

Name: Subhojit Ghimire

Roll. : 1912160

Section: C.S.E. 'K'

## PHYSICS ASSIGNMENT-II

1. Compute the intensity of the standing electromagnetic wave given by  $\vec{E}(x, t) = 2E_0 \cos kx \cos \omega t \hat{j}$ ,  
 $\vec{B}(x, t) = 2B_0 \sin kx \sin \omega t \hat{k}$ .

Soln: The Poynting vector is,

$$\vec{S} = \frac{\vec{E} \times \vec{B}}{\mu_0} = \frac{1}{\mu_0} (4E_0 B_0 \sin kx \sin \omega t \cos kx \cos \omega t \hat{j} \times \hat{k})$$

$$= \frac{E_0 B_0}{\mu_0} \sin 2kx \sin 2\omega t \hat{i}$$

∴ The time average of  $S$  is,

$$\langle S \rangle = \frac{E_0 B_0}{\mu_0} \sin 2K \langle \sin(2\omega t) \rangle = 0$$

$$\Rightarrow I = \langle S \rangle = 0$$

∴ The result is to be expected since the standing wave does not propagate.

2. Given  $\vec{E} = \hat{i} E_0 \cos \omega \left( \frac{z}{c} - t \right) + \hat{j} E_0 \sin \omega \left( \frac{z}{c} - t \right)$ ,  
determine the magnetic field  $\vec{B}$ .

Soln: From Maxwell's equation,

$$\vec{\nabla} \times \vec{E} = -\frac{d\vec{B}}{dt} \quad \text{--- (i)}$$

$$\text{Now, } \vec{\nabla} \times \vec{E} = \begin{vmatrix} \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ E_0 \cos \omega \left( \frac{z}{c} - t \right) & E_0 \sin \omega \left( \frac{z}{c} - t \right) & 0 \end{vmatrix}$$

$$= \hat{i} \left( 0 - \frac{E_0 \omega}{c} \cos \omega \left( \frac{z}{c} - t \right) \right) \\ + \hat{j} \left( 0 + \frac{E_0 \omega}{c} \sin \omega \left( \frac{z}{c} - t \right) \right) + 0$$

Putting in (i), we get,

$$\frac{-d\vec{B}}{dt} = -\frac{E_0 \omega}{c} \left[ \cos \omega \left( \frac{z}{c} - t \right) \hat{i} + \sin \omega \left( \frac{z}{c} - t \right) \hat{j} \right]$$

Integrating on both sides, we get,

$$B = \frac{E_0}{c} \left[ -\sin \omega \left( \frac{z}{c} - t \right) \hat{i} + \cos \omega \left( \frac{z}{c} - t \right) \hat{j} \right] + \frac{m}{\text{Constant}}$$

Q3 A uniform plane wave has a wavelength of 3cm in free space and 2cm in a dielectric for which  $\mu = 4.7 \times 10^{-7} \text{ NA}^{-2}$ . Determine the dielectric constant of the dielectric.

Soln:-

we know,  $c \propto \lambda$

$$\text{and, } c = \frac{1}{\sqrt{\epsilon_0 \mu_0}}$$

$$\text{Let, } \lambda_0 = 3 \text{ cm, } \lambda_m = 2 \text{ cm}$$

So,

$$\frac{\lambda_0}{\lambda_m} = \sqrt{\frac{\epsilon_m \mu_m}{\epsilon_0 \mu_0}}$$

$$\text{or, } \frac{3}{2} = \sqrt{\frac{\epsilon_m \mu_m}{\epsilon_0 \mu_0}}$$

$$\text{or, } \frac{9}{4} = \frac{\epsilon_m \mu_m}{\epsilon_0 \mu_0}$$

Let, dielectric constant  $\frac{\epsilon_m}{\epsilon_0} = k$  (say)

$$\text{Then, } K = \frac{9}{4} \times 2.7$$

$$\therefore K = 6.075$$

40 A plane electromagnetic wave propagates from one dielectric to another at normal incidence. Find the ratio of the indices of refraction of the two dielectrics for which the reflection and transmission coefficients are both equal to 0.5.

