



STRING THEORY:

AN INTRODUCTION

MARC SCOTT

SHEP

25 FEBRUARY 2013

Disclaimer:

I must have been crazy to suggest I'd do this!

I'll try and give an unbiased view.

“The moment you encounter string theory and realize that almost all of the major developments in physics over the last hundred years emerge -- and emerge with such elegance -- from such a simple starting point, you realize that this incredibly compelling theory is in a class of its own” M. Green

“What makes string theory so difficult to assess dispassionately is that it gains its support and chooses its directions of development almost entirely from aesthetic judgements guided by mathematical desiderata. I believe that it is important to record each of the turnings that the theory has undergone, and to point out that almost every turn has taken us further from observationally established facts.” R. Penrose

Contents

- ♦ Why do we need string theory
(or at least new physics...)
- ♦ Overview of the properties of string theory
- ♦ Classical string theory as a toy model



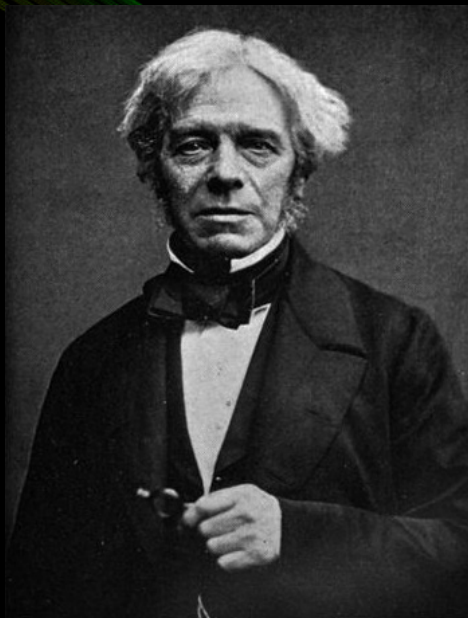
Why bother with string theory?

UNIFICATION OF PHYSICAL LAWS!

- In 1820 Biot, Savart and Ampère find electric currents produce B-fields
- In 1831 Faraday shows that a changing magnetic flux induces a current
- In 1865 James Clerk Maxwell derives his famous Maxwell equations, formally uniting Electricity and Magnetism into Electromagnetism.

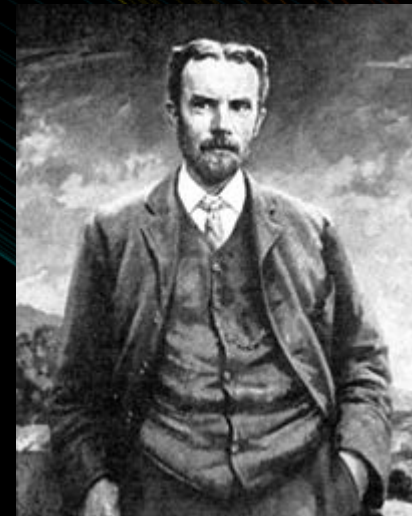
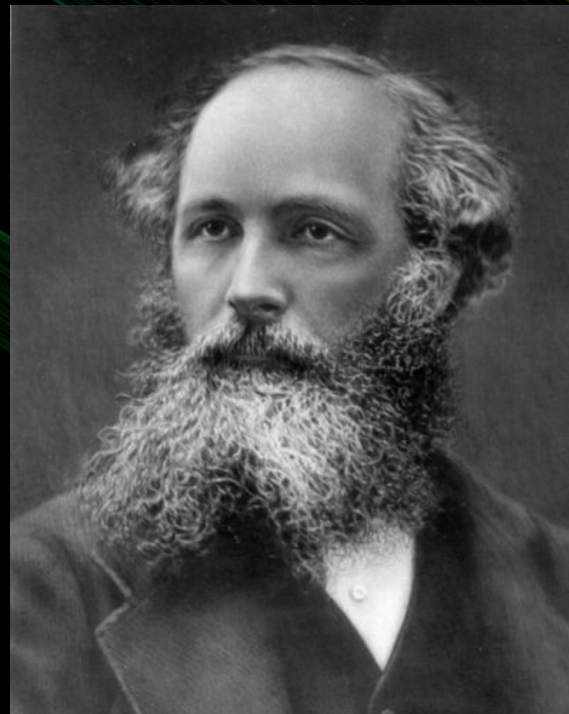


JB Biot



M Faraday

JCM



O Heaviside

Why bother with string theory?

UNIFICATION OF PHYSICS!

