter.

A common tool for analyzing these distortions on large scales is convergence mapping. Convergence ( $\kappa$ ) quantifies how much light from distant galaxies converge due to lensing, resulting in a magnification or distortion of their images. For a comprehensive review of weak gravitational lensing, please refer to ([Umetsu, 2020](#)). By mapping convergence across the sky, astronomers can identify areas with high mass concentration based on observed lensing data. Regions showing significant convergence but little visible matter likely indicate the presence of dark matter causing the lensing effect.

The **Shear Mapping in Python (SMPy)** package provides a standardized, well-documented, and open-source solution for creating convergence maps from weak lensing galaxy shear measurements. SMPy was initially developed to support the Superpressure Balloon-borne Imaging Telescope (SuperBIT), which completed its 45-night observing run in spring 2023 with over 30 galaxy cluster observations ([Gill et al., 2024](#)). SMPy has evolved into a general-purpose tool suitable for analyzing the weak lensing data from any source of galaxies.

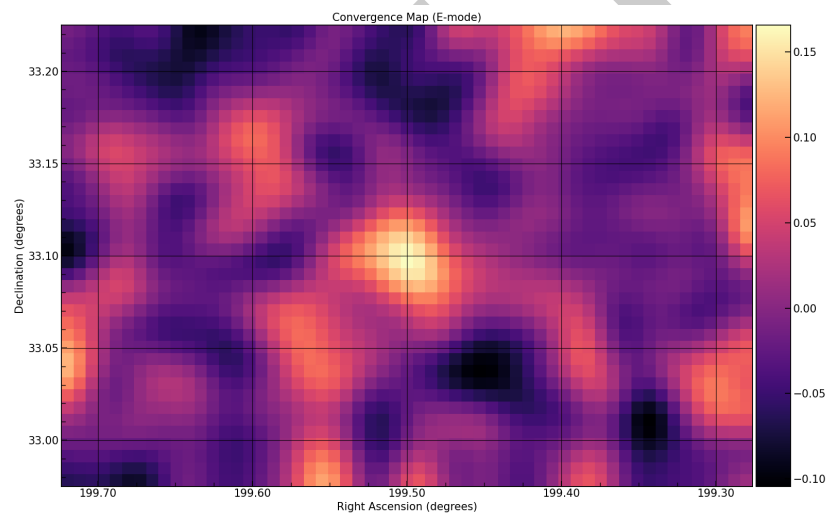
## Statement of Need

Mass maps are a critical and key part of many cosmological analyses ([Darwish et al., 2020](#)) ([Jeffrey et al., 2021](#)) ([Oguri et al., 2017](#)). SMPy addresses an outstanding need for the lensing community: A robust, well-documented, and open-source tool to construct publication quality mass maps from galaxy shear data. SMPy was built with multiple design directions in mind:

1. **Accessibility:** SMPy is written entirely in Python and deliberately relies only on widely-used scientific Python packages (numpy, scipy, pandas, astropy, matplotlib, and pyyaml). This choice of standard dependencies ensures that users can easily install the packages without complex dependency chains, and that the codebase is maintainable and familiar to the scientific Python community.
2. **Extensibility:** SMPy is built with a modular architecture that allows for easy implementation of new mass mapping techniques beyond the currently implemented Kaiser-Squires inversion algorithm ([Kaiser & Squires, 1993](#)). An example convergence map is shown in Figure 1, created from simulated galaxy cluster observations from SuperBIT ([McCleary](#)

et al., 2023). Aperture mass mapping (Leonard et al., 2012) and KS+ (Pires, 2020) algorithms are currently planned to be added to the codebase.

3. **Usability:** Creating convergence maps with SMPy requires minimal input - users need only provide a catalog of galaxies with their associated shears ( $g_1$  &  $g_2$ ) and coordinates. This straightforward input requirement makes the tool accessible to researchers at all levels. A flexible configuration system is integrated via a single YAML file that defines file paths, convergence map algorithm settings, plotting parameters, and more.
4. **Robustness:** Designed to be mathematically and algorithmically accurate, allowing the user to create convergence maps with any galaxy shear data. The coordinate system abstraction handles both RA/Dec celestial coordinates (with proper spherical geometry approximations) or pixel-based coordinates through a unified interface. Signal-to-noise maps can be generated using either spatial shuffling (randomizing galaxy positions while preserving shear values) or orientation shuffling (randomizing shear orientations while preserving positions) to distinguish real signals from noise.



**Figure 1:** Example convergence map created with SMPy showing the mass distribution of a simulated galaxy cluster. The map was generated using the Kaiser-Squires inversion method on simulated weak lensing data from SuperBIT. The color scale represents the dimensionless surface mass density (convergence), with brighter regions indicating higher mass concentrations.