Soln:-

$$R = T = 0.5 \quad (\text{given})$$

Also,

$$R = \left( \frac{n_1 - n_2}{n_1 + n_2} \right)^2 \quad - (i)$$

$$\text{and, } T = \frac{4n_1 n_2}{(n_1 + n_2)^2} \quad - (ii)$$

Using (i), we get

$$\frac{1}{2} = \left( \frac{n_1/n_2 - 1}{n_1/n_2 + 1} \right)^2$$

Let,  $\frac{n_1}{n_2} = k$ , then

$$\frac{k-1}{k+1} = \pm \frac{1}{\sqrt{2}}$$

$$\text{for } \frac{k-1}{k+1} = \frac{1}{\sqrt{2}} \quad \Rightarrow \quad k = \frac{2\sqrt{2}}{\sqrt{2}-1} = 3+2\sqrt{2}$$

$$\text{Similarly for } \frac{k-1}{k+1} = -\frac{1}{\sqrt{2}}$$

$$k = 3-2\sqrt{2}$$

$$\therefore k = 3 \pm 2\sqrt{2}$$

Exo A uniform plane wave whose electric field is given by  $E_I = 100 \cos(\omega t - 6\pi x) \hat{z} \text{ V m}^{-1}$  is incident from a region having  $\epsilon_1 = 4\epsilon_0$ ,  $\mu_1 = \mu_0$  normal to the plane surface of a material having  $\epsilon_2 = 9\epsilon_0$ ,  $\mu_2 = 4\mu_0$ . Write complete expressions for the incident, reflected and transmitted electric and magnetic fields.

Sol<sup>n</sup>

$$E_I = 100 \cos(\omega t - 6\pi x) \hat{z} \text{ V m}^{-1}$$

$$\Rightarrow E_I = 100 \cos\{- (6\pi x - \omega t)\} \hat{z}$$

$$\Rightarrow E_I = 100 \cos(6\pi x - \omega t) \hat{z} \quad \left\{ \because \cos(-\theta) = \cos\theta \right\}$$

$$\therefore \vec{V} = V(\hat{n})$$

$$\therefore B_I = \frac{1}{V_I} (\hat{n}) \times E(\hat{z}) = \frac{-1}{V_I} 100 \cos(6\pi x - \omega t) \hat{y}$$

for Medium I,

$$\epsilon_1 = 4\epsilon_0; \mu_1 = \mu_0$$

for Medium II,  $\epsilon_2 = 9\epsilon_0; \mu_2 = 4\mu_0$

Now,

$$100 = E_0 \sin$$

$$\therefore E_{\text{reflected}} = \left( \frac{1-\beta}{1+\beta} \right) E_0 \sin$$

where,

$$\beta = \frac{\mu_1 V_1}{\mu_2 V_2} = \left( \frac{\epsilon_2 \mu_1}{\epsilon_1 \mu_2} \right)^{\frac{1}{2}} = \left( \frac{9\epsilon_0 \mu_0}{4\epsilon_0 4\mu_0} \right)^{\frac{1}{2}}$$

$$\therefore \beta = \frac{3}{4}$$

$$\therefore E_{\text{reflection}} = \left( \frac{1 - \frac{3}{4}}{1 + \frac{3}{4}} \right) \times 100 = \frac{100}{7}$$

$$\therefore E_{\text{transmitted}} = \left( \frac{2}{1+\beta} \right) \times 100 = \frac{2}{7/4} \times 100 = \frac{8}{9} \times 100$$



∴ Equation for reflected wave;

$$E_I \text{ reflected} = \frac{100}{7} \cos(-6\pi x - \omega t) (-\hat{z})$$

$$B_I \text{ reflected} = \frac{1}{V_1} \left( \frac{100}{7} \right) \cos(-6\pi x - \omega t) (-\hat{y})$$

∴ Equation for transmitted wave:

$$E_I \text{ transmitted} = \left( \frac{800}{7} \right) \cos(6\pi x - \omega t) (\hat{z})$$

$$B_I \text{ transmitted} = \frac{1}{V_2} \left( \frac{800}{7} \right) \cos(6\pi x - \omega t) (\hat{y})$$

Now,

$$V_1 = \frac{1}{\sqrt{\epsilon_1 \mu_1}} = \frac{1}{\sqrt{4\epsilon_0 \mu_0}} = \frac{c}{2}$$

$$V_2 = \frac{1}{\sqrt{\epsilon_2 \mu_2}} = \frac{1}{3 \times 2 \sqrt{\epsilon_0 \mu_0}} = \frac{c}{6}$$

6. Find the momentum of (a) a 10.0-MeV gamma ray  
 (b) a 25-KeV x-ray (c) a 1.0- $\mu\text{m}$  infrared photon.  
 (d) a 150-MHz radio-wave photon. Express the momentum  
 in kg m/s and eV/c.