- Next big unification of fundamental forces: 1960s (Nobel prize 1979)
Glashow, **Weinberg** and **Salam** unite EM with the Weak interaction
 - Electroweak theory defines the electroweak scale, (~ 250 GeV) above which EM and the weak interaction unite.
 - W^+ , W^- , B^0 , A^0 become W^+ , W^- , Z^0 and γ via EWSB (\sim masses via Higgs mechanism)
 - Success of **QFT** = **SR** + **QM**
 - Birth of the **Standard Model**



Weinberg



Glashow



Salam

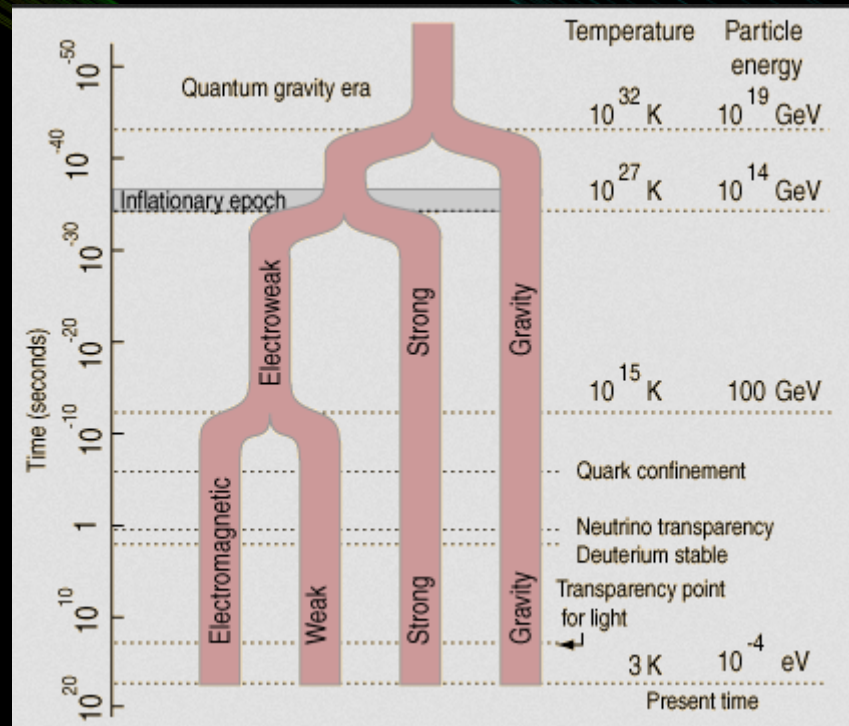
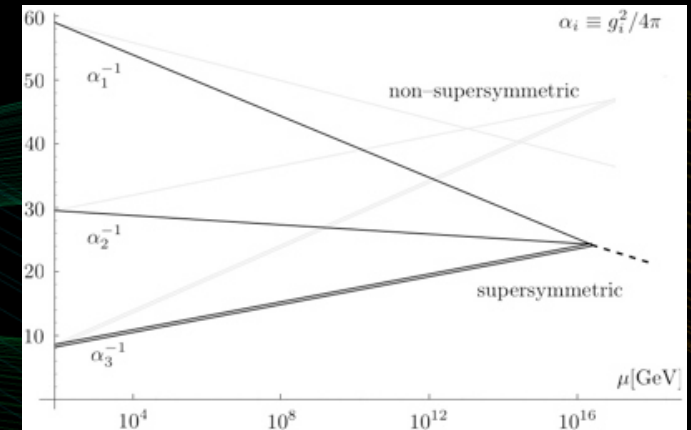
Why bother with string theory?

UNIFICATION OF PHYSICS!

- What's Next?

$$\text{GUTs} = \text{QCD} + \text{EW}$$

- Bigger picture: SM \supset QED + QCD + EW has problems at high energies



Why bother with string theory?

STANDARD MODEL PROBLEMS

- Standard Model Lagrangian... Large and Arbitrary?
- 18(25) Free Parameters
 - 12 Fermion Masses
 - 3 CKM mixing angles + 1 phase
 - 3 PMNS mixing angles + 1 phase
 - 3 coupling constants
 - Higgs mass
 - Z^0 mass

Plus more...

- Hierarchy Problems
- Neutrino Masses
- Dark Matter/ 'Energy'
- Imbalance of Particle to Antiparticle

