Geopotential

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This document describes the computations that are performed by the method GeneralSphericalHarmonicsAcceleration of class Geopotential to the determine the acceleration exerted by a non-spherical celestial on a point mass.

Let r be the vector going from the centre of the celestial to the point mass. Let (x, y, z) be a (direct) base whose z axis is along the axis of rotation of the celestial and whose x axis points toward a reference point on the celestial. In this base r has coordinates (x, y, z) which can be expressed in terms of the latitude β and the longitude λ :

$$r = \begin{pmatrix} x \\ y \\ z \end{pmatrix} = \begin{pmatrix} r \cos \beta \cos \lambda \\ r \cos \beta \sin \lambda \\ r \sin \beta \end{pmatrix}$$

where r is the norm of r.

The gravitational potential exerted by the celestial has the form:

$$U(\mathbf{r}) = -\frac{\mu}{r} \left(1 + \sum_{n=1}^{\infty} \sum_{m=0}^{n} \left(\frac{R}{r} \right)^{n} P_{mn} \left(\sin \beta \right) (C_{mn} \cos m\lambda + S_{mn} \sin m\lambda) \right)$$

where P_{mn} is the associated Legendre function defined as:

$$P_{mn} = (1 - t^2)^{\frac{m}{2}} \frac{d^m P_n(t)}{dt^m}$$

The terms that are relevant for evaluating the acceleration exerted by the spherical harmonics on the point mass have the form:

$$V_{nm}(\mathbf{r}) = \frac{1}{r} \left(\frac{R}{r}\right)^n P_{mn}(\sin\beta) (C_{mn}\cos m\lambda + S_{mn}\sin m\lambda)$$

and the acceleration itself is proportional to $\nabla V_{nm}(\mathbf{r})$.

 $V_{nm}(\mathbf{r})$ can be written as the product of a radial term, a latitudinal term and a longitudinal term, dependent of r, β , and λ , respectively:

$$V_{nm}(\mathbf{r}) = \Re(r) \Re(\beta) \Re(\lambda)$$

where:

$$\begin{cases} \Re(r) &= \frac{1}{r} \left(\frac{R}{r}\right)^n \\ \Re(\beta) &= P_{mn} (\sin \beta) \\ \Re(\lambda) &= C_{mn} \cos m\lambda + S_{mn} \sin m\lambda \end{cases}$$

The overall gradient can then be written as:

$$\nabla V_{nm}(\mathbf{r}) = \nabla \Re(\mathbf{r}) \Re(\beta) \Re(\lambda) + \Re(\mathbf{r}) \nabla \Re(\beta) \Re(\lambda) + \Re(\mathbf{r}) \Re(\beta) \nabla \Re(\lambda)$$

To compute this expression, we first determine the gradient of various elements appearing in the potential. Starting with the \Re term, we have trivially:

$$\nabla r^2 = \nabla (x^2 + y^2 + z^2) = 2r$$

from which we deduce:

$$\nabla r^n = \nabla ((r^2)^{\frac{n}{2}}) = \frac{n}{2} (r^2)^{\frac{n}{2} - 1} (2r) = nr^{n - 2} r$$

Noting that $\forall z = \mathbf{z}$ we can compute the following gradient needed for the $\mathfrak B$ term:

$$\nabla \sin \beta = \nabla \frac{z}{r} = \frac{r\mathbf{z} - zr^{-1}\mathbf{r}}{r^2} = \frac{\mathbf{z}}{r} - \mathbf{r}\frac{z}{r^3}$$

which can be written, in coordinates:

$$\nabla \sin \beta = \frac{1}{r^3} \begin{pmatrix} -xz \\ -yz \\ x^2 + y^2 \end{pmatrix} = \frac{1}{r^3} \begin{pmatrix} -r^2 \sin \beta \cos \beta \cos \lambda \\ -r^2 \sin \beta \cos \beta \sin \lambda \\ r^2 \cos^2 \beta \end{pmatrix} = \frac{\cos \beta}{r} \begin{pmatrix} -\sin \beta \cos \lambda \\ -\sin \beta \sin \lambda \\ \cos \beta \end{pmatrix}$$

For the $\mathfrak L$ term, we need to evaluate:

$$\begin{cases} \nabla \cos m\lambda &= -m\sin m\lambda \, \forall \lambda \\ \nabla \sin m\lambda &= m\cos m\lambda \, \forall \lambda \end{cases}$$

The angle λ is $\arctan \frac{y}{x}$ thus:

$$\nabla \lambda = \frac{1}{1 + \left(\frac{y}{x}\right)^2} \nabla \left(\frac{y}{x}\right) = \frac{1}{1 + \left(\frac{y}{x}\right)^2} \frac{xy - yx}{x^2} = \frac{xy - yx}{x^2 + y^2}$$

which can be written, in coordinates:

$$\forall \lambda = \frac{1}{x^2 + y^2} \begin{pmatrix} -y \\ x \\ 0 \end{pmatrix} = \frac{1}{r^2 \cos^2 \beta} \begin{pmatrix} -r \cos \beta \sin \lambda \\ r \cos \beta \cos \lambda \\ 0 \end{pmatrix} = \frac{1}{r \cos \beta} \begin{pmatrix} -\sin \lambda \\ \cos \lambda \\ 0 \end{pmatrix}$$