Soln we know,

$$E^2 = (pc)^2 + (mc^2)^2$$

Here,  $m=0$ , So,

$$E = pc \Rightarrow P = \frac{E}{c}$$

$$(a) \quad P = \frac{10 \times 10^6}{3 \times 10^8} \times 1.6 \times 10^{-19}$$

$$\Rightarrow P = 0.53 \times 10^{-20} \text{ kg.m/s}$$

$$\text{Again, } P = \frac{E}{c} = 10 \times 10^6 = 10^7 \frac{\text{eV}}{c}$$

$$(b) \quad P = \frac{E}{c} = \frac{25 \times 10^3}{c} = 25 \times 10^3 \frac{\text{eV}}{c}$$

$$\text{Again, } P = \frac{25 \times 10^3 \times 1.6 \times 10^{-19}}{3 \times 10^8} = 13.33 \times 10^{-24} \frac{\text{kg.m}}{\text{s}}$$

$$(c) \quad E = \frac{hc}{\lambda} = \frac{6.6 \times 10^{-34} \times 3 \times 10^8}{10^{-6}} \text{ J}$$

$$= \frac{6.6 \times 10^{-28} \times 3 \times 10^8}{1.6 \times 10^{-19}} \text{ eV}$$

$$\therefore E = 1.24 \text{ eV}$$

$$\therefore P = 1.24 \frac{\text{eV}}{c}$$

$$\text{Also, } P = \frac{6.6 \times 10^{-34} \times c}{10^{-6} \times c} = 6.6 \times 10^{-28} \frac{\text{kg.m}}{\text{s}}$$

$$(d) \quad E = \frac{h\nu}{c} = \frac{6.6 \times 10^{-34} \times 150 \times 10^6}{1.6 \times 10^{-19}}$$

$$\Rightarrow E = 6.188 \text{ eV}$$

$$\therefore P = 6.188 \frac{\text{eV}}{\text{e}}$$

$$\text{Again, } P = \frac{E}{c} = \frac{6.6 \times 10^{-34} \times 150 \times 10^6}{3 \times 10^8}$$

$$\therefore P = 3.3 \times 10^{-34} \text{ kg m/s}$$

Q70 A metal surface has a photoelectric cutoff wavelength of 325.6 nm. It is illuminated with light of wavelength 259.8 nm. What is the stopping potential?

Soln:- we know  $E = W_0 + eV$

$$\text{And, } V = \frac{hc}{e} \left( \frac{1}{\lambda} - \frac{1}{\lambda_0} \right)$$

$$= \frac{6.6 \times 10^{-34} \times 3 \times 10^8 \times 65.8 \times 10^{-9}}{1.6 \times 10^{-19} \times 325.6 \times 10^{-9} \times 259.8 \times 10^{-9}}$$

$$\therefore V = 0.968 \text{ eV} \quad \text{is the stopping potential}$$

Q80 The surface of the sun has temperature of 6000 K. At what wavelength does the sun emit its peak intensity? How does this compare with the peak sensitivity of the human eye?

$$\text{Soln:- } \lambda_{\text{peak(sun)}} = \frac{2.898 \times 10^{-3}}{T(\text{in K})} = \frac{2.898 \times 10^{-3}}{6000}$$

$$\therefore \lambda_{\text{peak}} = 4.83 \times 10^{-7} \text{ m} = 483 \text{ nm}$$

The peak sensitivity of human eye is 555 nm.

$$\therefore \lambda_{\text{peak(eye)}} \approx 1.15 \times \lambda_{\text{peak(sun)}}$$



9. A proton is accelerated from rest through a potential difference of  $-2.36 \times 10^5 \text{ V}$ . What is its de-Broglie wavelength.

Soln- We know,  $\lambda = \frac{h}{\sqrt{2mqV}}$

$$\Rightarrow \lambda = \frac{6.6 \times 10^{-34}}{\sqrt{2 \times 1.67 \times 10^{-27} \times 1.6 \times 10^{-19} \times 2.36 \times 10^5}}$$

$$\Rightarrow \lambda = 587.7 \times 10^{-12} \text{ m}$$

$$\therefore \lambda = 587.7 \text{ pm}$$

10. An electron moves in  $x$  direction with a speed of  $3.6 \times 10^6 \text{ m/s}$ . we can measure its speed to a precision of 1%. With what precision can we simultaneously measure its  $x$  coordinate?

Soln- From Heisenberg's uncertainty principle,

$$\Delta x \Delta p_x = \frac{h}{4\pi}$$

$$\text{or } \Delta p_x = 3.1 \times 10^{-31} \times 3.6 \times 10^6 \times \frac{1}{100}$$

$$\text{or } \Delta p_x = 3.3 \times 10^{-26} \frac{\text{kg m}}{\text{s}}$$

$$\Rightarrow \Delta x = \frac{6.6 \times 10^{-34}}{4 \times \left(\frac{22}{7}\right) \times 3.3 \times 10^{-26}}$$

$$\therefore \Delta x = 1.59 \text{ nm}$$