$$\begin{aligned}
 & -\frac{1}{2}\partial_\nu g_\mu^a \partial_\nu g_\mu^a - g_s f^{abc} \partial_\mu g_\nu^a g_\mu^b g_\nu^c - \frac{1}{4}g_s^2 f^{abc} f^{ade} g_\mu^a g_\nu^b g_\mu^c g_\nu^d + \\
 & \frac{1}{2}ig_s^2(\bar{q}_i^\mu \gamma^\mu q_j^\mu)g_\mu^a + G^a \partial^2 G^a + g_s f^{abc} \partial_\mu G^a G^b g_\mu^c - \partial_\nu W_\mu^+ \partial_\nu W_\mu^- - \\
 & M^2 W_\mu^+ W_\mu^- - \frac{1}{2}\partial_\nu Z_\mu^0 \partial_\nu Z_\mu^0 - \frac{1}{2c_w^2}M^2 Z_\mu^0 Z_\mu^0 - \frac{1}{2}\partial_\mu A_\nu \partial_\mu A_\nu - \frac{1}{2}\partial_\mu H \partial_\mu H - \\
 & \frac{1}{2}m_h^2 H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - M^2 \phi^+ \phi^- - \frac{1}{2}\partial_\mu \phi^0 \partial_\mu \phi^0 - \frac{1}{2c_w^2}M^2 \phi^0 \phi^0 - \beta_h \left[\frac{2M^2}{g^2} + \right. \\
 & \left. \frac{2M}{g}H + \frac{1}{2}(H^2 + \phi^0 \phi^0 + 2\phi^+ \phi^-) \right] + \frac{2M^4}{g^2}\alpha_h - igc_w [\partial_\nu Z_\mu^0 (W_\mu^+ W_\nu^- - \\
 & W_\nu^+ W_\mu^-) - Z_\nu^0 (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + Z_\mu^0 (W_\nu^+ \partial_\nu W_\mu^- - \\
 & W_\nu^- \partial_\nu W_\mu^+)] - igs_w [\partial_\nu A_\mu (W_\mu^+ W_\nu^- - W_\nu^+ W_\mu^-) - A_\nu (W_\mu^+ \partial_\nu W_\mu^- - \\
 & W_\mu^- \partial_\nu W_\mu^+) + A_\mu (W_\nu^+ \partial_\nu W_\mu^- - W_\nu^- \partial_\nu W_\mu^+)] - \frac{1}{2}g^2 W_\mu^+ W_\mu^- W_\nu^+ W_\nu^- + \\
 & \frac{1}{2}g^2 W_\mu^+ W_\nu^- W_\mu^+ W_\nu^- + g^2 c_w^2 (Z_\mu^0 W_\mu^+ Z_\nu^0 W_\nu^- - Z_\mu^0 Z_\nu^0 W_\mu^+ W_\nu^-) + \\
 & g^2 s_w^2 (A_\mu W_\mu^+ A_\nu W_\nu^- - A_\mu A_\nu W_\mu^+ W_\nu^-) + g^2 s_w c_w [A_\mu Z_\nu^0 (W_\mu^+ W_\nu^- - \\
 & W_\nu^+ W_\mu^-) - 2A_\mu Z_\mu^0 W_\nu^+ W_\nu^-] - g\alpha [H^3 + H\phi^0 \phi^0 + 2H\phi^+ \phi^-] - \\
 & \frac{1}{8}g^2 \alpha_h [H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4H^2 \phi^+ \phi^- + 2(\phi^0)^2 H^2] - \\
 & gMW_\mu^+ W_\mu^- H - \frac{1}{2}g\frac{M}{c_w^2}Z_\mu^0 Z_\mu^0 H - \frac{1}{2}ig[W_\mu^+ (\phi^0 \partial_\mu \phi^- - \phi^- \partial_\mu \phi^0) - \\
 & W_\mu^- (\phi^0 \partial_\mu \phi^+ - \phi^+ \partial_\mu \phi^0)] + \frac{1}{2}g[W_\mu^+ (H\partial_\mu \phi^- - \phi^- \partial_\mu H) - W_\mu^- (H\partial_\mu \phi^+ - \\
 & \phi^+ \partial_\mu H)] + \frac{1}{2}g\frac{1}{c_w}(Z_\mu^0 (H\partial_\mu \phi^0 - \phi^0 \partial_\mu H) - ig\frac{s_w^2}{c_w}MZ_\mu^0 (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \\
 & igs_w MA_\mu (W_\mu^+ \phi^- - W_\mu^- \phi^+) - ig\frac{1-2c_w^2}{2c_w}Z_\mu^0 (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) + \\
 & igs_w A_\mu (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) - \frac{1}{4}g^2 W_\mu^+ W_\mu^- [H^2 + (\phi^0)^2 + 2\phi^+ \phi^-] - \\
 & \frac{1}{4}g^2 \frac{1}{c_w^2}Z_\mu^0 Z_\mu^0 [H^2 + (\phi^0)^2 + 2(2s_w^2 - 1)^2 \phi^+ \phi^-] - \frac{1}{2}g^2 \frac{s_w^2}{c_w}Z_\mu^0 \phi^0 (W_\mu^+ \phi^- + \\
 & W_\mu^- \phi^+) - \frac{1}{2}ig^2 \frac{s_w^2}{c_w}Z_\mu^0 H (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \frac{1}{2}g^2 s_w A_\mu \phi^0 (W_\mu^+ \phi^- + \\
 & W_\mu^- \phi^+) + \frac{1}{2}ig^2 s_w A_\mu H (W_\mu^+ \phi^- - W_\mu^- \phi^+) - g^2 \frac{s_w}{c_w} (2c_w^2 - 1) Z_\mu^0 A_\mu \phi^+ \phi^- - \\