To our knowledge, SMPy is the first open-source, well-documented convergence mapping software that can compute convergence in either astrometric or pixel space. [NEEDS EDIT/EXPANSION]

## Software References

SMPy requires and uses the following packages: - NumPy - SciPy - Pandas - Astropy - Matplotlib - PyYAML

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Darwish, O., Madhavacheril, M. S., Sherwin, B. D., Aiola, S., Battaglia, N., Beall, J. A., Becker, D. T., Bond, J. R., Calabrese, E., Choi, S. K., Devlin, M. J., Dunkley, J., Dünner,

- 65 R., Ferraro, S., Fox, A. E., Gallardo, P. A., Guan, Y., Halpern, M., Han, D., ... Wollack, E.  
66 J. (2020). The atacama cosmology telescope: A CMB lensing mass map over 2100 square  
67 degrees of sky and its cross-correlation with BOSS-CMASS galaxies. *Monthly Notices of the*  
68 *Royal Astronomical Society*, 500(2), 2250–2263. <https://doi.org/10.1093/mnras/staa3438>
- 69 Gill, A. S., Benton, S. J., Damaren, C. J., Everett, S. W., Fraisse, A. A., Hartley, J. W.,  
70 Harvey, D., Holder, B., Huff, E. M., Jauzac, M., Jones, W. C., Lagattuta, D., Leung, J.  
71 S.-Y., Li, L., Luu, T. V. T., Massey, R., McCleary, J. E., Nagy, J. M., Netterfield, C. B., ...  
72 Vitorelli, A. Z. (2024). SuperBIT superpressure flight instrument overview and performance:  
73 Near-diffraction-limited astronomical imaging from the stratosphere. *The Astronomical*  
74 *Journal*, 168(2), 85. <https://doi.org/10.3847/1538-3881/ad5840>
- 75 Jeffrey, N., Gatti, M., Chang, C., Whiteway, L., Demirbozan, U., Kovacs, A., Pollina, G.,  
76 Bacon, D., Hamaus, N., Kacprzak, T., Lahav, O., Lanusse, F., Mawdsley, B., Nadathur, S.,  
77 Starck, J. L., Vielzeuf, P., Zeurcher, D., Alarcon, A., Amon, A., ... Collaboration, D. (2021).  
78 Dark energy survey year 3 results: Curved-sky weak lensing mass map reconstruction.  
79 *Monthly Notices of the Royal Astronomical Society*, 505(3), 4626–4645. <https://doi.org/10.1093/mnras/stab1495>
- 80 Kaiser, N., & Squires, G. (1993). Mapping the dark matter with weak gravitational lensing.  
81 404, 441–450. <https://doi.org/10.1086/172297>
- 82 Leonard, A., Pires, S., & Starck, J.-L. (2012). Fast calculation of the weak lensing aperture  
83 mass statistic. *Monthly Notices of the Royal Astronomical Society*, 423(4), 3405–3412.  
84 <https://doi.org/10.1111/j.1365-2966.2012.21133.x>
- 85 McCleary, J. E., Everett, S. W., Shaaban, M. M., Gill, A. S., Vassilakis, G. N., Huff, E.  
86 M., Massey, R. J., Benton, S. J., Brown, A. M., Clark, P., Holder, B., Fraisse, A.  
87 A., Jauzac, M., Jones, W. C., Lagattuta, D., Leung, J. S.-Y., Li, L., Luu, T. V. T.,  
88 Nagy, J. M., ... Tam, S. I. (2023). Lensing in the blue. II. Estimating the sensitivity of  
89 stratospheric balloons to weak gravitational lensing. *The Astronomical Journal*, 166(3),  
90 134. <https://doi.org/10.3847/1538-3881/ace7ca>
- 91 Oguri, M., Miyazaki, S., Hikage, C., Mandelbaum, R., Utsumi, Y., Miyatake, H., Takada,  
92 M., Armstrong, R., Bosch, J., Komiyama, Y., Leauthaud, A., More, S., Nishizawa, A. J.,  
93 Okabe, N., & Tanaka, M. (2017). Two- and three-dimensional wide-field weak lensing  
94 mass maps from the hyper supprime-cam subaru strategic program S16A data. *Publications*  
95 *of the Astronomical Society of Japan*, 70(SP1), S26. <https://doi.org/10.1093/pasj/psx070>
- 96 Pires, S. (2020). Euclid: Reconstruction of weak-lensing mass maps for non-gaussianity studies.  
97 *Astronomy & Astrophysics*, 638, A141. <https://doi.org/10.1051/0004-6361/201936865>
- 98 Umetsu, K. (2020). Cluster–galaxy weak lensing. *The Astronomy and Astrophysics Review*,  
99 28(1), 106. <https://doi.org/10.1007/s00159-020-00129-w>
- 100