 & g^1 s_w^2 A_\mu A_\mu \phi^+ \phi^- - \bar{e}^\lambda (\gamma \partial + m_e^\lambda) e^\lambda - \bar{\nu}^\lambda \gamma \partial \nu^\lambda - \bar{u}_j^\lambda (\gamma \partial + m_u^\lambda) u_j^\lambda - \bar{d}_j^\lambda (\gamma \partial + \\
 & m_d^\lambda) d_j^\lambda + igs_w A_\mu [-(\bar{e}^\lambda \gamma e^\lambda) + \frac{2}{3}(\bar{u}_j^\lambda \gamma u_j^\lambda) - \frac{1}{3}(\bar{d}_j^\lambda \gamma d_j^\lambda)] + \frac{ig}{4c_w} Z_\mu^0 [(\bar{\nu}^\lambda \gamma^\mu (1 + \\
 & \gamma^5) \nu^\lambda) + (\bar{e}^\lambda \gamma^\mu (4s_w^2 - 1 - \gamma^5) e^\lambda) + (\bar{u}_j^\lambda \gamma^\mu (\frac{4}{3}s_w^2 - 1 - \gamma^5) u_j^\lambda) + \\
 & (\bar{d}_j^\lambda \gamma^\mu (1 - \frac{8}{3}s_w^2 - \gamma^5) d_j^\lambda)] + \frac{ig}{2\sqrt{2}} W_\mu^+ [(\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) e^\lambda) + (\bar{u}_j^\lambda \gamma^\mu (1 + \\
 & \gamma^5) C_{\lambda\kappa} d_j^\kappa)] + \frac{ig}{2\sqrt{2}} W_\mu^- [(\bar{e}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{d}_j^\kappa C_{\lambda\kappa}^\dagger \gamma^\mu (1 + \gamma^5) u_j^\lambda)] + \\
 & \frac{ig}{2\sqrt{2}} \frac{m_e^\lambda}{M} [-\phi^+ (\bar{\nu}^\lambda (1 - \gamma^5) e^\lambda) + \phi^- (\bar{e}^\lambda (1 + \gamma^5) \nu^\lambda)] - \frac{g}{2} \frac{m_e^\lambda}{M} [H (\bar{e}^\lambda e^\lambda) + \\
 & i\phi^0 (\bar{e}^\lambda \gamma^5 e^\lambda)] + \frac{ig}{2M\sqrt{2}} \phi^+ [-m_d^\kappa (\bar{u}_j^\lambda C_{\lambda\kappa} (1 - \gamma^5) d_j^\kappa) + m_u^\lambda (\bar{u}_j^\lambda C_{\lambda\kappa} (1 + \\
 & \gamma^5) d_j^\kappa) + \frac{ig}{2M\sqrt{2}} \phi^- [m_d^\lambda (\bar{d}_j^\kappa C_{\lambda\kappa}^\dagger (1 + \gamma^5) u_j^\kappa) - m_u^\kappa (\bar{d}_j^\kappa C_{\lambda\kappa}^\dagger (1 - \gamma^5) u_j^\kappa) - \\
 & \frac{g}{2} \frac{m_u^\lambda}{M} H (\bar{u}_j^\lambda u_j^\lambda) - \frac{g}{2} \frac{m_d^\lambda}{M} H (\bar{d}_j^\lambda d_j^\lambda) + \frac{ig}{2} \frac{m_u^\lambda}{M} \phi^0 (\bar{u}_j^\lambda \gamma^5 u_j^\lambda) - \frac{ig}{2} \frac{m_d^\lambda}{M} \phi^0 (\bar{d}_j^\lambda \gamma^5 d_j^\lambda) + \\
 & \bar{X}^+ (\partial^2 - M^2) X^+ + \bar{X}^- (\partial^2 - M^2) X^- + \bar{X}^0 (\partial^2 - \frac{M^2}{c_w^2}) X^0 + \bar{Y} \partial^2 Y + \\
 & igc_w W_\mu^+ (\partial_\mu \bar{X}^0 X^- - \partial_\mu \bar{X}^+ X^0) + igs_w W_\mu^+ (\partial_\mu \bar{Y} X^- - \partial_\mu \bar{X}^+ Y) + \\
 & igc_w W_\mu^- (\partial_\mu \bar{X}^- X^0 - \partial_\mu \bar{X}^0 X^+) + igs_w W_\mu^- (\partial_\mu \bar{X}^- Y - \partial_\mu \bar{Y} X^+) + \\
 & igc_w Z_\mu^0 (\partial_\mu \bar{X}^+ X^- - \partial_\mu \bar{X}^- X^+) + igs_w A_\mu (\partial_\mu \bar{X}^+ X^- - \partial_\mu \bar{X}^- X^+) - \\
 & \frac{1}{2}gM[\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{c_w^2} \bar{X}^0 X^0 H] + \frac{1-2c_w^2}{2c_w} igM[\bar{X}^+ X^0 \phi^+ - \\
 & \bar{X}^- X^0 \phi^-] + \frac{1}{2c_w} igM[\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-] + igMs_w[\bar{X}^0 X^- \phi^+ - \\
 & \bar{X}^0 X^+ \phi^-] + \frac{1}{2}igM[\bar{X}^+ X^+ \phi^0 - \bar{X}^- X^- \phi^0]
 \end{aligned}$$

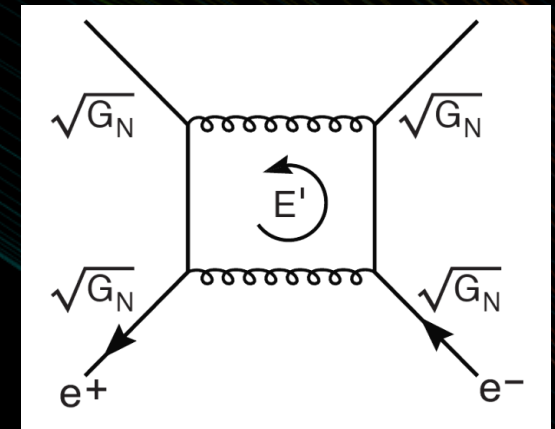
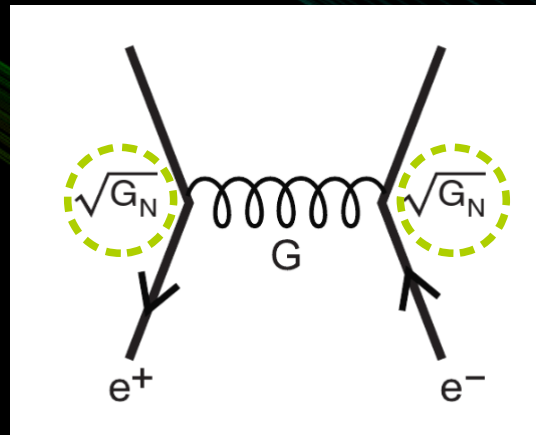
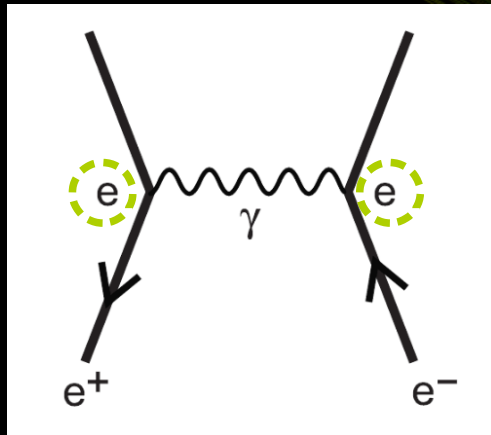
Why bother with string theory?

STANDARD MODEL PROBLEMS: GRAVITY

- Adapt (NR)QM for gravity : Schrödinger-Newton equation

$$i\frac{\partial\Psi}{\partial t} = -\frac{1}{2m}\nabla^2\Psi + V\Psi + m\Phi\Psi$$

- OK for low energies... how about QFT and Gravity? **Non-renormalisable!**



$$\mathcal{A} \propto e^2 = 4\pi\alpha_{\text{EM}}$$

$$\mathcal{A}_{\text{grav}}^{(0)} \propto G_N E^2$$

$$\mathcal{A}_{\text{grav}}^{(1)} \propto G_N^2 E^2 \int_0^\infty E' dE'$$

$$G_N \approx 7 \times 10^{-33} \text{TeV}^{-2}$$

Problem when $E \sim M_{\text{p}}$

“Strongly Divergent”

And now for some string theory...

Answer to all the problems? (Maybe) String Theory

WHAT EXACTLY IS STRING THEORY?

1 dimensional extension to the 'normal' point-particle (0-D)
representation of physics (QFT)

Admits open and closed strings

WHY IS THIS BETTER THAN QFT AND THE SM?

String theory combines all of the ideas of QFT and the SM into
a more general, tighter, unified structure.



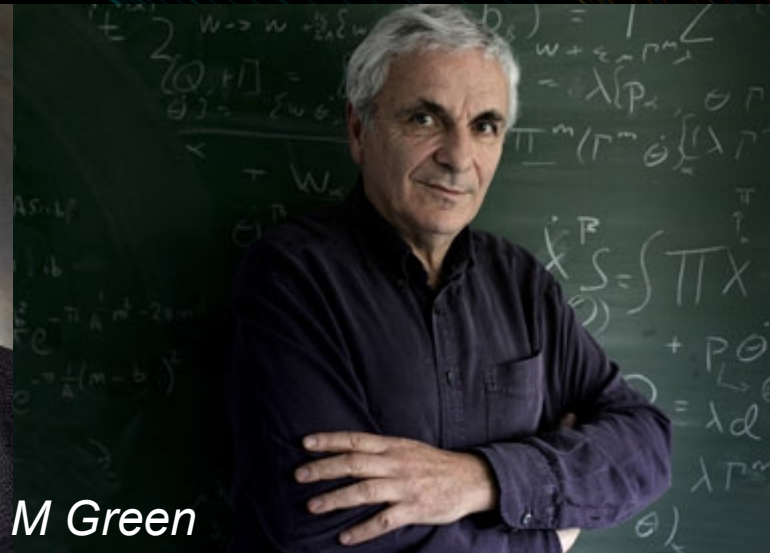
Closed string



Open string



E Witten



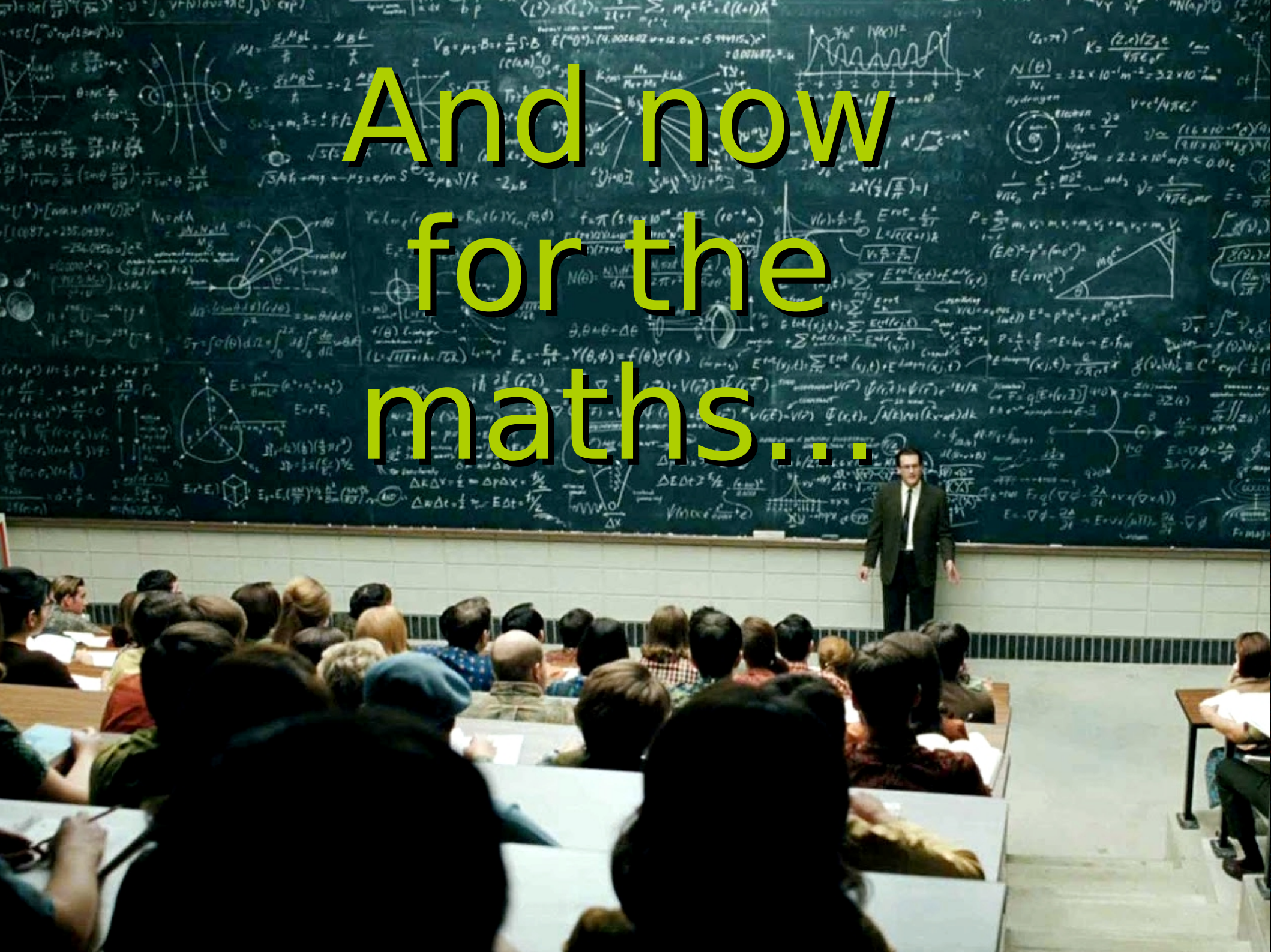
M Green

String Theory Ideas

STRING THEORY PROVIDES (for better or for worse)

- ♦ **Gravity** – All consistent STs contain massless spin-2 graviton whose interactions \rightarrow GR in low energy
- ♦ **Quantum Gravity** – In contradiction to QFT. **No infinities here!**
- ♦ **GUTs** – ST gauge groups large enough to swallow SM. Reduces to SM in low energy limit
- ♦ **Extra Dimensions** – ST requires definite # of EDs. 10 (26)
- ♦ **Supersymmetry** – All consistent STs require SUSY
- ♦ **Many Vacua** – Quantized ST \rightarrow many possible ground states
- ♦ **Free Parameters?** - ST has at most 1 free parameter
- ♦ **Uniqueness** – Lack of adjustable free parameters. Spacetime dimension # inherent.
- ♦ **Testing** – No decisive tests (?)

And now
for the
maths...



- General theory: p-brane theory
(this is an example of the humour you can expect from string theorists)

- Point Particle = 0-brane
- String = 1-brane
- Membrane = 2-brane

And so on...

The 0-brane

Action of a point particle

$$S_0 = -m \int ds = -m \int d\tau \sqrt{-g_{\mu\nu}(x) \dot{x}^\mu \dot{x}^\nu}$$

- Difficult to quantise → remove square root
- How about massless particles?

Introduce the **Einbein** auxiliary field... can reform into

$$\tilde{S}_0 = \frac{1}{2} \int d\tau \left(e(\tau)^{-1} \dot{x}^2 - m^2 e(\tau) \right)$$

Plus EoM
in $e(\tau)$

Generalising the 0-brane action

Can generalise \tilde{S}_0 to any p-brane

$$S_p = -T_p \int d\mu_p$$

Our (p+1)-dim volume element $d\mu_p$ generalises to

$$d\mu_p = \sqrt{-\det(G_{\alpha\beta})} d^{p+1}\sigma$$

With the pullback,

$$G_{\alpha\beta} = \frac{\partial x^\mu}{\partial \sigma^\alpha} \frac{\partial x^\nu}{\partial \sigma^\beta} g_{\mu\nu}(x)$$

$$\alpha, \beta \in \{0, 1, \dots, p\}$$

$$\mu, \nu \in \{0, 1, \dots, D \geq p\}$$

Comparison

$$S_0 = -m \int ds$$

For p=0, $\alpha, \beta=0$:

$$d\mu_p = \sqrt{-\det(G_{00})} d\tau$$

With

$$G_{00} = \dot{x}^\mu \dot{x}^\nu g_{\mu\nu}(x)$$

Regain

$$d\mu_0 = ds$$

1-brane or string action

- Worldline \rightarrow Worldsheet (parameterised by two coördinates)
- $\alpha, \beta = 0, 1$
- Assume background is Minkowski: $\eta_{\mu\nu}$

$$G_{\alpha\beta} = \frac{\partial x^\mu}{\partial \sigma^\alpha} \frac{\partial x^\nu}{\partial \sigma^\beta} \eta_{\mu\nu}(x) \rightarrow G_{\alpha\beta} = \begin{pmatrix} \dot{x}^2 & \dot{x}x' \\ \dot{x}x' & x'^2 \end{pmatrix}$$

Leads to **Nambu-Goto** action:

$$S_{NG} = -T \int d\tau d\sigma \sqrt{(\dot{x}x')^2 - \dot{x}^2 x'^2}$$

Add in auxillary field (analogue to $e(\tau)$) to get **POLYAKOV ACTION**

$$S_\sigma = -\frac{T}{2} \int d\tau d\sigma \sqrt{-h} h^{\alpha\beta} \frac{\partial x^\mu}{\partial \sigma^\alpha} \frac{\partial x^\nu}{\partial \sigma^\beta} g_{\mu\nu}$$

Finding the field equations

$$S_\sigma = -\frac{T}{2} \int d\tau d\sigma \sqrt{-h} h^{\alpha\beta} \frac{\partial x^\mu}{\partial \sigma^\alpha} \frac{\partial x^\nu}{\partial \sigma^\beta} g_{\mu\nu} \quad \text{Polyakov action}$$

- Can use two gauge symmetries of the action
- Reparameterisation Invariance $\sigma \rightarrow \sigma' = f(\sigma)$
- Weyl Invariance $h_{\alpha\beta}(\tau, \sigma) \rightarrow h'_{\alpha\beta}(\tau, \sigma) = e^{2\omega(\sigma)} h_{\alpha\beta}(\tau, \sigma)$
to gauge auxiliary field into Minkowski metric
- Weyl Inv is *local* \rightarrow cannot extend globally to worldsheet...
UNLESS... Worldsheet manifold has Euler characteristic = 0 (Topologically unobstructed)

A Polyakov

$$S_\sigma = \frac{T}{2} \int d\tau d\sigma (\dot{x}^2 - x'^2)$$



Finding the field equations

- Remarkably simple field equations result:

$$(\partial_\tau^2 - \partial_\sigma^2)x^\mu = 0$$

- Plus boundary conditions:

$$x^\mu(\tau, \sigma + n) = x^\mu(\tau, \sigma)$$

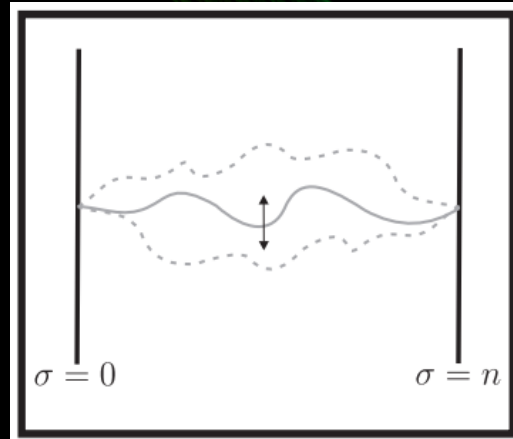
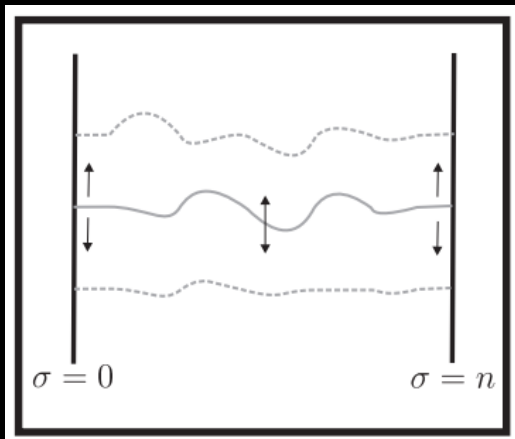
CLOSED STRING

$$\partial_\sigma x^\mu(\tau, \sigma) = \partial_\sigma x^\mu(\tau, \sigma + n) = 0$$

OPEN NEUMANN

$$x^\mu(\tau, \sigma = 0) = x_0^\mu \text{ AND } x^\mu(\tau, \sigma = n) = x_n^\mu$$

OPEN DIRICHLET



Finding the field equations

- Most general solution can be expanded in Fourier modes,

CLOSED STRING

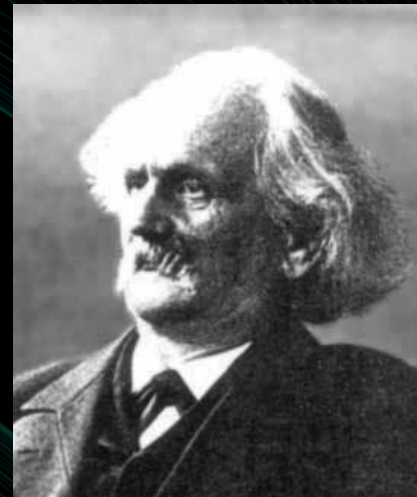
$$x^\mu = x_0^\mu + \tau l_s^2 p_0^\mu + \frac{i}{2} l_s \sum_{n \neq 0} \frac{1}{n} (\alpha_n^\mu e^{2in\sigma} + \tilde{\alpha}_n^\mu e^{-2in\sigma}) e^{-2in\tau}$$

OPEN STRING (NEUMANN)

$$x^\mu = x_0^\mu + l_s^2 \tau p_0^\mu + i l_s \sum_{m \neq 0} \frac{1}{m} \alpha_m^\mu e^{-im\tau} \cos(m\sigma)$$



J Fourier



C Neumann

$\ddot{T}h\tilde{\alpha}_N k \ \Upsilon \mathcal{O}(u)$